



Cherenkov Telescope Array: A next generation ground-based gamma-ray observatory



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for the
CTA consortium
ICRC 2007, Merida, Mexico*



Cherenkov Telescopes worldwide today



VERITAS



MAGIC + MAGIC II

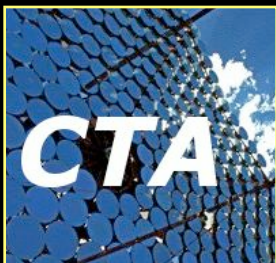


H.E.S.S. I and HESS II



CANGAROO III



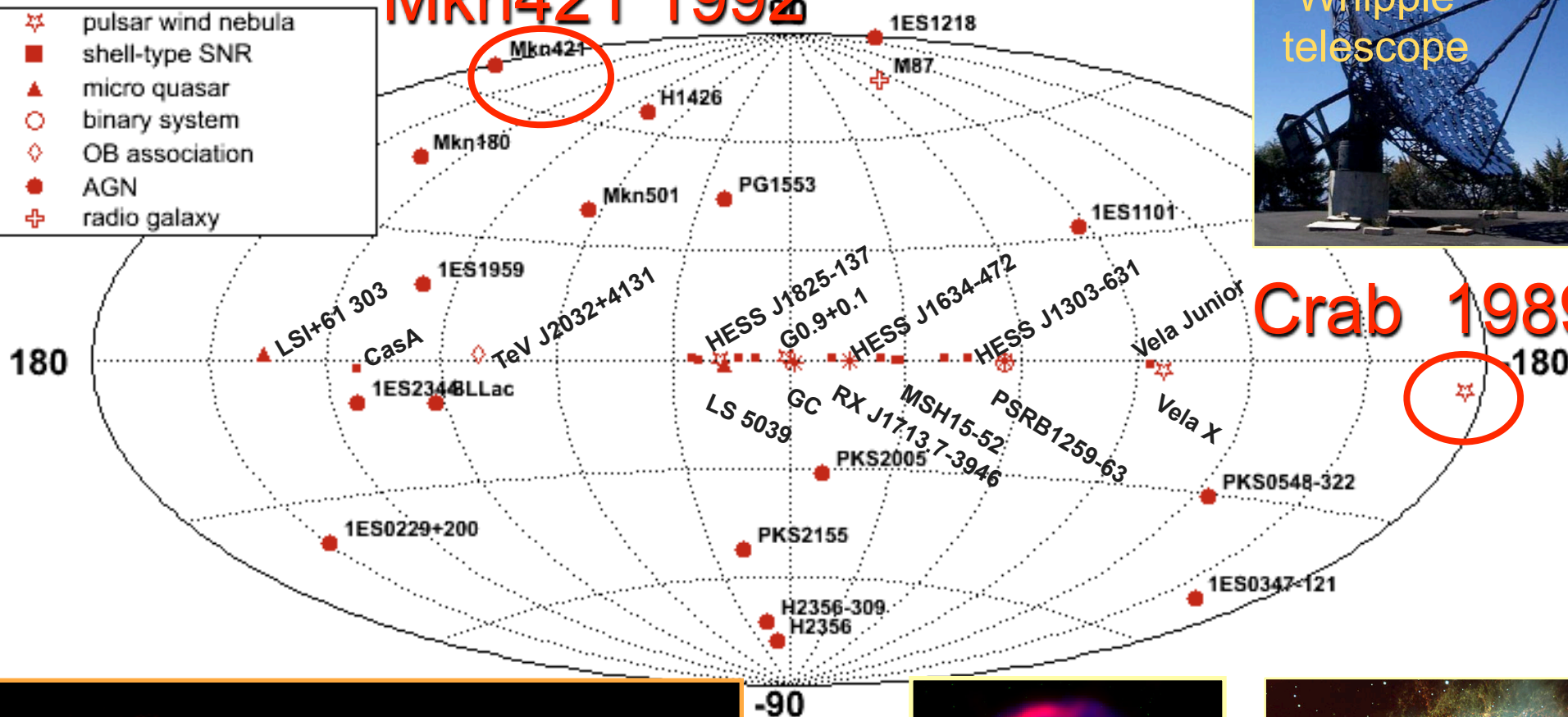


Gamma ray astronomy today: 47 sources



Mkn421 1992

- ☆ pulsar wind nebula
- shell-type SNR
- ▲ micro quasar
- binary system
- ◇ OB association
- AGN
- ⊕ radio galaxy



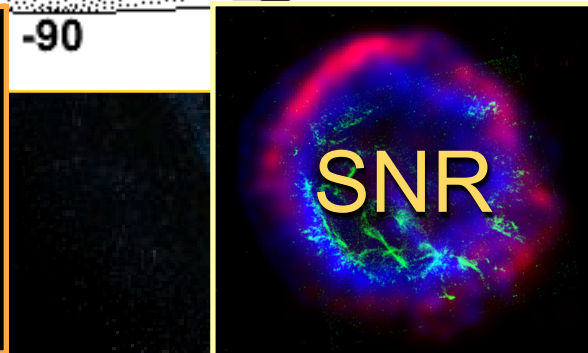
Whipple telescope

Crab 1989

Active galactic nuclei

SNR

PWN

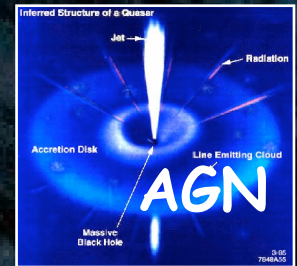




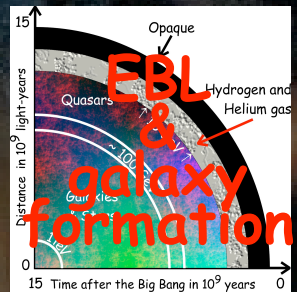
CTA characteristics and physics



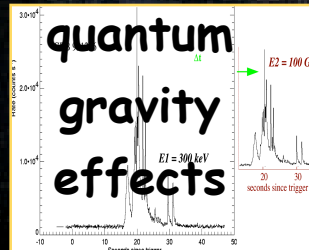
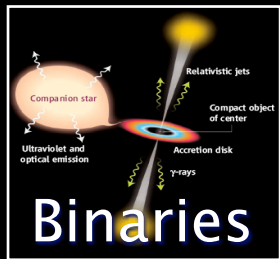
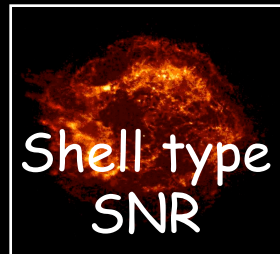
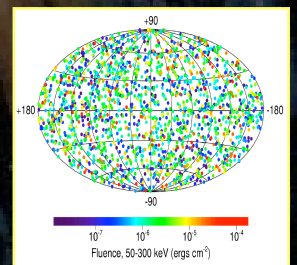
- ❖ Observation of very high energy photons
Some 10 GeV up to 100 TeV
with 1mCrab sensitivity and an improved angular resolution



- ❖ Full sky coverage (two installations)
- ❖ Design based on known technology
- ❖ Observatory open to external astronomers

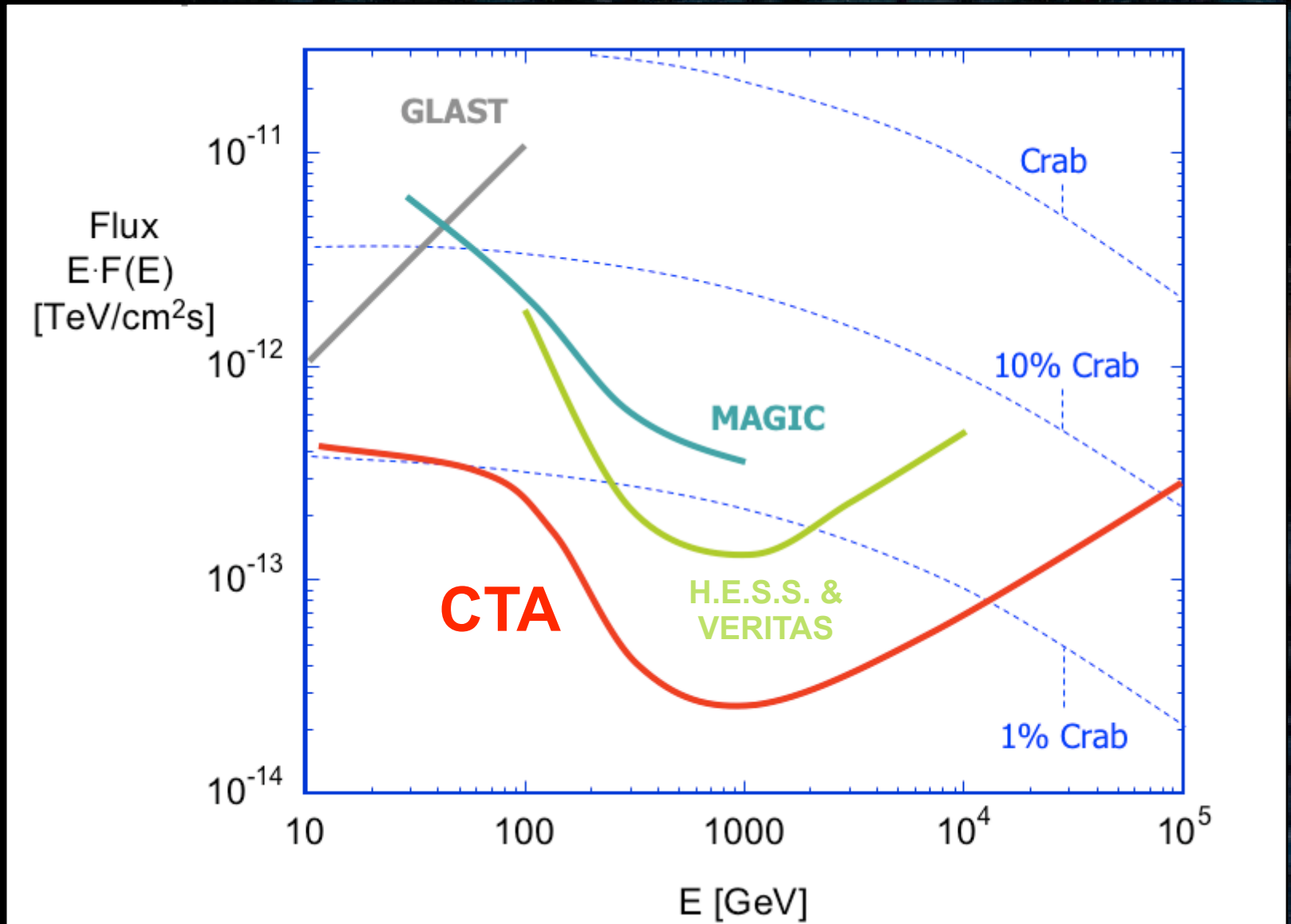


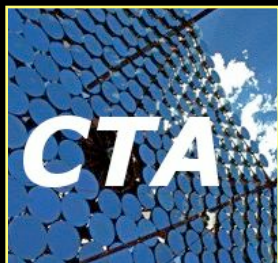
- ❖ Discovery of new source classes



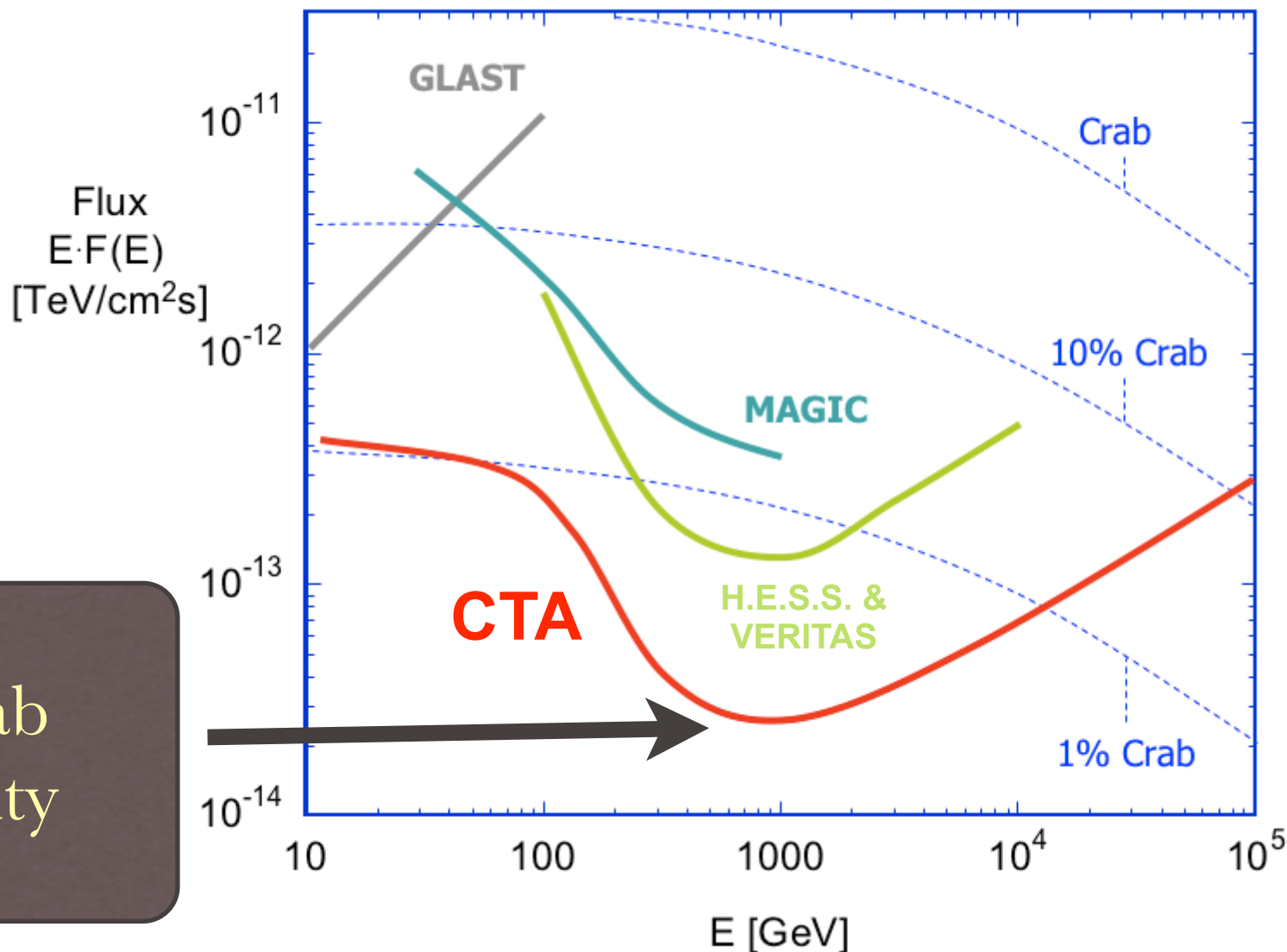


Aimed sensitivity





Aimed sensitivity



1 mCrab
sensitivity

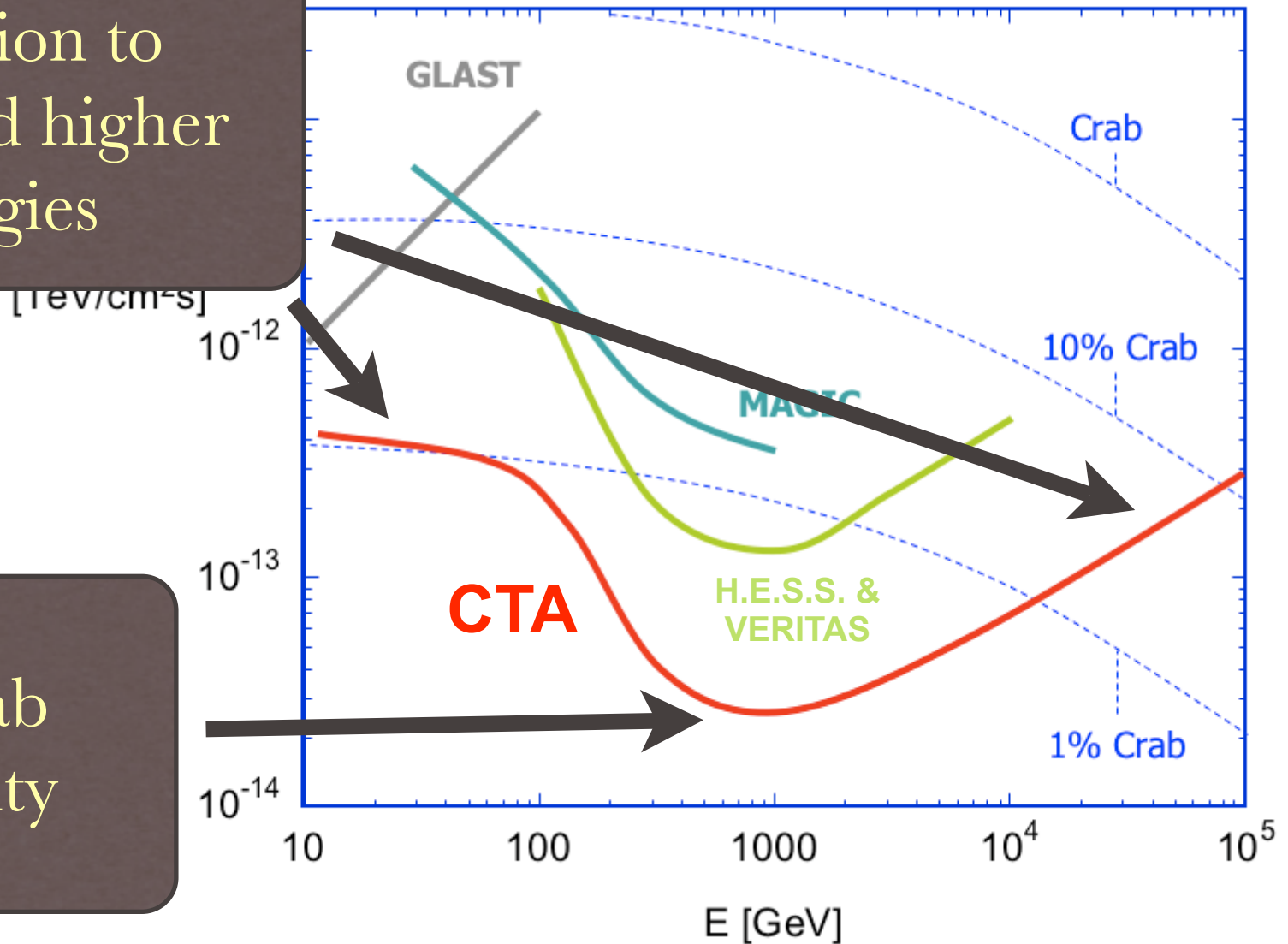


Aimed sensitivity



Extension to
lower and higher
energies

1 mCrab
sensitivity





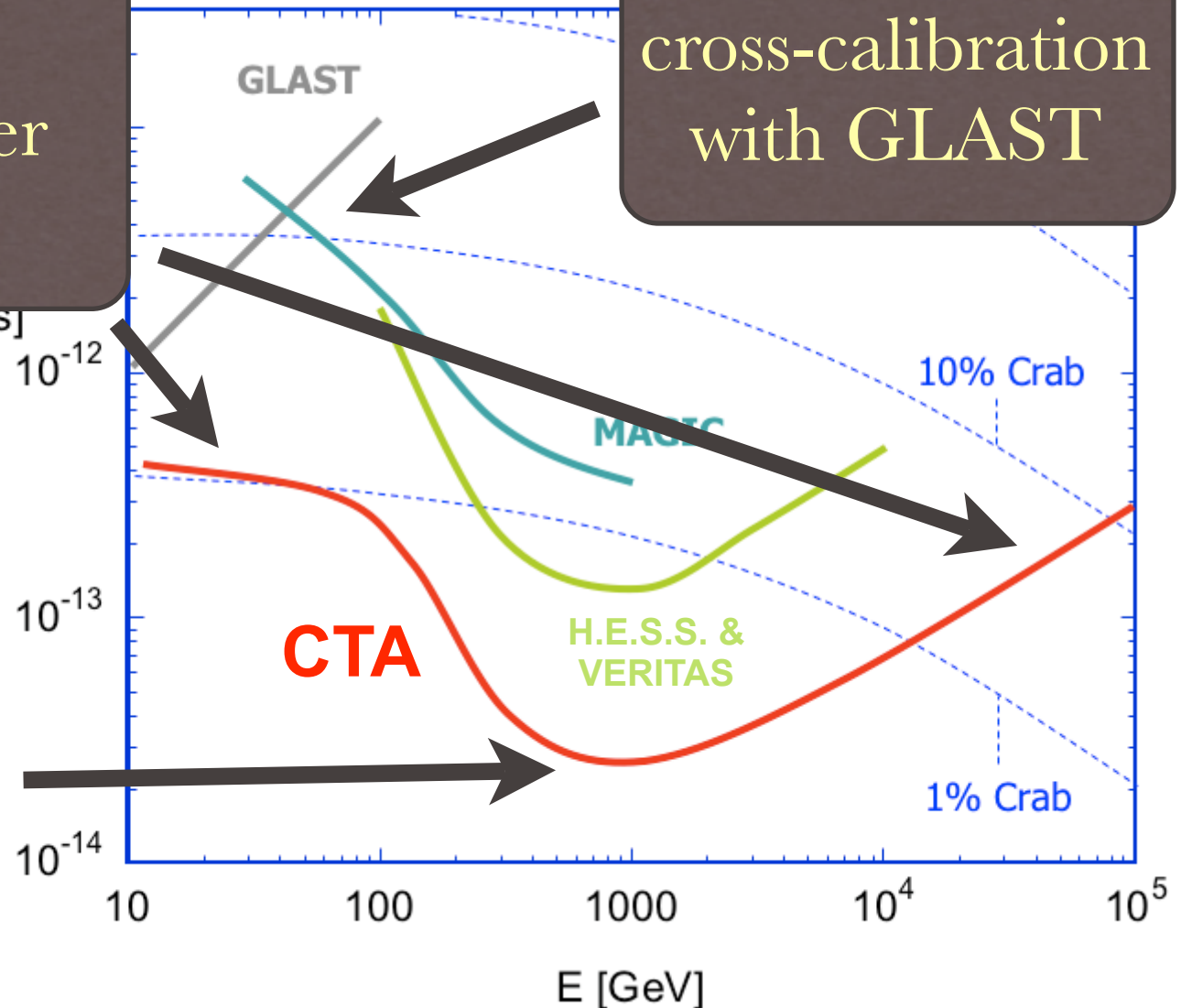
Aimed sensitivity



Extension to lower and higher energies

Overlap for cross-calibration with GLAST

1 mCrab sensitivity

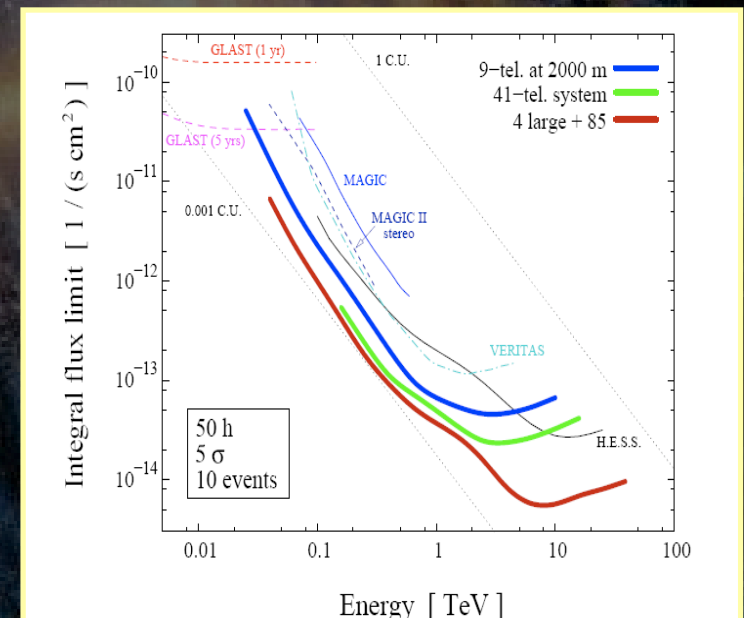
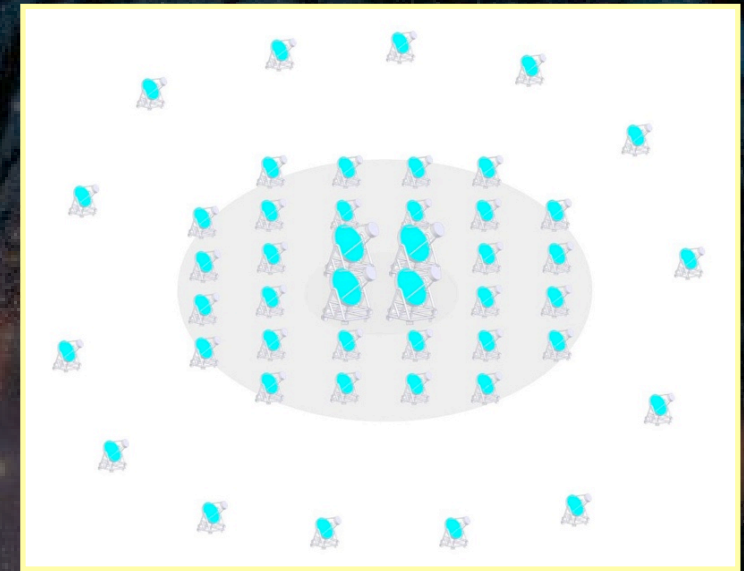




CTA configurations and first MC studies



- Different telescope sizes for different energy ranges
 - Large telescopes provide for low energy sensitivity
 - Many telescopes provide high sensitivity
 - Large area provides sensitivity at high energies
- Wide angle telescopes (FOV up to 12°)
- MC Simulation:
 - 23m diameter (5° FOV, 0.1° pixel size)
 - 10m diameter (7° FOV, 0.16° pixel size)
- Photodetectors
 - PMTs with 25% QE peak

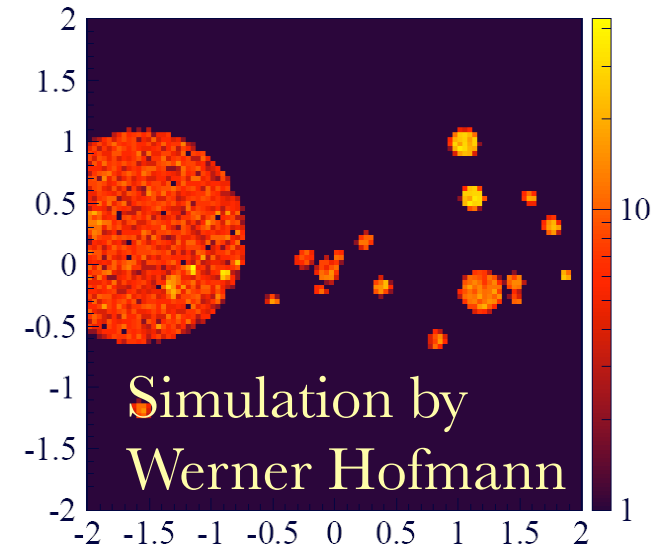
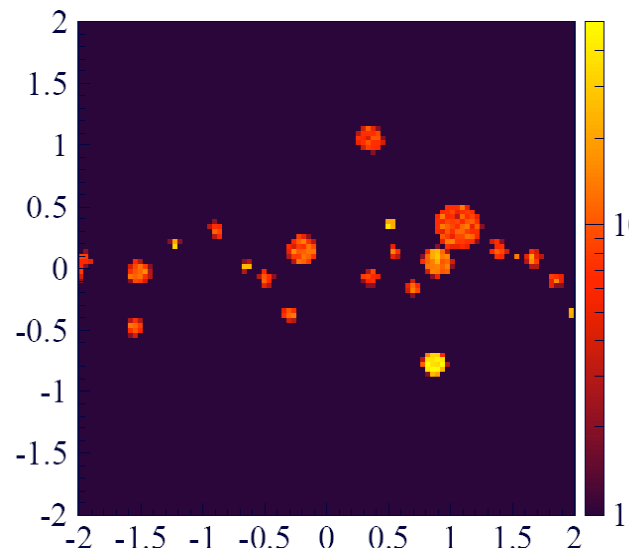
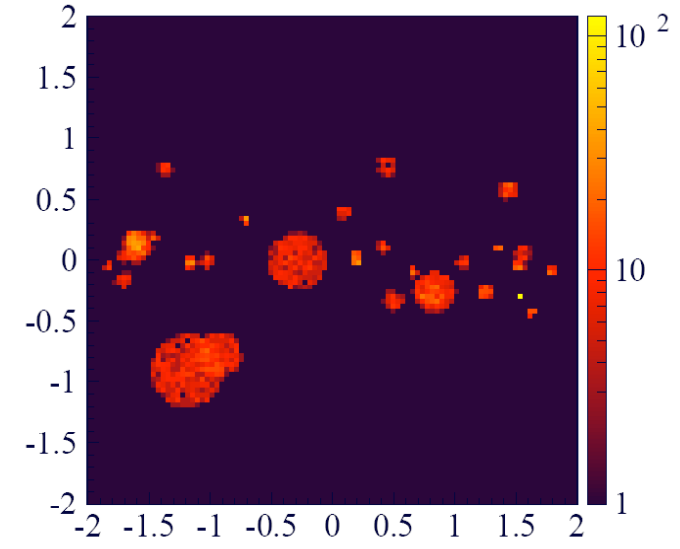
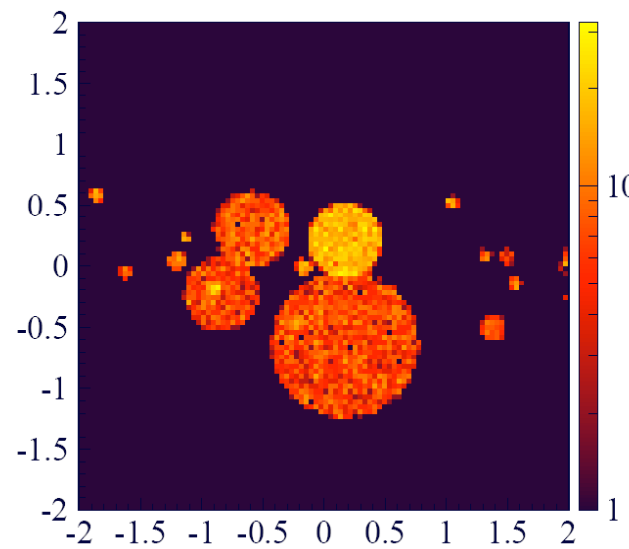


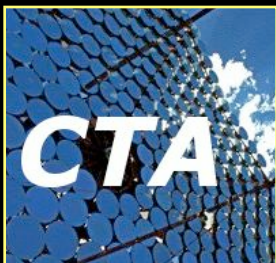


CTA field of view ($4^\circ \times 4^\circ$) (Simulation)



- SNR + PWN in galactic plane
- CTA sensitivity (1 mCrab)
- CTA angular resolution by a factor of 5 better
- FOV $4^\circ \times 4^\circ$

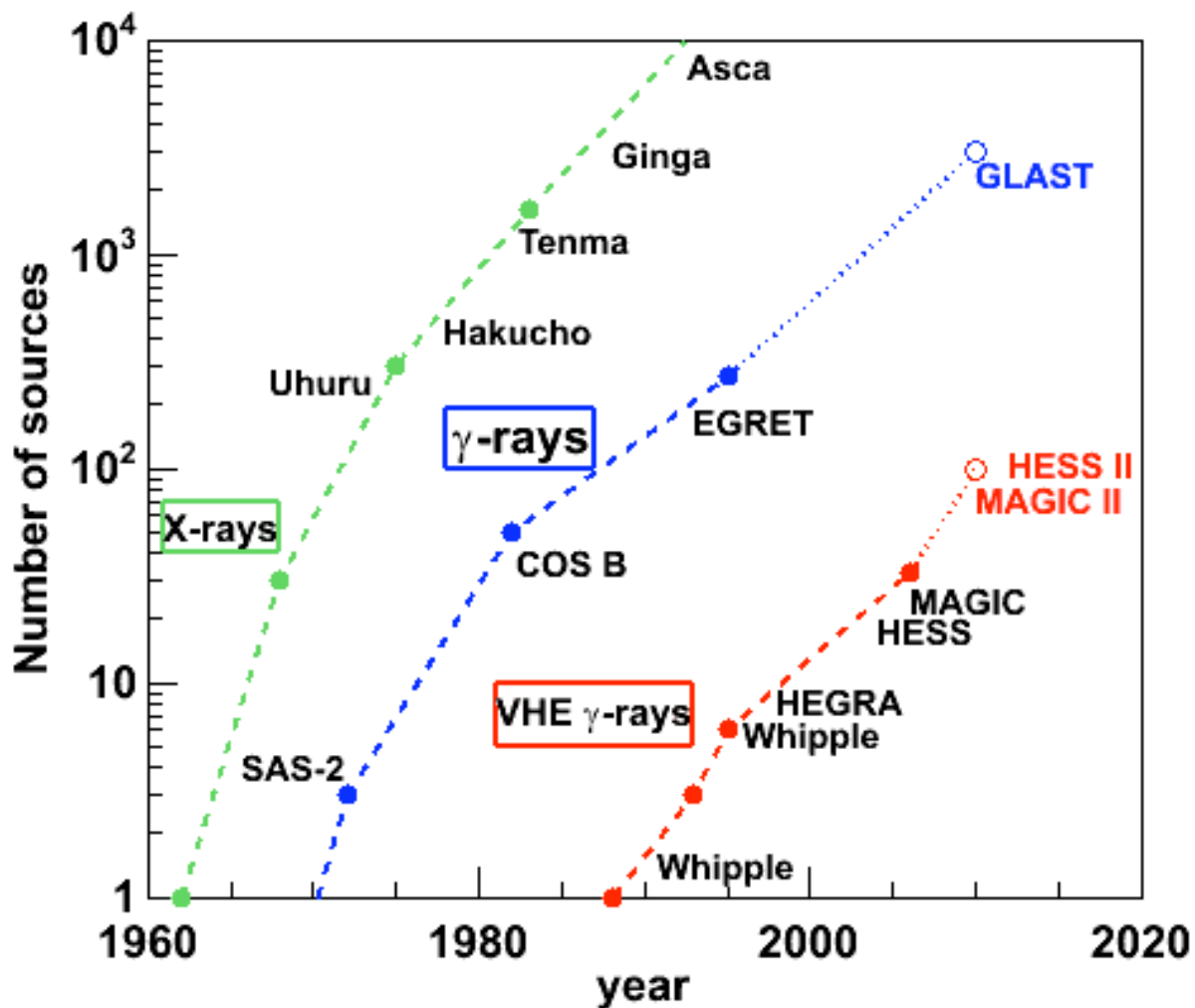




Optimistic expectation



Kifune Plot

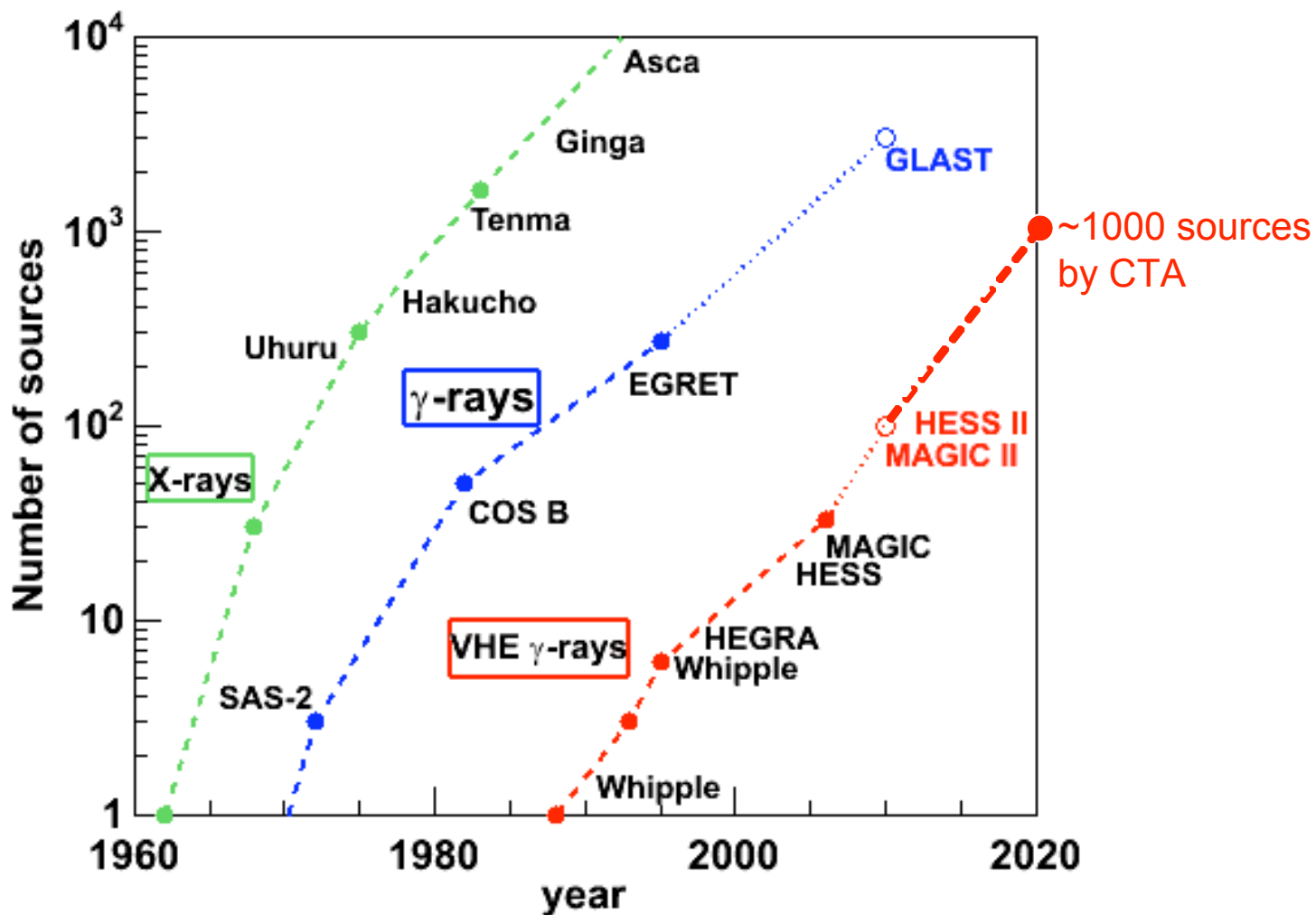




Optimistic expectation



Kifune Plot

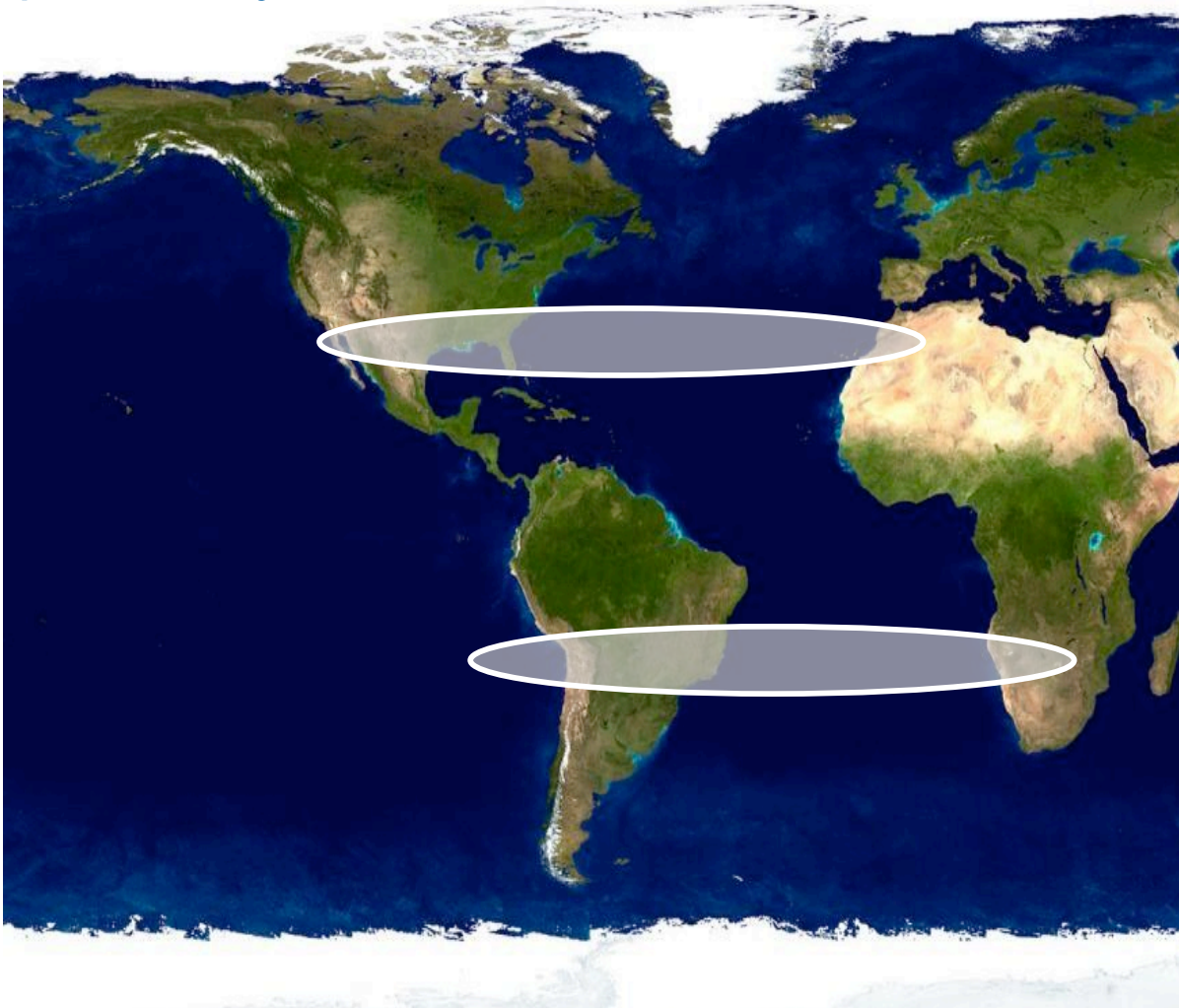




Two installations: full sky coverage



One observatory with two sites
operated by one consortium



Northern Array (50 ME)

- complementary to SA for full sky coverage
- Energy range
some 10 GeV ~1 TeV
- Small field of view
Mainly extragal. Sources

Southern Array (100 ME)

- Full energy and sensitivity coverage
some 10 GeV 100 TeV
- Angular resolution:
0.02 ... 0.2 deg
- Large field of view
Galactic + Extragal. Sources



Emerging proposal in ESFRI



> Emerging proposals

European Strategy Forum
on Research Infrastructures

ESFRI

Report 2006

CTA

is an advanced facility for ground based high-energy gamma ray astronomy, based on the observation of Cerenkov radiation. This approach has proven to be extremely successful for gamma rays of energies above few tens of GeV. The facility will consist in an array of telescopes enhancing the all sky monitoring capability.

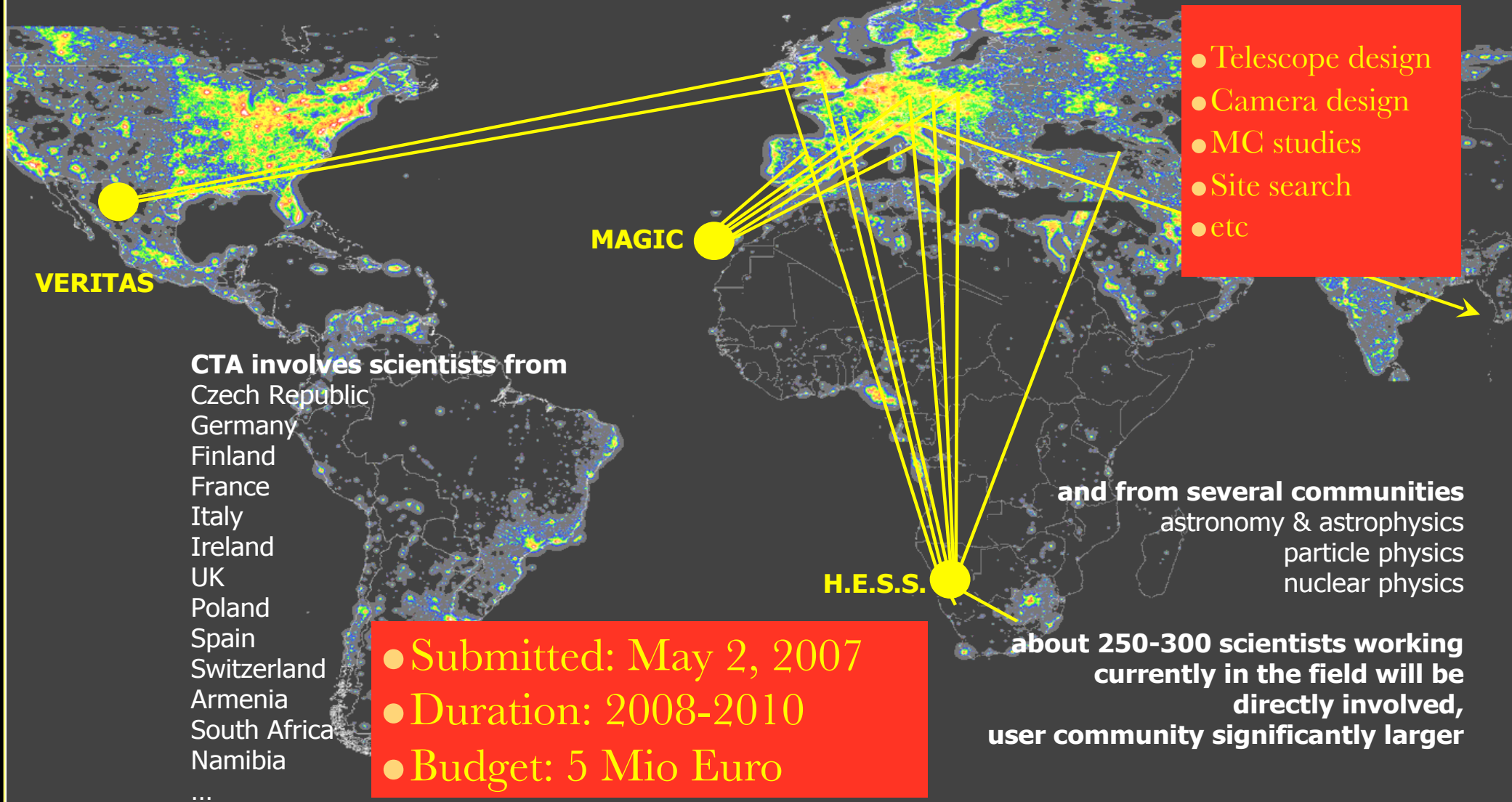
The Expert and Roadmap Working Groups focussed their discussions more on the mature projects, with the Emerging Proposals deserving attention for the future editions of the Roadmap.

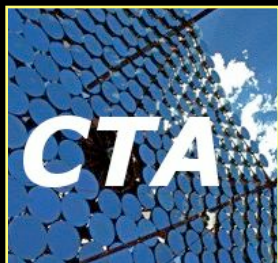


European FP7 design study



Participation of 34 institutes in 15 countries





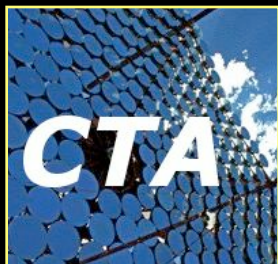
Preliminary schedule



FP 7 Design Study **Prep. Phase ?**

	06	07	08	09	10	11	12	13
Site exploration	█	█	█					
Array layout	█	█	█					
Telescope design		█	█	█				
Component prototypes			█	█	█			
Array prototype				█	█	█		
Array construction						█	█	█
Partial operation							█	█

↑ "Letter of Intent" ↑ Proposal ↑ Design Report GLAST



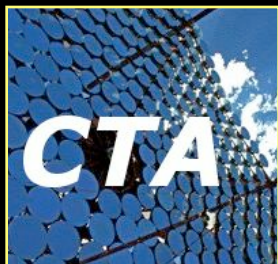
Conclusions



- CTA is a next generation Cherenkov observatory with one order of magnitude better sensitivity, larger FOV and about a factor of 5 better angular resolution
- There will be one station in the North and one in the South
- European initiative but possible collaboration with institutions from all over the world such as USA and Japan
- It will be run as an observatory, open to external astronomers
- CTA will be a very large project with around 50 institutions and 500 physicists
- Aiming for a budget of 150 Mio Euros

The end





EU FP7 Design study



Participation of 34 institutes in 15 countries

Participant no.	Participant organisation name	Part. short name	Country
1 (Coordinator)	Max-Planck-Gesellschaft	MPG	Germany
2	Centre Nationale de Recherche Scientifique	CNRS	France
3	Istituto Nazionale di Fisica Nucleare	INFN	Italy
4	Institut de Física d'Altes Energies	IFAE	Spain
5	Institut d'Estudis Espacials de Catalunya (IEEC-CSIC)	IEEC	Spain
6	Friedrich-Alexander Universität Erlangen-Nürnberg	UErlangen	Germany
7	University of Leeds	ULeeds	United Kingdom
8	University of Turku	UTurku	Finland
● ● ●			
29	Universiteit van Amsterdam	UvA	Netherlands
30	Utrecht University	UU	Netherlands
31	North-West University, Potchefstroom Campus	NWU	Republic of South Africa
32	Washington University	WashU	United States
33	Kavli Institute for Astroparticle Physics and Cosmology, Stanford University	StanU	United States
34	The Chancellor, Masters and Scholars of the University of Oxford	UOXF.DL	United Kingdom