

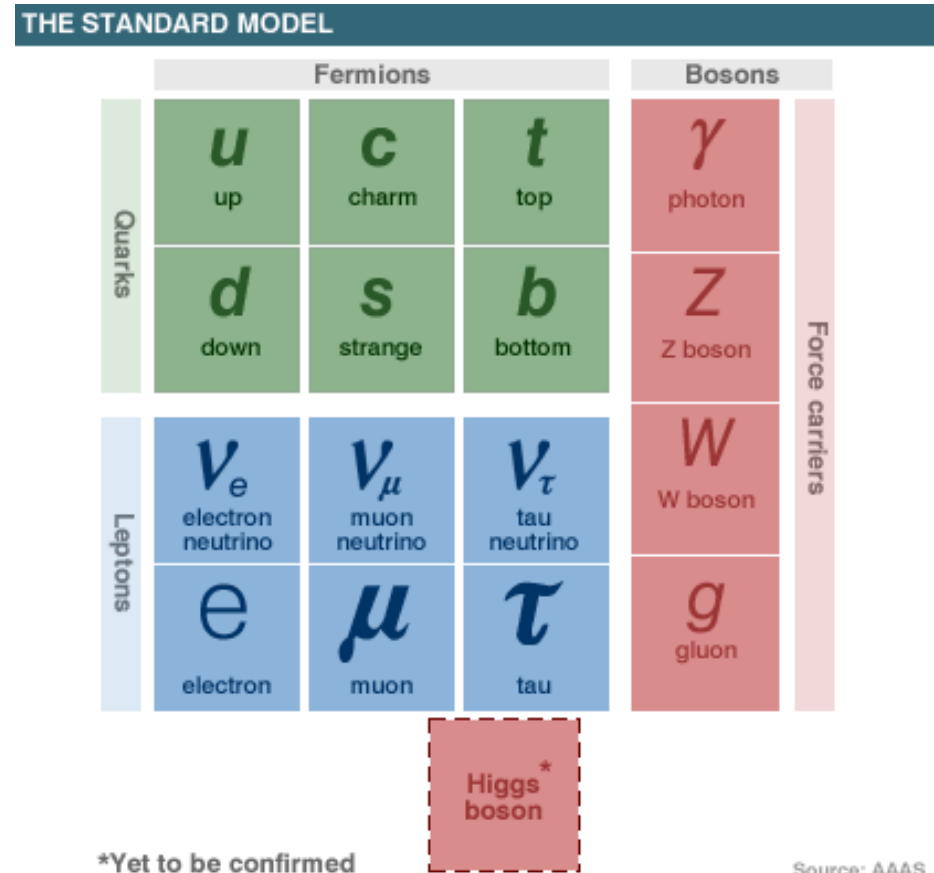
Searching for axions and ALPs from string theory.

Andreas Ringwald (DESY)

PASCOS 2012,
Merida, Mexico, 3-8 June 2012

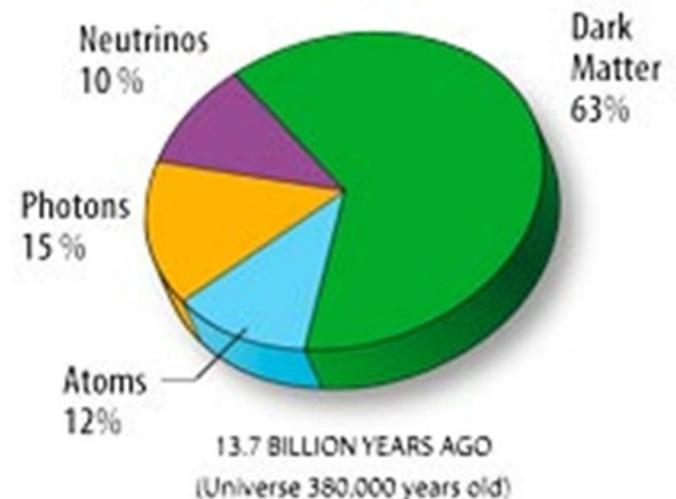
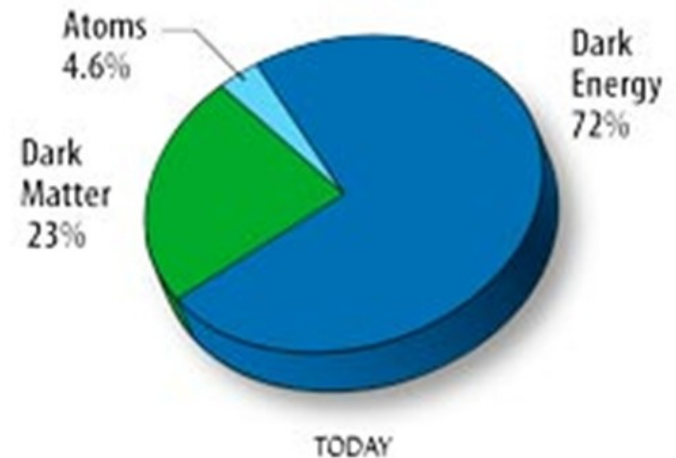
Particles beyond the Standard Model?!

- Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces



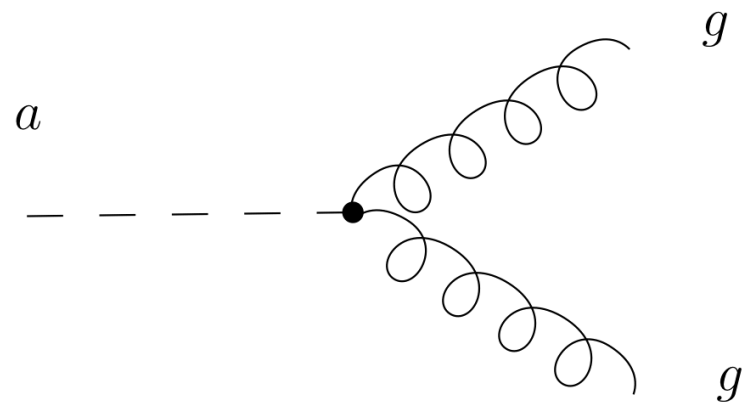
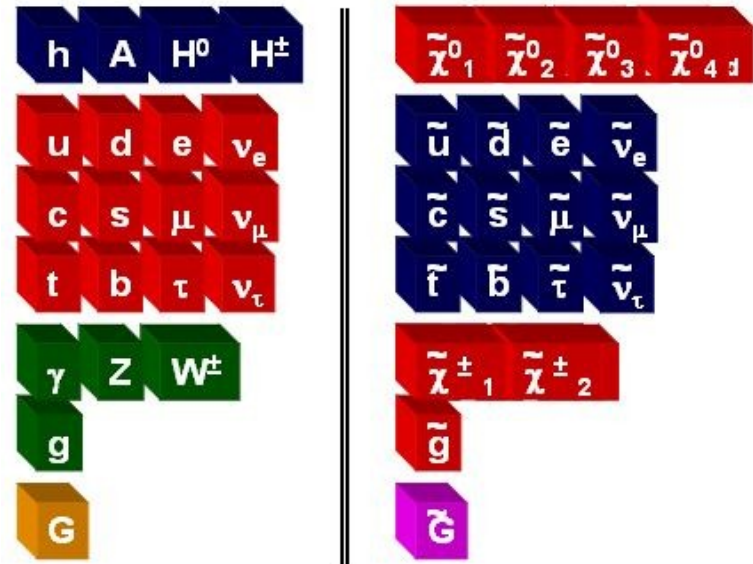
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- Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces
- Standard Model of Cosmology: only about 5 % of energy content of present universe consists of known particles



Particles beyond the Standard Model?!

- > Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces
- > Standard Model of Cosmology: only about 5 % of energy content of present universe consists of known particles
- > Extensions of the Standard Model of Particle Physics: several good motivated candidates for constituents of dark matter
 - SUSY: Neutralino, Gravitino
 - Peccei Quinn: Axion



Peccei-Quinn extension of the Standard Model

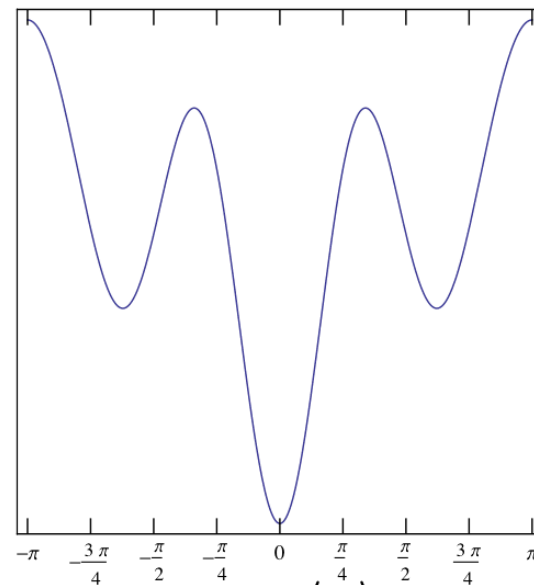
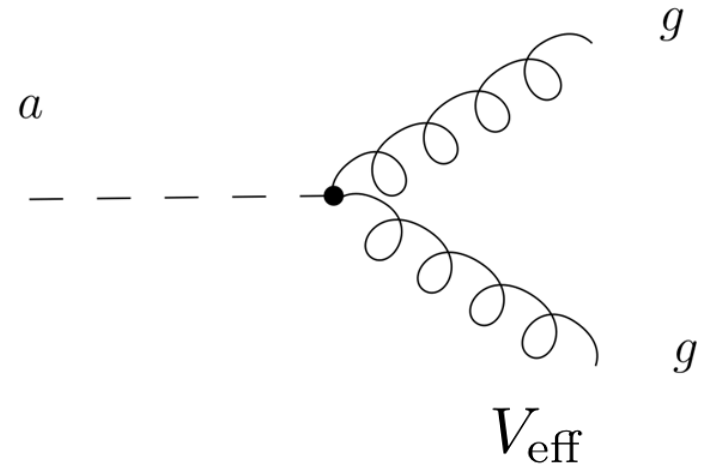
> Motivation:

Explanation of unnatural smallness, $\theta < 10^{-10}$, of CP-violating topological term in QCD Lagrangian.

$$\mathcal{L}_{\text{CP-viol.}} = \frac{\alpha_s}{4\pi} \theta \text{tr} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

> Axion field $\theta \rightarrow a(x)/f_a$

- $\langle a \rangle = 0$



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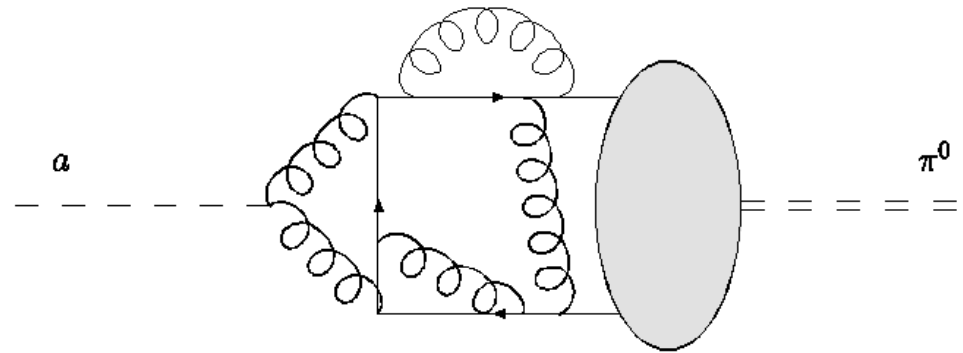
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- Axion: ultralight particle, cf. Weinberg '78; Wilczek '78

$$m_a = \frac{m_\pi f_\pi}{f_a} \frac{\sqrt{m_u m_d}}{m_u + m_d} \simeq 6 \text{ meV} \times \left(\frac{10^9 \text{ GeV}}{f_a} \right)$$



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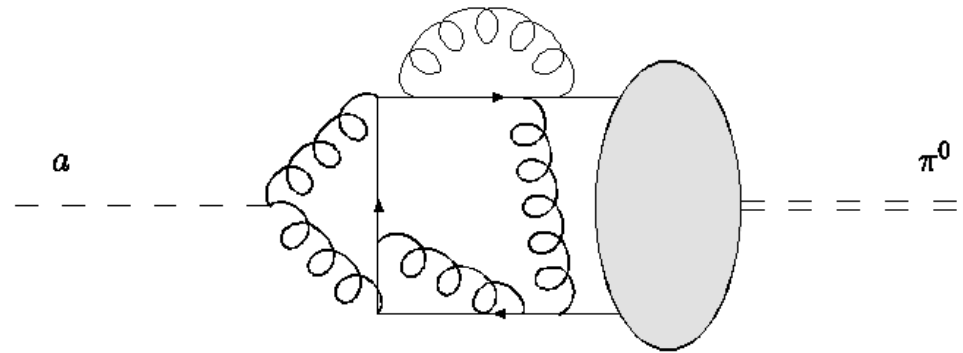
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- extremely weak interactions with Standard Model particles

$$\mathcal{L}_{a\gamma\gamma} = -\frac{1}{4} g_{a\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu} = g_{a\gamma} a \vec{E} \cdot \vec{B},$$

$$g_{a\gamma} \simeq \frac{\alpha}{2\pi f_a} \sim 10^{-12} \text{ GeV}^{-1} \left(\frac{10^9 \text{ GeV}}{f_a} \right)$$



Peccei-Quinn extension of the Standard Model

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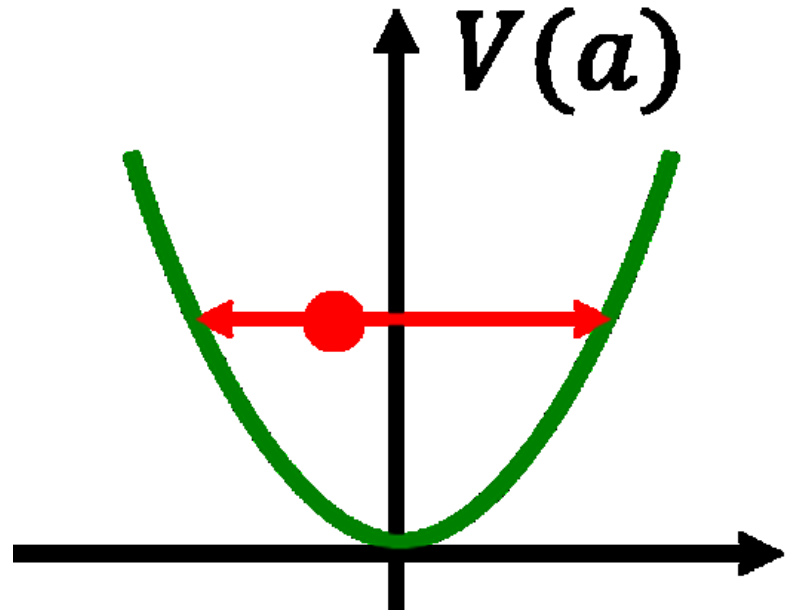
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> Welcome side effect:

- Axion: candidate for dark matter: created non-thermally via misalignment mechanism in form of coherent oscillations of axion field

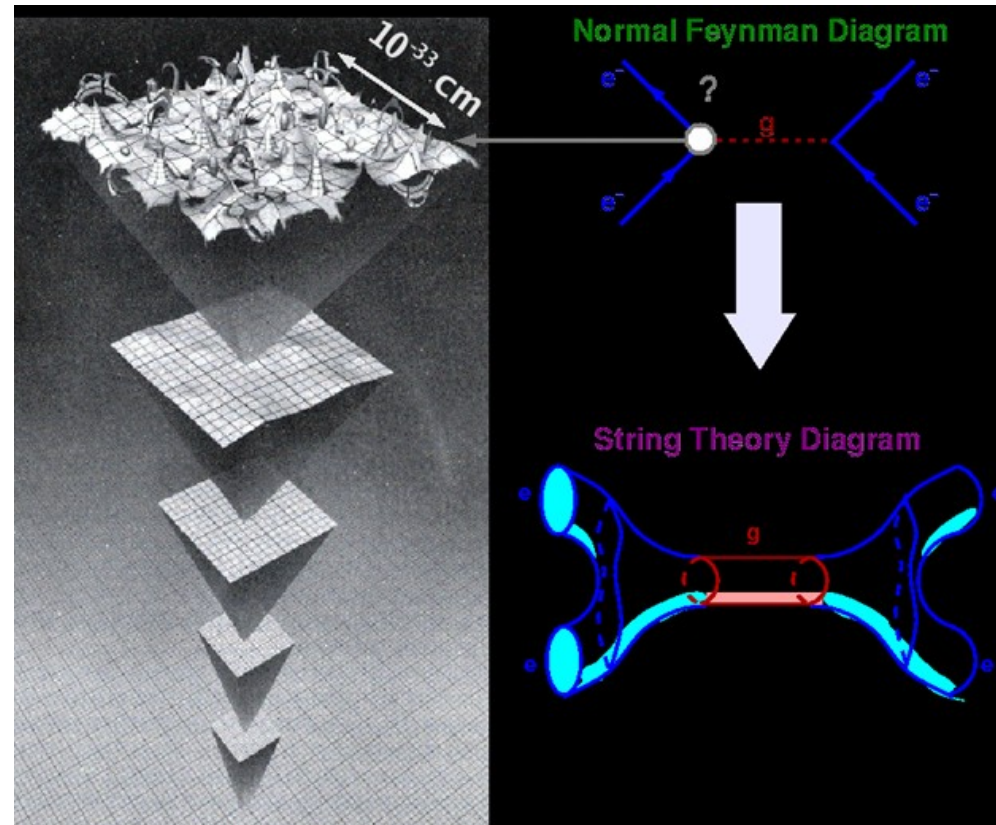


$$\Omega_a h^2 \approx 0.7 \left(\frac{f_a}{10^{12} \text{ GeV}} \right)^{7/6} \left(\frac{\theta_i}{\pi} \right)^2$$

Coexistence: SUSY and PQ extension in string theory

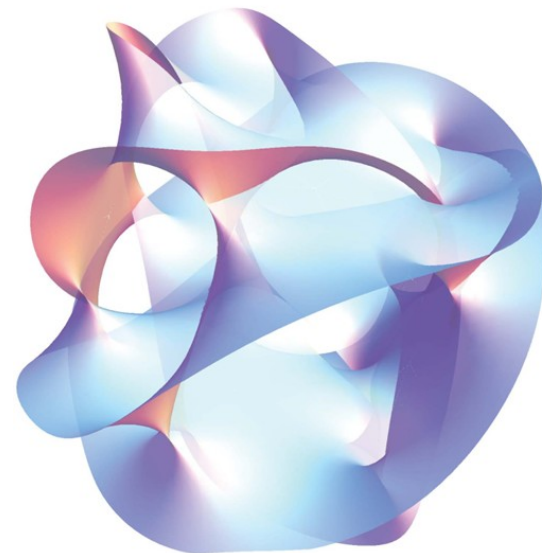
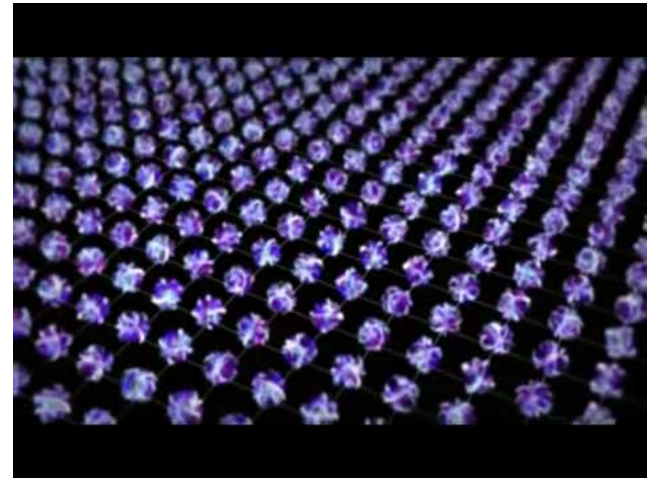
➤ Particularly strongly motivated extensions of Standard Model based on string theory:

- Unification of all forces
- Quantum gravity



Coexistence: SUSY and PQ extension in string theory

- Particularly strongly motivated extensions of Standard Model based on string theory:
 - Unification of all forces
 - Quantum gravity
- Spectrum of low-energy effective theory in (3+1)-dimensions is supersymmetric and possibly contains several kinds of very weakly interacting slim particles (WISPs): **Axion, ALPs (Axion-Like Particles)**
 - if the compact space comprised of the 6 extra dimensions has certain geometrical and topological properties (Calabi-Yau; several cycles)



Axions in string theory

- > String compactifications generically contain pseudo-scalar fields with axionic coupling to gauge fields and anomalous global shift symmetry

$$a_i F \tilde{F} \quad a_i \rightarrow a_i + \epsilon$$

- > These axion and axion-like particle (ALP) candidates arise in string compactifications as KK zero modes of antisymmetric tensor fields:
cf. [Witten '84](#)

heterotic string : B_2

IIB string : C_2, C_4



Axions in IIB string theory

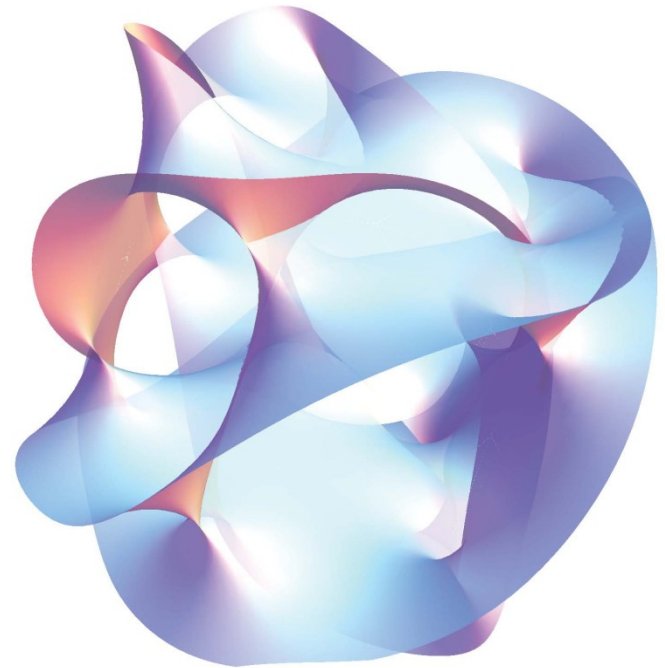
- Concentrate on IIB case (moduli stabilisation best understood):
Realisation of brane-world scenarios in string theory

- KK reduction (expansion in harmonic forms):

$$C_2 = c^a(x)\omega_a, \quad a = 1, \dots, h_-^{1,1}$$

$$C_4 = c_\alpha(x)\tilde{\omega}^\alpha + \dots, \quad \alpha = 1, \dots, h_+^{1,1}$$

- Number of axionic fields determined by topology of CY orientifold: number of topologically non-equivalent 2-cycles or 4-cycles



Axions in IIB string theory

➤ Number of cycles generically $O(100)$:

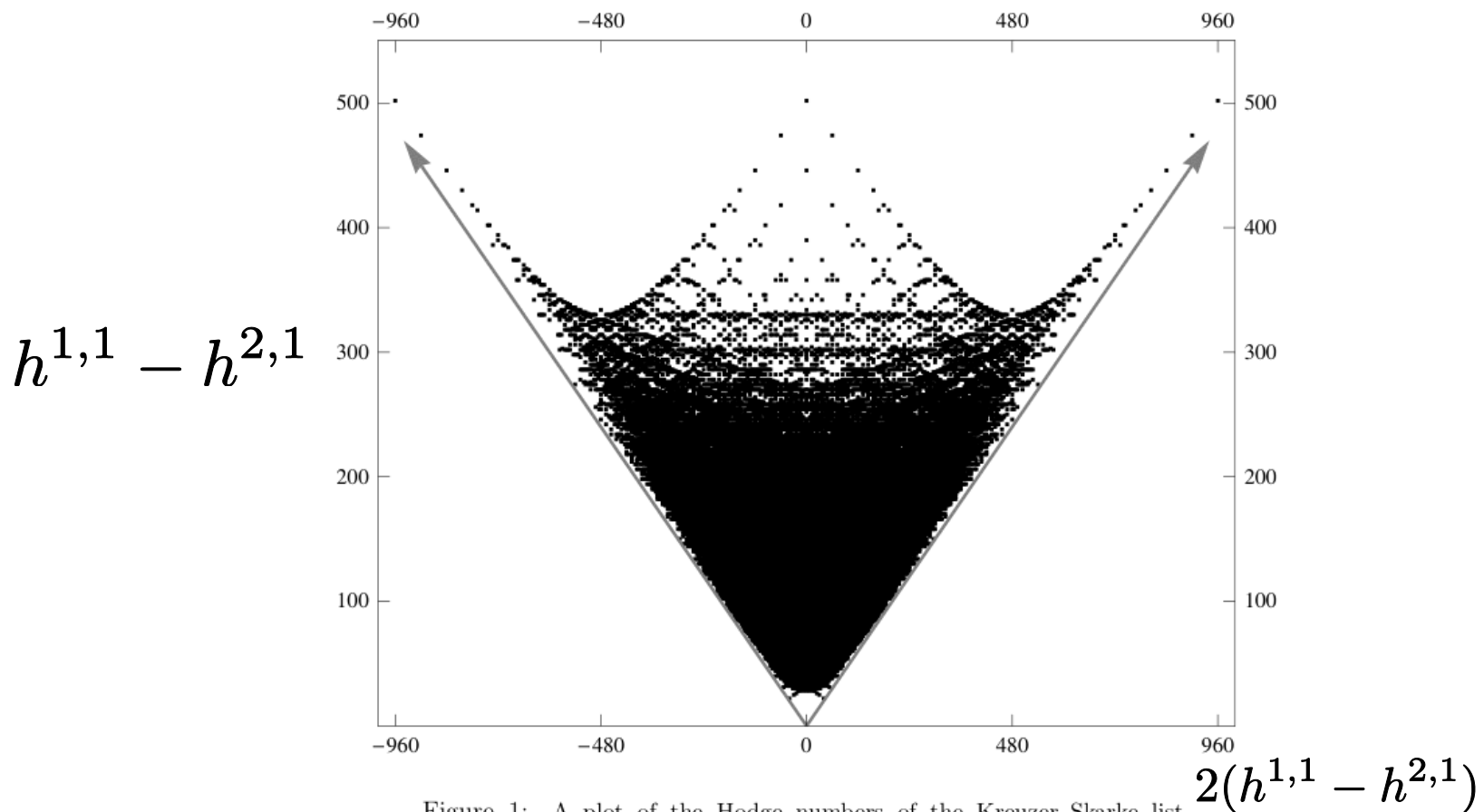
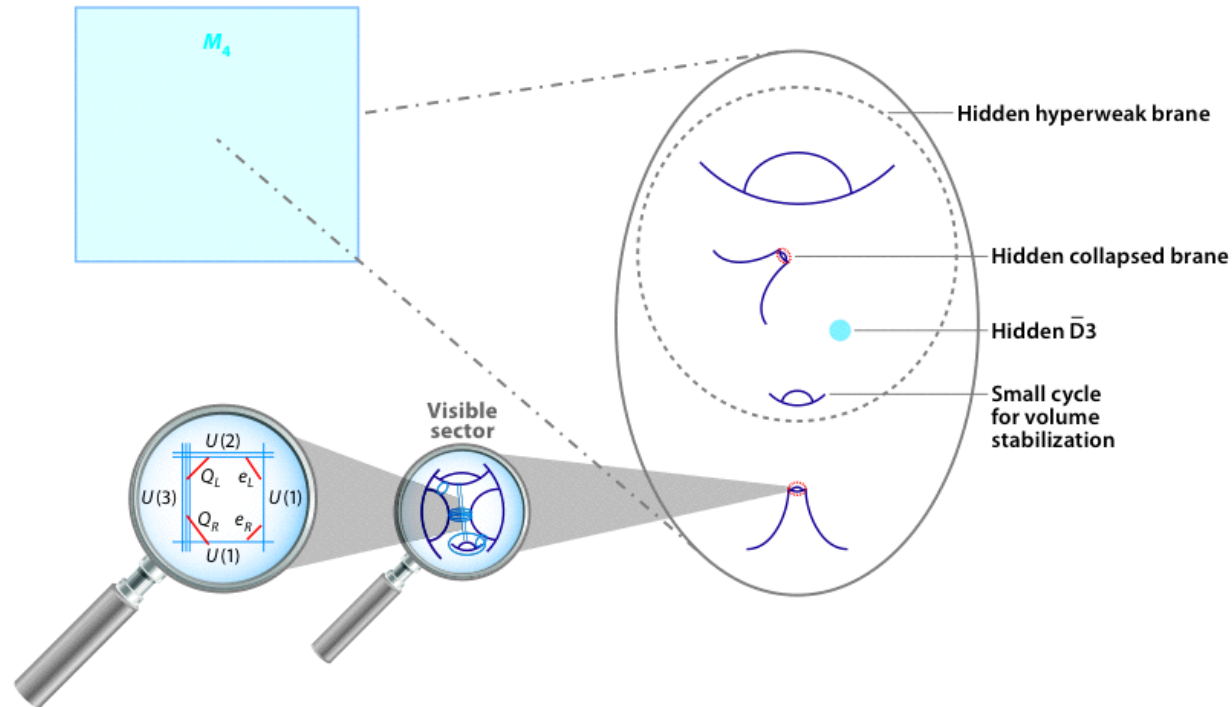


Figure 1: A plot of the Hodge numbers of the Kreuzer-Skarke list. $\chi = 2(h^{1,1} - h^{2,1})$ is plotted horizontally and $h^{1,1} + h^{2,1}$ is plotted vertically. The oblique axes bound the region $h^{1,1} \geq 0, h^{2,1} \geq 0$.



Axions in IIB string theory

- Cycles can be wrapped by space-time filling D-branes



- Each of these branes gives rise to a gauge theory at low energy

- Visible sector gauge theory realized by stacks of D7 branes wrapping small 4-cycles
- Hidden sectors, in particular hidden photons, realized by branes wrapping cycles not intersecting visible sector branes

Axions in IIB string theory

- Each axionic field comes along with a real scalar field – saxion – which is real part of lowest component of chiral superfield,

$$T_\alpha = \tau_\alpha + i c_\alpha$$

- τ_α ... Kähler modulus measuring the volume of 4-cycle α
- 4D EFT from KK reduction of D-brane action

$$S_p = \frac{-2\pi}{(2\pi\sqrt{\alpha'})^{p+1}} \left(\int_\Sigma d^{p+1}\xi e^{-\phi} \sqrt{\det(g + B + 2\pi\alpha'F)} + i \int_\Sigma e^{B+2\pi\alpha'F} \wedge \sum_q C_q \right)$$

- T_α is gauge kinetic function for theory on D7-brane:
 - volume measures gauge coupling, $\tau_\alpha \sim g^{-2}$
 - c_α has axionic coupling, $\sim c_\alpha F \wedge F$



Axions in IIB string theory

- 4D effective field theory, cf. [Jockers, Louis '05](#)

$$S \supset -dc_\alpha \frac{\mathcal{K}_{\alpha\beta}}{8} \wedge \star dc_\beta - \frac{r^{i\alpha} \tau_\alpha}{4\pi M_P} (F_i \wedge \star F_i) + \frac{r^{i\alpha} c_\alpha}{4\pi M_P} (F \wedge F),$$

$$\text{with } \mathcal{K}_{\alpha\beta} \equiv \frac{\partial^2 K}{\partial \tau_\alpha \partial \tau_\beta}, K = -2 \ln \mathcal{V}, r^{i\alpha} \equiv \ell_s^{-4} \int_{D_i} \tilde{\omega}^\alpha$$

- Decay constants and coupling to gauge bosons via canonical normalization of axion and gauge kinetic terms and matching to:

$$\begin{aligned} \mathcal{L} \supset & \frac{1}{2} \partial_\mu a_i \partial^\mu a_i - \frac{g_3^2}{32\pi^2} \left(\theta_0 + C_{i33} \frac{a_i}{f_{a_i}} \right) F_{3,\mu\nu}^b \tilde{F}_3^{b,\mu\nu} \\ & - \frac{g_2^2}{32\pi^2} C_{iWW} \frac{a_i}{f_{a_i}} F_{W,\mu\nu}^b \tilde{F}_W^{b,\mu\nu} - \frac{g_Y^2}{32\pi^2} C_{iYY} \frac{a_i}{f_{a_i}} F_{Y,\mu\nu} \tilde{F}_Y^{\mu\nu} \end{aligned}$$



- > An axiverse - QCD axion plus possibly many ultra-light ALPs whose mass spectrum is logarithmically hierarchical – may naturally arise from strings, cf. [Arvanitaki et al. `09](#)
- > Challenges to obtain an axiverse:
 - Only axions which are not projected out by orientifold projection appear in LE EFT
 - Only axions which do not get too heavy by Kähler moduli stabilisation can be candidates for QCD axion and other light ALPs, cf. [Conlon `06](#)
 - Only axions which are not eaten by Stückelberg mechanism to give masses to brane-localized anomalous U(1) gauge bosons will appear in LE EFT
- > [Acharya, Bobkov, Kumar `11](#): Moduli stabilisation via single non-perturbative correction to superpotential (cf. [Bobkov, Braun, Kumar, Raby `10](#)) fixes all Kähler moduli plus one axion combination: axiverse with $W_0 \ll 1$ and $f_a \sim 10^{16}$ GeV
- > [Cicoli, Goodsell, AR, 1206.0819](#): Moduli stabilisation of the so-called LARGE Volume Scenario (LVS) which exploits both perturbative and non-perturbative effects to fix Kähler moduli gives rise to an axiverse with $W_0 \sim 1$ and $f_a \sim 10^{10}$ GeV



> Strategy to fix moduli in LVS:
(cf. [Cicoli, Mayrhofer, Valandro '11](#))

- Exploit del Pezzo four-cycle supporting single non-perturbative effect; dP modulus fixed at small size, dP saxion and axion heavy; interplay with leading order alpha' correction yields exponentially large CY volume

- Visible sector with chirality built by wrapping magnetised D7-branes around rigid but not del Pezzo four cycles; D-term conditions stabilise d combinations of Kähler moduli, leaving

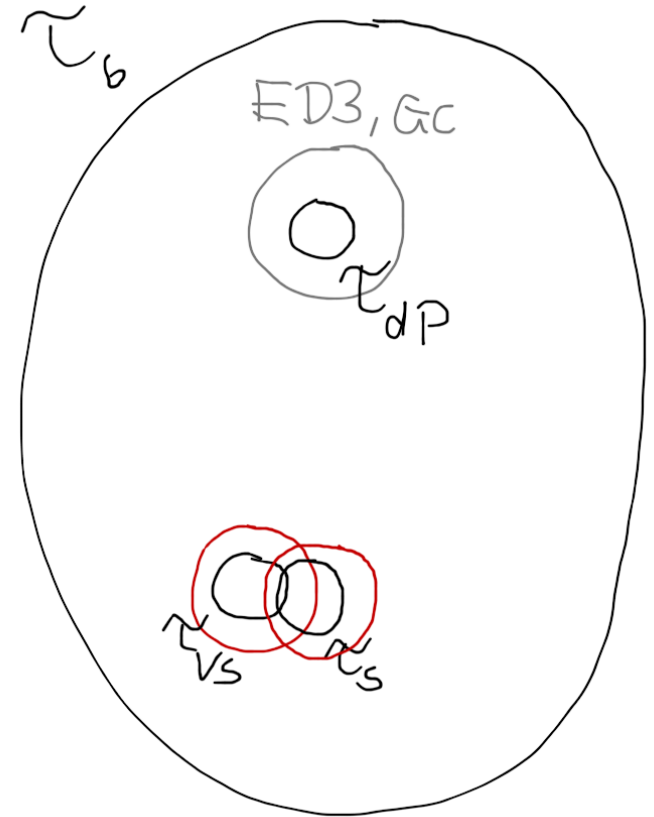
$$n_{\text{ax}} \equiv h^{1,1} - 1 - d \geq 1$$

flat directions; latter fixed by pert. corrections

- LVS requires $n_{\text{ax}} \geq 2$: one of the remaining cycles should be small to obtain correct value of $g_{\text{VS}}^2 \sim 1/\tau_{\text{VS}}$, while there should be at least one further which can be large; latter fixed at

$$\mathcal{V} \sim \tau_b^{3/2} \sim W_0 e^{2\pi\tau_{\text{dP}}}$$

> More on this: talk by [Michele Cicoli](#)



LVS axiverse

> Mass scales for $g_s \sim 0.1, W_0 \sim 1, \mathcal{V} \sim 10^{14}$:

$$M_s \sim \frac{M_P}{\sqrt{4\pi\mathcal{V}}} \sim 10^{10} \text{ GeV}$$

$$m_{\tau_s} \sim \frac{M_P}{\mathcal{V}^{1/2}} \sim 10^{10} \text{ GeV}$$

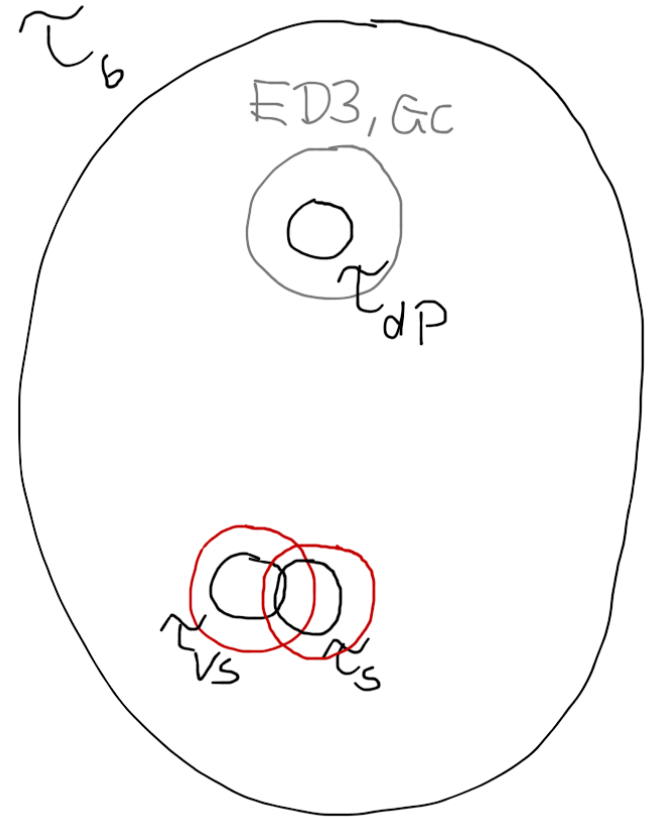
$$m_{\tau_{dP}} \sim \frac{M_P}{\mathcal{V}} \ln \mathcal{V} \sim 30 \text{ TeV}$$

$$m_{3/2} \sim \sqrt{g_s/(4\pi)W_0} \frac{M_P}{\mathcal{V}} \sim 1 \text{ TeV}$$

$$m_{\tau_{vs}} \sim \alpha_{vs} m_{3/2} \sim 40 \text{ GeV}$$

$$m_{\tau_b} \sim \frac{M_P}{\mathcal{V}^{3/2}} \sim 0.1 \text{ MeV}$$

- No cosmological moduli problem since τ_b diluted by entropy production due to decay of τ_{vs} reheating universe to $O(\text{GeV})$



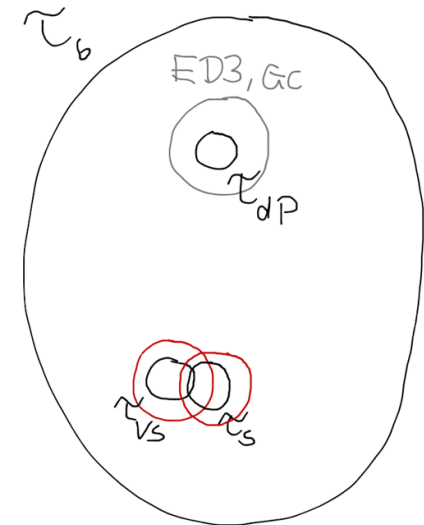
➤ Scaling of axion decay constants and couplings:

$$f_{a_b} = \frac{\sqrt{3}}{4\pi} \frac{M_P}{\tau_b} \simeq \frac{M_P}{4\pi \mathcal{V}^{2/3}} \simeq \frac{M_{\text{KK}}^{10D}}{4\pi}, \quad f_{a_s} = \frac{1}{\sqrt{6} (2\tau_s)^{1/4}} \frac{M_P}{4\pi \sqrt{\mathcal{V}}} \simeq \frac{M_s}{\sqrt{4\pi \tau_s}^{1/4}},$$

$$C_{bbb} \simeq g_b^{-2} \frac{f_{a_b}}{M_P} \simeq \mathcal{O}(1), \quad C_{sbb} \simeq g_b^{-2} \frac{f_{a_s} \tau_s^{3/4}}{\mathcal{V}^{1/2} M_P} \simeq \mathcal{O}(\epsilon) \simeq \mathcal{O}(\mathcal{V}^{-1/3}),$$

$$C_{bss} \simeq g_s^{-2} \frac{f_{a_b}}{M_P} \simeq \mathcal{O}(\epsilon^2) \simeq \mathcal{O}(\mathcal{V}^{-2/3}), \quad C_{sss} \simeq g_s^{-2} \frac{f_{a_s}}{\tau_s^{3/4} M_s} \simeq \mathcal{O}(1).$$

- \mathbf{a}_{VS} : QCD axion with $f_{a_{\text{VS}}} \sim f_{a_s} \sim 10^{10}$ GeV
- \mathbf{a}_b : essentially massless ALP with $f_{a_b} \sim 10^8$ GeV, but nearly decoupled, $C_{bss} \sim 10^{-10}$



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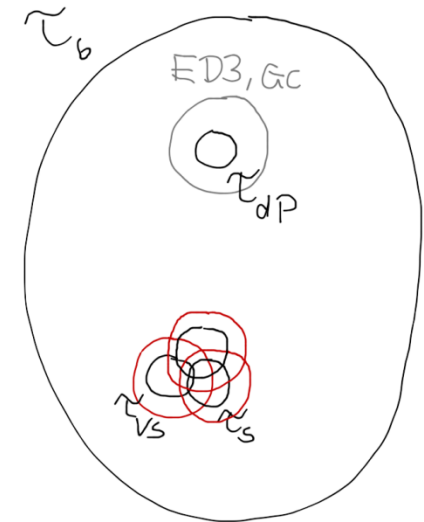
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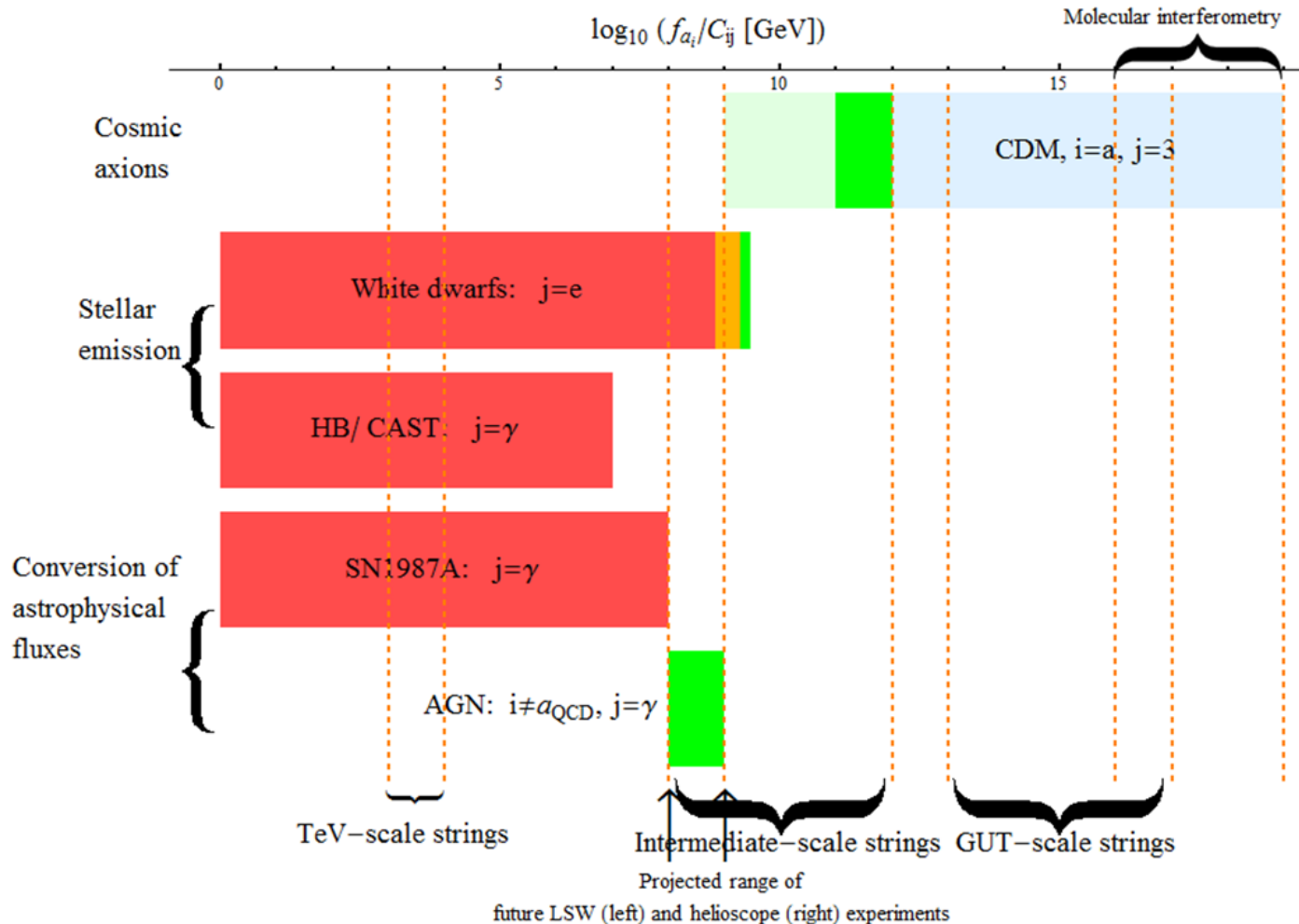
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- Possibly more ultralight, $m_{a_b} \ll m_a$, ALPs with

$$f_{a_s} \sim 10^{10} \text{ GeV}, C_{sss} \sim 1$$



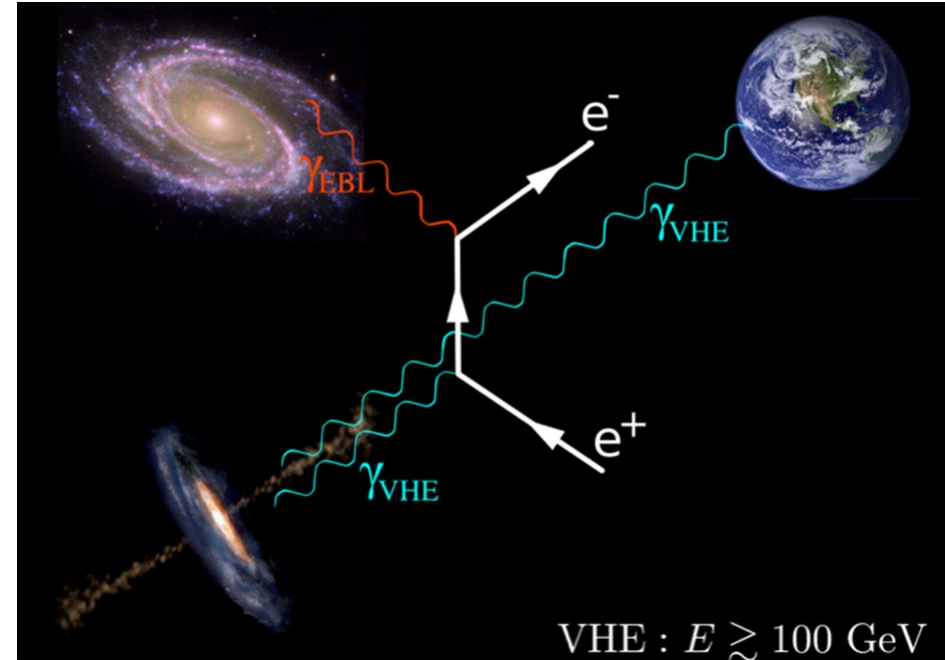
Axion and ALPs with intermediate scale decay constant?

- Current limits and possible hints from astrophysics and cosmology:



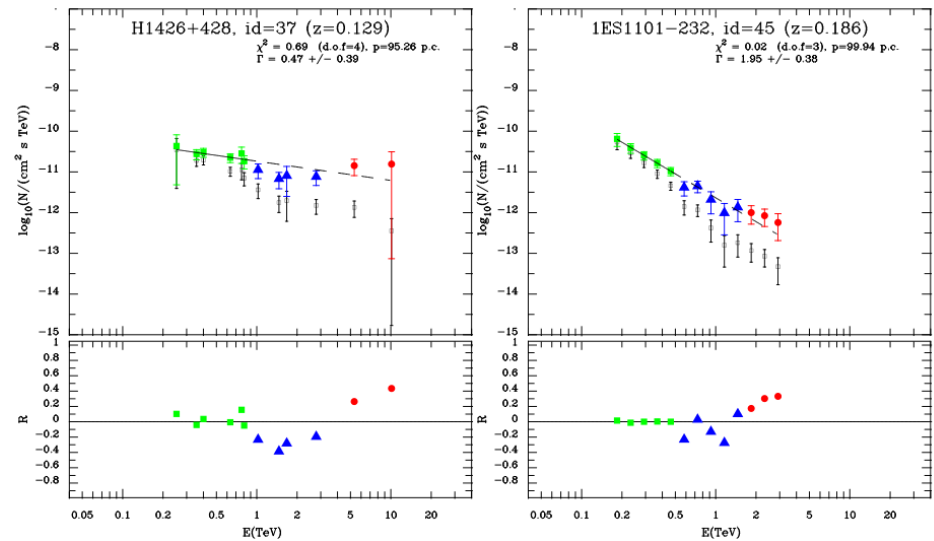
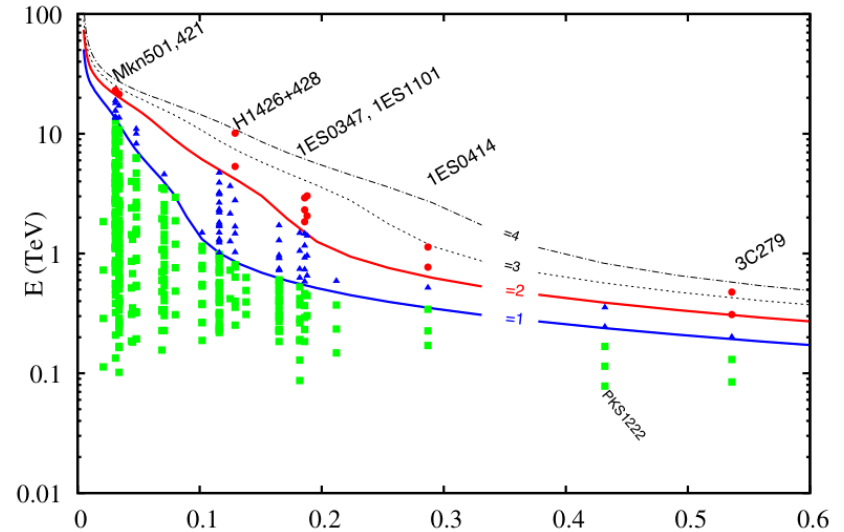
Hints for axion or ALP effects in propagation of TeV photons through universe?

- TeV photons of distant Active Galactic Nuclei (AGN) should feature absorption breaks due to electron-positron pairproduction on the Extragalactic Background Light (EBL)



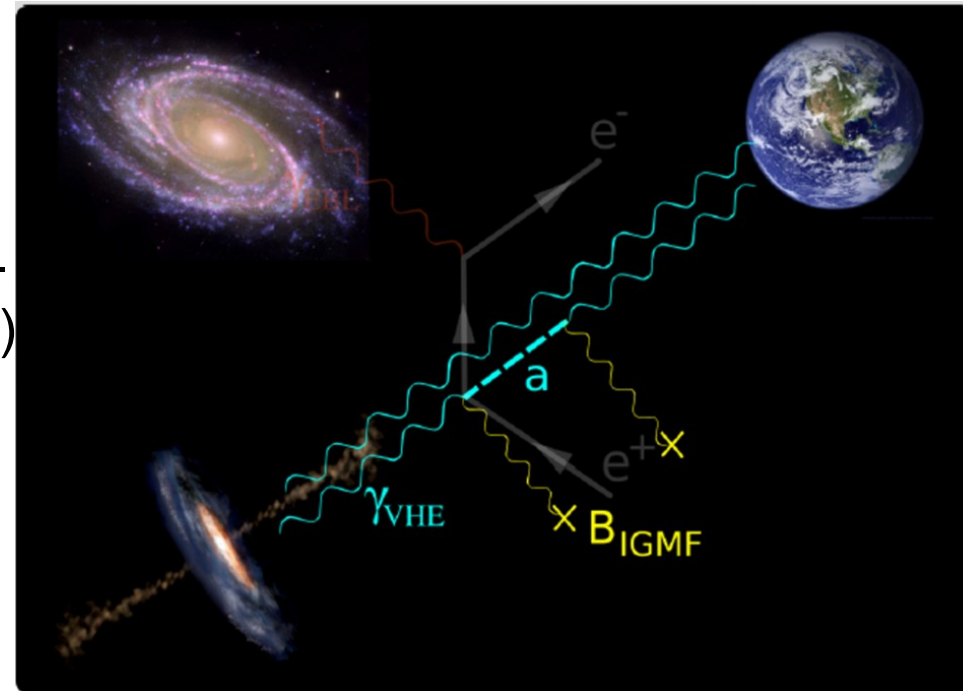
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 - 50 spectra (HESS, MAGIC, Veritas), assumption: minimal EBL; absorption ruled out by more than 4 sigma



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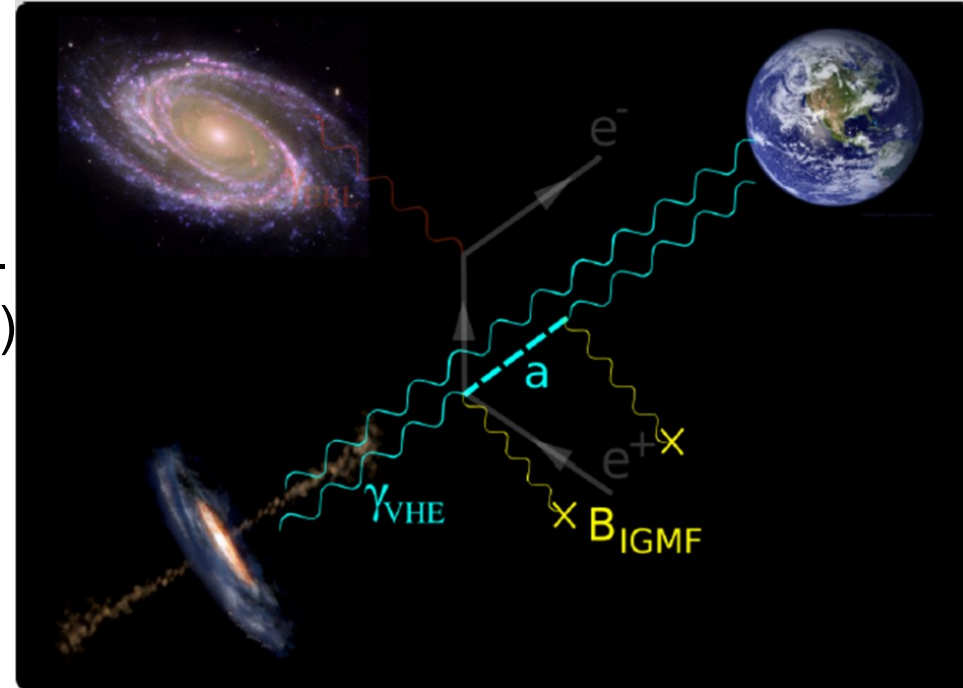
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$$P(a \leftrightarrow \gamma) = 4 \frac{(g_{a\gamma} \omega B)^2}{m_a^4} \sin^2 \left(\frac{m_a^2}{4\omega} L_B \right)$$

Hints for axion or ALP effects in propagation of TeV photons through universe?

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- Possible explanation: Photon \leftrightarrow ALP conversion in astrophysical magnetic fields ([Roncadelli et al.](#), [Sanchez-Conde et al.](#))
- Has to be an ALP: too light for a QCD axion with such a decay constant



$$g_{i\gamma} \equiv \frac{\alpha C_{i\gamma}}{2\pi f_{a_i}} = 10^{-12} \div 10^{-11} \text{ GeV}^{-1}$$
$$\rightarrow \frac{f_{a_i}}{C_{i\gamma}} \simeq 10^8 \div 10^9 \text{ GeV}$$

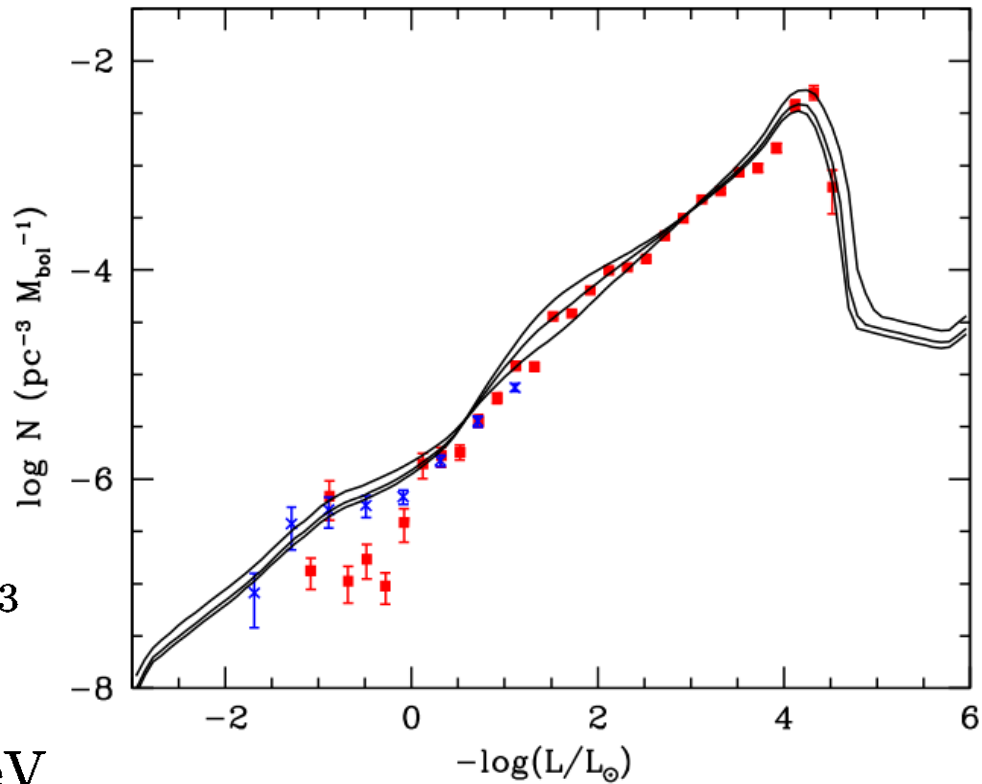
for $m_{a_i} \lesssim \text{neV}$

Hint for axion/ALP-production in white dwarfs?

- Non-standard energie-loss mechanism in white dwarfs, cf. [Isern et al.](#)
- Compatible with axion or ALP production in electron-bremsstrahlung

$$g_{ie} \equiv \frac{C_{ie} m_e}{f_{a_i}} = (2.0 \div 7.0) \times 10^{-13}$$
$$\rightarrow \frac{f_{a_i}}{C_{i,e}} \simeq (0.7 \div 2.6) \times 10^9 \text{ GeV}$$

for $m_{a_i} \lesssim \text{keV}$



Can we explain these hints within IIB axiverse?

- Anomalous transparency of universe and anomalous energy loss of white dwarfs could be explained by

$$C_{i\gamma}/C_{ie} \simeq 10, \quad f_{a_i}/C_{i\gamma} \simeq 10^8 \text{ GeV}, \quad m_{\text{ALP}} \lesssim 10^{-9} \div 10^{-10} \text{ eV}.$$

- Model where visible sector build from intersecting branes in geometric regime

$$\frac{C_{i\gamma}}{C_{ie}} \sim \frac{8\pi\tau_*}{3}, \quad \frac{f_{a_i}}{C_{i\gamma}} = \frac{1}{8\pi N_{i\gamma}\tau_*^{1/4}} \frac{M_P}{\sqrt{\mathcal{V}}} = \frac{1}{8\pi N_{i\gamma}\tau_*^{1/4}} \sqrt{\frac{g_s M_P m_{3/2}}{W_0}}.$$

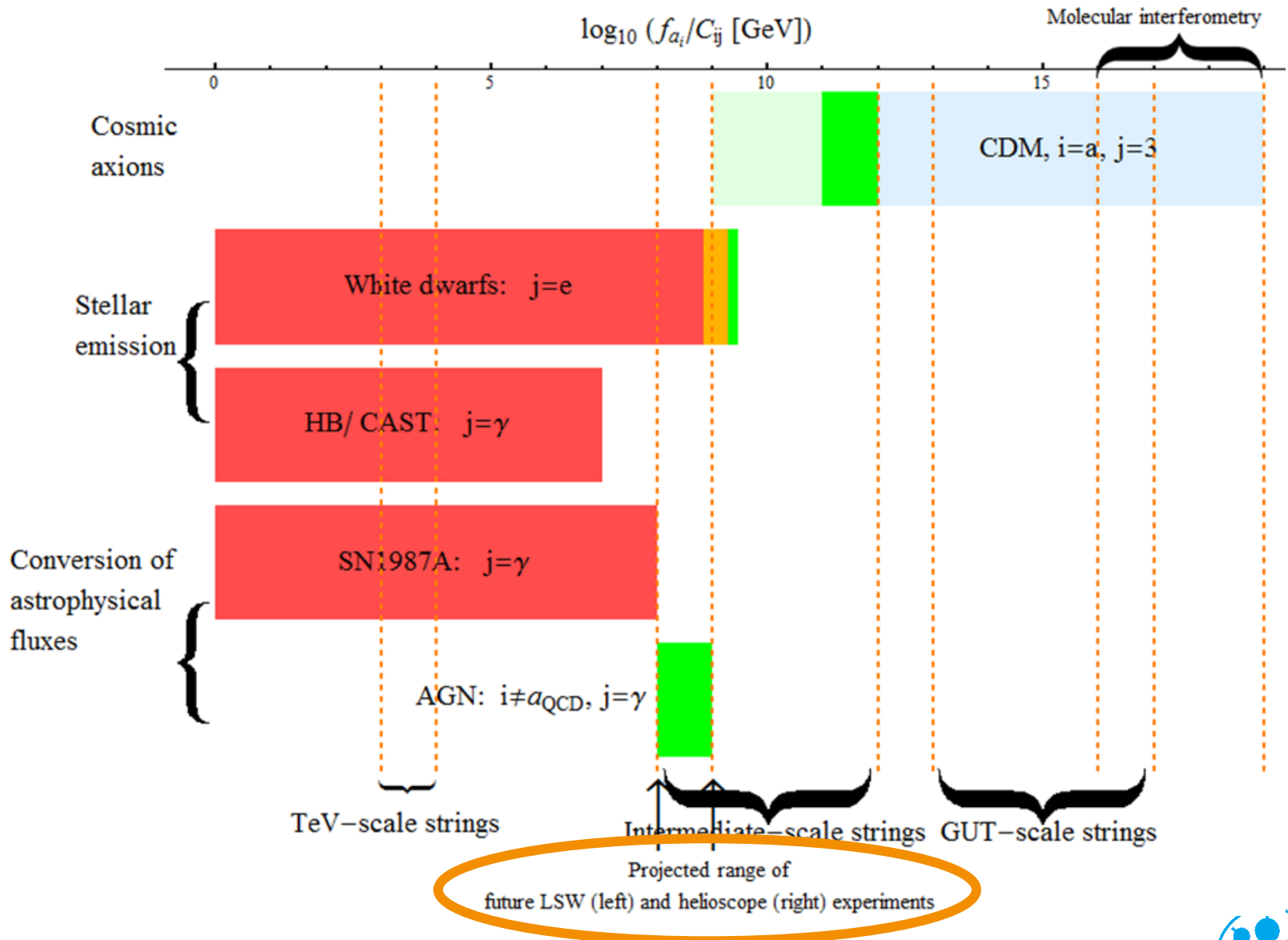
- Yields required values for $m_{3/2} = 10 \text{ TeV}$, $g_s \simeq 0.1$ and $W_0 \sim 10$
- ALP mass could be generated by single Kähler potential instanton, e.g. $m_{\text{ALP}} \sim m_{3/2} e^{-\pi\tau_*} \sim 10^{-10} \text{ eV}$ requires

$$\tau_* \sim \frac{1}{\pi} \ln \left(\frac{g_s m_{3/2}}{m_{\text{ALP}}} \right) \sim 16.$$

- Astrophysical hints compatible with intermediate string-scale scenario

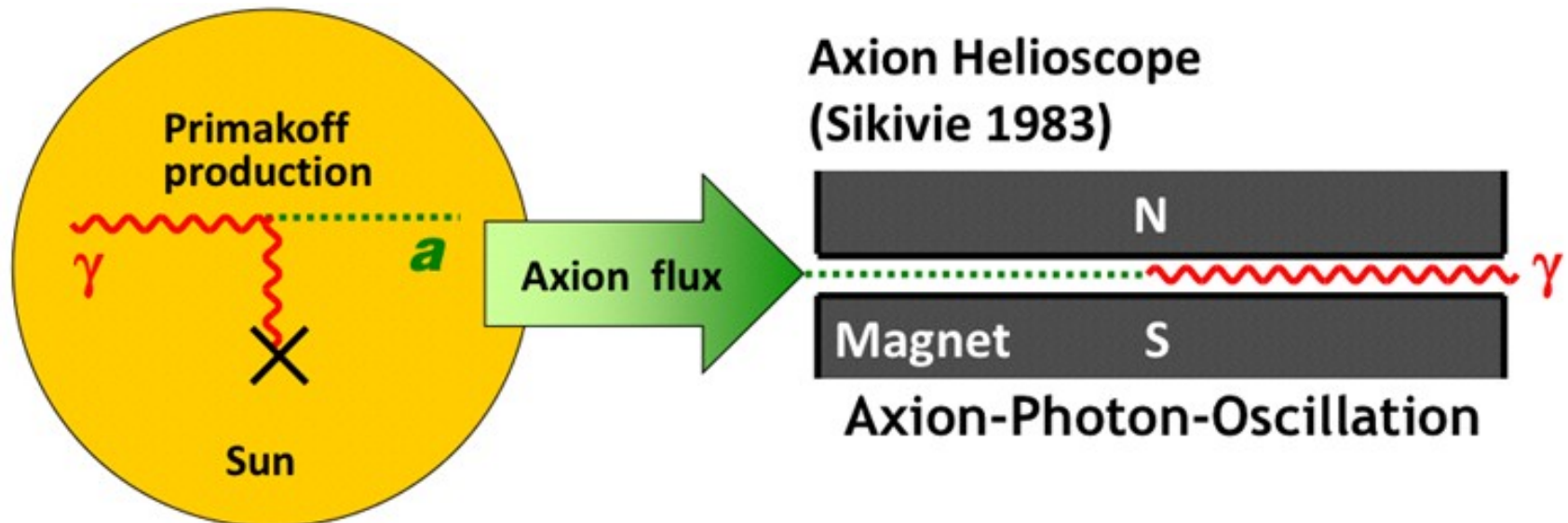


Laboratory probes of axion and ALPs with intermediate scale decay constant?



Search for solar axions and ALPs

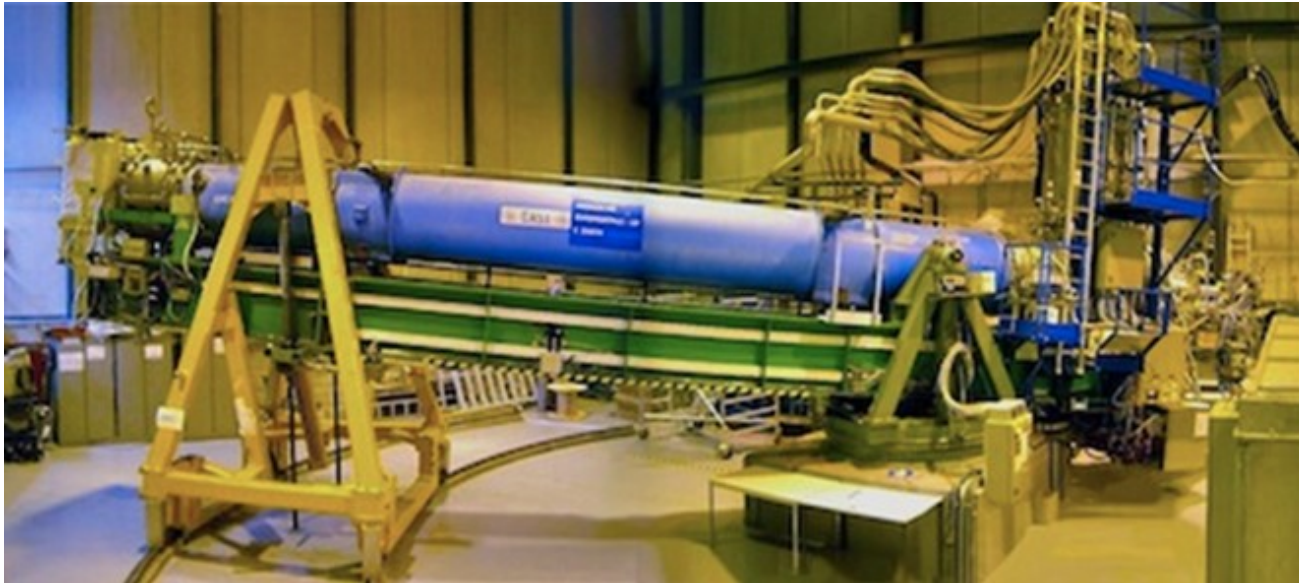
- Sun strong source of axions and ALPs
- Helioscope searches for axions/ALPs and HPs



$$P(a \leftrightarrow \gamma) = 4 \frac{(g_{a\gamma} \omega B)^2}{m_a^4} \sin^2 \left(\frac{m_a^2}{4\omega} L_B \right)$$

Search for solar axions and ALPs

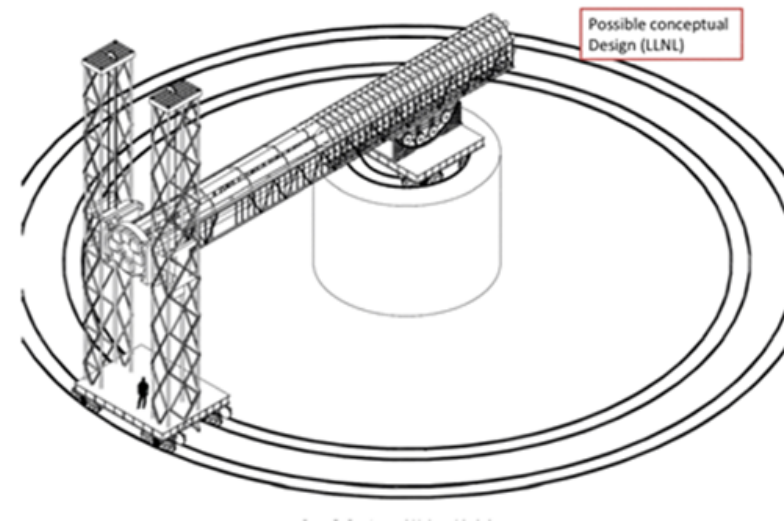
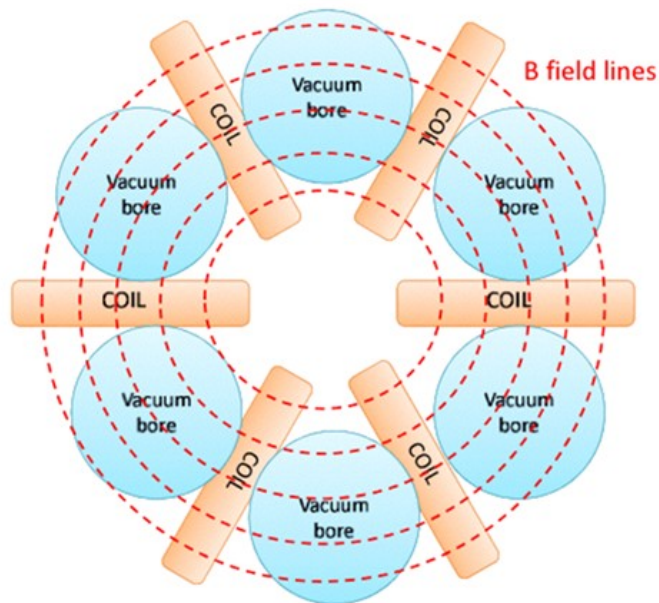
- Sun strong source of axions and ALPs
- Helioscope searches for axions/ALPs and HPs



- CAST ... CERN Axion Solar Telescope

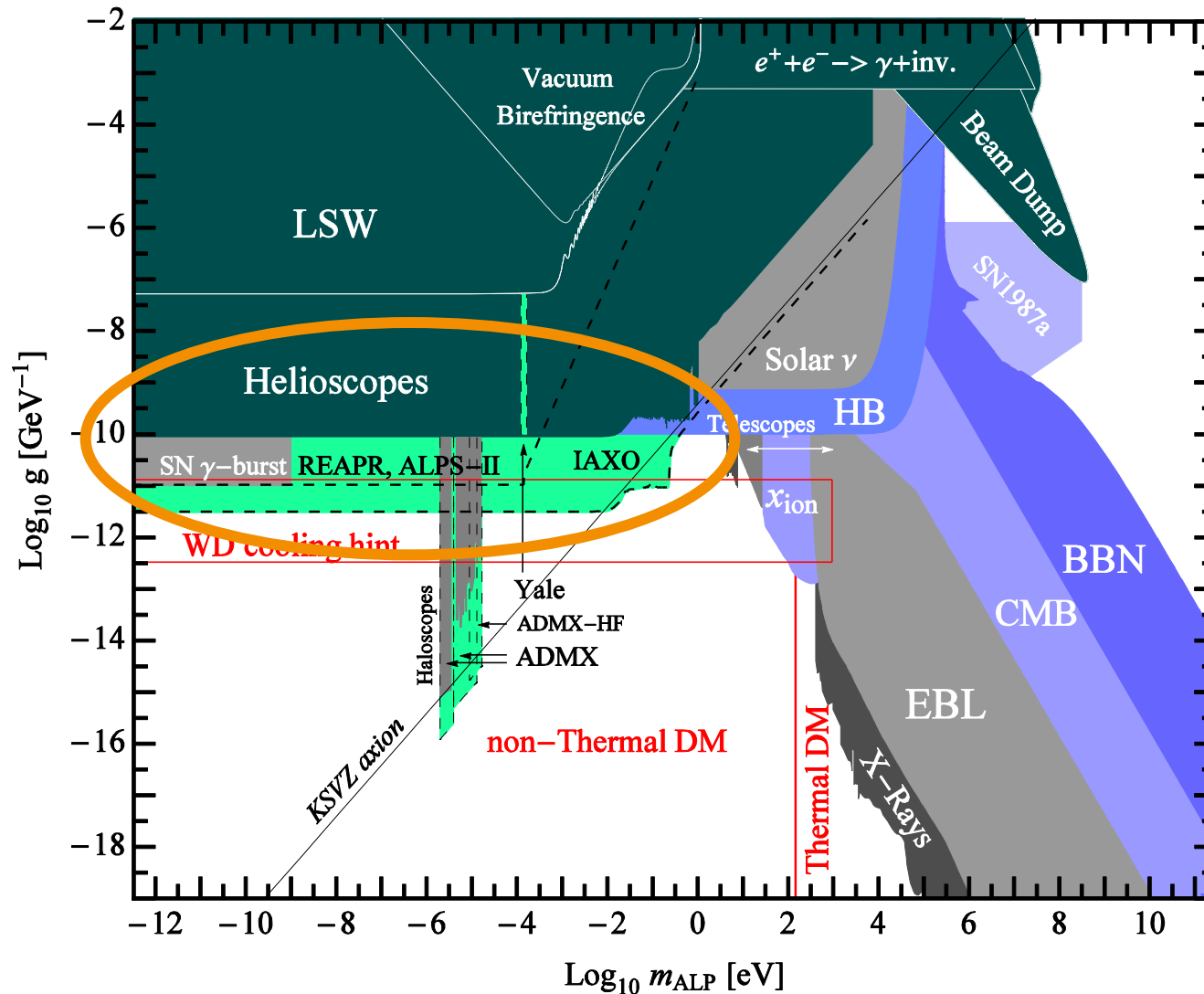
Search for solar axions and ALPs

- Sun strong source of axions and ALPs
- Helioscope searches for axions/ALPs and HPs



- CAST ... CERN Axion Solar Telescope
- IAXO ... International Axion Observatory (under investigation)

Search for solar axions and ALPs



Search for ALPs via light-shining-through-walls

- ALPs can pass walls
- Light-shining-through-walls experiments: (here ALPS (@DESY)):

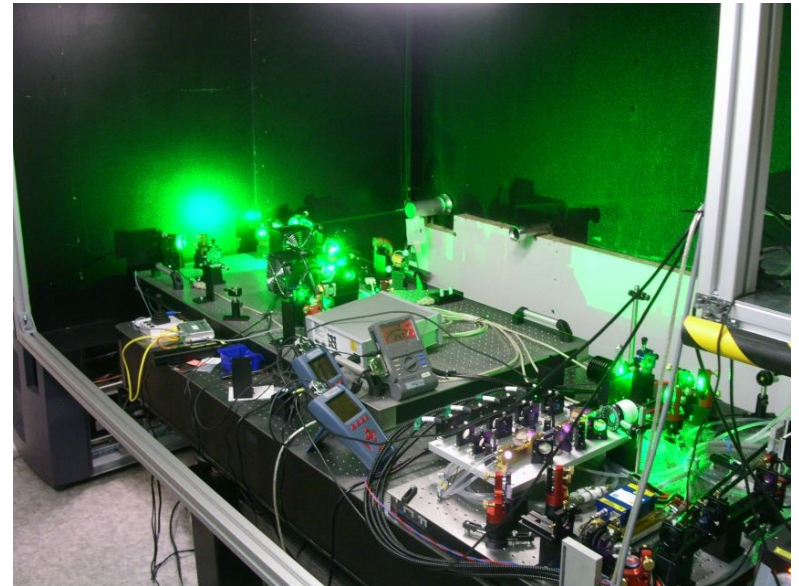
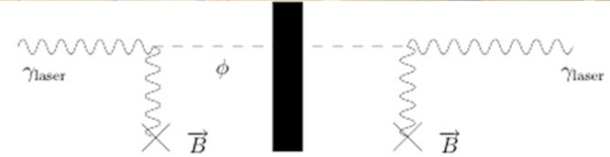


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Search for ALPs via light-shining-through-walls

> ALPS:

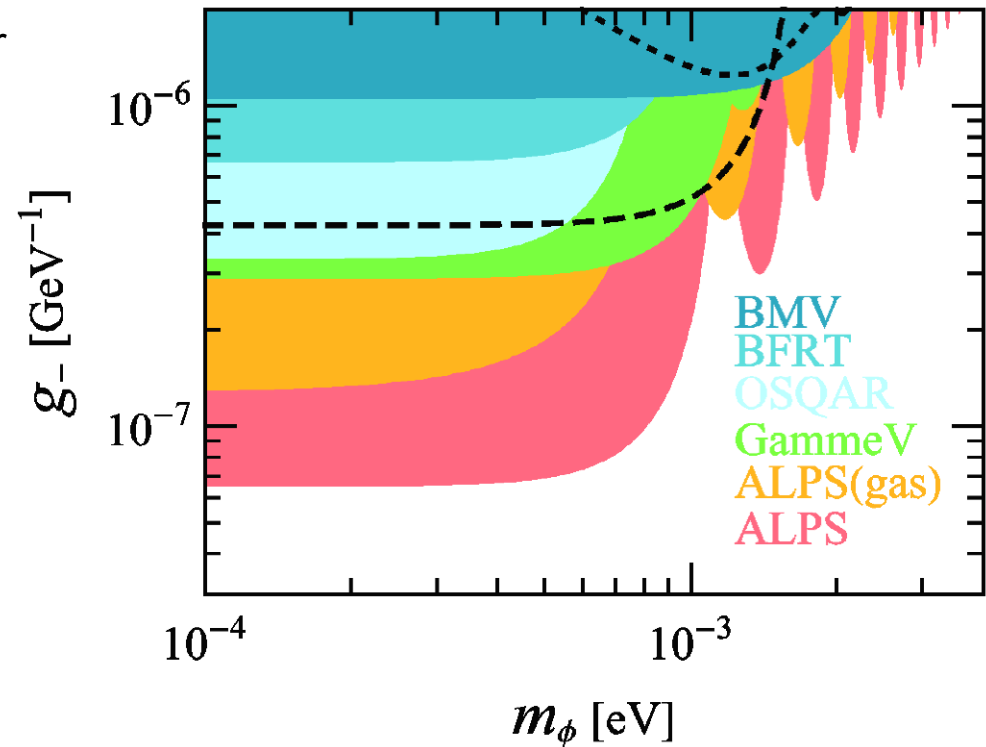
- HERA dipole (8.4 m, 5 T)
- Primary laser: enhanced LIGO laser (1064 nm, 35 W)
- Frequency doubled: 523 nm
- 300-fold power build-up in cavity



Search for ALPs via light-shining-through-walls

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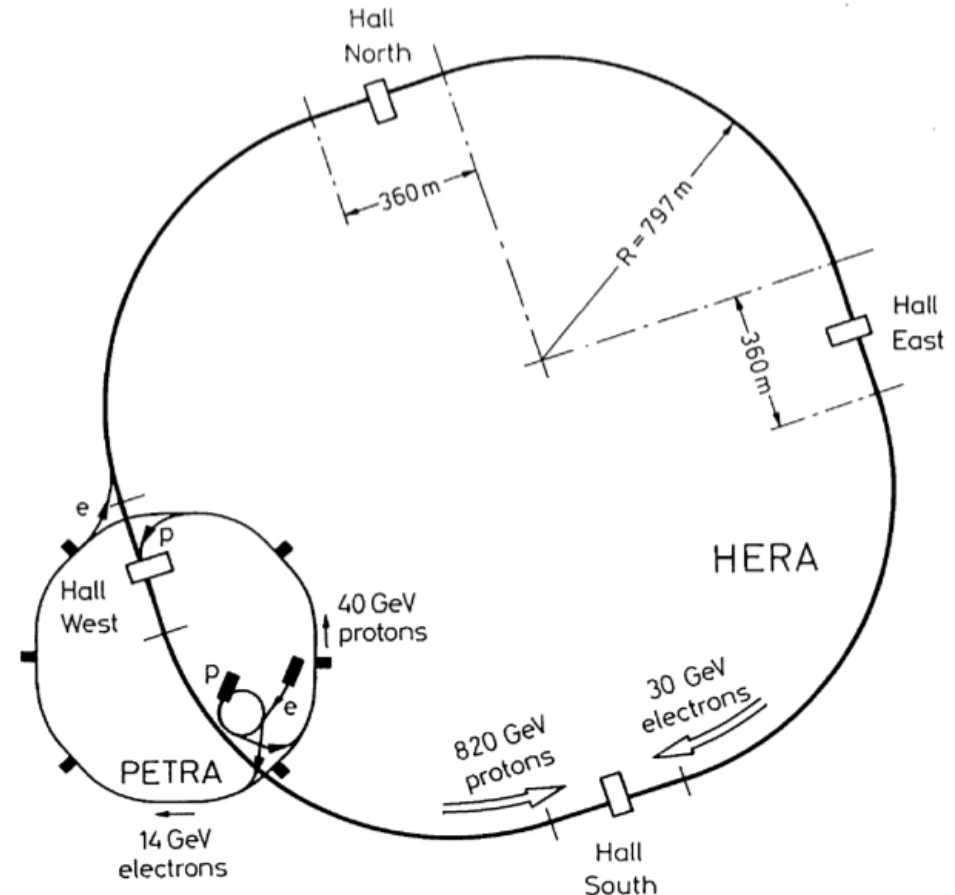
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> ALPS-II plans (2016+):

- 12 + 12 HERA dipoles
- Increased power build-up (~5000)
- Cavity also on regeneration part



Search for ALPs via light-shining-through-walls

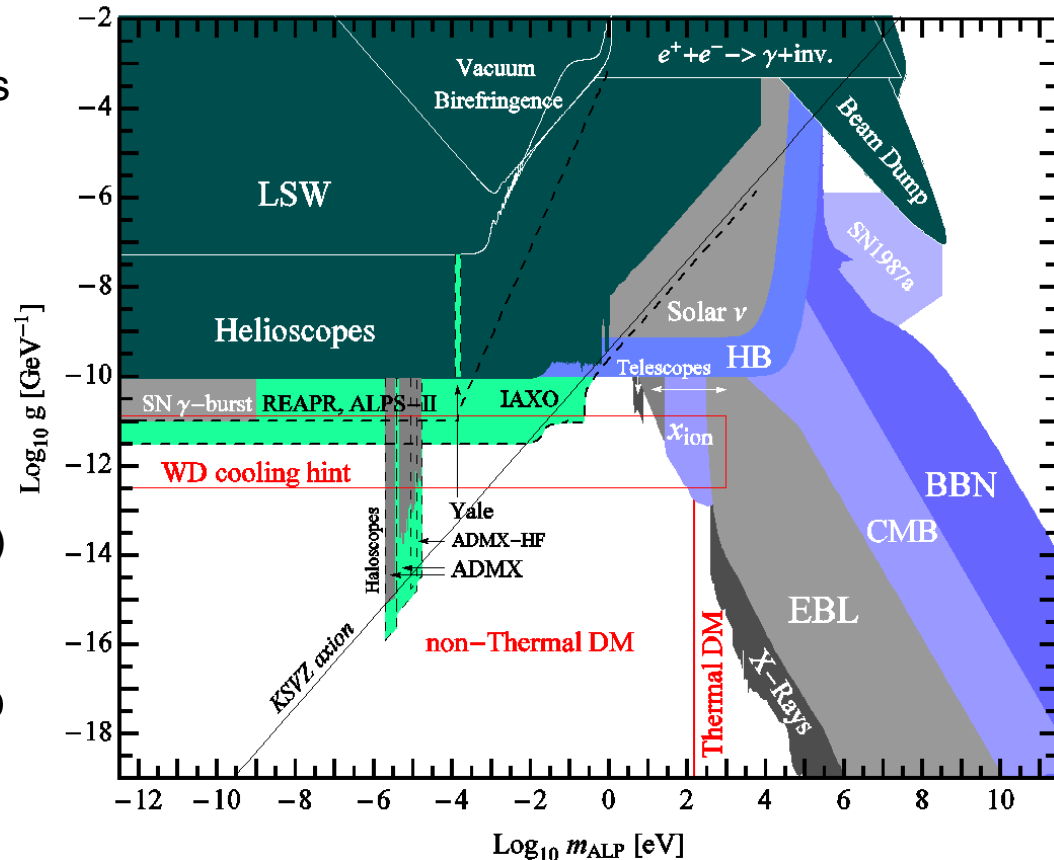
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> ALPS-II plans (2016+):

- 12 + 12 HERA dipoles
- Increased power build-up (~5000)
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> Similar plans also at Fermilab (REAPR)

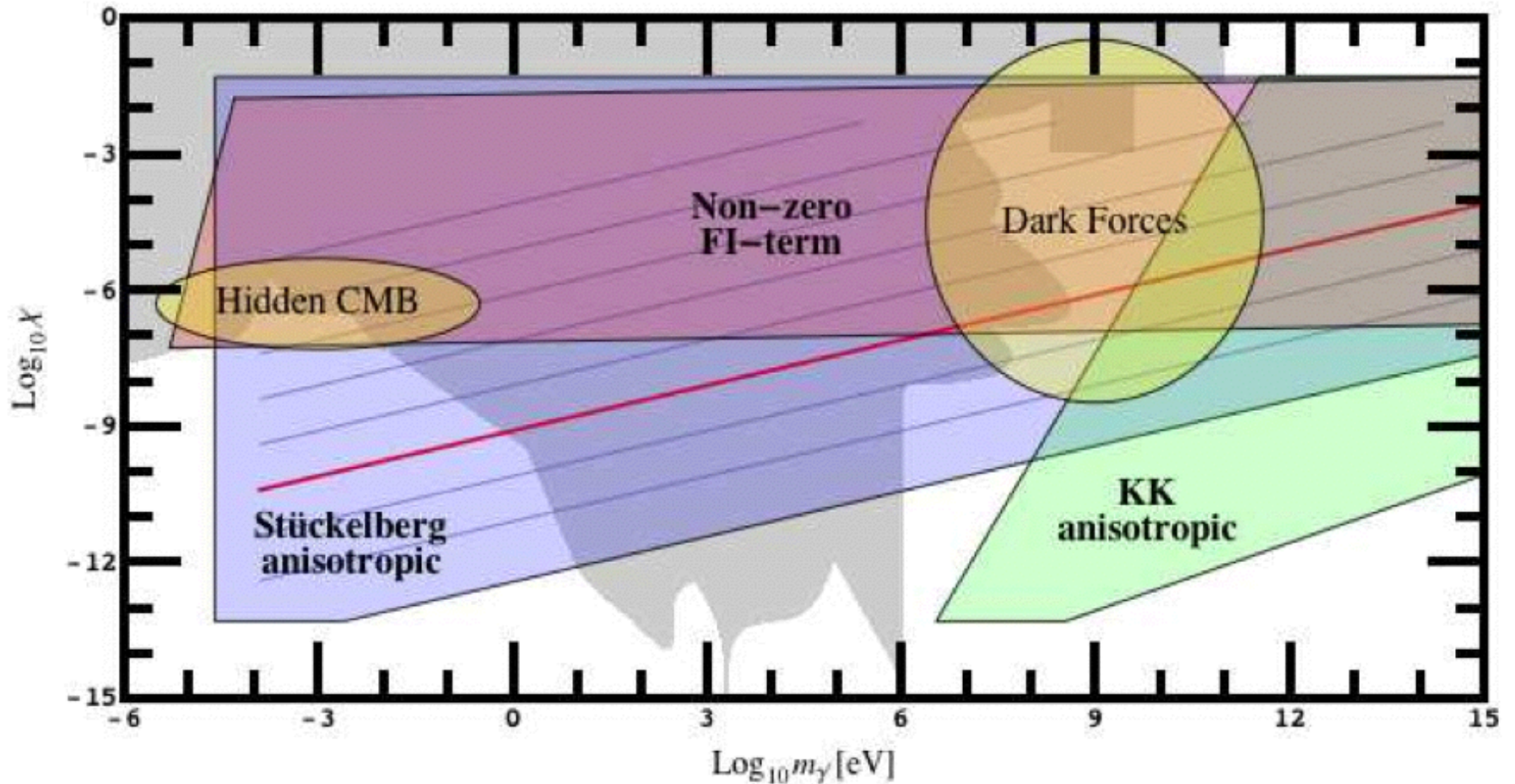


Summary and conclusions

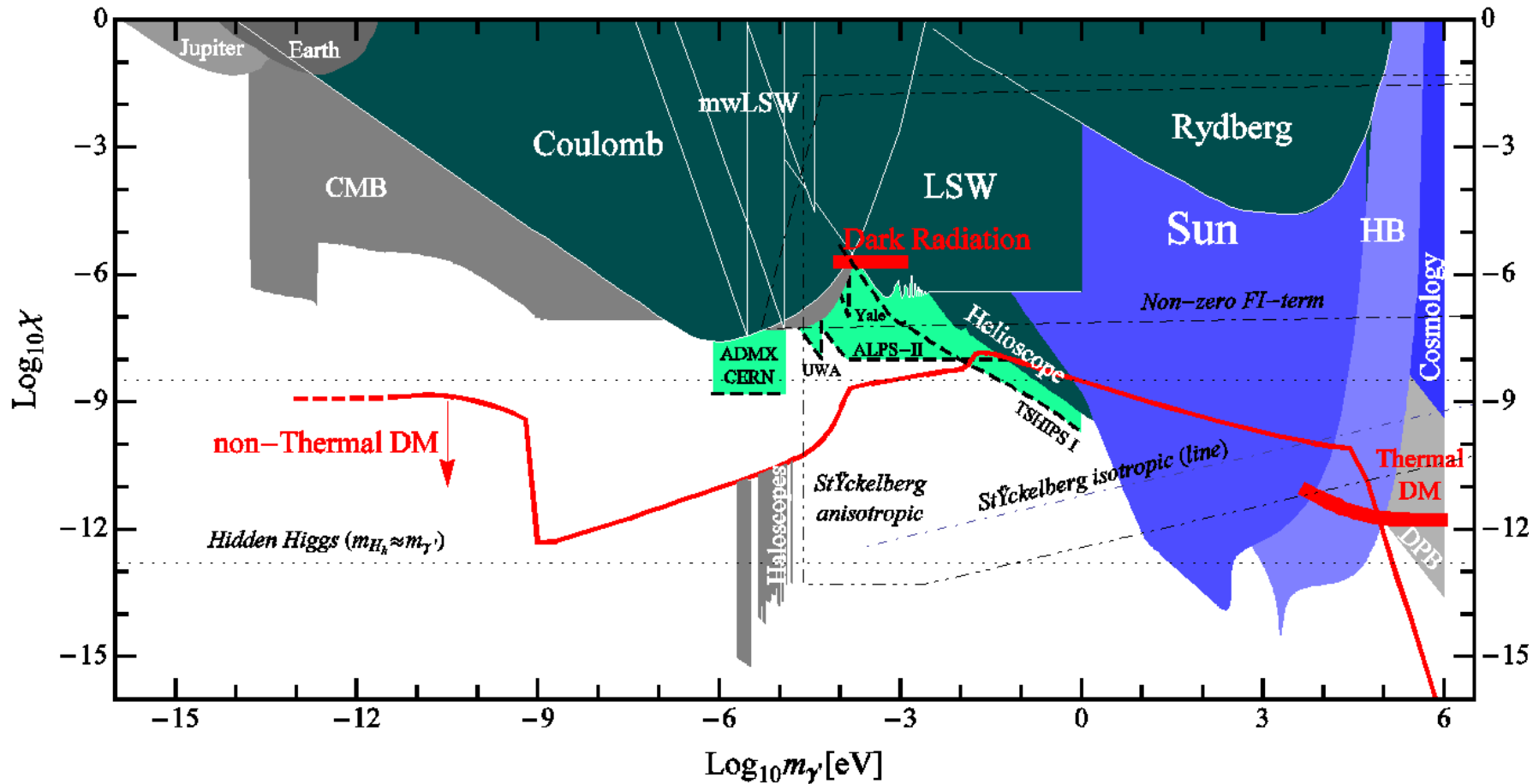
- String phenomenology holds the promise of an axiverse – the QCD axion plus a (possibly large) number of further ultralight axion-like particles, possibly populating each decade of mass down to the Hubble scale
- Promise fulfilled in LARGE Volume Scenario of IIB string compactifications
- Models that exhibit a QCD axion with an intermediate-scale decay constant and additional even lighter axion-like particles having the same decay constant and coupling to the photon can explain astrophysical anomalies and be tested in the next generation of helioscopes and light-shining-through-walls experiments
- Cosmology of LVS axiverse still to be investigated in detail



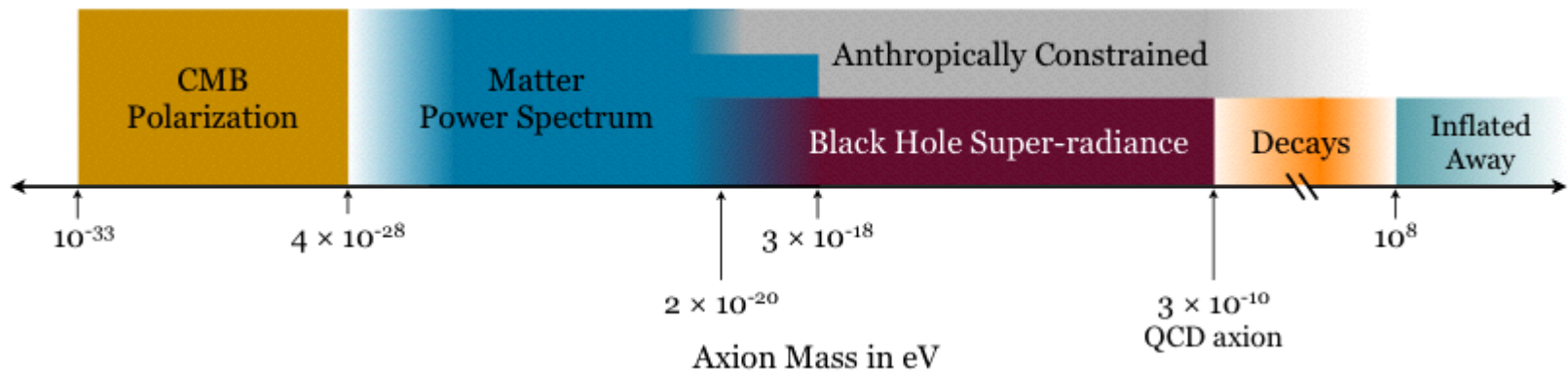
Backup: Hidden photons



Backup: Hidden photons



Backup: More probes of ultralight axions and ALPs



cf. Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell '09

