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Calibration of the EAS Radio Pulse Height

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for the LOPES Collaboration



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für
Radioastronomie



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in der Helmholtz-Gemeinschaft

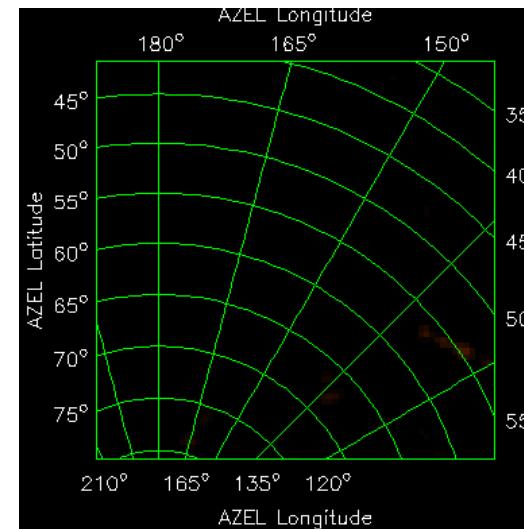
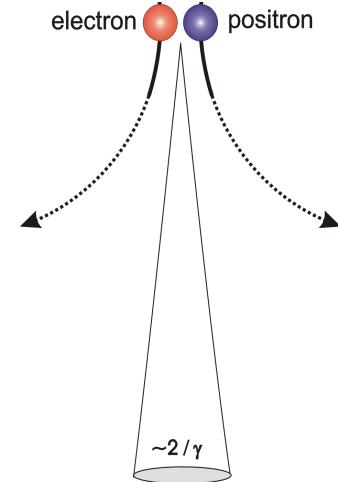


Radio Emission from Air Showers

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- Air showers emit short, intense radio pulses
- Radiation due to geomagnetic emission process e.g. geosynchrotron
- Coherent emission at low frequencies
- Measuring the radio emission from air showers could give several benefits:
 - Higher duty cycle than fluorescence telescopes
 - Effective RFI suppression allows measuring in polluted (populated) areas
 - Data integrated over the shower evolution, can be complementary to particle detectors
 - High angular resolution possible





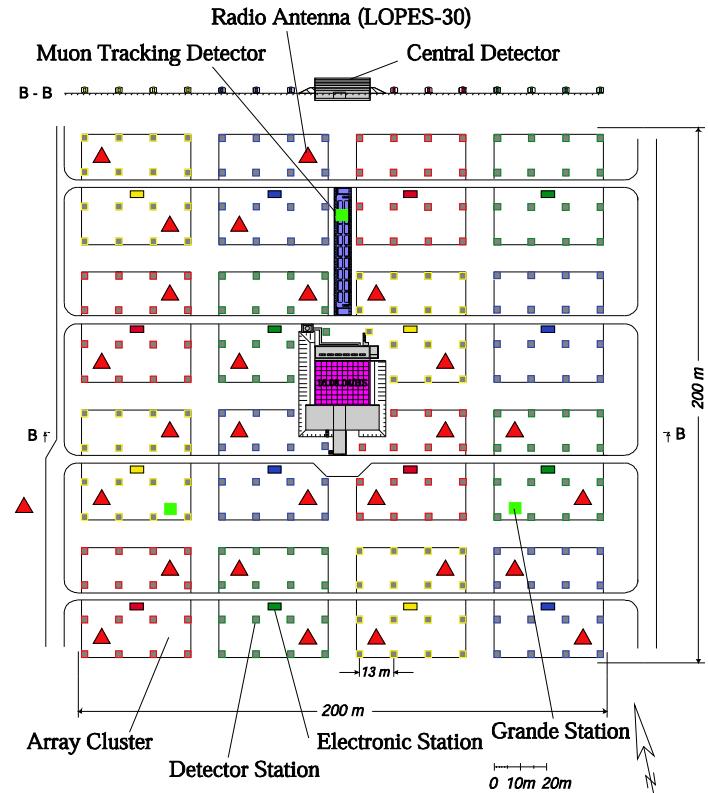
LOPES

(LOFAR Prototype Station)

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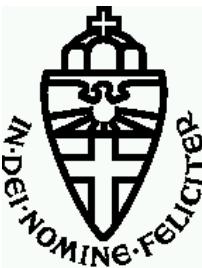
- Set up at the KASCADE-Grande site
- Frequency range of 40 – 80 MHz
- Triggered by large event trigger
- 10 antennas in the first phase, 30 antennas in second phase
- Goals:
 - Develop techniques to measure the radio emission from air showers
 - Determine the radiation mechanism of air showers
 - Calibrate the radio data with theoretical and experimental values from an existing air shower array



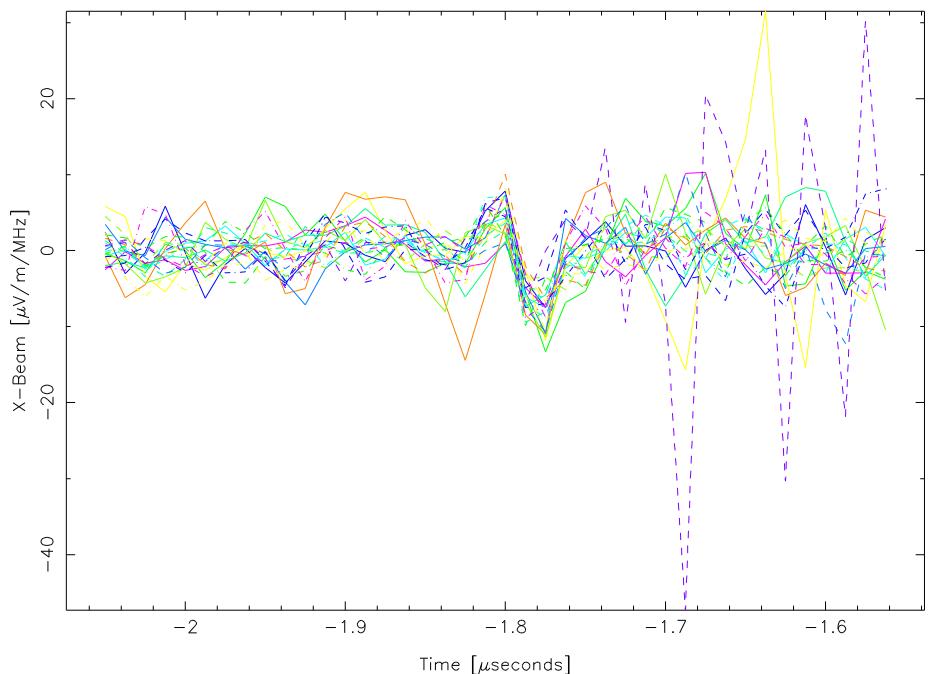


Radio Data

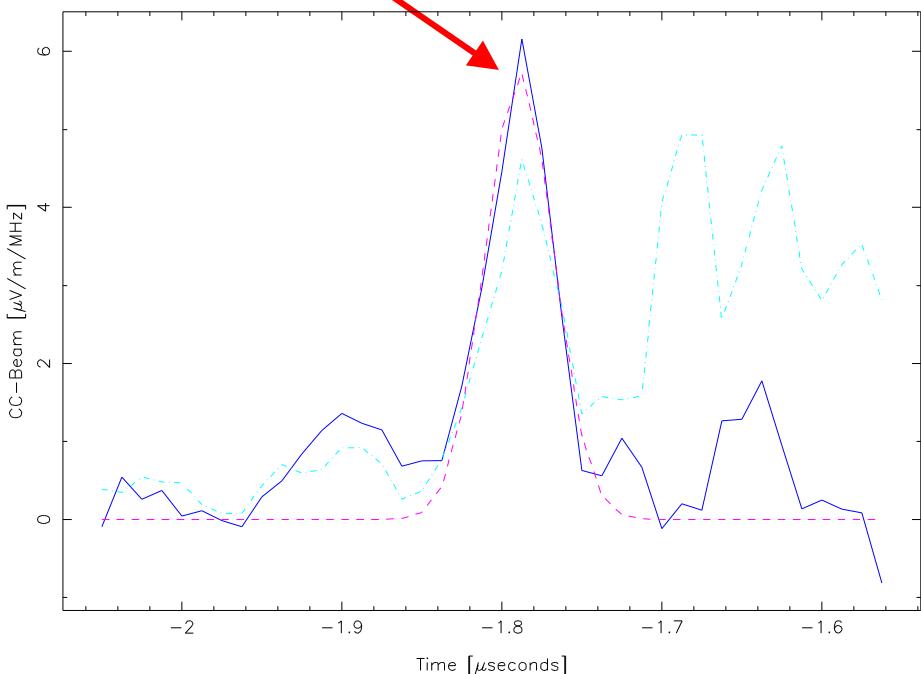
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This height is what we want



Single antenna traces



After beam-forming



Field Strength Calculation

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- The field strength is calculated by:

$$\varepsilon = \sqrt{\frac{4\pi\nu\mu_0}{G_{(\theta, \phi, \nu)}c} \frac{1}{A_{ele(\nu)}R_{ADC}}} V_{ADC}$$

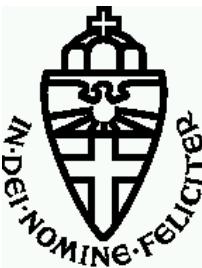
(With: ε : field strength, ν : observation frequency, G : antenna gain, A_{ele} : amplification (gain) of the electronics, R_{ADC} : ADC impedance, and V_{ADC} : voltage at the ADC.)

- V_{ADC} is the measured value
- A_{ele} and G are calibration values that need to be measured
- All other values are either constants or determined by the experiment



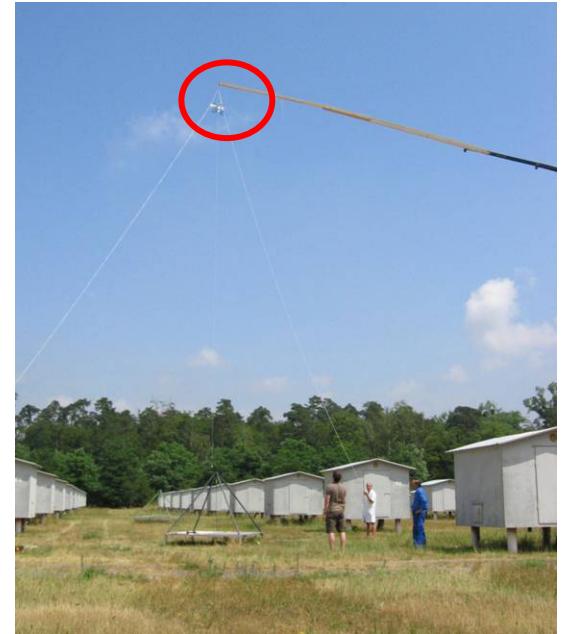
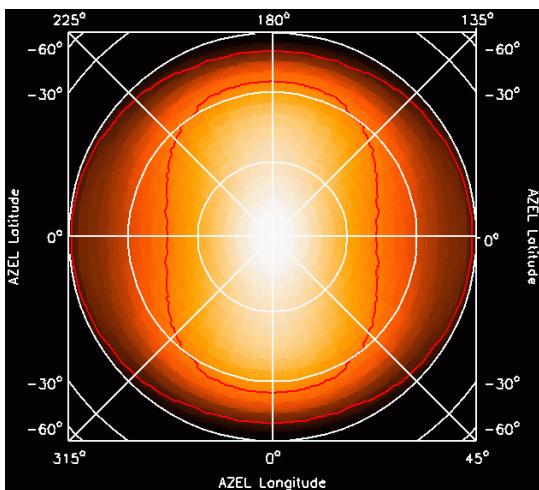
Calibration Measurements

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- Antenna gain from simulations
- Electronic Gain from measurements with reference source
 - Also mitigates errors of the antenna simulations

$$\epsilon = \sqrt{\frac{4\pi\nu\mu_0}{G_{(\theta, \phi, \nu)}c}} \frac{1}{A_{ele(\nu)}R_{ADC}} V_{ADC}$$





LOPES30 Data

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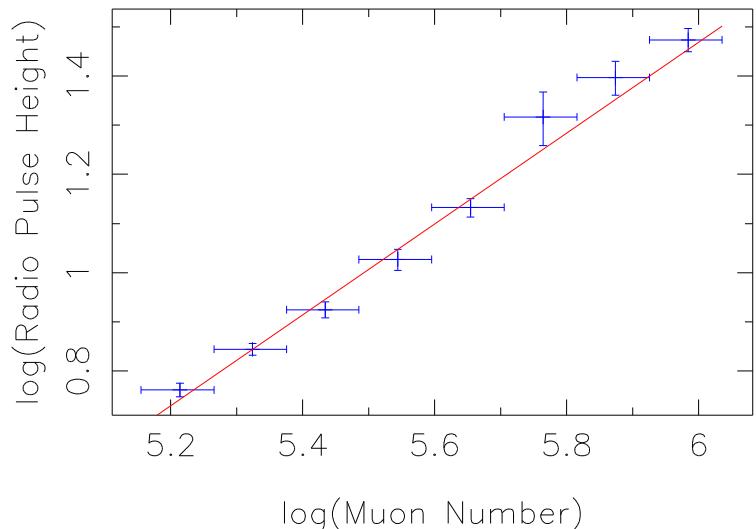
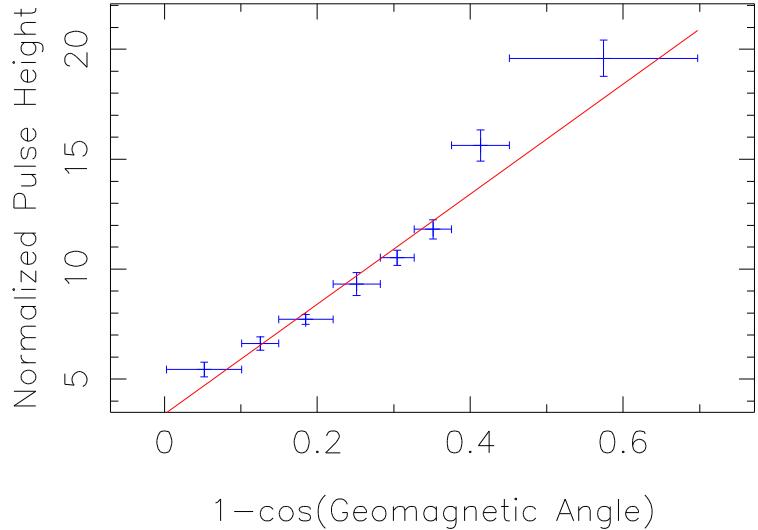
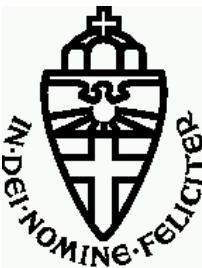
This slide was not in my presentation, but it should have been...

- Used data from November 2005 to September 2006
- used selection for further study:
 - *good KASCADE data*
 - KASCADE array processor didn't fail
 - distance of the core to the array center < 91m
 - good "age parameter"
 - truncated muon number $> 10^{5.2}$
 - not during thunderstorm
 - zenith angle $< 50^\circ$

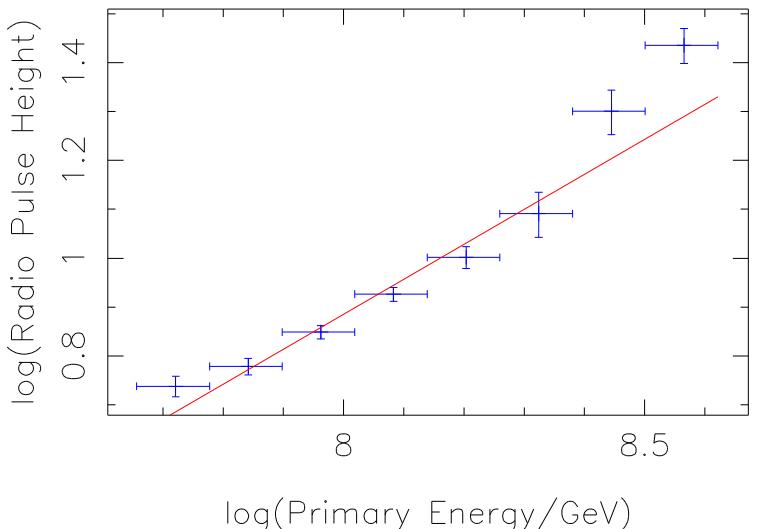
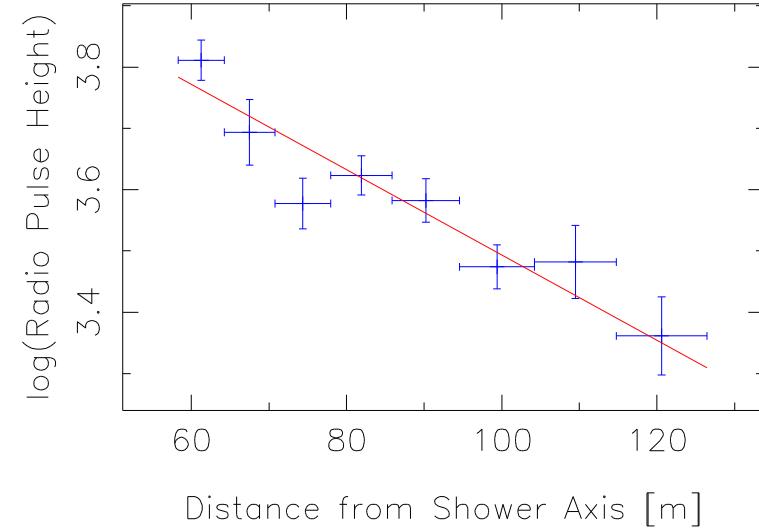


LOPES30 Pulse Height Dependence

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or





LOPES30: Field Strength Parameterization

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$$\begin{aligned}\varepsilon_{est, N_\mu} &= (55.5 \pm 5.8) \left[\frac{\mu V}{m \text{ MHz}} \right] (1 + (0.08 \pm 0.013) - \cos(\alpha)) \\ &\quad \times \exp\left(\frac{-R_{SA}}{(145 \pm 31)m} \right) \left(\frac{N_\mu}{10^6} \right)^{(0.98 \pm 0.03)} \\ \varepsilon_{est, E_p} &= (10.9 \pm 1.1) \left[\frac{\mu V}{m \text{ MHz}} \right] (1 + (0.16 \pm 0.02) - \cos(\alpha)) \cos(\theta) \\ &\quad \times \exp\left(\frac{-R_{SA}}{(202 \pm 64)m} \right) \left(\frac{E_p}{10^{17} eV} \right)^{(0.94 \pm 0.03)}\end{aligned}$$

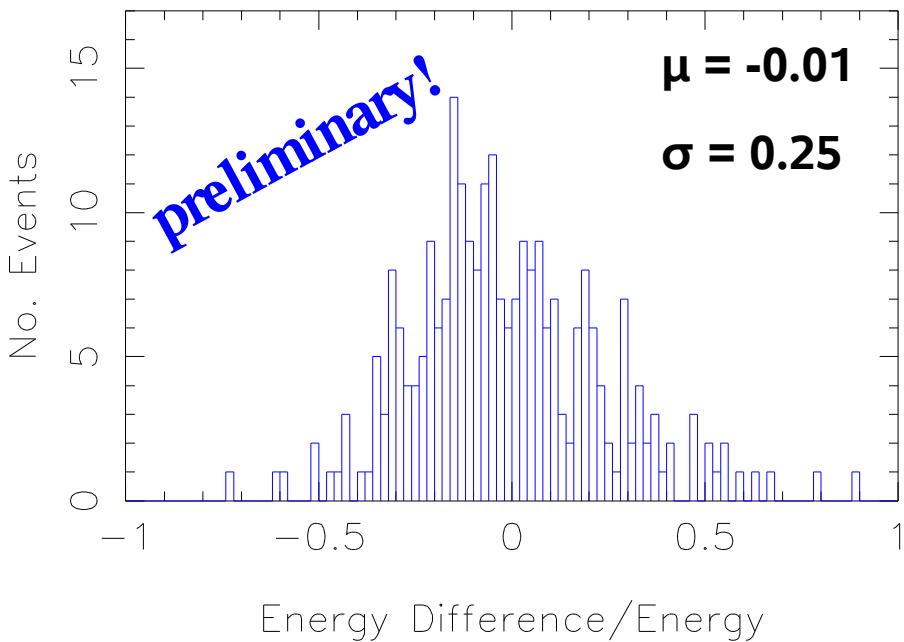
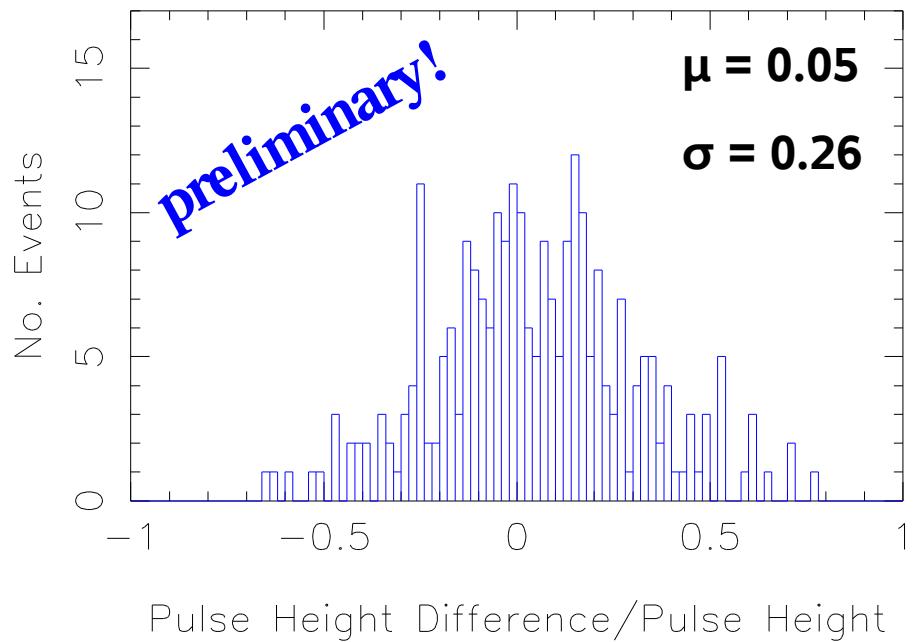
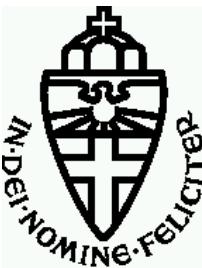
(ε_{est} : EW-pol field strength per unit bandwidth, α : geomagnetic angle, θ : zenith angle, R_{SA} : mean distance antennas \leftrightarrow shower axis, N_μ : truncated muon number, E_p : primary particle energy)

preliminary!



Statistical Spread

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Summary

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- LOPES has demonstrated that radio measurement of air showers is a viable method
- Full end to end calibration of the electronics is the key to good field strength determination
- The radio pulse height can be parameterized as a function of α , θ , R_{SA} , and N_μ or E_p
- This parameterization can be used to determine the primary particle energy from radio data alone.



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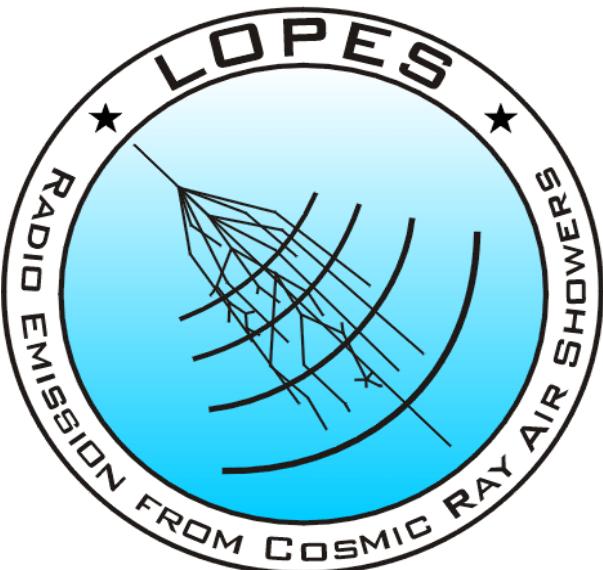
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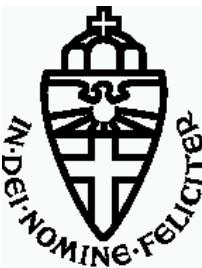
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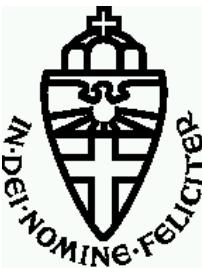
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LOFAR

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- Digital radio interferometer for the frequency range of 10 - 270 MHz
- Array of 77+ stations of 96 simple antennas
- Fully digital: received waves are digitized and sent to a central computer cluster
 - Digital radio interference suppression
 - Ability to store the complete radio data for a short amount of time
 - This allows to form beams after a transient event has been detected, combining the advantages of low gain and high gain antennas
- LOFAR will be a good tool to measure the radio emission from air showers

