

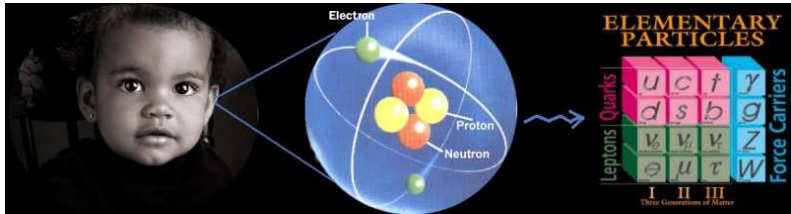
# Challenges of string theory: particle physics and cosmology

Saúl Ramos-Sánchez

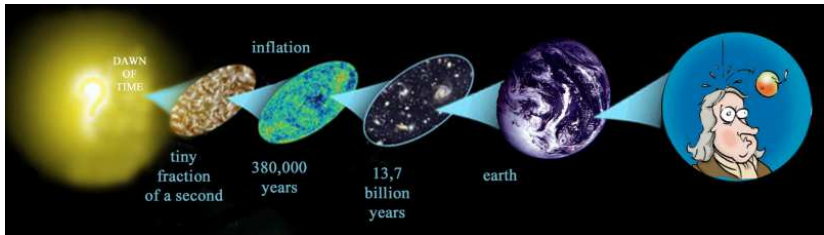
XIII Mexican Workshop on Particles and Fields

October 20, 2011

# What do we know?



SM = QCD (SU(3)) + EW (SU(2) × U(1)<sub>Y</sub>) ☺



General relativity: Cosmology ☺

# Urge to go beyond SM...



## HIGGS BOSON

# H



The **HIGGS BOSON** is the theoretical particle of the Higgs mechanism, which physicists believe will reveal how all matter in the universe get its mass. Many scientists hope that the Large Hadron Collider in Geneva, Switzerland will detect the elusive Higgs Boson when it begins colliding particles at 99.99% the speed of light.

*Wool felt with gravel fill for maximum mass.*



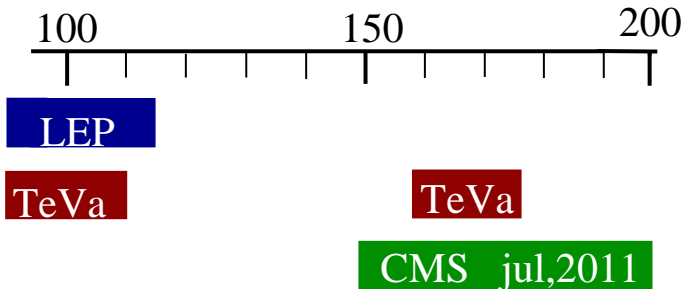
LIGHT

HEAVY

**\$9.75** PLUS SHIPPING

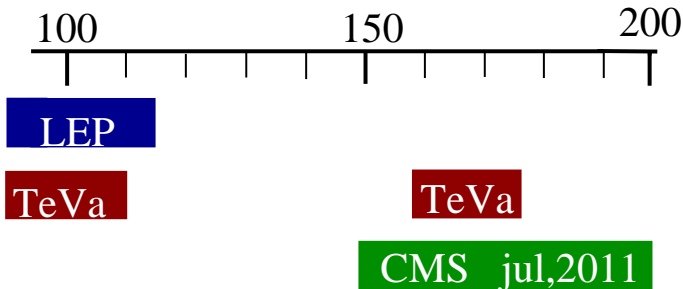
GLUON PHOTON NEUTRINO TACHYON ELECTRON UP QUARK DOWN QUARK TAU NEUTRINO MUON UP QUARK  
NEUTRON DOWN QUARK TAU GLUON **HIGGS BOSON** NEUTRINO TACHYON ELECTRON UP QUARK DOWN  
NEUTRINO MUON UP QUARK PROTON NEUTRON DOWN QUARK TAU GLUON PHOTON NEUTRINO TACHY  
UP QUARK DOWN QUARK TAU NEUTRINO MUON UP QUARK PROTON NEUTRON DOWN QUARK TAU GLU  
The **PARTICLE ZOO** ELECTRON UP QUARK DOWN QUARK TAU NEUTRINO MUON UP QUARK PROTON  
NEUTRON DOWN QUARK TAU GLUON PHOTON NEUTRINO TACHYON ELECTRON UP QUARK DOWN QUARK TAU NEU

# Urge to go beyond SM...



$$115 < m_H < 145 \text{ GeV}$$

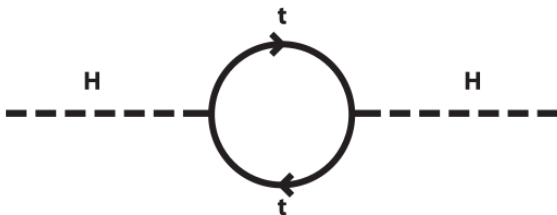
# Urge to go beyond SM...



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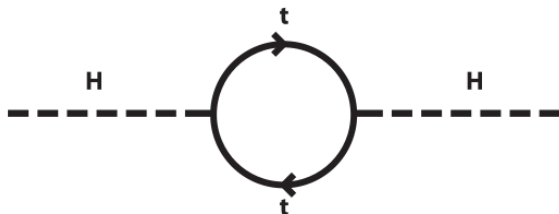
Why?

# Urge to go beyond SM...



$$\Rightarrow \Delta m_H^2 \sim \Lambda^2$$

# Urge to go beyond SM...



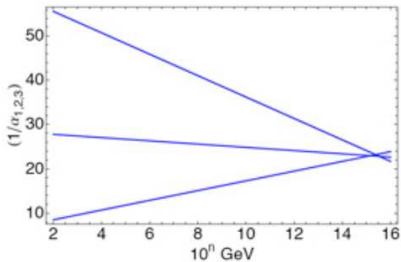
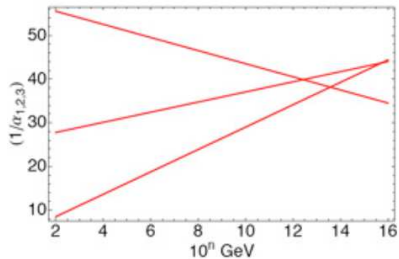
+ SUSY @  $E \sim \text{TeV}$  ?



$$\Rightarrow \Delta m_H^2 \sim \Lambda^2 - \Lambda^2$$

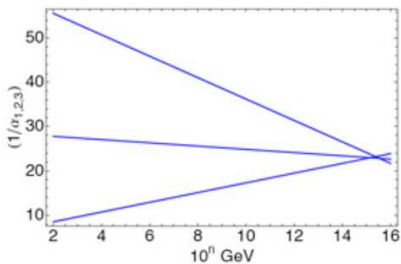
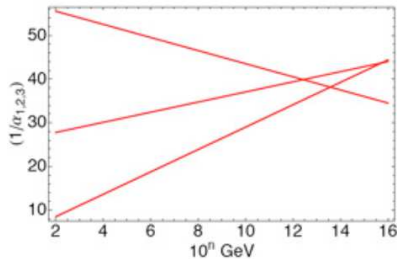


# Urge to go beyond SM...



Why?

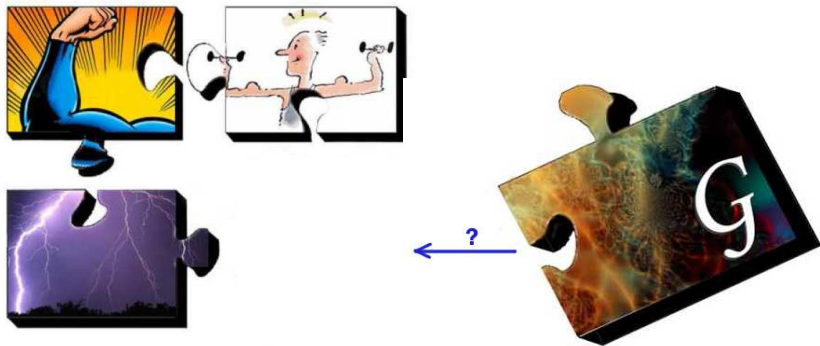
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Why?

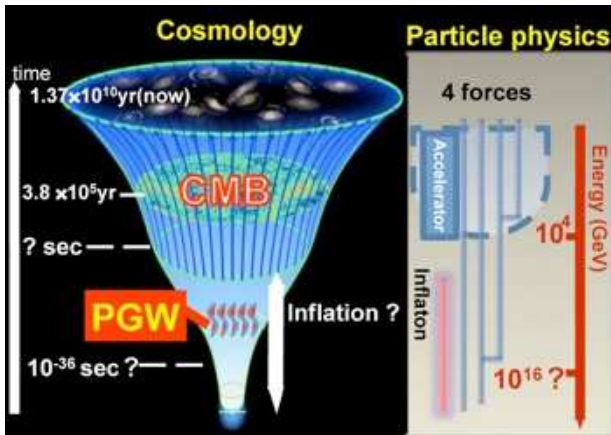
GUTs?

# Urge to go beyond SM...



Quantum Gravity ?

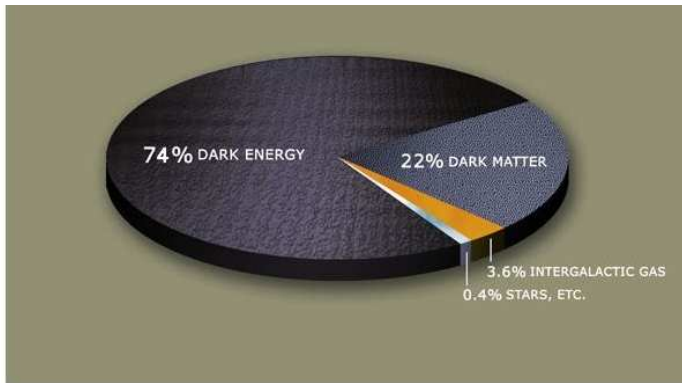
# Urge to go beyond $\Lambda$ CDM...



$$V(\varphi) = ?$$

What is  $\varphi$  ?

# Urge to go beyond $\Lambda$ CDM...



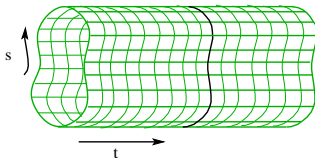
dark matter: neutralino, gravitino, ... ?

dark energy:  $\Lambda \sim 10^{-120}$ , chameleon, ... ?

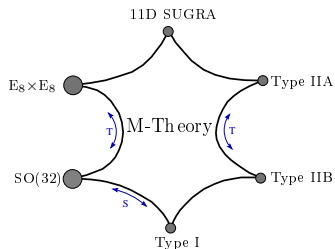


# Cuerdas

1970's: particles  $\rightarrow$  strings



80-90's: 5 theories of superstrings (+branes)



quantum consistency

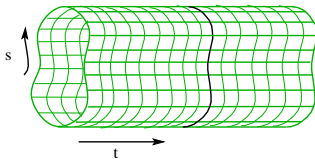
(no anomalies, "ghosts", tachyons):



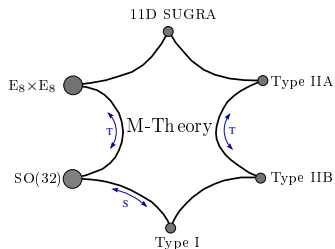
- \* graviton included
- \* gauge bosons
- \* supersymmetry
- \* 10 dimensions

# Cuerdas

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- \* graviton included ☺
- \* gauge bosons ☺
- \* supersymmetry ☺ ☹
- \* 10 dimensions ☹

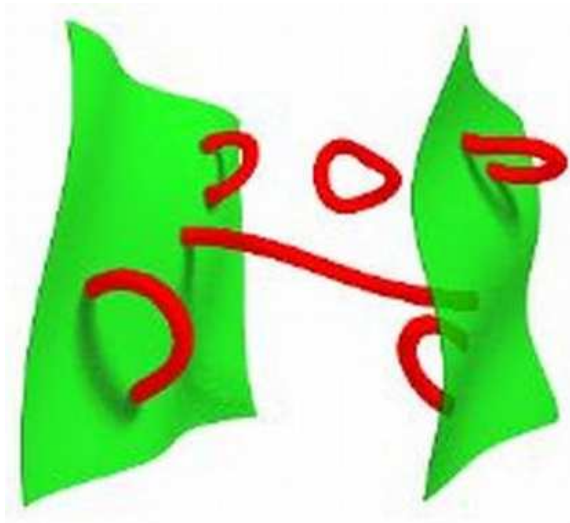


# D-Branes in type I/II



Open strings  $\rightarrow$   $U(1)$  gauge symmetry (90's revolution!)

# Stacks D-Branes in type I/II



Open strings  $\rightarrow$   $U(2) \simeq SU(2) \times U(1)$  gauge symmetry

Can we reproduce our universe?

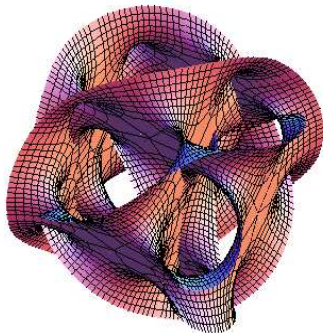
# String Phenomenology Challenges

# First challenge: $10 \neq 4$

- **Compactifications**

$$X_{10} = \mathbb{M}^4 \otimes X_6 \quad \text{size}(X_6) \sim \ell_{Pl}^6, \quad \mathcal{N} = 1$$

- ①  $X_6$ : Calabi-Yau ( $CY_3$ ) manifolds Candelas, Horowitz, Strominger, Witten (1985)

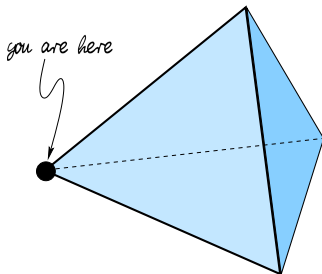


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- 3 Generalized, half-flat,  $SU(3) \times SU(3)$  manifolds, ... Oscar Loaiza-Brito

## Second challenge: some matter!

- Heterotic strings  $E_8 \times E_8$  or  $SO(32)$

$$\begin{array}{rcl} E_8 & \xrightarrow{\text{compact.}} & E_6 \times SU(3) \\ 248 & \rightarrow & (78, 1) + (1, 8) + (27, 3) + (\overline{27}, \overline{3}) \end{array}$$

$E_6$  GUTs ☺

# Second challenge: some matter!

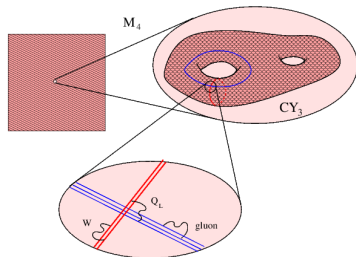
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$$248 \rightarrow (78, 1) + (1, 8) + (27, 3) + (\overline{27}, \overline{3})$$

$E_6$  GUTs ☺

- Type II A/B

Berkooz et al. (1996)

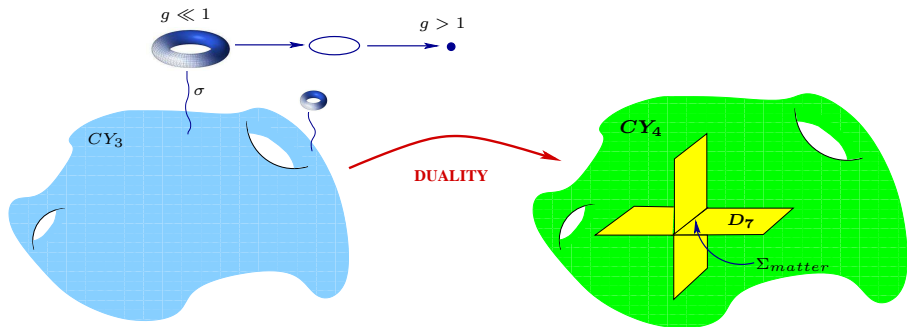




# Second challenge: some matter beyond traditional strings!

- F-theory

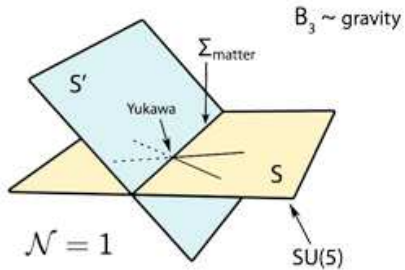
Beasley, Heckman, Vafa (2008-2010)



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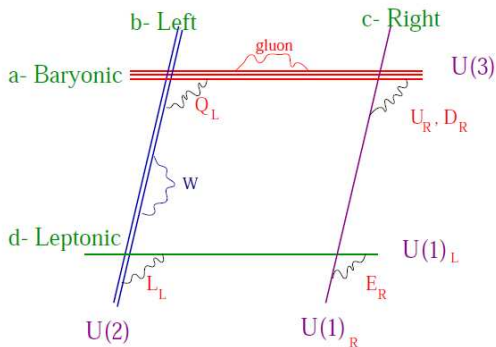
Beasley, Heckman, Vafa (2008-2010)



# Third challenge: (MS)SM matter & interactions

- Type IIA : Madrid Model

Cremades, Ibáñez, Marchesano (2001-2003)



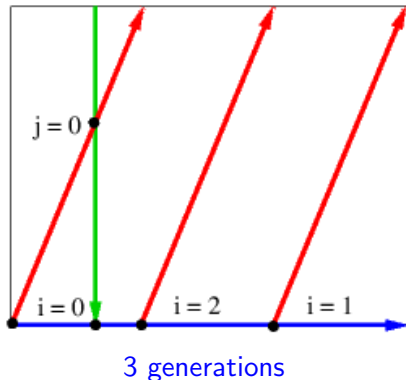
$$X_6 = T^2 \times T^2 \times T^2 / \mathbb{Z}_2 \times \mathbb{Z}_2$$

with  $D6$  branes  $\Rightarrow$   $SU(3) \times SU(2) \times U(1)_Y \times U(1)^4$

# Third challenge: (MS)SM matter & interactions

- Type IIA : Madrid Model

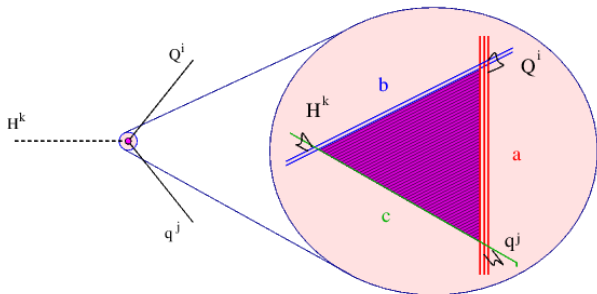
Cremades, Ibáñez, Marchesano (2001-2003)



# Third challenge: (MS)SM matter & interactions

- Type IIA : Madrid Model

Cremades, Ibáñez, Marchesano (2001-2003)



Yukawa couplings from instantons

$$Y_u^{33} \propto e^{-A} \neq 0, \quad A \sim \text{area triangle}$$

# Third challenge: (MS)SM matter & interactions

- Type IIA : Madrid Model hurdles and solutions

Cremades, Ibáñez, Marchesano (2001-2003)

- Anomalies: tadpoles do not cancel
- ~~SUSY~~ @  $M_{str}$   $\Rightarrow$  hierarchy problem! ☹️

Solution:  $X_6 = CY_3$  or orbifold

Blumenhagen (2002), Honecke, Gmeiner (2004-2008)

- Only  $Y_{b,t} \neq 0$  non-perturbatively  
 $\Rightarrow$  4 quarks massless ☹️  
 $Y_t \ll 1$  ☹️

Solution: different intersections

$\Rightarrow Y_{u,d,c,s} \neq 0$  perturbatively

- Other issues:  $Y_t < Y_u$ , chiral exotics, ... ☹️

# Third challenge: (MS)SM matter & interactions

- Type IIA : An intersecting D-brane model

Gmeiner, Honecker (2008)

$$SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_{B-L} \times G_{hidden}$$

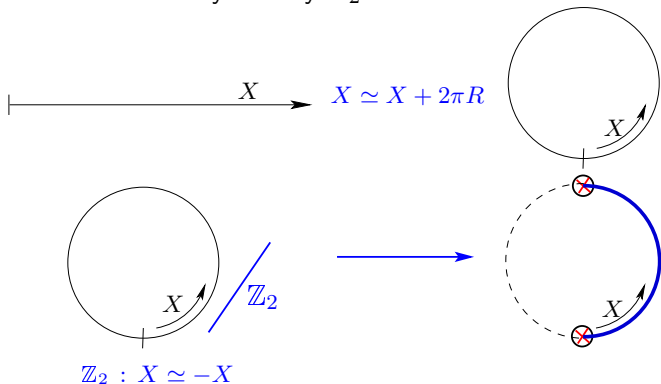
$$\begin{aligned} [C] &= 3 \times \left[ (\mathbf{3}, \mathbf{2})_{1/6, 1/3}^{(0,0)} + (\overline{\mathbf{3}}, \mathbf{1})_{1/3, -1/3}^{(1,0)} + (\overline{\mathbf{3}}, \mathbf{1})_{-2/3, -1/3}^{(-1,0)} \right. \\ &\quad + (\mathbf{1}, \mathbf{1})_{1,1}^{(1,1)} + (\mathbf{1}, \mathbf{1})_{0,1}^{(-1,1)} + 2 \times (\mathbf{1}, \mathbf{2})_{-1/2, -1}^{(0,-1)} \\ &\quad \left. + (\mathbf{1}, \mathbf{2})_{1/2, 1}^{(0,1)} + 3 \times (\mathbf{1}, \overline{\mathbf{2}})_{-1/2, 0}^{(-1,0)} + 3 \times (\mathbf{1}, \overline{\mathbf{2}})_{1/2, 0}^{(1,0)} \right] \\ &\equiv 3 \times \left[ Q_L + d_R + u_R + e_R + \nu_R + 2 \times L + \overline{L} \right] + 9 \times \left[ H_d + H_u \right] \end{aligned}$$

- plus  $\sim 100$  vectorlike exotics
- Yukawa couplings for 2 generations allowed
- Majorar neutrino masses allowed

# Third challenge: (MS)SM matter & interactions

- $E_8 \times E_8$  heterotic orbifolds

1D Orbifold with symmetry  $\mathbb{Z}_2$  in 5D



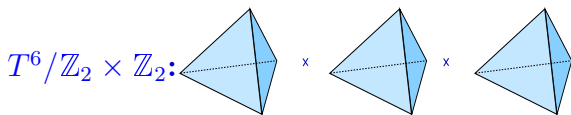
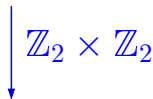
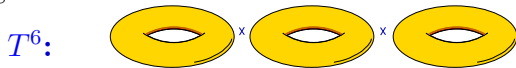
Very small singular space  $R \ll 1\text{mm} \rightarrow$  we do not see it!!

Kaluza, Klein (1920s)



# Third challenge: (MS)SM matter & interactions

- $E_8 \times E_8$  heterotic orbifolds



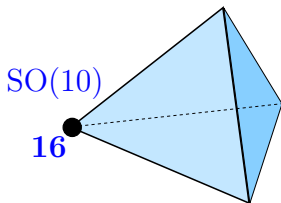
Strings in the 'bulk' : **gravity** and  
 $SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_{B-L} \times G_{hidden}$  ☺

Can we compute stuff?

Lebedev, Nilles, Ramos-Sanchez, Ratz, Vaudrevange (2006-2008)

# Third challenge: (MS)SM matter & interactions

- $E_8 \times E_8$  heterotic orbifolds



$$SU(3)_c \times SU(2)_L \times U(1)_Y \times U(1)_{B-L} \times G_{hidden} \quad \text{☺}$$

Strings @ 3 singularities (by construction):

**16** = complete family

- Many other appealing features: neutrino masses  $m_\nu \sim 10^{-2} \text{eV}$ ,  $M_{SUSY} \sim \text{TeV}$ , proton stability, no strong CP problem,...

Can we compute stuff?

Lebedev, Nilles, Ramos-Sanchez, Ratz, Vaudrevange (2006-2008)

# Third challenge: (MS)SM matter & interactions

- $E_8 \times E_8$  heterotic orbifolds. Quark sector

$$Y_u \sim \begin{pmatrix} 0.0316272 & 0.0316272 & 0.0085564 \\ 0.0316272 & 0.031 & 0.00841811 \\ 0.0183063 & 0.0183171 & 1.14437 \end{pmatrix},$$

$$Y_d \sim \begin{pmatrix} 0.000483779 & 0.000664073 & 6.158592 \times 10^{-7} \\ 0.000664073 & 0.000483779 & 1.026432 \times 10^{-7} \\ 0.0000867024 & 0.000036018 & 0.0357596 \end{pmatrix},$$

$$|Y_u^{\text{diag}}| \sim \text{diag}(0.00032, 0.06265, 1.14466),$$

$$|Y_d^{\text{diag}}| \sim \text{diag}(0.00018, 0.00115, 0.03576).$$

semirealistic quark masses!

Lebedev, Raby, Ramos-Sanchez

# Fourth challenge: moduli stabilization (towards cosmology)

Potential problem: the geometry of space is not fixed!!

Allowed deformations: position of branes, size and shape of  $X_6$

⇒ moduli:  $\varphi_j$

Perturbatively  $V(\varphi_j) = 0$  ☹️

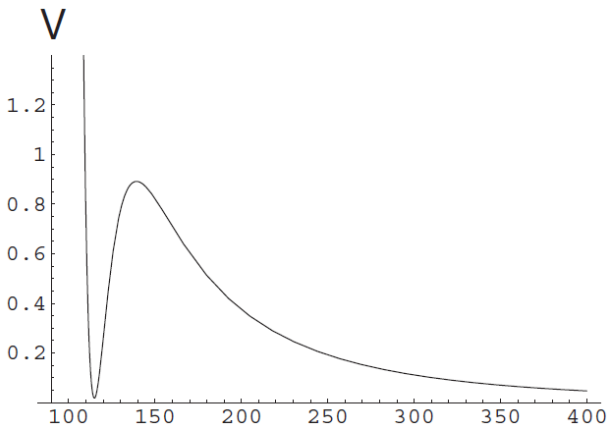
fifth forces, cosmological overclosure

**BUT** non-perturbative effects (instantons, gaugino condensation, fluxes...) and possible field-VEVs can induce

$$V(\varphi_j) \sim -\frac{1}{\varphi_j} + e^{-\alpha\varphi_j}$$

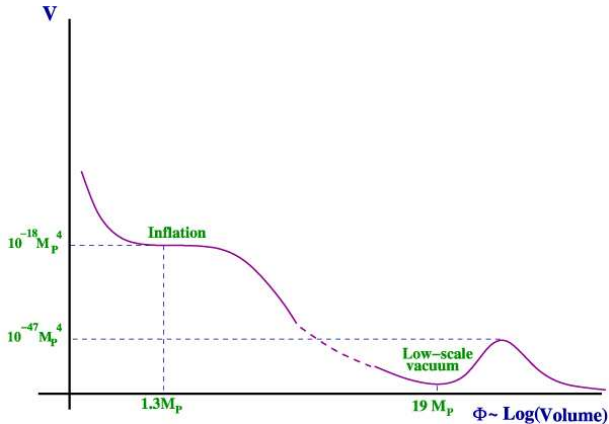
Kachru, Kallosh, Linde, Trivedi (2003)

# Fourth challenge: moduli stabilization (towards cosmology)



Kachru, Kallosh, Linde, Trivedi (2003)

# Fourth challenge: moduli stabilization and inflation



Conlon, Kallosh, Linde, Quevedo (2008s)

## To take home...

- String phenomenology from all corners ✓
- Very close to the MSSM ✓
- In some scenarios, even computability available ✓
- Still much work to do for cosmology