

A Prototype Device for Acoustic Neutrino Detection in Lake Baikal

Baikal Collaboration



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Abstract

In April 2006, a 4-channel acoustic antenna has been put in long-term operation on Lake Baikal. The detector was installed at a depth of about 100m on the instrumentation string of the Baikal Neutrino Telescope NT200+. This detector may be regarded as a prototype of a subunit for a future underwater acoustic neutrino telescope. We describe the design of the acoustic detector and present first results obtained from data analysis. The device runs also in 2007, in common mode with NT200+.

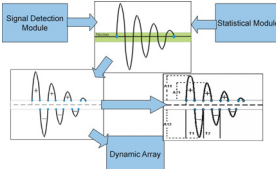
Regimes of operation

The module is designed for operation together with the Baikal Neutrino Telescope NT200+. There are 3 regimes of operation of the instrument:

- Transmitting of one second sample of data from all hydrophones to the shore computer centre after trigger signals from NT200 and outer strings.
- Online search for short acoustic pulses of definite shape, which can be interpreted as signals from quasi- local sources.
- An autonomous analysis of acoustic background statistics.

Signal Detection and Classification Procedure

The program performs online a search for signals, that exceed the threshold, using calculated statistical characteristics and the calculated thresholds.



The program distinguishes the following properties of signals: time of registration, duration, number of periods, the maximal amplitude and others. Impulses are classified by these parameters and the information about them is stored in corresponding dynamic arrays.

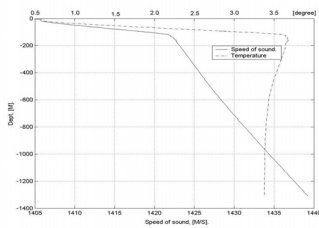
Reconstruction of Direction

For signals from distances >50m the acoustic front is flat. We measure time coordinates of each signal: $(t_1, t_2, t_3, t_4)_{exp}$. The procedure of obtaining θ, φ consists in minimization of the functional

$$S = \sum_{i=1}^4 [t_i - t_i^{exp}]^2 \frac{1}{\sigma_i^2} \quad \text{where} \quad t_i = t_0 + \frac{(r_i - r_0) \cos \theta}{V_s}$$

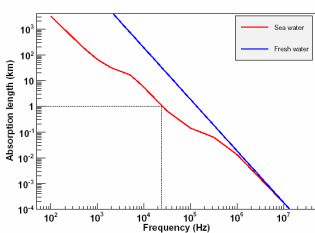
- t_i - Start point of a signal in time series
- r_0 - Position of center of the pyramid
- r_i - Positions of hydrophones
- t_0 - The moment of passage of sound front through r_0
- σ_i - Time measurement error
- V_s - Speed of sound in water

Speed of Sound in Baikal Water



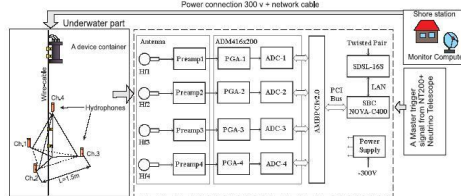
Sound speed and temperature as function of depth in Baikal water.

Acoustic Absorption Length



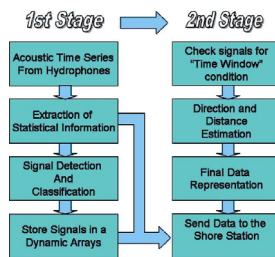
Absorption length of sound in freshwater (Baikal) and seawater. Signals from UHE showers are mainly from 20-30kHz.

The Acoustic Device



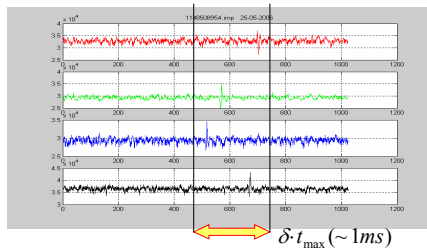
Schematic view of the underwater 4-channel digital device for detection of acoustic signals from high energy neutrinos. Installed at depth of 100m below surface, operated in common with the Baikal Neutrino Telescope NT200+.

Online Data Handling



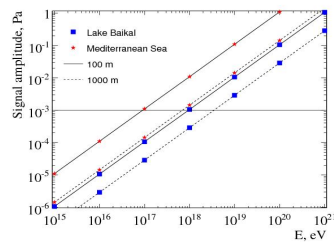
Block scheme of data handling, performed by the PC underwater.

Pulse Selection



Example of experimental bipolar pulses. Those signals which satisfy to $t_1 < \delta t_{max}$ will be accepted. Here $t_{max} = L_2/V_s$ is the propagation time of a signal from one hydrophone to another, L_2 is the distance between two hydrophones, V_s - sound velocity at the depth where antenna is placed. We set $\delta=1.2$.

Acoustic signals from UHE cascades

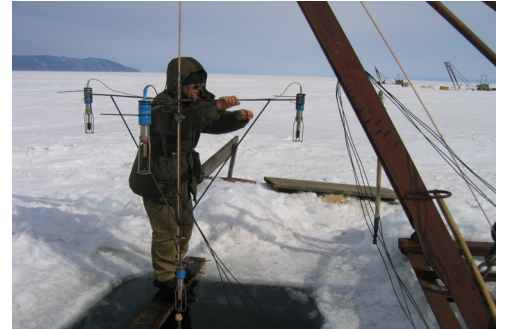


Dependence of expected acoustic signal amplitudes generated by high-energy showers versus shower energy. Seawater compared to Baikal freshwater, for 100m and 1000m distance to shower. Indicated is a noise threshold of 1mPa for Lake Baikal.

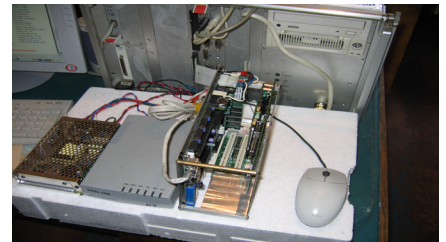
Conclusion

The results of the experiment have demonstrated the feasibility of the proposed acoustic pulse detection technique in searching for signals from UHE cascade. Although the Baikal water temperature is close to the temperature of its maximal density, the absence of strong acoustic noise sources in the lake's deep zone, the very low absorption and direction background suppression ("top" vs. "down") may allow for neutrino detection in Lake Baikal with a threshold of order 10^{18} - 10^{19} eV. In 2007, this prototype device was again deployed, and operates in common with the Neutrino Telescope NT200+.

Acoustic Antenna of the device

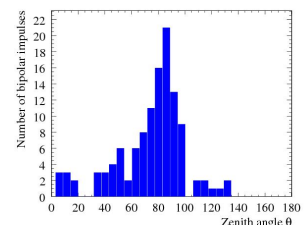


Electronics



The underwater components: Power supply, DSL-modem, PC104-Computer with PCI/ADC Card (left to right).

Reconstructed Signal Direction



Distribution of reconstructed zenith angles triggered events (April - May 2006). Note: Acoustic signals from the lake surface and the upper 100m above the device have zenith angles between 0° and 90°.

References

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