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## CAN $^{59}\text{Ni}$ SYNTHESIZED IN OB ASSOCIATIONS DECAY TO $^{59}\text{Co}$ BEFORE BEING ACCELERATED TO COSMIC-RAY ENERGIES?

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### Abstract content

Observations from the Cosmic Ray Isotope Spectrometer (CRIS) aboard NASA's Advanced Composition Explorer (ACE) have shown that all relevant galactic cosmic ray isotopic ratios measured are consistent with an OB-Association origin of galactic cosmic rays (GCRs). Additionally CRIS measurements of the isotopic abundances of  $^{59}\text{Ni}$  and  $^{59}\text{Co}$  have shown that the  $^{59}\text{Ni}$  has completely decayed into  $^{59}\text{Co}$ , indicating a delay of

$10^5$  years between nucleosynthesis and acceleration. However, it has been suggested that shocks generated from high-velocity Wolf-Rayet winds in the OB-Association environment must accelerate nuclei synthesized in nearby core-collapse SNe on a time scale short compared to the  $^{59}\text{Ni}$  lifetime of  $7.6 \times 10^4$  years. If this were the case, it would imply that OB Associations could not be the source of most galactic cosmic rays. In this paper, we describe the OB-Association history and environment and show that the time scales for acceleration are such that most  $^{59}\text{Ni}$  should be expected to decay naturally in that setting, strengthening the argument that OB associations are the likely source of a substantial fraction of galactic cosmic rays.

**If this papers is presented for a collaboration, please specify the collaboration**

### Summary

### Reference

Proceedings of the 30th International Cosmic Ray Conference; Rogelio Caballero, Juan Carlos D'Olive, Gustavo Medina-Tanco, Lukas Nellen, Federico A. Sánchez, José F. Valdés-Galicia (eds.); Universidad Nacional Autónoma de México, Mexico City, Mexico, 2008; Vol. 2 (OG part 1), pages 27-30

**Primary author(s) :** BINNS, W. R. (Washington University in St. Louis)

**Co-author(s) :** ISRAEL, M. H. (Washington University in St. Louis); CUMMINGS, A. C. (California Institute of Technology); LESKE, R. A. (California Institute of Technology); MEWALDT, R. A. (California Institute of Technology); STONE, E. C. (California Institute of Technology); WIEDENBECK, M. E. (Jet Propulsion Laboratory, California Institute of Technology); DE NOLFO, G. A. (NASA/Goddard

Space Flight Center); VON ROSENVINGE, T. T. (NASA/Goddard Space Flight Center); ARNOULD, M. (Institut d'Astronomie et d'Astrophysique, Bruxelles); GORIELY, S. (Institut d'Astronomie et d'Astrophysique, Bruxelles); MEYNET, G. (Geneva Observatory, Sauverny)

**Presenter(s) :** BINNS, W. R. (Washington University in St. Louis)

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