# Phase structure of λφ<sup>4</sup> model in the presence of a magnetic background

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# Motivation

Strong interactions under magnetic fields can be found in nature in:

- Magnetars
- The early universe
- Non-central heavy ion collisions



The earth's magnetic field

0.6 Gauss



A common hand-held magnet

100 Gauss



The strongest steady magnetic fields  $4.5 \times 10^5$  Gauss achieved so far in the laboratory

Surface field of magnetars

10<sup>15</sup> Gauss



Heavy ion collisions: the strongest magnetic field ever achieved in the 10<sup>18</sup> Gauss laboratory

Kharzeev, QM2009

# Heavy ion collision: why?



High temperature plasma – above 4 trillion Kelvin

# Heavy ion collision: where?

# Relativistic heavy ion collider (RHIC)

Large Hadron Collider (LHC) (~ 1 month a year)





200 GeV/nucleon

2.76 TeV/nucleon



VS

## Proton-proton



~ 20 particles

Heavy ions



~ 4800 particles in central collisions

# High temperatures and densities yield to phase transitions Confinement – Deconfinement

Chiral symmetry restoration



## Phase transitions

## Broken x restored phase:

The order parameter carries information about the symmetries of the system.

Chiral symmetry: 
$$\begin{cases} \langle \sigma \rangle \neq 0 &, & \text{low } T \\ \langle \sigma \rangle = 0 &, & \text{high } T \end{cases}$$
  
Confinement: 
$$\begin{cases} \langle L \rangle = 0 &, & \text{low } T \\ \langle L \rangle \neq 0 &, & \text{high } T \end{cases}$$



## Magnetic fields in heavy ion collisions





Au-Au, 62 GeV



Au-Au, 200 GeV

# THE CHIRAL MAGNETIC EFFECT

non-trivial gauge field configurations



the second second second

[Kharzeev, Warringa, McLerran]

#### strong magnetic field

Charge separation

• How does the QCD diagram look like including another external control parameter, the magnetic field B?

- Are there modifications in the nature of the phase transition?
- How do the chiral and deconfinement transitions react to this magnetic field?
- How is the interplay between this two transitions?



# Quark-gluon plasma: how to deal with it?

QCD:

$$\mathcal{L} = -\frac{1}{4} F^a_{\mu\nu} F^{\mu\nu}_a + \sum_f \overline{\psi}_f (i\gamma^\mu \partial_\mu - m_f - ig\gamma^\mu A^a_\mu \tau^a) \psi_f$$

$$F^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + g f^{abc} A^b_\mu A^c_\nu$$

Dynamic fields:  $N_f$  flavors of quarks,  $N_c$  colors and  $(N_c^2-1)$  gluons; SU(N) symmetry: non-Abelian, gluons interact with themselves.

Lattice QCD: full numerical simulations. Sign problem, ambiguities in fermion definitions, computationally expensive.

Asymptotic freedom  $\Rightarrow$  perturbation theory

AdS/CFT: formal solutions of a `cousin' theory.

Effective models: not the whole story. Mimic of a sector of the theory.

Combination of all this can lead to a more robust characterization of the plasma!



## Gauge sector

#### Polyakov loop models:

$$\ell(x) = \frac{1}{N} \operatorname{tr} \left( \mathcal{P} \exp \left( ig \int_0^{1/T} A_0(x,\tau) \, d\tau \right) \right)$$

#### Parameters fix demanding:

- Stefan-Boltzmann limit reached at
   T → ∞
- first order transition happens at  $T=T_0$

 the potential fits lattice data for thermodynamical quantities (pressure, energy density and entropy)



 $\frac{V_L(L,T)}{T^4} = -\frac{1}{2}a(T)L^*L + b(T)\ln\left[1 - 6L^*L + 4\left(L^{*3} + L^3\right) - 3\left(L^*L\right)^2\right]$ 

# Finite temperature formalism

## Quantum mechanics

$$\hat{\rho} = e^{-\beta(\hat{H} - \mu_i \hat{N}_i)}$$
  $Z = Z(V, T, \mu_1, \mu_2, ...) = Tr \hat{\rho}$ 

Field theory – Matsubara frequencies

$$Z(\beta) = \int [\mathcal{D}q(\tau)] e^{-S_E(\beta)} \left[ S_E = \int_0^\beta d\tau \int d^3x \left[ \frac{1}{2} \phi \left( -\frac{\partial^2}{\partial \tau^2} - \nabla^2 + m^2 \right) \phi + V(\phi) \right] \right]$$

$$S_E = \frac{1}{2}\beta^2 \sum_{n=-\infty}^{\infty} \int \frac{d^3k}{(2\pi)^3} [(\omega^2 + \omega_n^2)\phi_n(k)\phi_{-n}(-n)]$$

$$\log(Z(\beta)) = -\frac{1}{2} \sum_{n} \int \frac{d^{3}k}{(2\pi)^{3}} \log(\omega^{2} + \omega_{n}^{2})$$

# Ring diagrams

In thermal field theory the contribution after 1-loop order is not 2-loops. A sum of a class of diagrams converge to a contribution of order  $\lambda^{3/2}$ 





IR regularization for massless theories!

## Including an external magnetic field

For simplicity we assume a magnetic field that is constant and homogeneous:

$$\vec{B} = B\hat{z}$$
Gauge choice
$$A^{\mu} = (A^{0}, \vec{A}) = (0, -By, 0, 0)$$

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$$(\partial^{2} + m^{2})\phi = 0$$

$$\partial_{\mu} \to \partial_{\mu} + iqA_{\mu}$$
Fermionic  $\Rightarrow$ 

$$\begin{pmatrix} i\gamma^{\mu}\partial_{\mu} - m)\psi = 0\\ \partial_{\mu} \to \partial_{\mu} + iqA_{\mu} \end{pmatrix}$$

• charged scalars: propagator in the presence of a magnetic field

$$D = \int_0^\infty \frac{ds}{\cosh(qBs)} e^{-s\left[(\omega_n - i\mu)^2 + p_z^2 + m^2 + p_\perp^2 \frac{\tanh(qBs)}{qBs}\right]}$$

The  $\lambda \phi^4$  model

$$L = \frac{1}{2} \Big[ \partial_{\mu} \phi \partial^{\mu} \phi \Big] - m^2 |\phi|^2 + \lambda |\phi|^4 + \delta L$$

Kinetic term

Spontaneously symmetry breaking Counterterms

## Weak field approximation

$$iD^{B}(k) = \frac{i}{k_{\parallel}^{2} - k_{\perp}^{2} - m^{2}} \left\{ 1 - \frac{(eB)^{2}}{(k_{\parallel}^{2} - k_{\perp}^{2} - m^{2})^{2}} - \frac{2(eB)^{2}k_{\perp}^{2}}{(k_{\parallel}^{2} - k_{\perp}^{2} - m^{2})^{3}} \right\}$$

# Effective potential

1-loop: 
$$V^{(1)} = -\frac{\pi^2 T^4}{90} + \frac{m^2 T^2}{24} - \frac{m^3 T}{12\pi} - \frac{m^4}{32\pi^2} ln\left(\frac{m}{4\pi T}\right) + \mathcal{O}(m^4)$$

# Ring diagrams contribution

$$V^{(ring)} = \frac{T}{12\pi} \left[ m^3 - (m^2 + \Pi_{\beta}^B)^{3/2} \right] +$$

$$\frac{T}{8\pi}(eB)^2 \frac{\Pi_\beta}{(\Pi^B_\beta)^2} \left[ (m^2 + \Pi^B_\beta)^{1/2} - \frac{m^2 + \Pi^B_\beta}{m} + \cdots \right]$$

## Vacuum contribution: IR divergences

$$= \int dm^2 \ T \sum_{n=-\infty}^\infty \int rac{d^d p}{(2\pi)^d} D(m{k},\omega_n)$$

## Renormalization!

$$\begin{split} V_B &= -\left[\frac{\mu^2}{4} - \frac{18\lambda^2 v_0^2}{32\pi^2}\psi\left(\frac{1}{2} + \frac{3\lambda v_0^2 - \mu^2}{2qB}\right) \\ &+ \frac{2qB}{(4\pi)^2} 3\lambda \left\{\ln\Gamma\left(\frac{1}{2} + \frac{3\lambda v_0^2}{2qB}\right) - \frac{1}{2}\ln(2\pi)\right\}\right]v^2 \\ &+ \left[\frac{\mu^2}{8v_0^2} - \frac{9\lambda^2}{32\pi^2}\psi\left(\frac{1}{2} + \frac{3\lambda v_0^2 - \mu^2}{2qB}\right)\right]v^4 \\ &+ \left(\frac{2qB}{4\pi}\right)^2 \int dm^2 \left(\ln\Gamma\left(\frac{1}{2} + \frac{3\lambda v_0^2}{2qB}\right) - \frac{1}{2}\ln(2\pi)\right) \end{split}$$

Put all together and study the parameter space

## Final remarks

QCD in the presence of magnetic fields is a subject that has been calling attention in the last years.

Contributions from different fields: effective models, lattice, AdS/CFT, data analysis, etc.

Magnetic backgrounds can change remarkably the QCD phase diagram: 2<sup>nd</sup> to 1<sup>st</sup> phase transitions, change of the critical temperature, split of the transition lines, etc.

Mix of different areas and techniques: thermal field theory, electrodynamics, heavy ion phenomenology, phase structures, etc...

Recent area: still a lot to explore and to be understood.