

# IceTop Tank Response To Muons

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## The IceTop Air Shower Array

IceTop is an air shower array of ice— extend the DOM's dynamic range (for de-Cherenkov counters [1, 2]. Each of its curtails, cf. [3]). rent 26 stations shown in Fig. 1 is made up of two IceTop tanks. The tank ice is viewed by two standard IceCube digital optical modules (DOMs) (see Fig. 2). They consist of a 10" Hamamatsu R7081-02 photo multiplier tube (PMT) and processing and readout electronics. Two different types of digitizers are used to process the PMT signal: a fast pipelined ADC (FADC) with 255 samples of 25 ns each, and two Analog Transient Wave Digitizer (ATWD) chips, with three channels of up to 128 samples of about 3.6 ns each. The three channels are configured with different pre–amplification factors to

Surface map of IceCube 2007 (as built)

#### Simulating The Tank Response

The DOM responses are simulated by • Generating showers with CORSIKA [5]: – hydrogen and helium primaries  $-10 \, GeV \le E_{primary} < 415 \, GeV$  $-\Theta_{primary} < 70 \, deg$ • Tank simulations (GEANT4 based, [6]):

- generate and track Cherenkov light/secondaries in the tank.
- account for tank properties: reflectivities of sides and top, ice quality. - account for PMT quantum efficiency.

The total spectrum peak position is at 247 pe, the best estimate for the VEM is determined as the mean of a Gaussian fit to the black histogram, 236 pe. The ratio gives a correction factor of about five percent, which is in agreement with the correction factor obtained from the muon tagger measurements. Currently, this correction factor is assumed to be the same for all IceTop tanks.



#### Calibration using stopping muons

Stopping muons in IceTop tanks:

• 210 MeV  $\leq E_{kin,max}(\mu) \leq 430$  MeV

•  $E_{kin,max}(e) < 53 \,\mathrm{MeV}$  $\Rightarrow$  electron range less than 25 cm  $\Rightarrow$  most decay electrons contained in tank!

Are there decay signals?



Figure 6 shows the results: • light grey: Total simulated charge spectrum • blue: Single muons charge contribution • black: Same, but with  $\Theta_{\mu} < 17 \, deg$ 



• red triangles: Total measured charge spec-**Figure 6:** MC simulated charge spectrum for trum from DOM 21–63. DOM 21–63.

#### VEM Calibration Run Parameters

Periodic special IceTop calibration runs serve two purposes: one, to calibrate the conversion from integrated waveform to vertical equivalent muon (VEM) for each DOM in each tank, and two, to monitor the DOM response's time dependence. The calibration run configuration differs from the regular one used for air shower data runs:

#### air shower mode:

• the high gain DOMs of both tanks in a sta-

• the simple majority trigger requires a mul-

tion are in local coincidence (LC).

tiplicity of six.

• the two DOMs in the same tank are set to different gains (in 2006,  $5 \cdot 10^6$  and  $5 \cdot 10^4$ , respectively).

singles mode (= calibration mode):

• every DOM is set to the same nominal gain of  $5 \cdot 10^{6}$ .

• local coincidence (LC) between DOMs is turned off.

COLUMN TWO IS NOT

• the simple majority trigger is disabled.

For the DOMs that are operated at the lower gain, the VEM might differ due to changes in the collection efficiency of the PMT. Currently, that effect is not taken into account.

• Two signal FADC traces were selected.

• Figure 11 shows  $\Delta t$  in blue  $\Rightarrow$  Suppress background with stringent cuts! Cuts were determined tuned with MC studies.

•  $\Delta t$  for remaining events in red, fit yields a lifetime of  $\tau = 2.06 \pm 0.16 \,\mu s$ .

The Michel spectrum is calculated by following the method outlined in [7] ("Auger method"):

1. Define a "decay" window between 1 and  $2\,\mu s$ .

2. Define a "crossing" window between 5 and  $6 \,\mu s$ .

3. Collect FADC's second signal charges for both time windows.

4. Subtract "crossing" from "decay" spectrum  $\Rightarrow$  Michel spectrum

• Figure 12 compares measured spectrum (red symbols) with simulated spectrum.

• Qualitatively agreement between mea sured and simulated spectra.

 $0.511 \le E_e(MeV) \le 52.8$ Figure 10: Feynman diagram of muon decay.

 $\tau = 2.19703 \,\mu s$ 

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### What is a VEM?

A DOM's response to a vertical muon passing an IceTop tank is defined to be one Vertical Equivalent Muon (VEM). Such a muon deposits around 200 MeV in the tank ice [4]. By finding the vertical muon signal in the measured total charge spectrum, the DOM-dependent chargeto-VEM conversion factor is determined. However, single IceTop tanks cannot discriminate between different particles or incident angles. Therefore, the relation between the measured peak position of the total charge spectrum and the VEM must be determined with simulations and a tagging telescope:

- . Determine the charge contribution of nearly-vertical muons to the overall charge spectrum with the help of a tagging telescope for muons.
- 2. By using the zenith angle acceptance of the tagging telescope, simulate the tank response
- 3. Compare both results and define the VEM contribution to the total DOM charge spectrum.

Measure charge spectra with all DOMs and extract the VEM, monitoring its stability over time.

### The Muon Tapping Telescope

tern • correcting the signal droop (implemented, but not tested)

• adjusting any residual baseline

The calibration data is processed by

### Charge Spectrum 21\_63, 2007-03-15 03:24:08 ndau plus two exponential Gauss plus one exponentia

• calibrating the charge in photo electrons • summing the charge and plotting the charge distribution • applying fits (e.g. [4]) to extract the full

- spectrum peak position
- using the conversion of that peak position value to VEM.

Charge Spectrum 21\_64, 2007-03-15 03:24:08

plus two exponential

Gauss plus one exponential

# Conclusions

#### VEM calibration:

- Procedures and tools have been established and well understood.
- Periodic calibration runs keep track of IceTop DOM's response stability.
- Muon telescope measurements and tank simulations agree well on the nearly vertical muon contribution of the total charge spectrum.
- More muon tagger measurements will help fine tune the relation between total charge spectrum peak position and VEM.
- Simulations with higher statistics will make more realistic muon tagger cuts possible.
- Significant decrease in the VEM of several DOMs is not understood yet, but at current level does not limit the array's performance.

# • correcting each raw waveform for the specific, ATWD chip-dependent pedestal pat-

VEM Calibration And Monitoring Results



Figure 3: Tagging Telescopes getting Figure 4: ... and being on duty. ready...

A portable, solar-powered muon telescope was developed to tag muons that have angles close to vertical (< 17 deg)and pass through the center of the tank. It measures signals in coincidence between two scintillator slabs 70 cm apart and records the GPS clock time stamp. Measurements were taken during the po-

200 300 400 Charge (PE)

300

250

150

**H** 400

350

200

150

200 300 40 Charge (PE)

lar season 2005/2006 on tanks deployed one year earlier. The charge spectra of the tagged muon sample are shown in Fig. 5, superimposed in blue over the total charge spectra for DOMs 39–63 and 39–64 in black. The VEM is determined to be about 95% of the full spectrum peak position.

> Figure 5: Total charge spectra (black) for tank 39b with tagged muon spectrum (blue) superimposed



Stopping muons: • Feasibility study was completed. • Comparison between data and simulation looks promising. • Further improvements in both the analysis and the simulation needed  $\Rightarrow$  supplementary calibration method Acknowledgments

This work is supported by the U.S. National Science Foundation. Grants No. OPP-0236449 and OPP-0602679. Photographs by J. Roth and L. Demirörs

#### References

[1] T. Gaisser et al. Performance of the IceTop array. In Proc. 30th Int. Cosmic Ray Conf., Mérida, Mexico, 2007. [2] Todor Stanev and Ralf Ulrich. Nucl. Phys. Proc. Suppl., 145:327–330, 2005. [3] A. Achterberg et al. Astropart. Phys., 26:155–173, 2006.

[4] M. Beimforde. Diploma Thesis, Humboldt–Universität zu Berlin, 2006. (http://www-zeuthen.desy.de/nuastro/publications/diploma/arbeiten/ ThesisBeimforde.pdf).

[5] D. Heck, G. Schatz, T. Thouw, J. Knapp, and J. N. Capdevielle. FZKA-6019. [6] P. Clem, J. Nießen and S. Stoyanov. Response of IceTop tanks to low-energy particles. In Proc. 30th Int. Cosmic Ray Conf., Mérida, Mexico, 2007. [7] P. Allison et al. Proc. 29th Int. Cosmic Ray Conf., 8:299, 2005.