

The S_3 flavour symmetry and the reactor neutrino mixing angle

arXiv:1205.4755

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PASCOS 2012

Experimental data of neutrino oscillations at $1 \sigma^1$:

- * The neutrino squared mass differences:

$$\Delta m_{21}^2 = 7.62 \pm 0.19 \times 10^{-5} \text{ eV}^2, \quad \Delta m_{31}^2 = \begin{cases} -2.40_{-0.07}^{+0.10} \times 10^{-3} \text{ eV}^2, \\ +2.53_{-0.10}^{+0.08} \times 10^{-3} \text{ eV}^2. \end{cases}$$

- * The solar and atmospheric mixing angles:

$$\sin^2 \theta'_{12} = 0.320_{-0.017}^{+0.015}, \quad \sin^2 \theta'_{23} = \begin{cases} 0.53_{-0.07}^{+0.05} \\ 0.49_{-0.05}^{+0.08} \end{cases},$$

- The reactor mixing angle

$$\sin^2 \theta'_{13} = \begin{cases} 0.027_{-0.004}^{+0.003} \\ 0.026_{-0.004}^{+0.003} \end{cases},$$

the upper (lower) row corresponds to inverted (normal) neutrino mass hierarchy.

¹Pilar Coloma *et al* arXiv:1206.0475 [hep-ph]

Implications of neutrino oscillations

- The simple fact that neutrinos oscillate is a signal of physics beyond Standard Model.
 - Neutrinos massive.
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- Impose the symmetry directly in the mass matrix and break sequentially according to the chain $S_3 \supset S_3^{\text{diag}} \supset S_2^{\text{diag}}$.

[J. Barranco, FGC and A Mondragón PRD 82 073010]

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- The S_3 symmetry is left unbroken, and the concept of flavour is extended to the Higgs sector by introducing in the theory three Higgs fields which are $SU(2)$ doublets.

[J. Kubo, A. Mondragón, M. Mondragón and E. Rodriguez-Jauregui Prog. Theor. Phys. 109 795-807]

Irreducible representations of S_3

The group S_3 has two one-dimensional irreps (singlets) and one two-dimensional irrep (doublet)

- one dimensional:
 - $\mathbf{1}_A$ antisymmetric singlet,
 - $\mathbf{1}_S$ symmetric singlet.
- Two - dimensional:
 - $\mathbf{2}$ doublet

Direct product of irreps of S_3

$$\mathbf{1}_S \otimes \mathbf{1}_S = \mathbf{1}_S, \mathbf{1}_S \otimes \mathbf{1}_A = \mathbf{1}_A,$$

$$\mathbf{1}_A \otimes \mathbf{1}_A = \mathbf{1}_S, \mathbf{1}_S \otimes \mathbf{2} = \mathbf{2},$$

$$\mathbf{1}_A \otimes \mathbf{2} = \mathbf{2}$$

$$\mathbf{2} \otimes \mathbf{2} = \mathbf{1}_S \oplus \mathbf{1}_A \oplus \mathbf{2}$$

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Direct product of irreps of S_3

$$\begin{aligned}\mathbf{1}_s \otimes \mathbf{1}_s &= \mathbf{1}_s, & \mathbf{1}_s \otimes \mathbf{1}_A &= \mathbf{1}_A, \\ \mathbf{1}_A \otimes \mathbf{1}_A &= \mathbf{1}_s, & \mathbf{1}_s \otimes \mathbf{2} &= \mathbf{2}, \\ \mathbf{1}_A \otimes \mathbf{2} &= \mathbf{2}\end{aligned}$$

$$\mathbf{2} \otimes \mathbf{2} = \mathbf{1}_s \oplus \mathbf{1}_A \oplus \mathbf{2}$$

[J. Kubo, A. Mondragón, M. Mondragón and E. Rodriguez-Jauregui Prog. Theor. Phys. 109 795-807]

The direct (tensor) product of two doublets

$$\mathbf{p}_D = \begin{pmatrix} p_{D1} \\ p_{D2} \end{pmatrix} \text{ and } \mathbf{q}_D = \begin{pmatrix} q_{D1} \\ q_{D2} \end{pmatrix}$$

has two singlets, r_s and r_A , and one doublet r_D^T

$$r_s = p_{D1}q_{D1} + p_{D2}q_{D2} \quad \text{is invariant,}$$

$$r_A = p_{D1}q_{D2} - p_{D2}q_{D1} \quad \text{is not invariant}$$

$$r_D^T = \begin{pmatrix} p_{D1}q_{D2} + p_{D2}q_{D1} \\ p_{D1}q_{D1} - p_{D2}q_{D2} \end{pmatrix}$$

A Minimal S_3 invariant extension of the Standard Model

The Higgs sector is modified,

$$\Phi \rightarrow H = (\Phi_1, \Phi_2, \Phi_3)^T$$

H is a reducible $\mathbf{1}_s \oplus 2$ rep. of S_3

$$H_s = \frac{1}{\sqrt{3}} (\Phi_1 + \Phi_2 + \Phi_3)$$

$$H_D = \begin{pmatrix} \frac{1}{\sqrt{2}} (\Phi_1 - \Phi_2) \\ \frac{1}{\sqrt{6}} (\Phi_1 + \Phi_2 - 2\Phi_3) \end{pmatrix}$$

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Quark, lepton and Higgs fields

$$Q^T = (u_L, d_L), u_R, d_R,$$

$$L^\dagger = (\nu_L, e_L), e_R, \nu_R, H$$

All these fields have three species (flavours) and belong to a reducible $\mathbf{1} \oplus 2$ rep. of S_3

Leptons' Yukawa interactions

$$\mathcal{L}_{Y_E} = -Y_1^e \bar{L}_I H_S e_{iR} - Y_3^e \bar{L}_3 H_S e_{3R} - Y_2^e [\bar{L}_I \kappa_{IJ} H_1 e_{jR} + \bar{L}_I \eta_{IJ} H_2 e_{jR}] \\ - Y_4^e \bar{L}_3 H_I e_{iR} - Y_5^e \bar{L}_I H_I e_{3R} + \text{h.c.},$$

$$\mathcal{L}_{Y_\nu} = -Y_1^\nu \bar{L}_I (i\sigma_2) H_S^* \nu_{iR} - Y_3^\nu \bar{L}_3 (i\sigma_2) H_S^* \nu_{3R} - Y_4^\nu \bar{L}_3 (i\sigma_2) H_I^* \nu_{iR} \\ - Y_2^\nu [\bar{L}_I \kappa_{IJ} (i\sigma_2) H_1^* \nu_{jR} + \bar{L}_I \eta_{IJ} (i\sigma_2) H_2^* \nu_{jR}] - Y_5^\nu \bar{L}_I (i\sigma_2) H_I^* \nu_{3R} + \text{h.c.},$$

$$\kappa = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \text{and} \quad \eta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

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$$\kappa = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \text{and} \quad \eta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

Furthermore, the Majorana mass terms for the right handed neutrinos are

$$\mathcal{L}_M = -\nu_R^T C M \nu_R$$

where C is the charge conjugation matrix.

Mass matrices

We will assume that

$$\langle H_{D1} \rangle = \langle H_{D2} \rangle \neq 0 \quad \text{and} \quad \langle H_3 \rangle \neq 0$$

and

$$\langle H_3 \rangle^2 + \langle H_{D1} \rangle^2 + \langle H_{D2} \rangle^2 \approx \left(\frac{246}{2} \text{GeV} \right)^2$$

Then, the Yukawa interactions yield mass matrices of the general form

$$\mathbf{M} = \begin{pmatrix} \mu_1 + \mu_2 & \mu_2 & \mu_5 \\ \mu_2 & \mu_1 - \mu_2 & \mu_5 \\ \mu_4 & \mu_4 & \mu_3 \end{pmatrix}$$

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The Majorana masses for ν_L are obtained from the see-saw mechanism

$$M_\nu = M_{\nu D} \tilde{M}^{-1} (M_{\nu D})^T \quad \text{with} \quad \tilde{M} = \text{diag}(M_1, M_2, M_3)$$

The leptonic sector

To achieve a further reduction of the number of parameters, in the leptonic sector, we introduce an additional discrete Z_2 symmetry

-	+
H_I, ν_{3R}	$H_S, L_3, L_I, e_{3R}, e_{IR}, \nu_{IR}$

then, $Y_1^e = Y_3^e = Y_1^\nu = Y_5^\nu = 0$. Hence, the leptonic mass matrices are

$$M_e = \begin{pmatrix} \mu_2^e & \mu_2^e & \mu_5^e \\ \mu_2^e & -\mu_2^e & \mu_5^e \\ \mu_4^e & \mu_4^e & 0 \end{pmatrix} \quad M_{\nu D} = \begin{pmatrix} \mu_2^\nu & \mu_2^\nu & 0 \\ \mu_2^\nu & -\mu_2^\nu & 0 \\ \mu_4^\nu & \mu_4^\nu & \mu_3^\nu \end{pmatrix}$$

The unitary matrix that diagonalized the mass matrix of the charged leptons as function of its eigenvalues

$$U_{eL} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & e^{i\delta_e} \end{pmatrix} \begin{pmatrix} \frac{1}{\sqrt{2}} \frac{\tilde{m}_e}{\tilde{m}_\mu} & -\frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ -\frac{1}{\sqrt{2}} \frac{\tilde{m}_e}{\tilde{m}_\mu} & \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ -1 & -\frac{\tilde{m}_e}{\tilde{m}_\mu} & 0 \end{pmatrix} + \mathcal{O}(10^{-5})$$

$$\tilde{m}_\mu = m_\mu/m_\tau \text{ and } \tilde{m}_e = m_e/m_\tau.$$

There are no free parameters in \mathbf{M}_e other than the Dirac Phase $\delta!!$.

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There are no free parameters in \mathbf{M}_e other than the Dirac Phase $\delta_{!!}$.

The neutrino mass matrix \mathbf{I}

The Majorana masses for ν_L are obtained from the see-saw mechanism

$$\mathbf{M}_\nu = \mathbf{M}_{\nu D} \tilde{\mathbf{M}}_R^{-1} \mathbf{M}_{\nu D}^T$$

with

$$\tilde{\mathbf{M}}_R = \text{diag}[M_1, M_2, M_3] \quad M_1 \neq M_2 \neq M_3$$

The Majorana masses for ν_L are obtained from the see-saw mechanism

$$\mathbf{M}_{\nu_L} = \begin{pmatrix} \frac{2(\mu_2^\nu)^2}{M} & \frac{2\lambda(\mu_2^\nu)^2}{\bar{M}} & \frac{2\mu_2^\nu\mu_4^\nu}{M} \\ \frac{2\lambda(\mu_2^\nu)^2}{M} & \frac{2(\mu_2^\nu)^2}{\bar{M}} & \frac{2\mu_2^\nu\mu_4^\nu\lambda}{M} \\ \frac{2\mu_2^\nu\mu_4^\nu}{M} & \frac{2\mu_2^\nu\mu_4^\nu\lambda}{M} & \frac{2(\mu_4^\nu)^2}{M} + \frac{(\mu_3^\nu)^2}{M_3} \end{pmatrix}, \quad \lambda = \frac{1}{2} \left(\frac{M_2 - M_1}{M_1 + M_2} \right),$$

$$\bar{M} = 2 \frac{M_1 M_2}{M_2 + M_1}.$$

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when $\lambda = 0$, we recover the mass matrix given by Kubo, A.Mondragón, M. Mondragón y E. Rodríguez-Jauregui Prog.Theor.Phys. 109 (2003) 795-807. In this case two of the right-handed neutrino masses are degenerate and θ_{13} is different from zero but very small.

The mass matrix with two texture zeros displaced

For simplify the analysis we will consider the case $\arg\{\mu_2^\nu\} = \arg\{\mu_3^\nu\}$.

$$\mathbf{M}_{\nu L} = \mathbf{Q} \mathcal{U}_{\frac{\pi}{4}} \left(\mu_0 \mathbf{I}_{3 \times 3} + \widehat{\mathbf{M}} \right) \mathcal{U}_{\frac{\pi}{4}}^\dagger \mathbf{Q},$$

where $\mathbf{Q} = e^{i\phi_2} \text{diag}\{1, 1, e^{i\delta_\nu}\}$, $\delta_\nu = \phi_4 - \phi_2 = \arg\{\mu_4^\nu\} - \arg\{\mu_2^\nu\}$,

$$\mathcal{U}_{\frac{\pi}{4}} = \begin{pmatrix} \frac{1}{\sqrt{2}} & 0 & \frac{1}{\sqrt{2}} \\ -\frac{1}{\sqrt{2}} & 0 & \frac{1}{\sqrt{2}} \\ 0 & 1 & 0 \end{pmatrix}, \quad \mu_0 = \frac{2|\mu_2^\nu|^2}{|\widehat{M}|} (1 - |\lambda|), \quad \widehat{\mathbf{M}} = \begin{pmatrix} 0 & A & 0 \\ A & B & C \\ 0 & C & 2d \end{pmatrix}$$

with

$$A = \sqrt{2} \frac{|\mu_2^\nu| |\mu_4^\nu|}{|\widehat{M}|} (1 - |\lambda|), \quad B = \frac{2|\mu_4^\nu|^2}{|\widehat{M}|} + \frac{|\mu_3^\nu|^2}{M_3} - \frac{2|\mu_2^\nu|^2}{|\widehat{M}|} (1 - |\lambda|),$$

$$C = \sqrt{2} \frac{|\mu_2^\nu| |\mu_4^\nu|}{|\widehat{M}|} (1 + |\lambda|) \quad \text{and} \quad d = \frac{2|\lambda| |\mu_2^\nu|^2}{|\widehat{M}|}.$$

The unitary matrix that diagonalizes the mass matrix of the neutrinos as function of its eigenvalues

$$\mathbf{U}_\nu = \mathbf{Q}^\nu \mathcal{U}_{\frac{\pi}{4}} \mathbf{O}_\nu^{N[I]}$$

where

$$\left(\begin{array}{c} \sqrt{\frac{[-1](m_{\nu 3} - \mu_0)(m_{\nu 2} - \mu_0) f_1}{\mathcal{D}_1^{N[I]}}} \\ \sqrt{\frac{[-1]2d(\mu_0 - m_{\nu 1}) f_1}{\mathcal{D}_1^{N[I]}}} \\ - \sqrt{\frac{[-1](\mu_0 - m_{\nu 1}) f_2 f_3^{N[I]}}{\mathcal{D}_1^{N[I]}}} \\ \sqrt{\frac{(m_{\nu 3[1]} - \mu_0)(\mu_0 - m_{\nu 1[3]}) f_2^{N[I]}}{\mathcal{D}_2^{N[I]}}} \\ \sqrt{\frac{2d(m_{\nu 2} - \mu_0) f_2^{N[I]}}{\mathcal{D}_2^{N[I]}}} \\ \sqrt{\frac{(m_{\nu 2} - \mu_0) f_1 f_3^{N[I]}}{\mathcal{D}_1^{N[I]}}} \\ - \sqrt{\frac{[-1](\mu_0 - m_{\nu 1})(m_{\nu 2} - \mu_0) f_3^{N[I]}}{\mathcal{D}_3^{N[I]}}} \\ \sqrt{\frac{[-1]2d(m_{\nu 3} - \mu_0) f_3^{N[I]}}{\mathcal{D}_3^{N[I]}}} \\ - \sqrt{\frac{(m_{\nu 3} - \mu_0) f_1 f_2^{N[I]}}{\mathcal{D}_1^{N[I]}}} \end{array} \right)$$

$$\mathcal{D}_1^{N[I]} = 2d(m_{\nu 2} - m_{\nu 1})(m_{\nu 3[1]} - m_{\nu 1[3]}), \quad \mathcal{D}_2^{N[I]} = 2d(m_{\nu 2} - m_{\nu 1})(m_{\nu 3[2]} - m_{\nu 2[3]}),$$

$$\mathcal{D}_3^{N[I]} = 2d(m_{\nu 3[1]} - m_{\nu 1[3]})(m_{\nu 3[2]} - m_{\nu 2[3]}), \quad f_1 = (2d + \mu_0 - m_{\nu 1}),$$

$$f_2^{N[I]} = [-1](2d + \mu_0 - m_{\nu 2}), \quad f_3^{N[I]} = [-1](m_{\nu 3} - \mu_0 - 2d),$$

$m_{\nu 2[1]} > \mu_0 > m_{\nu 1[3]}$ and $m_{\nu 3[2]} > 2d + \mu_0 > m_{\nu 2[1]}$. The superscripts N and I denote the normal and inverted hierarchies respectively.

The neutrino mixing matrix II

From a comparison of V_{PMNS}^{th} with V_{PMNS}^{PDG} , we obtain the neutrino mixing angles as function of the lepton masses

$$\sin^2 \theta'_{12} = \frac{|(V_{PMNS})_{12}|^2}{1 - |(V_{PMNS})_{13}|^2}, \quad \sin^2 \theta'_{23} = \frac{|(V_{PMNS})_{23}|^2}{1 - |(V_{PMNS})_{13}|^2}, \quad \sin^2 \theta'_{13} = |(V_{PMNS})_{13}|^2$$

$$\sin^2 \theta'_{12} = \frac{\left(\frac{\tilde{m}_e}{\tilde{m}_\mu}\right)^2 \left(O_{11}^{N[I]}\right)^2 + \left(O_{21}^{N[I]}\right)^2 - 2 \frac{\tilde{m}_e}{\tilde{m}_\mu} O_{11}^{N[I]} O_{21}^{N[I]} \cos \delta_I}{1 - \left(\frac{\tilde{m}_e}{\tilde{m}_\mu}\right)^2 \left(O_{23}^{N[I]}\right)^2 - \left(O_{33}^{N[I]}\right)^2 + 2 \frac{\tilde{m}_e}{\tilde{m}_\mu} O_{23}^{N[I]} O_{33}^{N[I]} \cos \delta_I}$$

$$\sin^2 \theta'_{23} = \frac{\left(O_{13}^{N[I]}\right)^2 + \left(\frac{\tilde{m}_e}{\tilde{m}_\mu}\right)^2 \left(O_{23}^{N[I]}\right)^2 + 2 \frac{\tilde{m}_e}{\tilde{m}_\mu} O_{13}^{N[I]} O_{23}^{N[I]} \cos \delta_I}{1 - \left(\frac{\tilde{m}_e}{\tilde{m}_\mu}\right)^2 \left(O_{23}^{N[I]}\right)^2 - \left(O_{33}^{N[I]}\right)^2 + 2 \frac{\tilde{m}_e}{\tilde{m}_\mu} O_{23}^{N[I]} O_{33}^{N[I]} \cos \delta_I}$$

$$\sin^2 \theta'_{13} = \left(\frac{\tilde{m}_e}{\tilde{m}_\mu}\right)^2 \left(O_{23}^{N[I]}\right)^2 + \left(O_{33}^{N[I]}\right)^2 - 2 \frac{\tilde{m}_e}{\tilde{m}_\mu} O_{23}^{N[I]} O_{33}^{N[I]} \cos \delta_I$$

The Reactor Mixing Angle

In a first, preliminary analysis for the reactor mixing angle θ'_{13} and for an normal neutrino mass hierarchy

$$m_{\nu_1} = 3.22 \times 10^{-3} \text{ eV}, m_{\nu_2} = 9.10 \times 10^{-3} \text{ eV}, m_{\nu_3} = 4.92 \times 10^{-2} \text{ eV}.$$

the parameter values $\delta_l = \pi/2$, $\mu_0 = 0.049 \text{ eV}$ and $d = 8 \times 10^{-5} \text{ eV}$, we get

$$\sin^2 \theta'_{13} \approx 0.029 \longrightarrow \theta'_{13} \approx 10.8^\circ,$$

in good agreement with experimental data.

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The solar and atmospheric mixing angles:

$$\theta_{12}^{lth} = 35^\circ, \quad \theta_{23}^{lth} = 46^\circ,$$

FCNC I

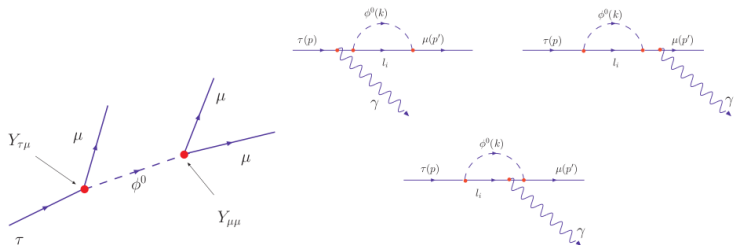
In the Standard Model the FCNC at tree level are suppressed by the GIM mechanism.

Models with more than one Higgs $SU(2)$ doublet have tree level FCNC due to the exchange of scalar fields.

The mass matrix written in terms of the Yukawa couplings is

$$\mathcal{M}_Y^e = Y_w^{E1} H_1^0 + Y_w^{E2} H_2^0,$$

FCNC processes: The left contributes to the process $\tau^- \rightarrow 3\mu$. The right contribute to the process $\tau \rightarrow \mu\gamma$.



Yukawa matrices in the mass representation²

The Yukawa matrices in the mass basis defined by

$$\tilde{Y}_m^{EI} = U_{eL}^\dagger Y_w^{EI} U_{eR}$$

and

$$\tilde{Y}_m^{E1} \approx \frac{m_\tau}{v_1} \begin{pmatrix} 2\tilde{m}_e & -\frac{1}{2}\tilde{m}_e & \frac{1}{2}x \\ -\tilde{m}_\mu & \frac{1}{2}\tilde{m}_\mu & -\frac{1}{2} \\ \frac{1}{2}\tilde{m}_\mu x^2 & -\frac{1}{2}\tilde{m}_\mu & \frac{1}{2} \end{pmatrix}_m, \quad \tilde{Y}_m^{E2} \approx \frac{m_\tau}{v_2} \begin{pmatrix} -\tilde{m}_e & \frac{1}{2}\tilde{m}_e & -\frac{1}{2}x \\ \tilde{m}_\mu & \frac{1}{2}\tilde{m}_\mu & \frac{1}{2} \\ -\frac{1}{2}\tilde{m}_\mu x^2 & \frac{1}{2}\tilde{m}_\mu & \frac{1}{2} \end{pmatrix}_m$$

$$x = m_e/m_\mu.$$

All off diagonal terms give rise to FCNC processes!!

²A. Mondragon *et al* Phys. Rev. D 76 076003

Branching ratios

We define the partial branching ratio (only leptonic decays)

$$Br(\tau \rightarrow \mu e^+ e^-) = \frac{\Gamma(\tau \rightarrow \mu e^+ e^-)}{\Gamma(\tau \rightarrow e \nu \bar{\nu}) + \Gamma(\tau \rightarrow \mu \nu \bar{\nu})}, \quad \Gamma(\tau \rightarrow \mu e^+ e^-) \approx \frac{m_\tau^5}{32^{10} \pi^3} \frac{(Y_{\tau\mu}^{1,2} Y_{ee'}^{1,2})^2}{M_{H_{1,2}}^4}$$

thus

$$Br(\tau \rightarrow \mu e^+ e^-) \approx \frac{9}{4} \left(\frac{m_e m_\mu}{m_\tau^2} \right)^2 \left(\frac{m_\tau}{M_{H_{1,2}}} \right)^4,$$

Similar computations lead to

$$Br(\tau \rightarrow e \gamma) \approx \frac{3\alpha}{8\pi} \left(\frac{m_\mu}{M_H} \right)^4, \quad Br(\tau \rightarrow \mu \gamma) \approx \frac{3\alpha}{128\pi} \left(\frac{m_\mu}{m_\tau} \right)^2 \left(\frac{m_\tau}{M_H} \right)^4,$$

$$Br(\tau \rightarrow 3\mu) \approx \frac{9}{64} \left(\frac{m_\mu}{M_H} \right)^4, \quad Br(\mu \rightarrow 3e) \approx 18 \left(\frac{m_e m_\mu}{m_\tau^2} \right)^2 \left(\frac{m_\tau}{M_H} \right)^4,$$

$$Br(\mu \rightarrow e \gamma) \approx \frac{27\alpha}{64\pi} \left(\frac{m_e}{m_\mu} \right)^4 \left(\frac{m_\tau}{M_H} \right)^4.$$

Leptonic processes via FCNC

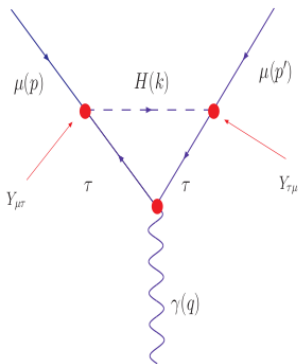
FCNC processes	Theoretical BR	Experimental upper bound BR	References
$\tau \rightarrow 3\mu$	8.43×10^{-14}	5.3×10^{-8}	B. Aubert <i>et al.</i> (2007)
$\tau \rightarrow \mu e^+ e^-$	3.15×10^{-17}	8×10^{-8}	B. Aubert <i>et al.</i> (2007)
$\tau \rightarrow \mu\gamma$	9.24×10^{-15}	6.8×10^{-8}	B. Aubert <i>et al.</i> (2005)
$\tau \rightarrow e\gamma$	5.22×10^{-16}	1.1×10^{-11}	B. Aubert <i>et al.</i> (2006)
$\mu \rightarrow 3e$	2.53×10^{-16}	1×10^{-12}	U. Bellgardt <i>et al.</i> (1998)
$\mu \rightarrow e\gamma$	2.42×10^{-20}	1.2×10^{-11}	M. L. Brooks <i>et al.</i> (1999)

Small FCNC processes mediating non-standard quark-neutrino interactions could be important in the theoretical description of the **gravitational core collapse and shock generation** in the explosion stage of a supernova

Muon Anomalous Magnetic Moment

The anomalous magnetic moment of the muon is related to the gyroscopic ratio by

$$a_\mu = \frac{\mu_\mu}{\mu_B} - 1 = \frac{1}{2}(g_\mu - 2)$$



In models with more than one Higgs $SU(2)$ doublet, the exchange of flavour changing neutral scalars also contribute to the anomalous magnetic moment of the muon

$$\delta a_\mu^{(H)} = \frac{Y_{\mu\tau} Y_{\tau\mu}}{16\pi^2} \frac{m_\mu m_\tau}{M_H^2} \left(\log \left(\frac{M_H^2}{m_\tau^2} \right) - \frac{3}{2} \right)$$

From our results: $Y_{\mu\tau} Y_{\tau\mu} = \frac{m_\mu m_\tau}{4v_1 v_2}$

$$\delta a_\mu^{(H)} = \frac{m_\tau^2}{(246 \text{ GeV})^2} \frac{(2 + \tan^2 \beta)}{32\pi^2} \frac{m_\mu^2}{M_H^2} \left(\log \left(\frac{M_H^2}{m_\tau^2} \right) - \frac{3}{2} \right),$$

$$\tan \beta = \frac{v_2}{v_1}$$

From the experimental upper bound on $(\mu \rightarrow 3e)$, we get $\tan \beta \leq 14$,
Hence

$$\delta a_\mu = 1.7 \times 10^{-10}$$

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Hence

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Contribution to the anomaly of the muon's magnetic moment

The difference between the experimental value and the Standard Model prediction for the anomaly is

$$\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (28.79.1)10^{-10}$$

$\Delta a_\mu \sim 3\sigma$ (three standard deviations) !!

But, the uncertainty in the computation of higher order hadronic effects is large

$$\delta a_\mu^{\text{LBL}}(3, \text{had}) \approx 1.5910^{-9}; \quad \delta a_\mu^{\text{VP}}(3, \text{had}) \approx -1.8210^{-9}$$

$$\frac{\delta a_\mu^{(H)}}{\Delta a_\mu} \approx \frac{1.7}{28} \approx 6\% \quad \text{and} \quad \delta a_\mu^{(H)} < \delta a_\mu(3, \text{had})$$

The contribution of the exchange of flavour changing scalars to the anomaly of the muon's magnetic moment, $\delta a_\mu^{(H)}$, is small but not negligible, and it is compatible with the best, state of the art, measurements and theoretical predictions.

Summary

- By introducing three $SU(2)_L$ Higgs doublet fields, in the theory, we extended the concept of flavour and generations to the Higgs sector and formulated a minimal S_3 -invariant Extension of the SM
- The neutrino mixing angles θ_{12} , θ_{23} and θ_{13} , are determined by an interplay of the S_3Z_2 symmetry, the see-saw mechanism and the lepton mass hierarchy
- The fit of, $\sin^2 \theta_{13}^{th}$ to $\sin^2 \theta_{13}^{exp}$ breaks the mass degeneracy of the right handed neutrinos.
- The branching ratios of all flavour changing neutral processes in the leptonic sector are strongly suppressed by the S_3Z_2 symmetry and powers of the small mass ratios m_e/m_τ , m_μ/m_τ , and $(m_\tau/M_{H_{1,2}})^4$, but could be important in astrophysical processes
- The anomalous magnetic moment of the muon gets a small but non-negligible contribution from the exchange of flavor changing scalar fields