

Radio-detection of UHECR by the CODALEMA experiment

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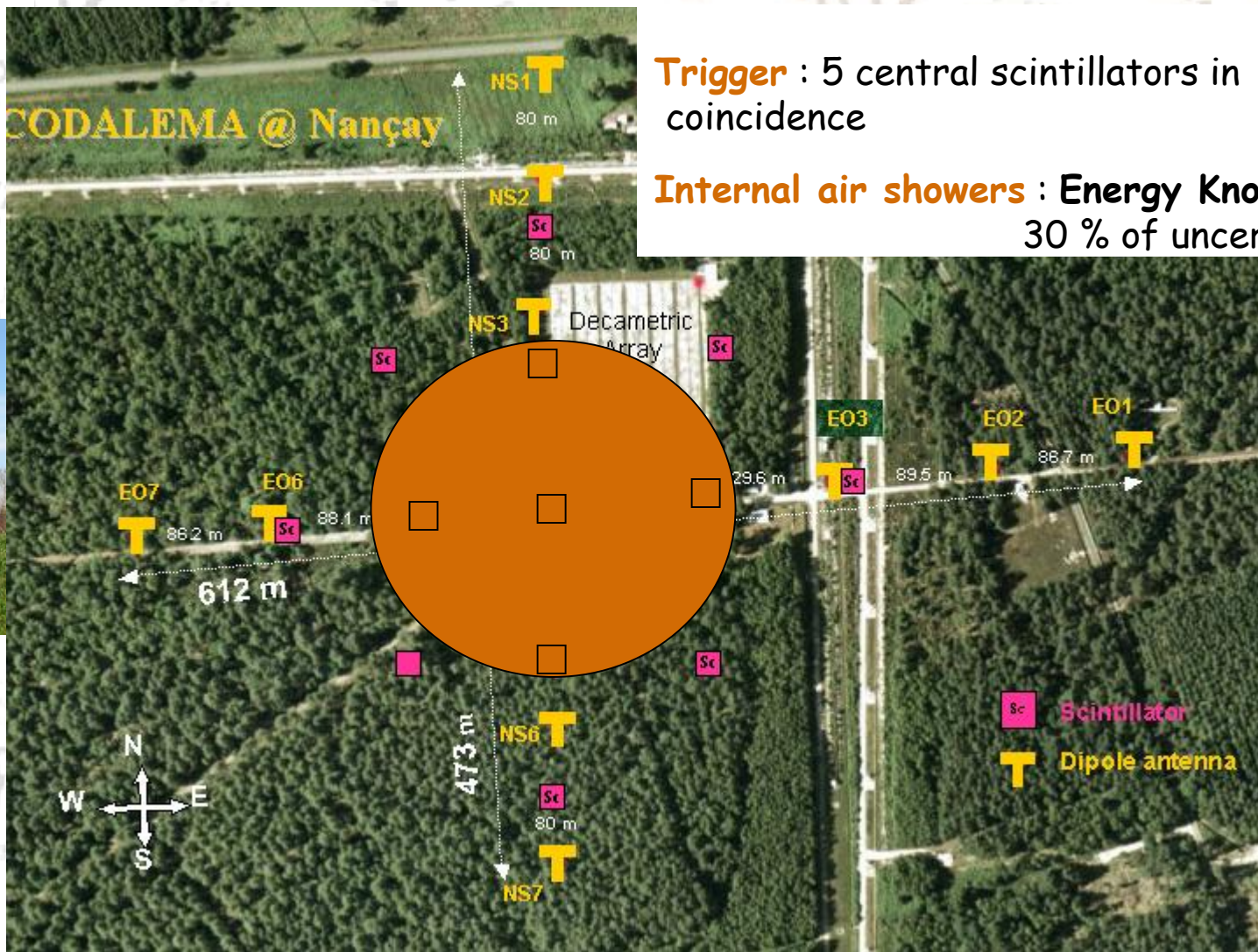
SUBATECH, Nantes, France

and the CODALEMA collaboration

Scintillator
array
(trigger)



Dipole
array



Trigger : 5 central scintillators in coincidence

Internal air showers : Energy Known
30 % of uncertainty

EW polarization of the Electric field

Statistics

since December 2006, new setup

Effective time : 170 days

Multiplicity ≥ 3 (3 antennas tagged at least) : 613 events

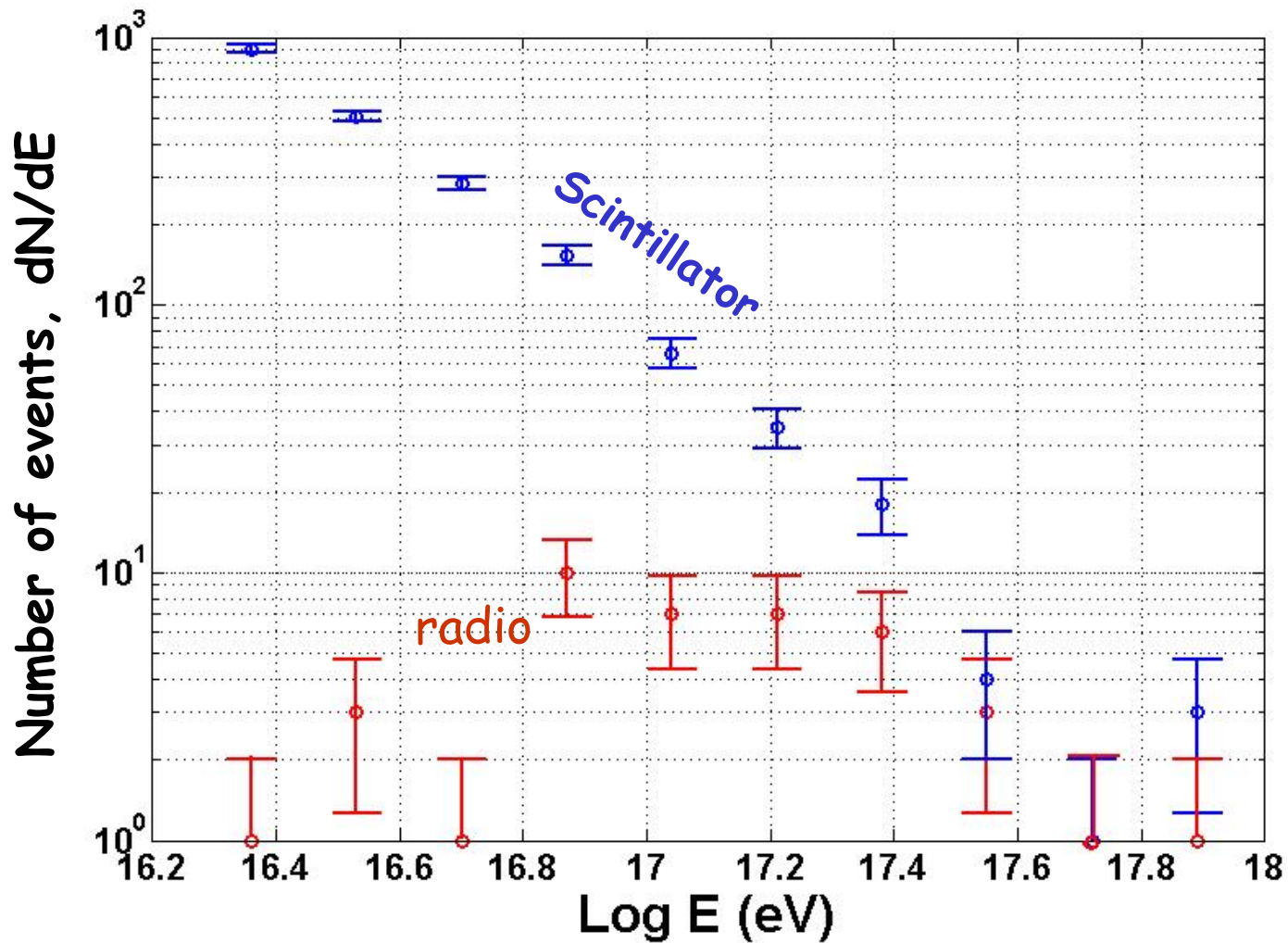
Multiplicity ≥ 3 + time and angular coincidence between both arrays

141 cosmic ray showers radio-detected

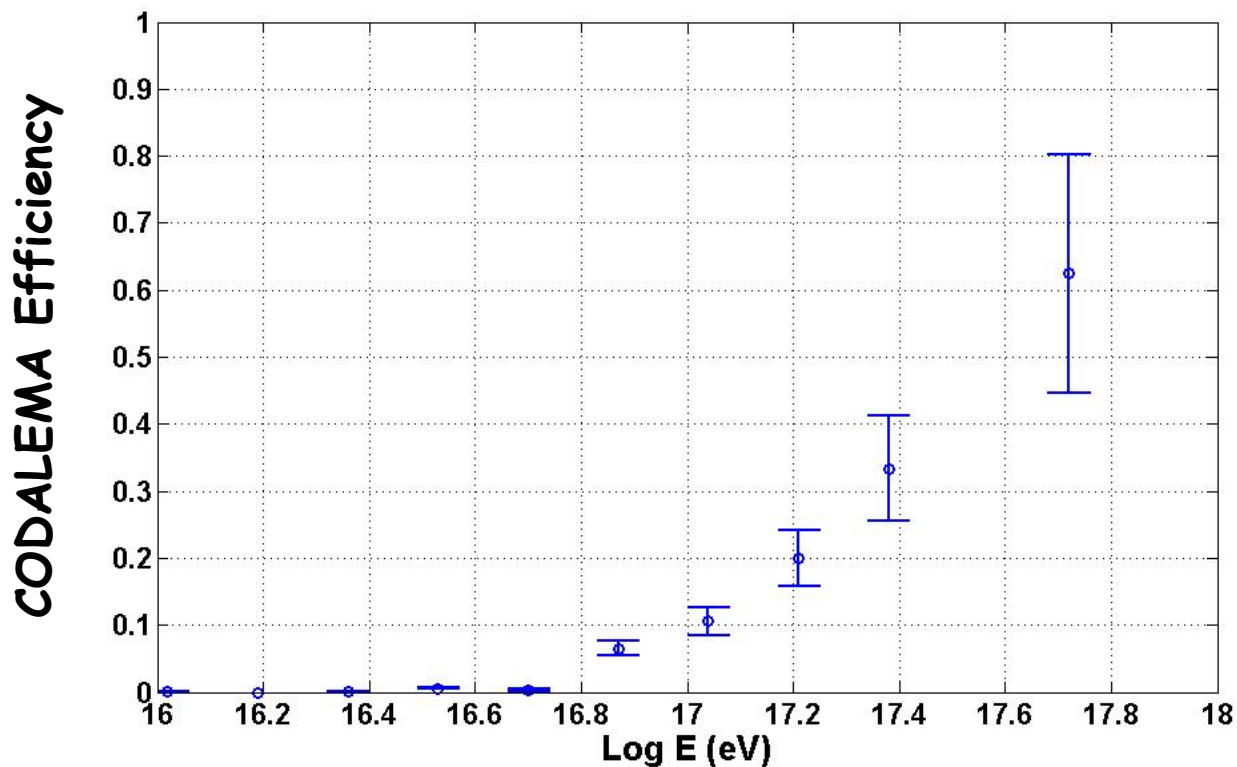
Counting rate : **0.8 events/day**

43 showers with energy known (Internal)

Energy distribution

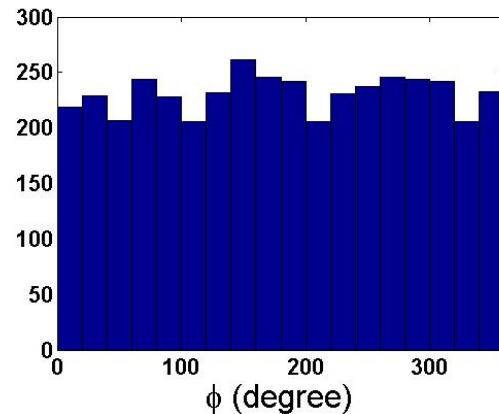
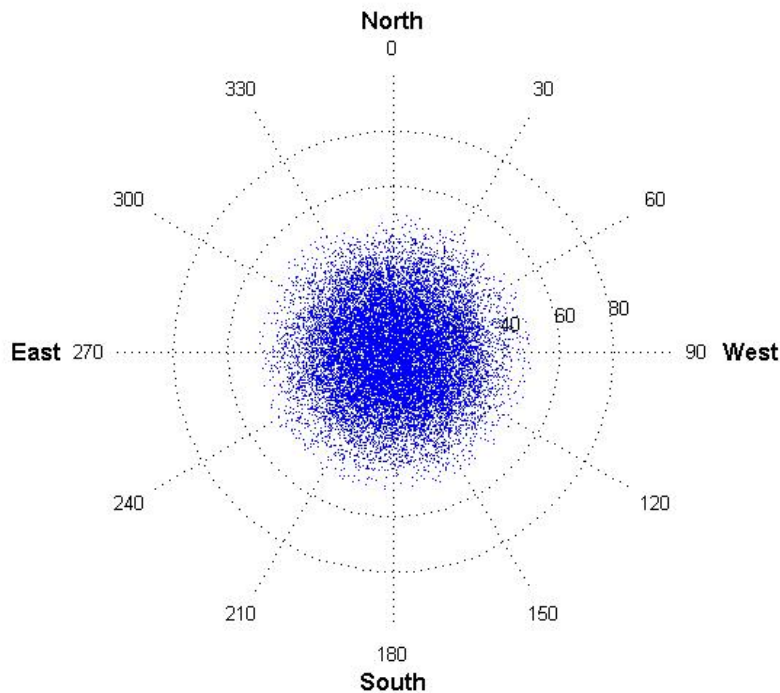


CODALEMA radio-detection efficiency (with one Electric field polarization measured)

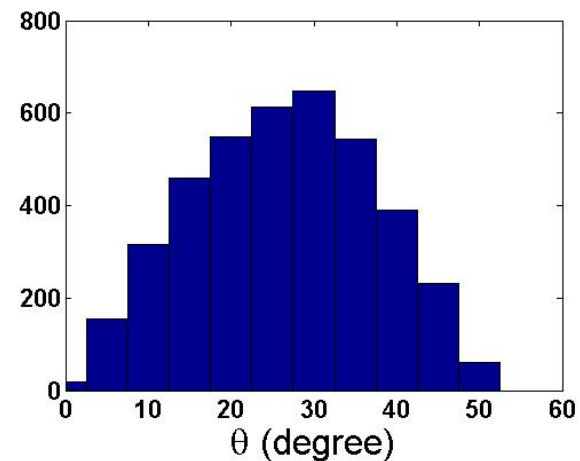


Scintillator distributions (internal showers)

Shower arrival directions calculated with the
scintillator data



Azimuthal distribution



Zenithal distribution
Limited at $\theta < 50^\circ$

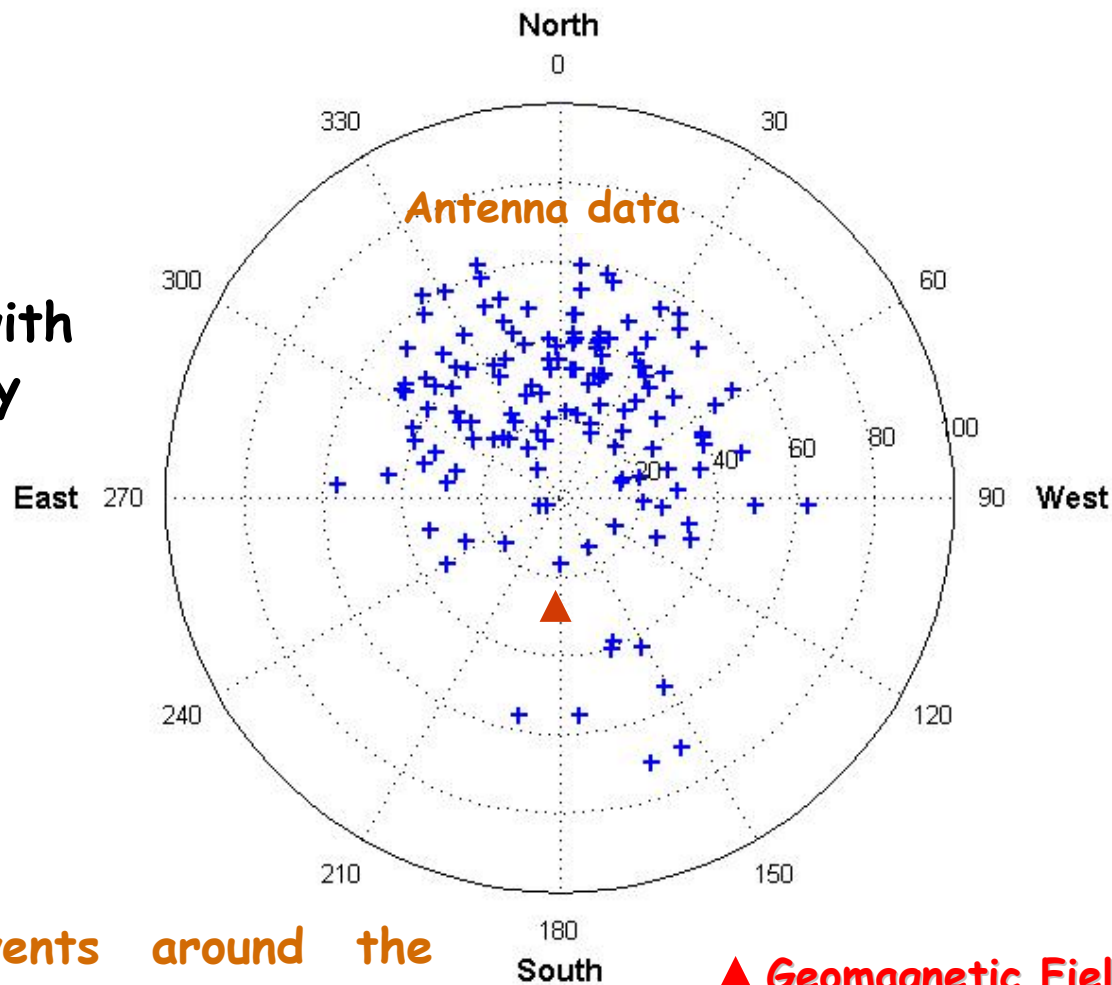
Energy threshold for scintillator array $\sim 10^{15}$ eV

Shower arrival directions

141 radio-detected
showers in coincidence with
particle detector array

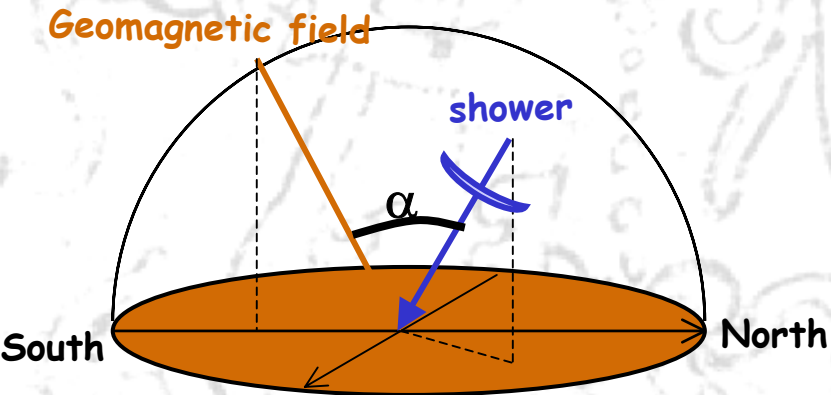
★ Shower deficit in the
South direction

★ Small number of events around the
Geomagnetic field direction

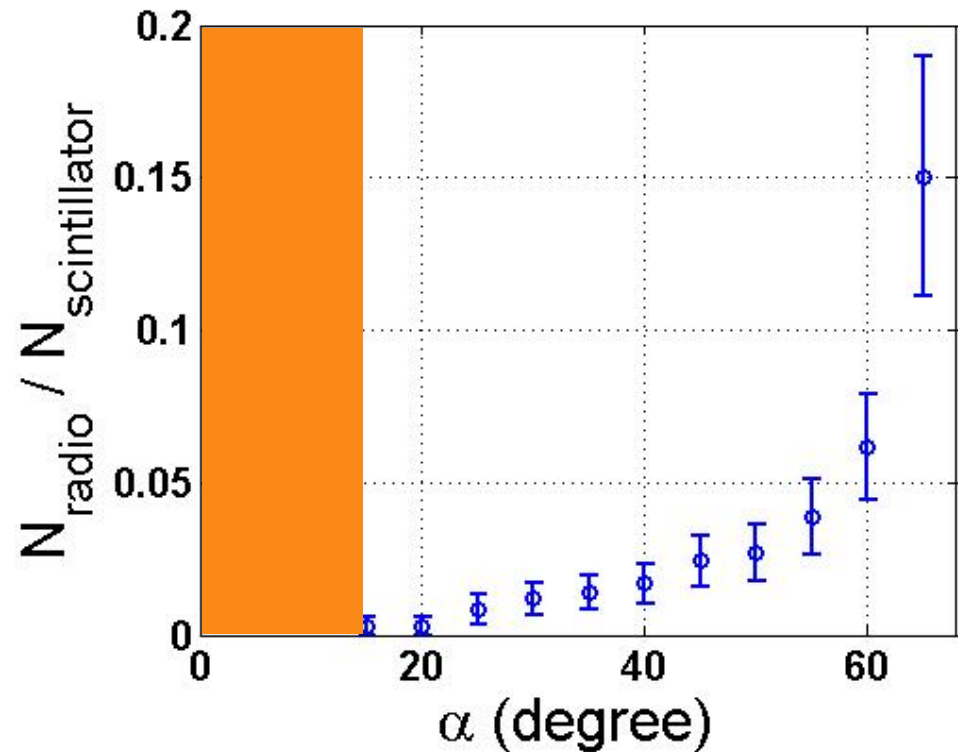


▲ Geomagnetic Field
@ Nançay
 $\theta = 27^\circ, \phi = 180^\circ$

Geomagnetic effect



α : angle between Geomagnetic field and cosmic ray arrival direction



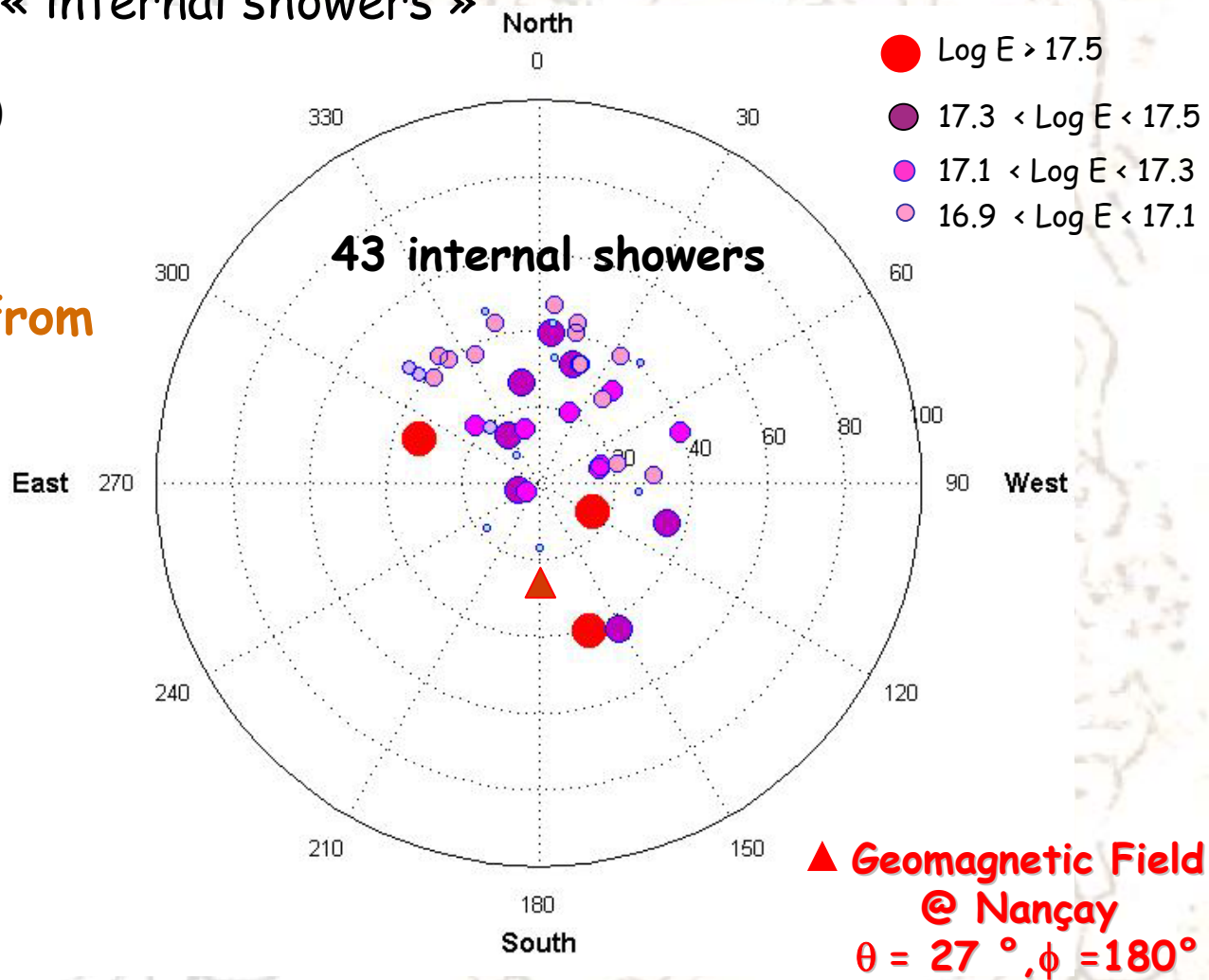
Corrected for the triggered events distribution $dN/d\alpha$

Geomagnetic effect / Energy

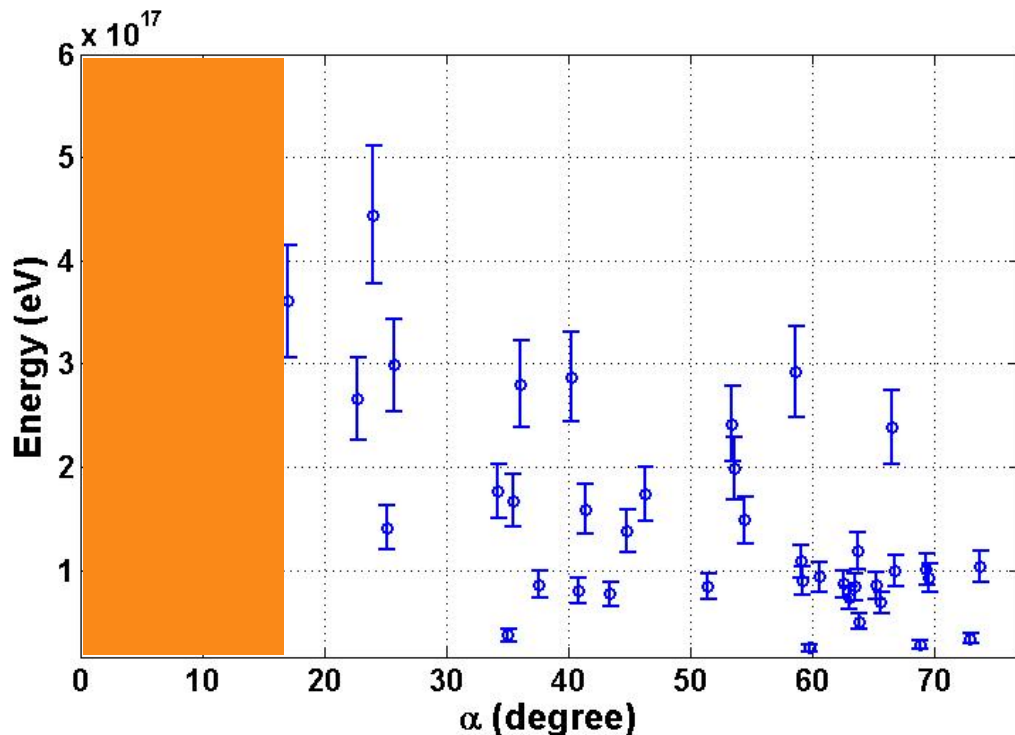
Energy known only for « internal showers »

(CIC method, precision 30 %)

Air showers detected from the South are more energetic



Energy known only for « internal showers » (CIC method, precision 30 %)



Low counting rate around
the Geomagnetic field

Deficit of low energy events
for small α



**Evidence for a geomagnetic effect
in the radio emission process
(not only geosynchrotron)**

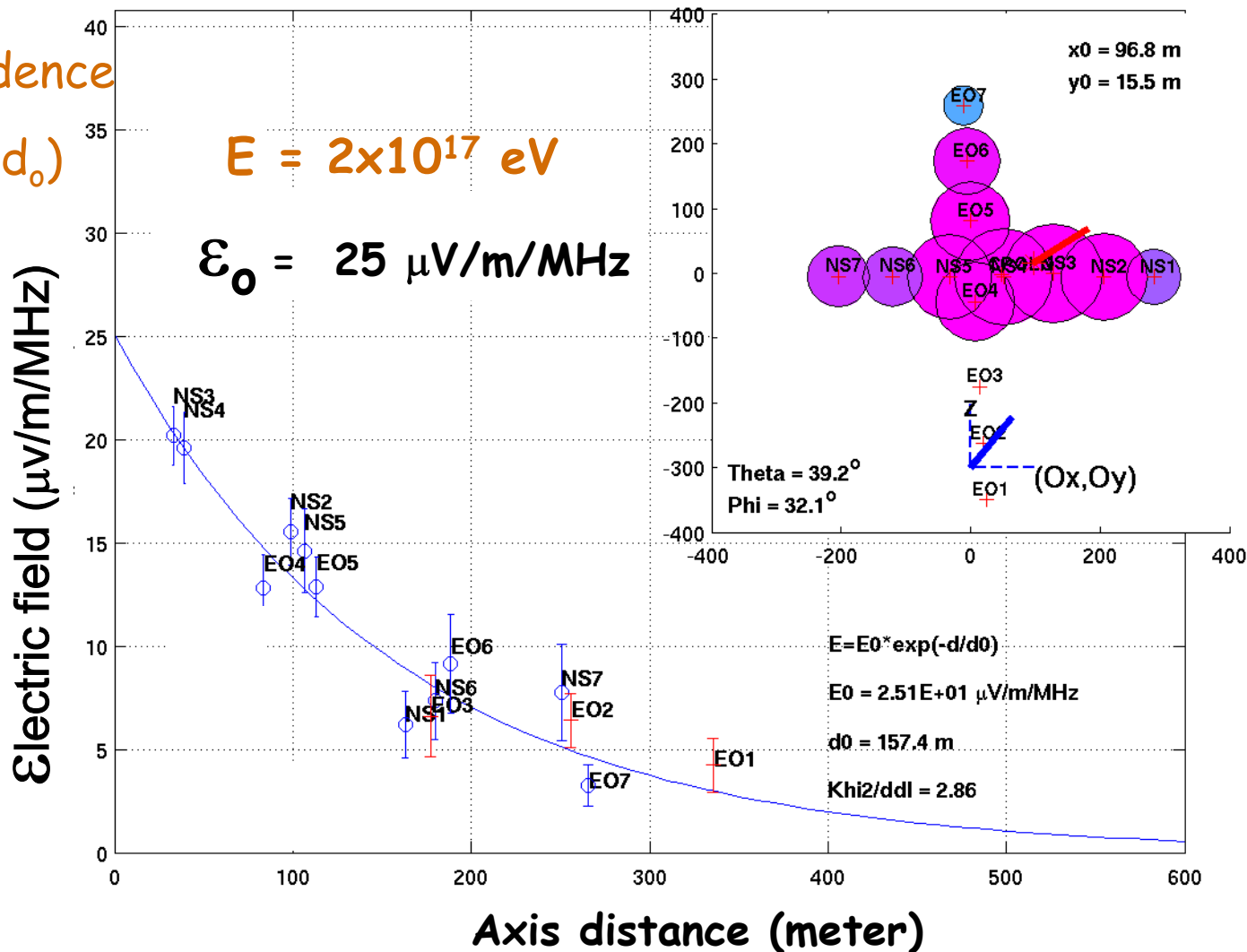
Run=789 Evt=1169 Theta=39.2° Phi=32.1°

Exponential dependence

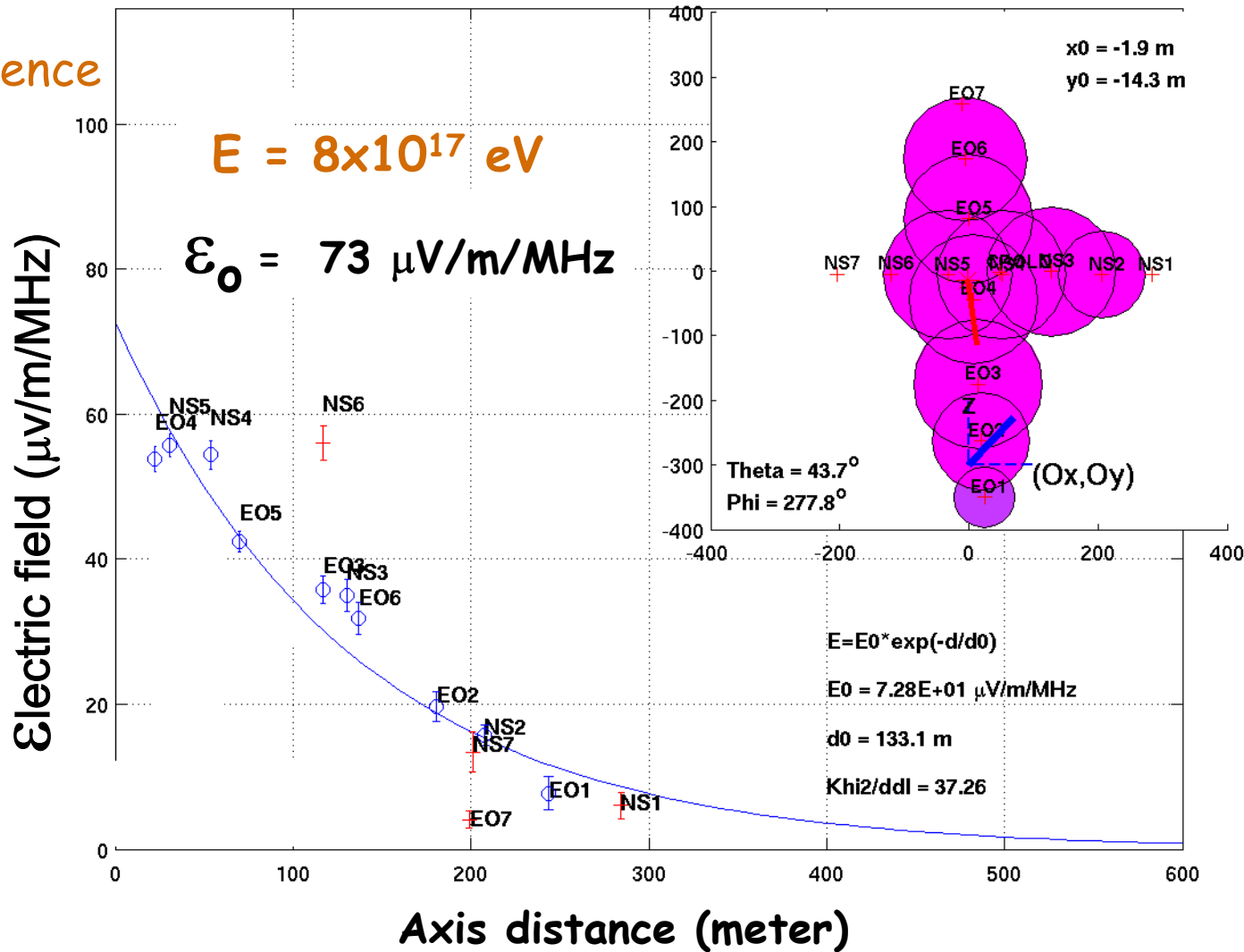
$$\epsilon = \epsilon_0 \exp(-d/d_0)$$

$$E = 2 \times 10^{17} \text{ eV}$$

$$\epsilon_0 = 25 \text{ } \mu\text{V/m/MHz}$$



Run=789 Evt=446 Theta=43.7° Phi=277.8°



Exponential dependence

$$\epsilon = \epsilon_0 \exp(-d/d_0)$$

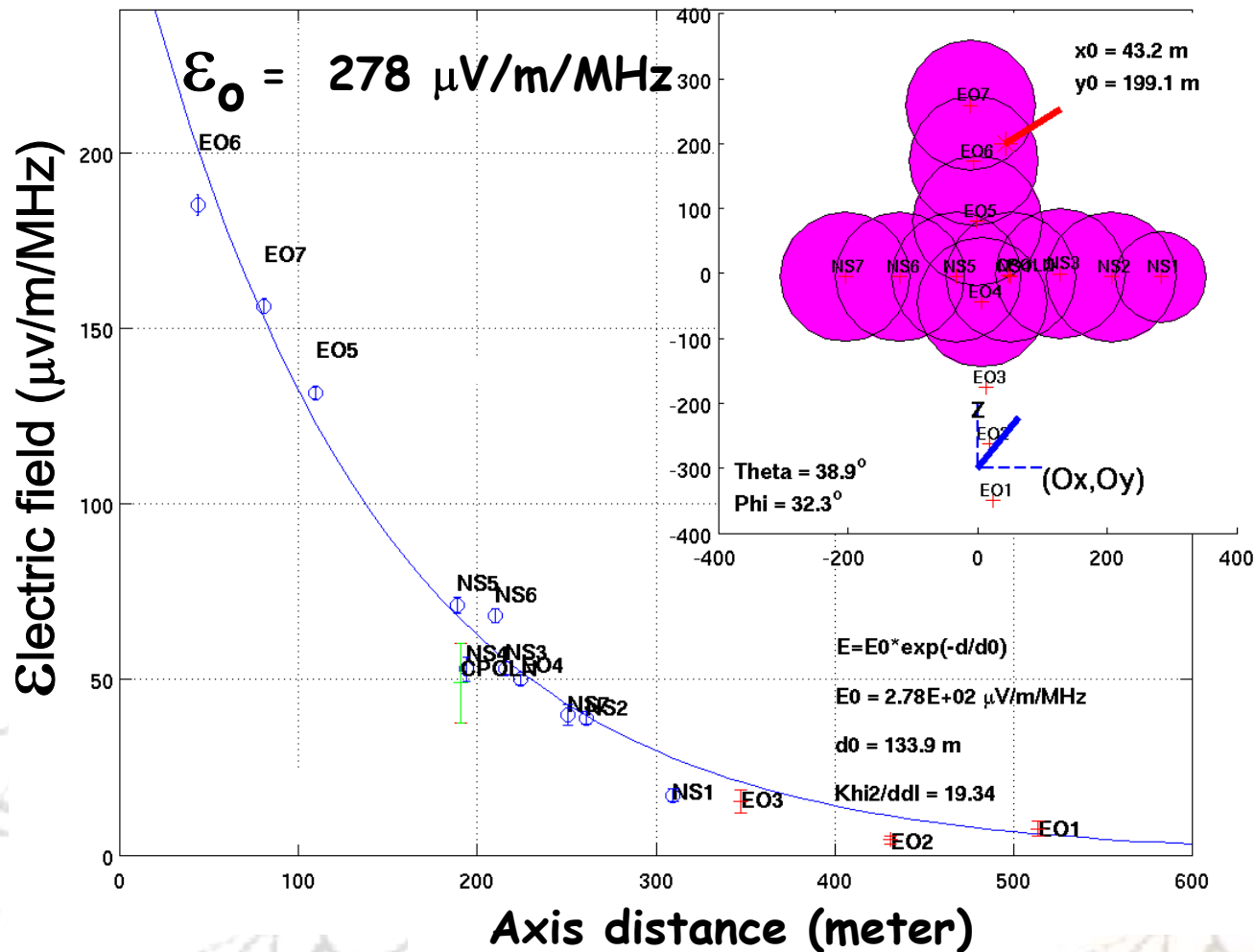
More statistic
needed
to correlate

ϵ_0 and Energy

"Giant event" (Energy above 10^{18} eV)

Unfortunately
not internal

HUGE electric
field !





CONCLUSION



- ★ CODALEMA radio detection efficiency increases with energy

Evidence for a Geomagnetic effect

radio-detection deficit close to the Geomagnetic field direction

➡ effect on the radio-detection efficiency around 10^{17} eV

➡ constraint on the emission process

Detection of all polarization could help

- ★ At the present time, we do not see clear correlation between the cosmic ray Energy and the measured electric field

➡ Larger autonomous antennas array (in 2008 @ Nançay)

CODALEMA @ ICRC 2007

see also 3 posters

- ★ Radiodetection of astronomical phenomena in the cosmic ray dedicated CODALEMA experiment
Jacob Lamblin
- ★ Design and performance of a fully autonomous antenna for radio detection of extensive air showers
Benoît Revenu
- ★ Radio detection of High-Energy cosmic rays at the Pierre Auger Observatory
A Van Den Berg