

# EAS-TOP: THE COSMIC RAY ANISOTROPY IN THE ENERGY REGION $E_0 = 10^{14} \div 10^{15}$ EV

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## OUTLINE

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  - The EAS-TOP array and the data
  - The East-West method
- 3 RESULTS
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  - The wave shapes

## LARGE SCALE ANISOTROPIES

- Essential tool for understanding CR origin and propagation
- At  $10^{14} - 10^{15}$  eV can provide insight on the origin of the “knee” of the spectrum (source effect or propagation result)

## FROM THE EXPERIMENTAL POINT OF VIEW (EAS ARRAYS)

- Main problems are both statistical and systematical:
  - large number of events are required,
  - collected by long duration experiments
  - over large areas.
- Detector stability is a crucial issue, and atmospheric effects (e.g., variations of pressure and temperature). Data have to be corrected for, and, for high sensitivities, they could induce systematical effects difficult to evaluate.

## LARGE SCALE ANISOTROPIES and EAS-TOP PAST RESULTS

### AT 100 TeV

- Data correction for atmospheric effects
- Harmonic analysis on corrected data
- Observation of significant anisotropy (10 “ $\sigma$ ” level) in sidereal time ( $A_{sid} = (3.4 \pm 0.3) 10^{-4}$ ,  $\phi_{sid} = (3.3 \pm 0.4)$  h LST)
- Observation of expected Compton Getting effect (due to Earth revolution around Sun) in solar time and absence of anti-sidereal signal insured the reliability of the observation

The EAS-TOP result extended the anisotropy measurement up to 100 TeV, showing the constancy (in amplitude and phase) with respect to lower energy ones (amplitude  $((3 \div 6) 10^{-4})$  and phase  $((0 \div 4)$  h LST) between  $10^{11}$  and  $10^{14}$  eV

## LARGE SCALE ANISOTROPIES and EAS-TOP PAST RESULTS

### **AT HIGHER ENERGIES, i.e., ABOVE 400 TeV**

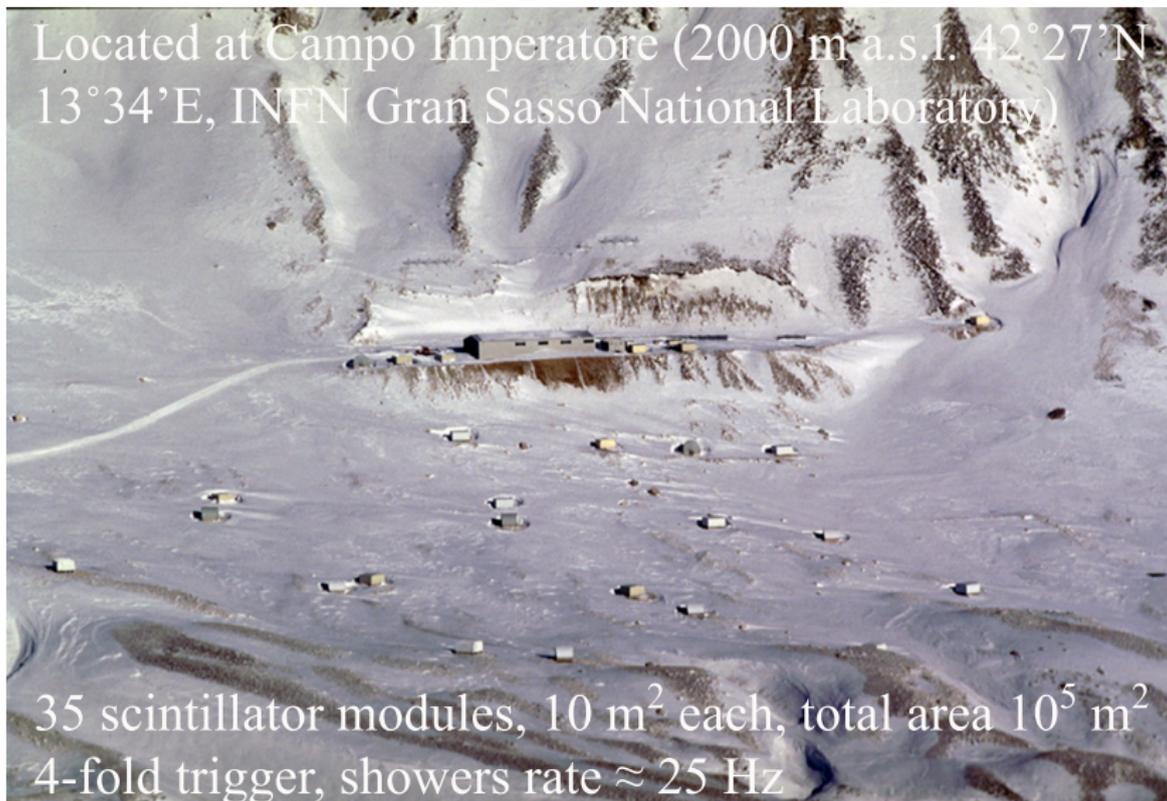
- Same analysis technique as for 100 TeV (i.e., correction for pressure effects on EAS)
- Observation of increasing amplitudes in sidereal harmonic analysis, but with a limited significance level
- Upper limits given

## Present analysis

- We use the full EAS-TOP data set (i.e., eight full years of data)
- We apply cuts to select two primary energies (100 and 400 TeV)
- We adopt a differential method (EW), based on the counting rate differences between East-ward and West-ward directions, to remove spurious counting rate variations (no need for atmospheric corrections)

## THE EAS-TOP ARRAY

Located at Campo Imperatore (2000 m a.s.l. 42°27'N  
13°34'E, INFN Gran Sasso National Laboratory)



35 scintillator modules, 10 m<sup>2</sup> each, total area 10<sup>5</sup> m<sup>2</sup>  
4-fold trigger, showers rate  $\approx$  25 Hz

## THE DATASET

| Class | $N_{modules}$ | $E_0$ [TeV] | $N_{EW}$         |
|-------|---------------|-------------|------------------|
| I     | $\geq 4$      | 100         | $1.5 \cdot 10^9$ |
| II    | $\geq 12$     | 400         | $1.7 \cdot 10^8$ |

- 1431 full days between January 1992 and December 1999
- Counting rates every 20 min
- $\phi$  inside  $\pm 45^\circ$  around the East and West directions
- $\theta < 40^\circ$
- Two primary energies: cuts in number of triggered modules ( $E_0$  evaluated for primary protons and QGSJET hadron interaction model in CORSIKA)

## THE EAST-WEST METHOD

$$\frac{dI}{dt} \simeq \frac{C_E(t) - C_W(t)}{\delta t}$$

$C_{E,W}(t)$  = n. of counts from the East, West sectors in  $\Delta t = 20$  min

$I$  = total intensity

$\delta t = 1.7$  h = average hour angle between the vertical and each of the two sectors

The integrated wave shape:

$$I(t_{N_{int}}) = \frac{\Delta t}{N_{int}} \sum_{i=1}^{N_{int}} i \frac{C_E(i) - C_W(i)}{\delta t} + \langle I \rangle$$

where  $N_{int} = 72$  intervals of solar / sidereal / anti-sidereal time and  $t_{N_{int}} = N_{int} \cdot \Delta t$ .

- Harmonic analysis performed on the differences  
 $D(i) = C_E(i) - C_W(i)$
- Differential amplitude and phase are transformed into the integral ones:  $r_I = \frac{r_D}{\delta t}$  and  $\phi_I = \phi_D + \frac{\pi}{2}$
- Uncertainties on  $r_I$  and  $\phi_I$ :  $\sigma_{r_I} = \frac{1}{\delta t} \sqrt{\frac{2}{N_{EW}}}$  and  $\sigma_{\phi_I} = \frac{\sigma_{r_I}}{r_I}$
- Rayleigh imitation probability:  $P = \exp\left(-\frac{r_I^2}{2\sigma_{r_I}^2}\right)$

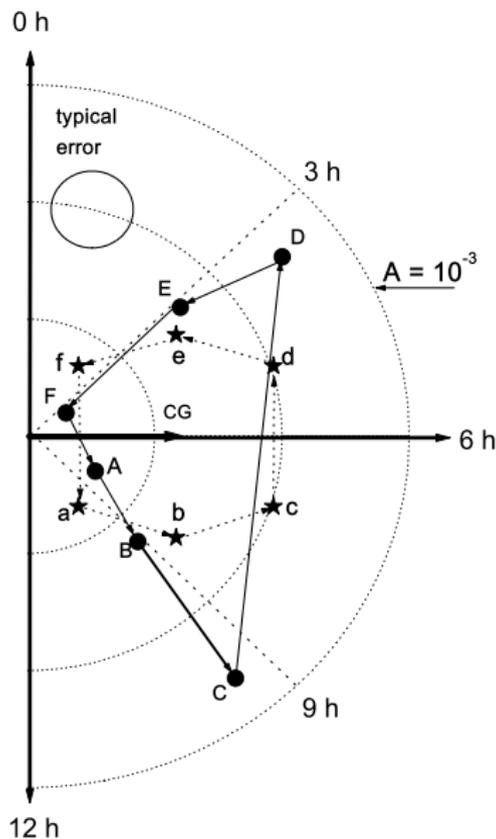
## 100 TeV

| $A_{sol} 10^4$ | $\phi_{sol} [h]$ | $P(\%)$ | $A_{sid} 10^4$ | $\phi_{sid} [h]$ | $P(\%)$ | $A_{asid} 10^4$ | $\phi_{asid} [h]$ | $P(\%)$ |
|----------------|------------------|---------|----------------|------------------|---------|-----------------|-------------------|---------|
| $2.8 \pm 0.8$  | $6.0 \pm 1.1$    | 0.2     | $2.6 \pm 0.8$  | $0.4 \pm 1.2$    | 0.5     | $1.2 \pm 0.8$   | $23.9 \pm 2.8$    | 32.5    |

- **Solar time analysis:** amplitude and phase in excellent agreement with expected Compton-Getting effect at our latitude,  $A_{sol,CG} = 3.0 \cdot 10^{-4}$ ,  $\phi_{sol,CG} = 6.0$  h.
- **Sidereal time analysis:** amplitude and phase (chance probability 0.5%) consistent with our previous results
- **Anti-sidereal time analysis:** no significant amplitude.

### FROM THE ASTRONOMICAL POINT OF VIEW

Higher counting rate when looking inside the Galaxy and a lack of events from directions corresponding to northern galactic latitudes.



*Bi-monthly solar vectors of the I harmonic (black dots), and expected ones (black stars) from the measured solar and sidereal amplitudes.*

- Expected anti-clockwise rotation of the solar vector clearly visible
- Instantaneous observed anisotropy = combination of solar and sidereal vectors
- Expected and measured rotations fully compatible within the statistical uncertainties

## 400 TeV

| $A_{sol} 10^4$ | $\phi_{sol} [h]$ | $P(\%)$ | $A_{sid} 10^4$ | $\phi_{sid} [h]$ | $P(\%)$ | $A_{asid} 10^4$ | $\phi_{asid} [h]$ | $P(\%)$ |
|----------------|------------------|---------|----------------|------------------|---------|-----------------|-------------------|---------|
| $3.2 \pm 2.5$  | $6.0 \pm 3.4$    | 44.1    | $6.4 \pm 2.5$  | $13.6 \pm 1.5$   | 3.8     | $3.4 \pm 2.5$   | $22.3 \pm 3.2$    | 39.7    |

- **Solar time analysis:** significance of the first harmonic rather marginal, but amplitude and phase still consistent with the Compton-Getting effect.
- **Sidereal time analysis:** indication of change of phase (from 0.4 to 13.6 h) and increase in amplitude by a factor 2.5 (limited significance, chance probability 3.8%)
- **Anti-sidereal time:** no significant amplitude

### FROM THE ASTRONOMICAL POINT OF VIEW

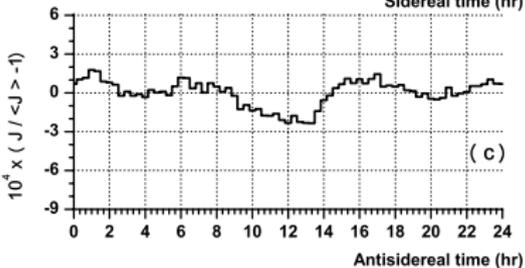
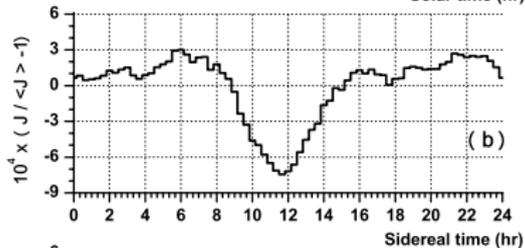
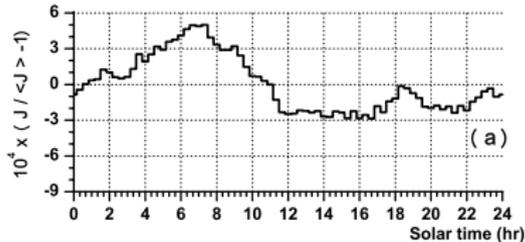
The change of phase would correspond to an excess of events towards northern galactic latitudes, inside the local galactic arm.

## SECOND HARMONIC ANALYSIS

| $E_0$<br>(TeV) | $A_{sol} 10^4$ | $\phi_{sol}[\text{h}]$ | $P(\%)$ | $A_{sid} 10^4$ | $\phi_{sid}[\text{h}]$ | $P(\%)$ | $A_{asid} 10^4$ | $\phi_{asid}[\text{h}]$ | $P(\%)$ |
|----------------|----------------|------------------------|---------|----------------|------------------------|---------|-----------------|-------------------------|---------|
| 100            | $1.4 \pm 0.8$  | $7.0 \pm 1.2$          | 21.6    | $2.3 \pm 0.8$  | $6.3 \pm 0.7$          | 1.6     | $0.6 \pm 0.8$   | -                       | 75.5    |
| 400            | $1.7 \pm 2.5$  | -                      | 79.4    | $1.5 \pm 2.5$  | -                      | 83.5    | $1.2 \pm 2.5$   | -                       | 89.1    |

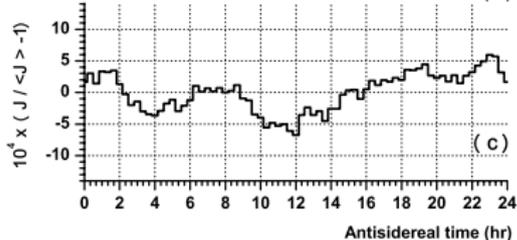
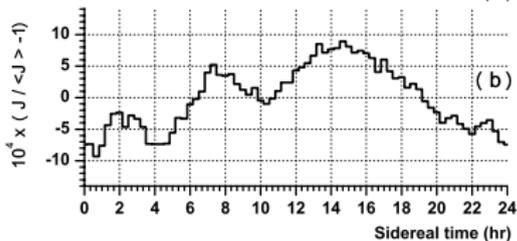
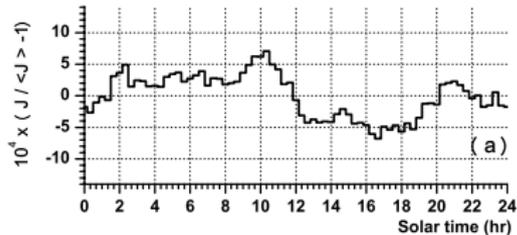
- Significant amplitude in sidereal time at 100 TeV (comparable with the first harmonic one:  $A^{II} = (2.3 \pm 0.8) 10^{-4}$ ,  $\phi^{II} = (6.3 \pm 0.7) \text{ h LST}$ ,  $P = 1.6\%$ )
- No other effects observed

## WAVE SHAPES AT 100 TeV



- **Solar wave:** Compton-Getting effect clearly seen
- **Sidereal wave:** shape in remarkable agreement with previous measurements (EAS and underground muon detectors)
- **Anti-sidereal wave:** no significant structure.

## WAVE SHAPES at 400 TeV



- **Solar wave:** Compton-Getting effect still marginally visible
- **Sidereal wave:** rather different from the one at 100 TeV: a signal is rising with max around 13-16 h LST, and increased amplitude
- **Anti-sidereal wave:** no significant structure

## Summary

- A new measurement of the CR anisotropy by EAS-TOP:
  - full data-set (8 years of data)
  - cuts in energy at 100 and 400 TeV
  - East-West analysis method
    - Includes systematical uncertainties into statistical ones
    - Loss of significance but not affected by systematics
- From the analysis of the first harmonic at 100 TeV:
  - in solar time: CG effect observed
  - in sidereal time: consistency with previous measurements
  - in anti-sidereal time: no significant effect
- From the analysis of the first harmonic at 400 TeV:
  - in solar time: CG effect marginally observed
  - in sidereal time: indication of change of phase (from  $0.4 \pm 1.2$  to  $13.6 \pm 1.5$ ) and increase in amplitude (from  $2.6 \pm 0.8 \cdot 10^{-4}$  to  $6.4 \pm 2.5 \cdot 10^{-4}$ )
  - in anti-sidereal time: no significant effect