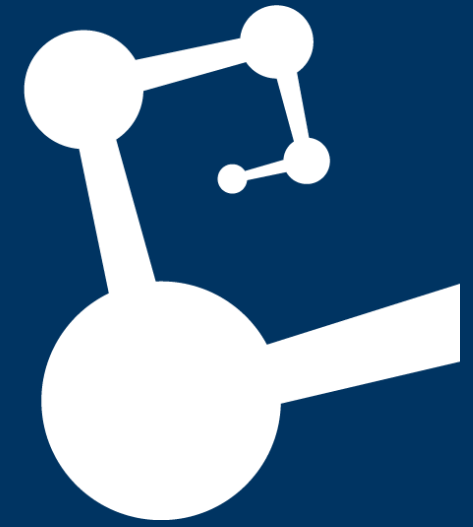




Instituto de  
Ciencias  
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# Signs of criticality in Heavy-ion collisions via two-pion correlations

**Santiago Bernal Langarica**

WONPAQCD, Valparaíso, December 1st, 2025

# Acknowledgments

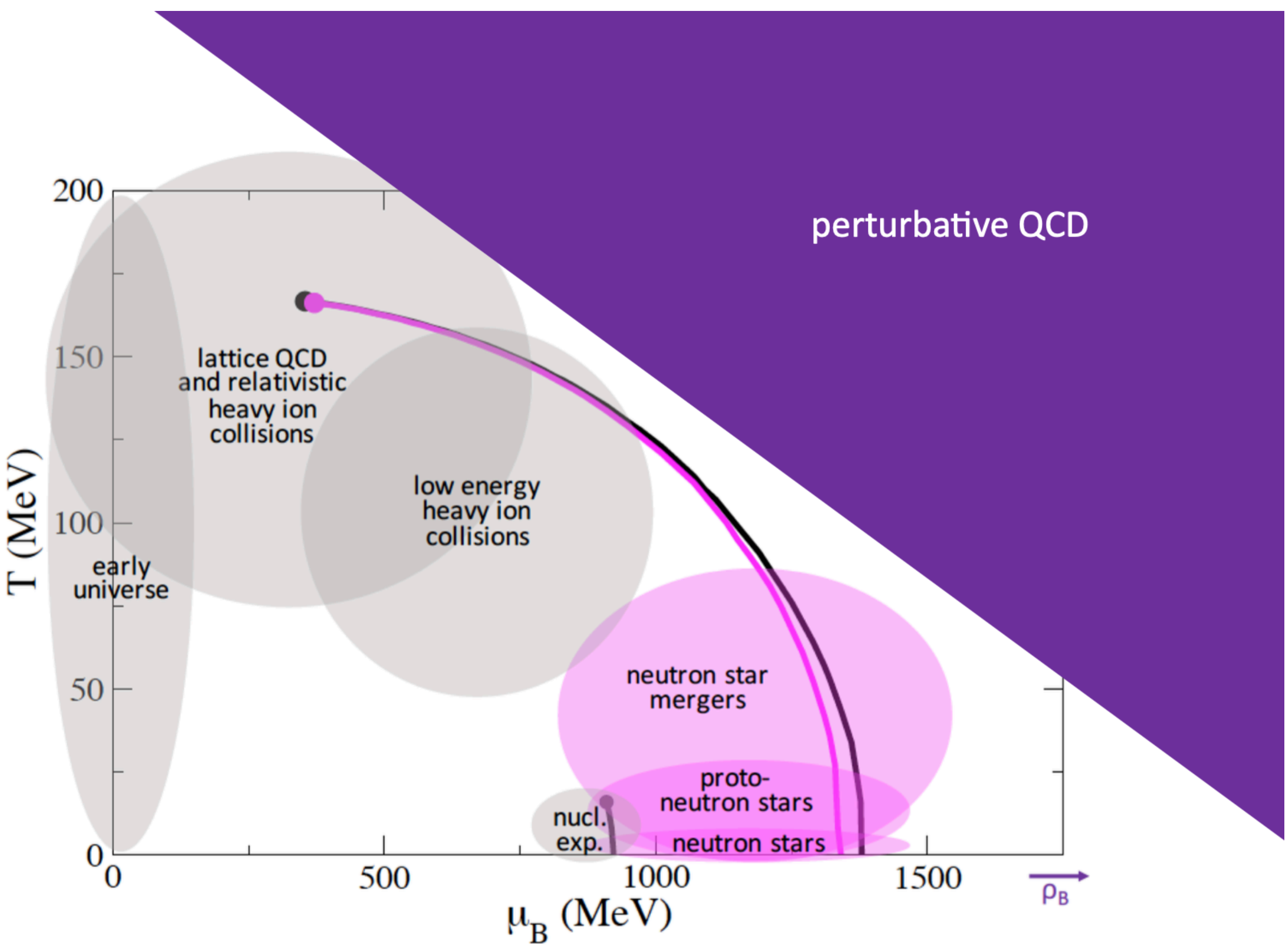
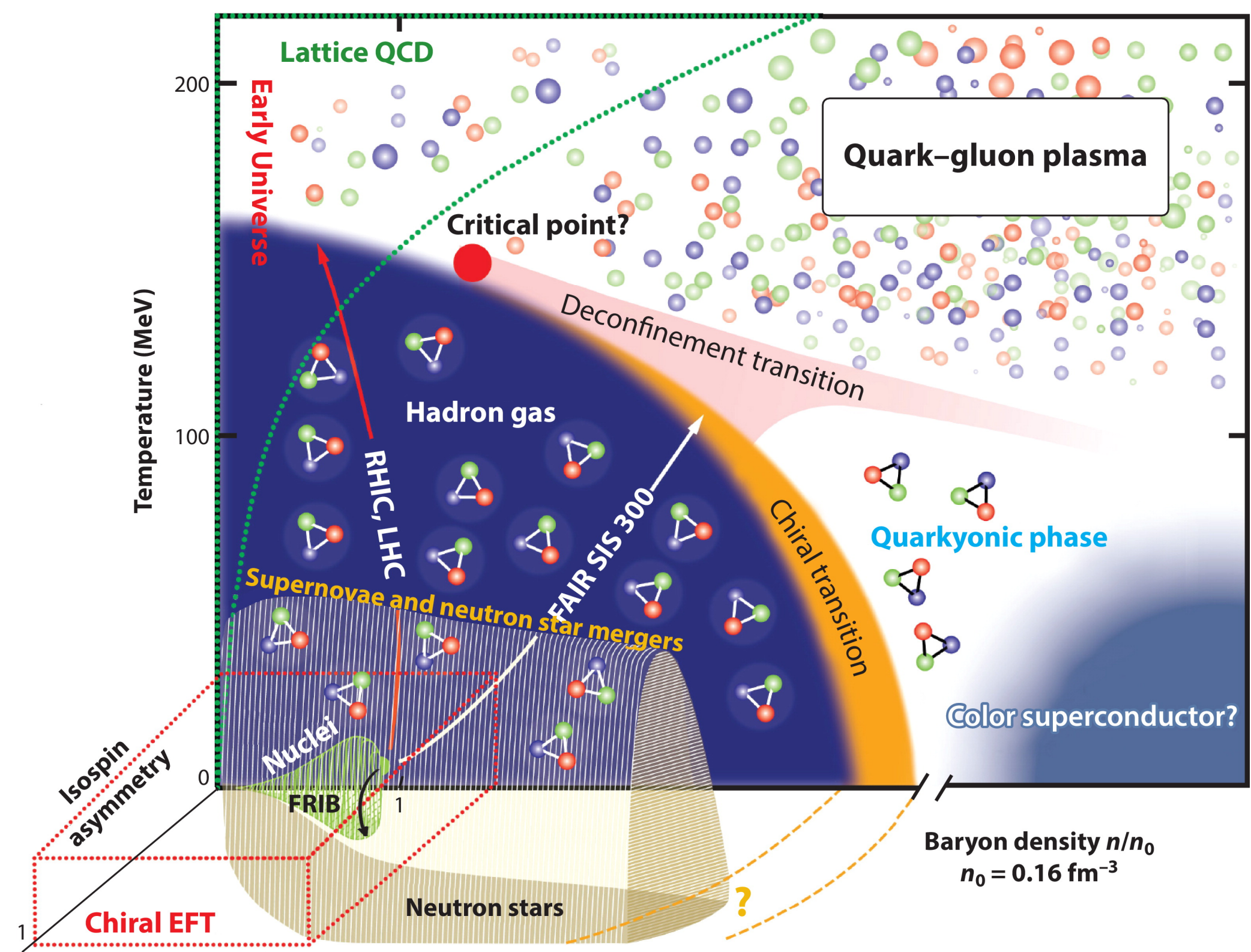
- Alejandro Ayala (UNAM)
- Isabel Domínguez (UAS)
- Irving Gaspar (UAM-I)
- Luis Hernández (UAM-I)
- Johannes Jahan (UH)
- Michael Kahangirwe (KSU)
- Jamie Karthein (MIT)
- Grégoire Pihan (UH)
- Tomas Podelnicek (CTU)
- Jordi Salinas (UIUC)
- Malena Tejeda (UCol)
- Boris Tomasik (CTU)

# Outline of this talk

- QCD phase diagram at finite baryonic density and its critical exponents
- LQCD-based finite density EoS
- Femtoscopic measurements and Lévy sources
- Preliminary results
- Summary



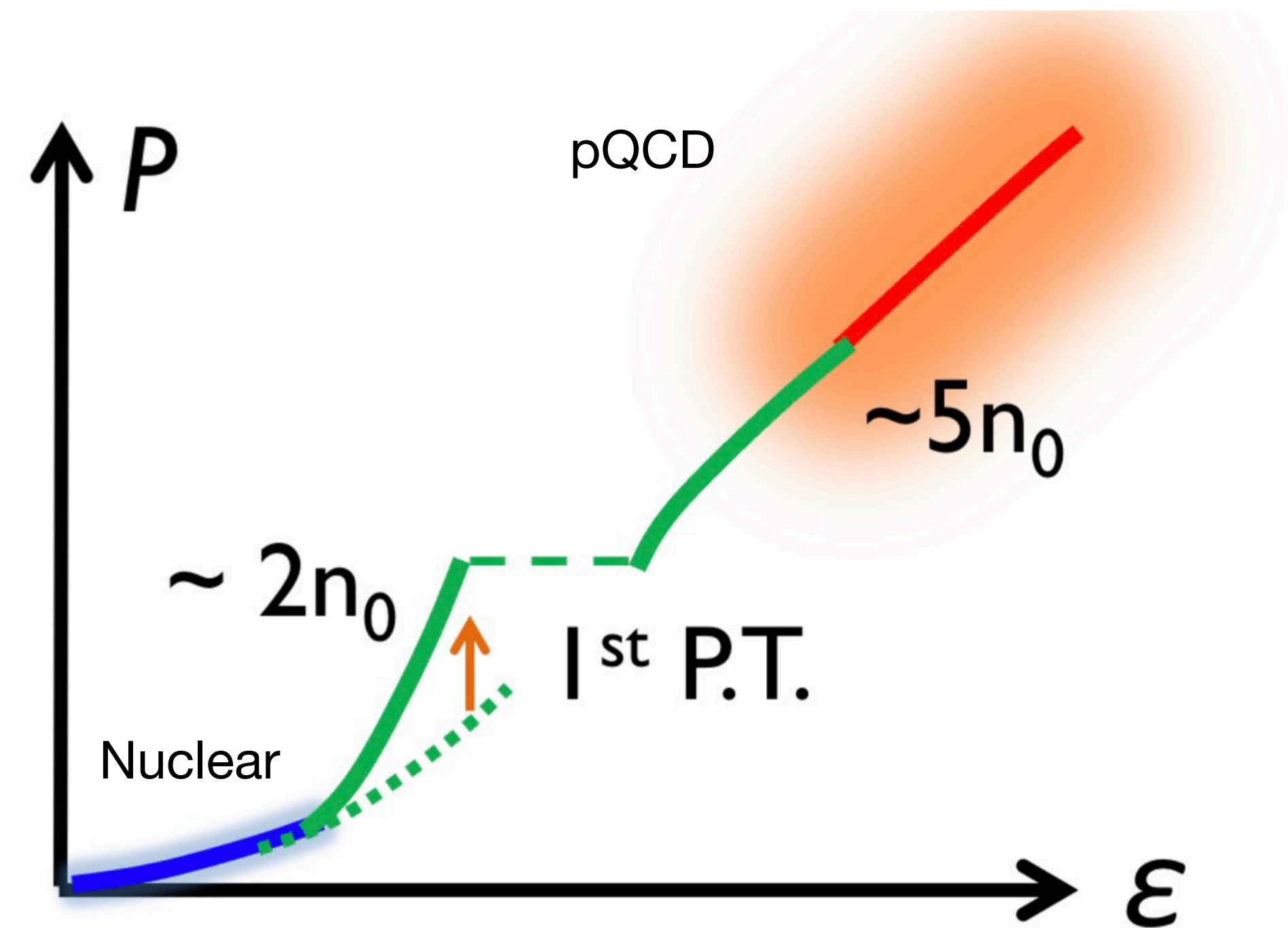
# QCD phase diagram



# Critical exponents of QCD

## And its universality class

- In the vicinity of the CEP, where a second order phase transition occurs, critical phenomena occur
- To characterize this critical behavior, critical exponents are introduced
- In QCD, there exist 6 critical exponents, 2 of which are independent



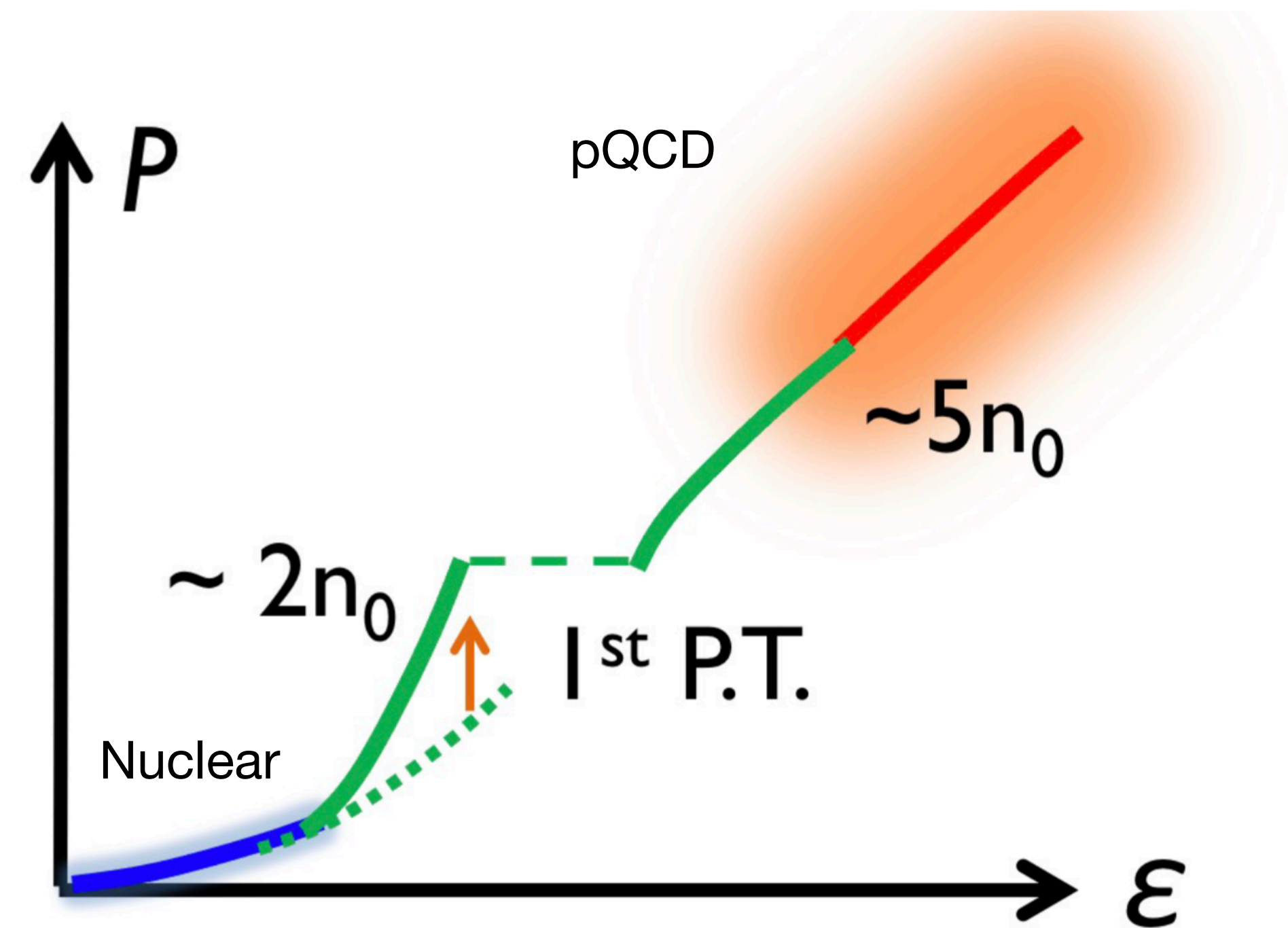
arXiv:2011.10940



# Critical exponents of QCD

## And its universality class

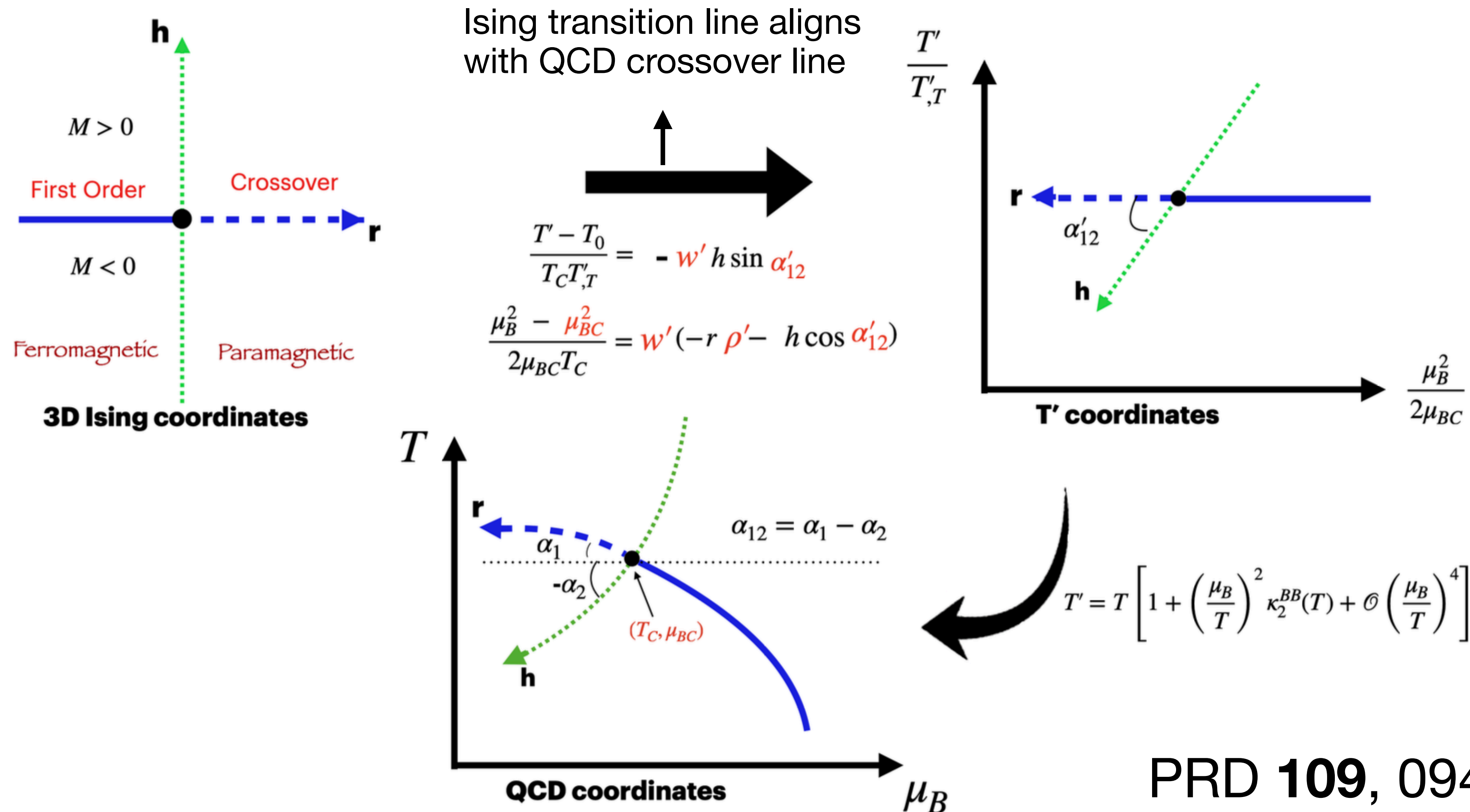
- In Nucl. Phys. B **399**, 395 (1993), a two-flavour model with massless quarks, the critical exponents were computed and it was shown that the universality class of this QCD is the same as that of the 3d-Ising model
- Due to the nature of HIC, the true universality class is the same as in the random 3d-Ising model [PRB **52**, 6659 (1995)] :
- $\eta = 0.5 \pm 0.05$ ,  $\nu = 1.1 \pm 0.2$ ,  
 $\alpha = -1.3 \pm 0.6$ ,  $\beta = 0.6 \pm 0.1$ ,  
 $\gamma = 2.2 \pm 0.4$ ,  $\delta = 4.7 \pm 0.3$



arXiv:2011.10940

# Finite density QCD equation of state

## Lattice-based T expansion and critical point

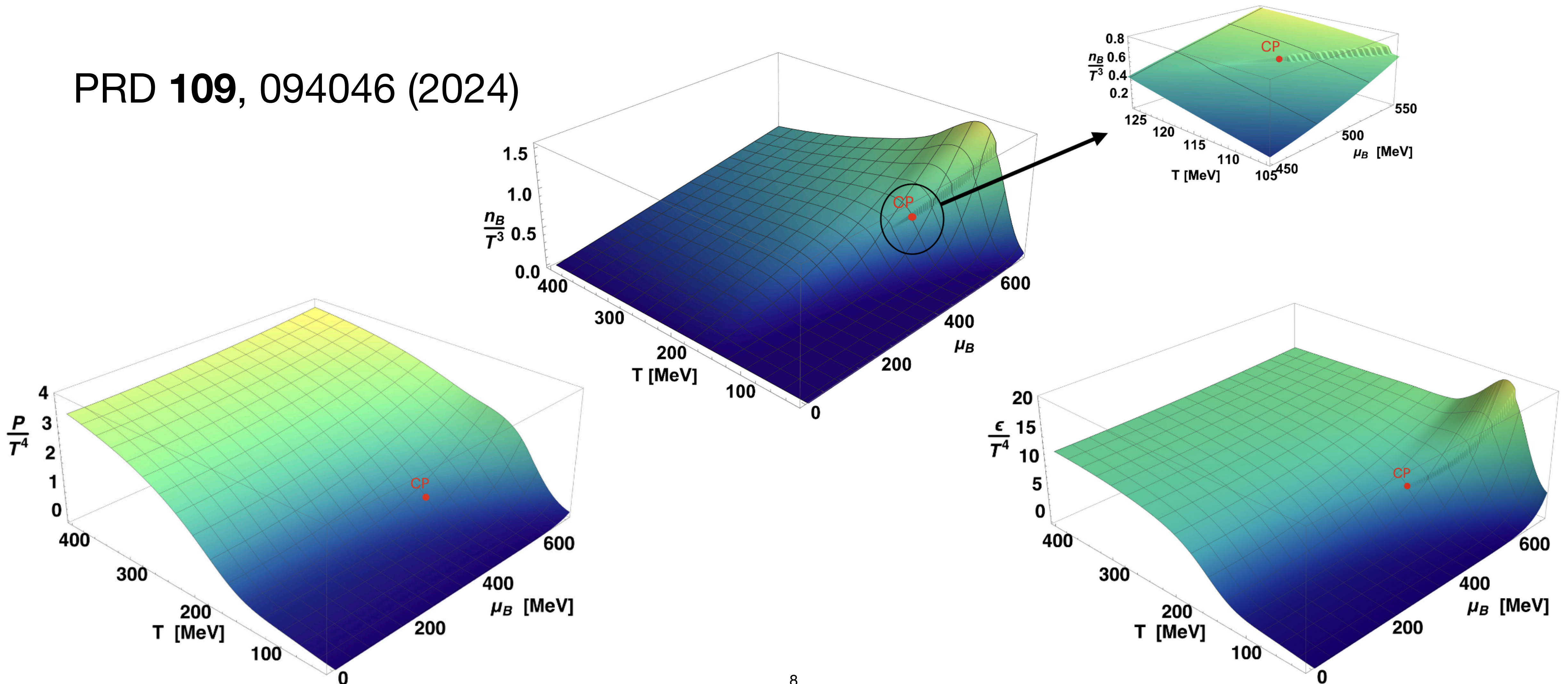




# Finite density QCD equation of state

## Lattice-based T expansion and critical point

PRD 109, 094046 (2024)





# Two-identical particle correlations

## Or HBT femtoscopy

- After a collision, the one-particle momentum distribution can be found as

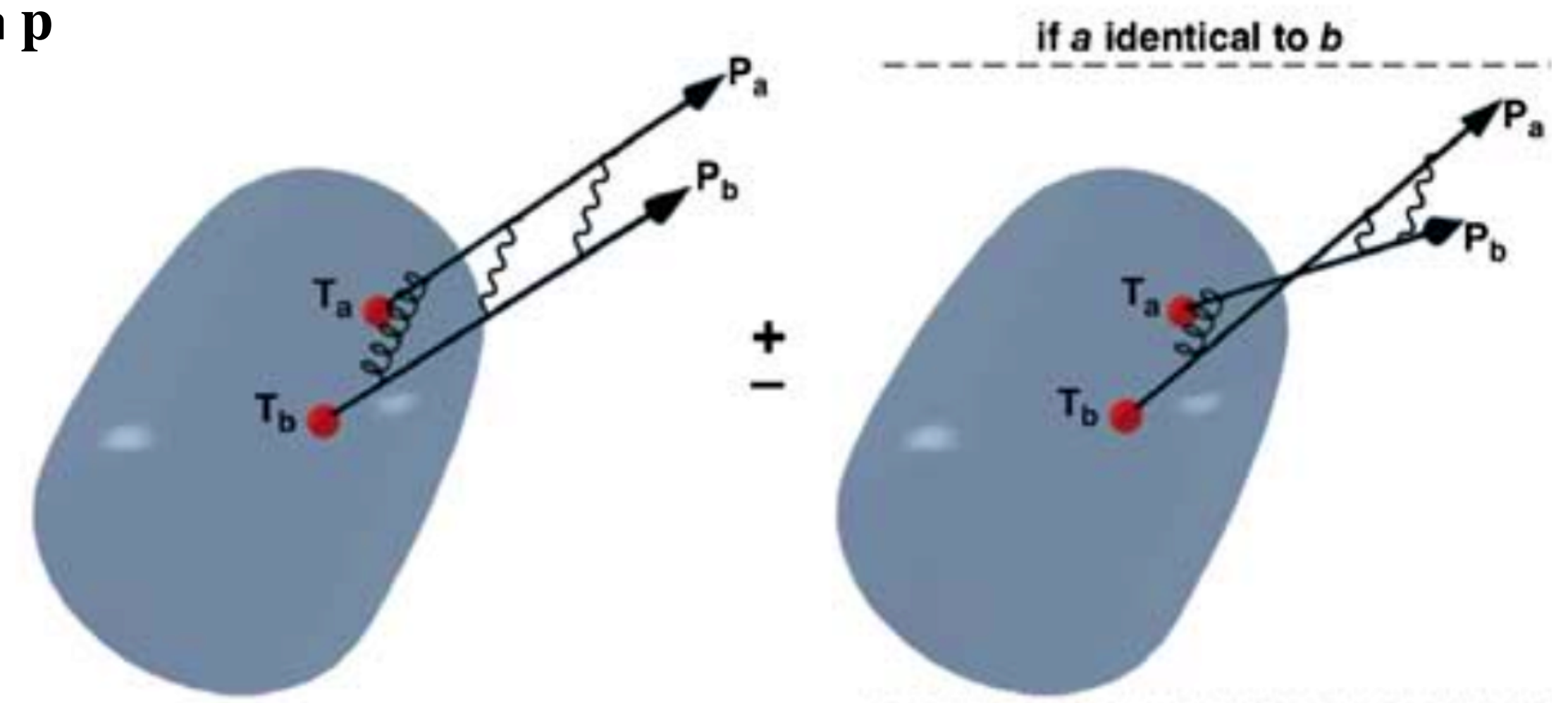
- $\mathcal{P}_1(\mathbf{p}) \equiv \frac{d^3N}{dp^3} \longrightarrow$  Probability of emission of a particle with momentum  $\mathbf{p}$

- In a similar way, the two-particle momentum distribution

- $\mathcal{P}_2(\mathbf{p}_1, \mathbf{p}_2) \equiv \frac{d^6N}{d^3p_1 d^3p_2} \longrightarrow$  Probability simultaneous emission of particles with momenta  $\mathbf{p}_1$  and  $\mathbf{p}_2$

- When the emission process of the particles are independent of each other,  $\mathcal{P}_2$  can be factorized as the product of  $\mathcal{P}_1$ 's

- $\mathcal{P}_2(\mathbf{p}_1, \mathbf{p}_2) = \mathcal{P}_1(\mathbf{p}_1)\mathcal{P}_1(\mathbf{p}_2)$

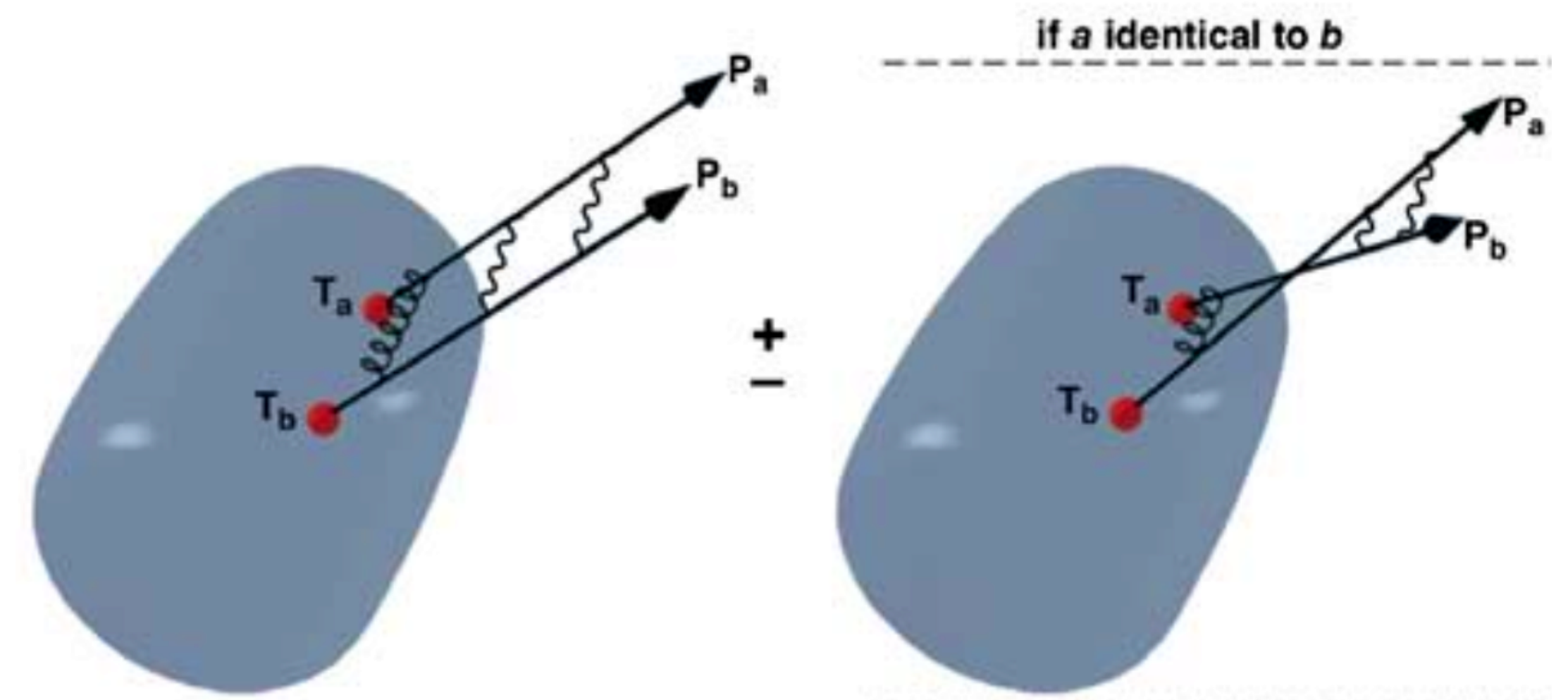


Annu. Rev. Nucl. Part. Sci. **55**, 357 (2005)

# Two-identical particle correlations

## Or HBT femtoscopy

- If this factorization is not valid, then it is because the emission processes are not independent of each other, but correlated due to:
- Conservation laws
- Decays
- Quantum nature of the particles
- Long etcetera ...



Annu. Rev. Nucl. Part. Sci. **55**, 357 (2005)

# Two-particle correlation functions

## Femtoscopy

- From the theoretical point of view, this function is defined as

$$C_2(\mathbf{p}_1, \mathbf{p}_2) = \frac{\mathcal{P}_2(\mathbf{p}_1, \mathbf{p}_2)}{\mathcal{P}_1(\mathbf{p}_1) \cdot \mathcal{P}_1(\mathbf{p}_2)}$$

Two-particle momentum distribution

One-particle momentum distribution

- Correlation functions are usually described as functions of the **pair relative momentum**,  $q = p_1 - p_2$ , and the **pair average momentum**,  $K = \frac{p_1 + p_2}{2}$ .  
On-shell conditions imply that  $C_2$  is only a function of  $\mathbf{q}$  and  $\mathbf{K}$



# Two-particle correlation functions

## Femtoscopy

- Two-particle correlation functions can also be related to the particle emission source in phase-space  $S(x, p)$ . Assuming that the emission is not initially correlated, that particles are bosons and that they not interact in their final state,

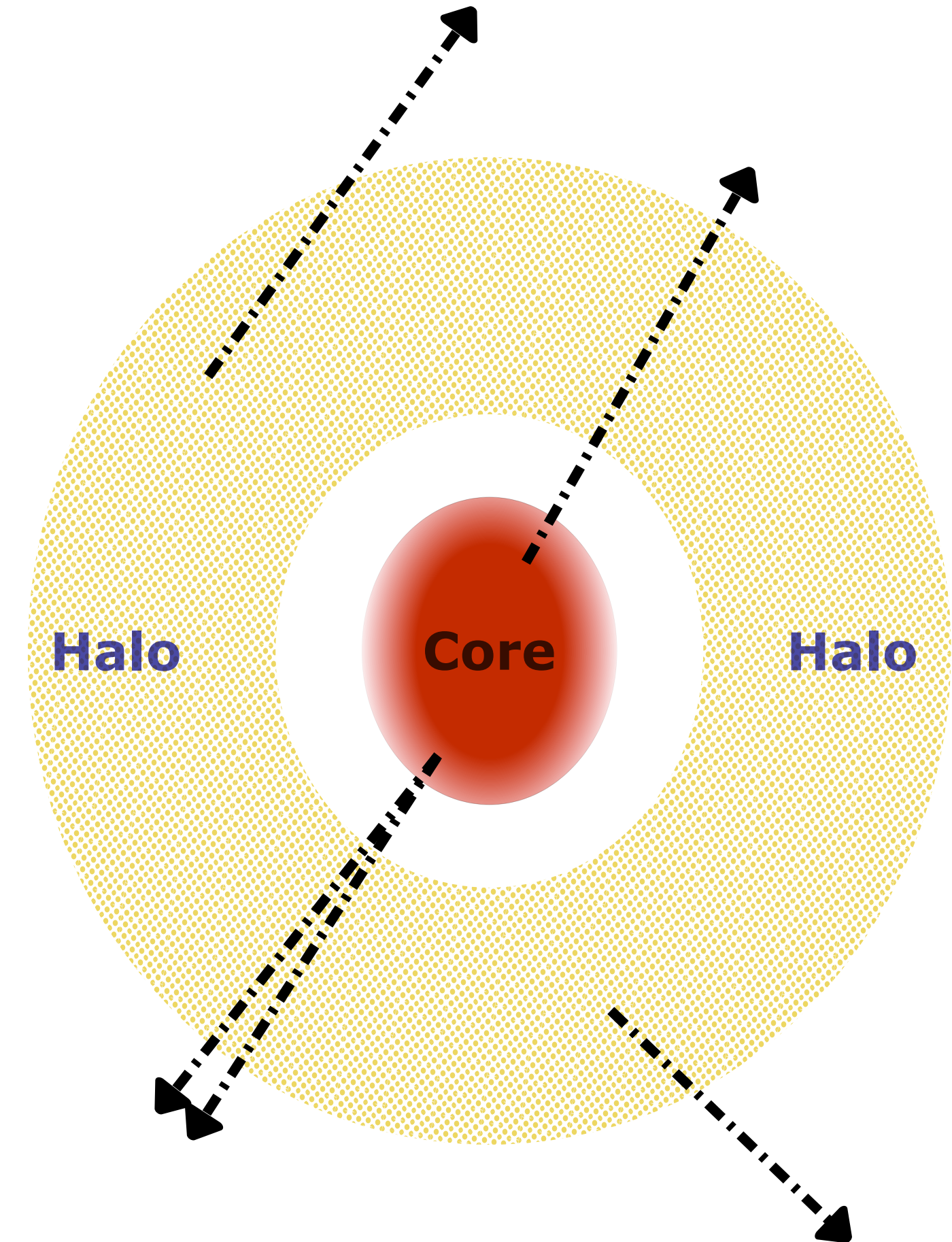
- If the Fourier transformation of the source is  $\tilde{S}(q, p) = \int d^4x e^{iq \cdot x} S(x, p)$ ,  
then

- $$C_2(q, K) = 1 + \frac{\left| \tilde{S}(q, K) \right|^2}{\left| \tilde{S}(0, K) \right|^2}$$

# Correlation functions

## Core - halo model [PRD 47, 3860 (1993), Z. Phys. C 71, 491 (1996)]

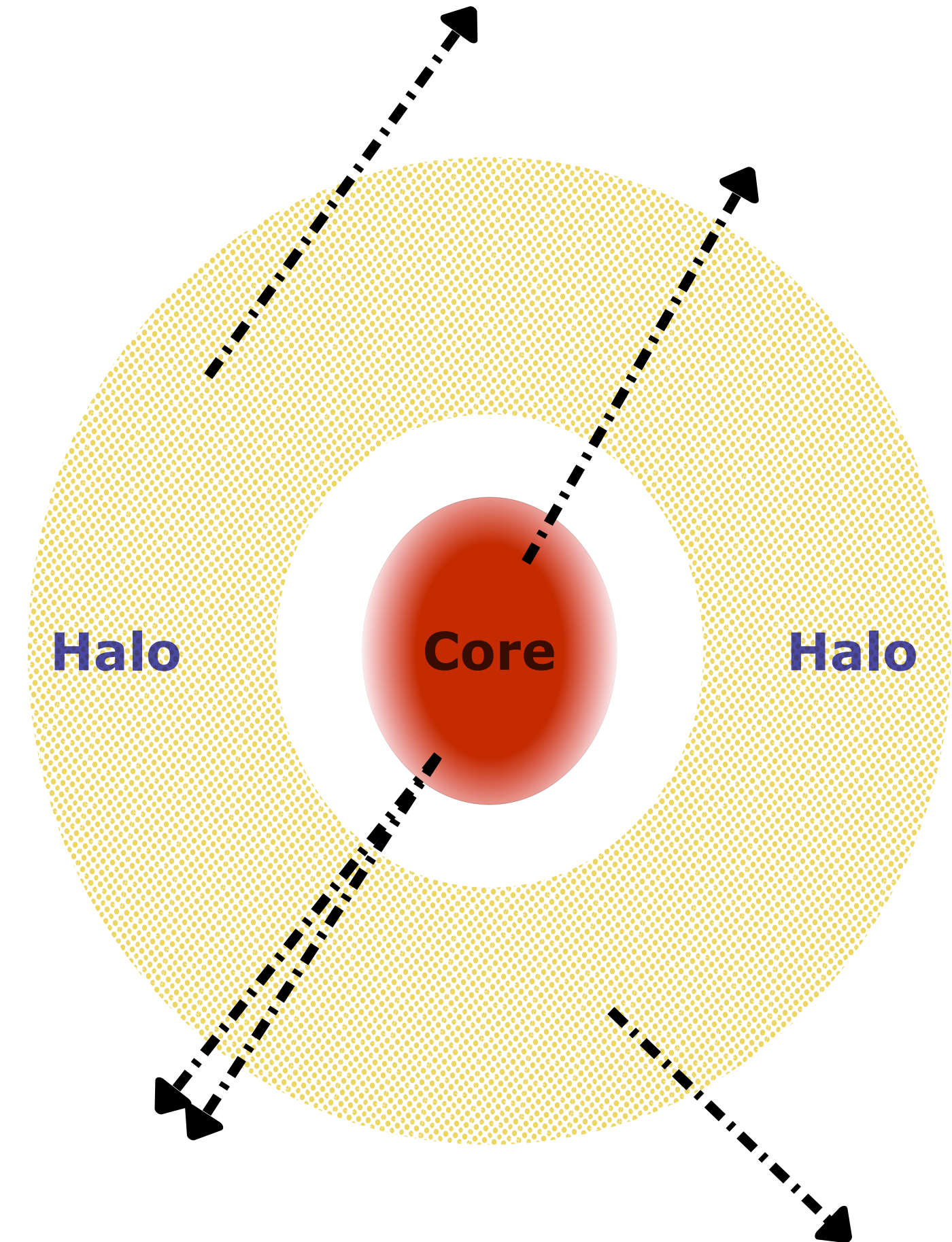
- The correlation function for  $q_{inv} \rightarrow 0$ , reaches a value of  $1 + \lambda$ , with  $0 \leq \lambda \leq 1$ , this is due to the resonances that decay and form a halo around the interaction region
- A detector with finite resolution will not be able to solve the halo if its characteristic size is larger than the momentum resolution



# Correlation functions

## Core - halo model [PRD 47, 3860 (1993), Z. Phys. C 71, 491 (1996)]

- Let us assume that the detector has a momentum resolution  $\Delta_q$  and that the source is made of two components, such that
- $S = S_{\text{core}} + S_{\text{halo}}$
- Where the halo is composed of particles produced through the decays of resonance
- The characteristic scale of the halo is  $R_{\text{halo}} \gtrsim \frac{1}{\Delta_q}$ , then
- $R_{\text{core}} < \frac{1}{\Delta_q} \lesssim R_{\text{halo}}$





# Correlation functions

## Core - halo model [PRD 47, 3860 (1993), Z. Phys. C 71, 491 (1996)]

- The core and halo distributions can be written as
- $\mathcal{P}_{1,\text{core}}(K) = \int d^4x S_{\text{core}}(x, K) = \tilde{S}_{\text{core}}(0, K) \equiv N_{\text{core}}$
- $\mathcal{P}_{1,\text{halo}}(K) = \int d^4x S_{\text{halo}}(x, K) = \tilde{S}_{\text{halo}}(0, K) \equiv N_{\text{halo}}$
- Since, for the halo, the region with  $q < \Delta_q$  cannot be solved, then  $\tilde{S}_{\text{halo}}(q, K) \approx 0$ , and  $\tilde{S}(q, K) \simeq \tilde{S}_{\text{core}}(q, K)$ . Hence
- $C_2(q, K) = 1 + \left( \frac{N_{\text{core}}}{N_{\text{core}} + N_{\text{halo}}} \right)^2 \frac{\left| \tilde{S}_{\text{core}}(q, K) \right|^2}{\left| \tilde{S}_{\text{core}}(0, K) \right|^2}$ , with  $\lambda = \left( \frac{N_{\text{core}}}{N_{\text{core}} + N_{\text{halo}}} \right)^2$

# Correlation functions

## And its parameterizations

- Once the correlation function is obtained, it can be fitted to extract the source characteristics
- As an example, assume a 1D source that can be factorized into a space-time distribution and a momentum distribution  $S(x, p) = f(x) \cdot g(p)$ , with
- $\int dx f(x) = 1$  and  $\int dp g(p) = N$
- Then
- $C_2(q, K) = 1 + \left| \tilde{f}(q) \right|^2 \approx 2 - q^2 (\langle x^2 \rangle - \langle x \rangle^2) + \dots \approx 1 + \exp(-q^2 R^2)$
- Where  $R^2 = \langle x^2 \rangle - \langle x \rangle^2$  and  $\tilde{f}(q) = \int dx \exp(iqx) f(x)$  is known as the characteristic function

# Correlation functions

## And its parameterizations

- Under which conditions are these Gaussian assumptions valid?
- The emission is a superposition of several independent processes whose emission coordinate is separated by  $\delta x_i$
- If the variance that characterize this separation is finite, then by means of the central limit theorem, the probability distribution will tend to a Gaussian
- Since the Fourier transformation of a Gaussian is also a Gaussian, the correlation function will also be a Gaussian



# Correlation functions

## And its parameterizations

- In the neighborhood of a CEP, where the correlation length diverges, and the distributions that characterize physical quantities behave as power-laws
- These kind of distributions have non-finite variance (and even non-finite mean) and hence have a non-analytic behavior for certain values
- Then the probability distribution will be different from a Gaussian
- These special case distributions are called Lévy distributions

# Correlation functions

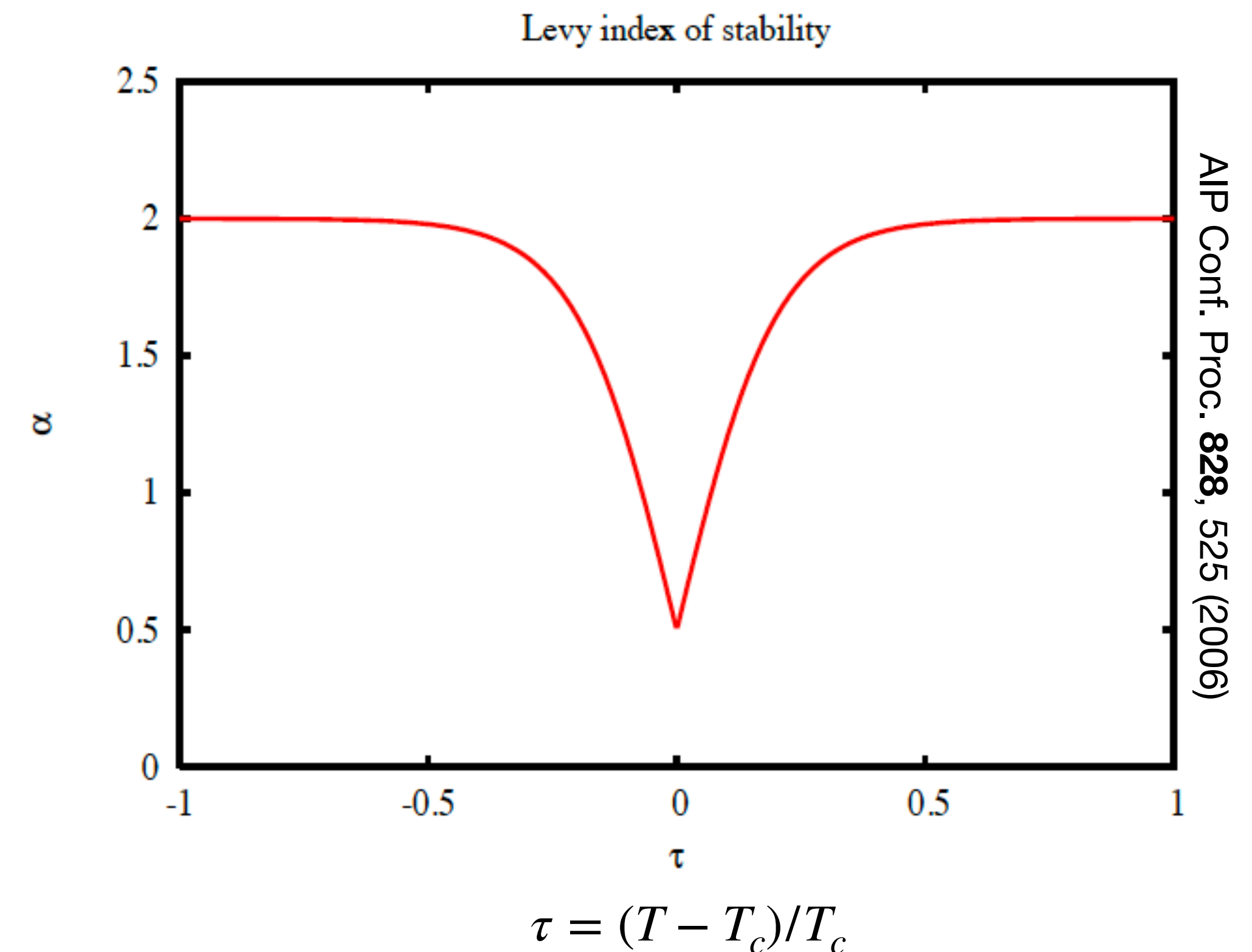
## And its parameterizations

- The Lévy characteristic function is
- $\tilde{f}(q) = \exp \left( -\gamma^\alpha |q|^\alpha + i \beta \gamma^\alpha \text{sign}(q) \tan \left( \frac{\alpha\pi}{2} \right) + i q \delta \right)$ , with  $0 \leq \alpha \leq 2$ ,  
 $-1 \leq \beta \leq 1$ ,  $\gamma > 0$  and  $-\infty < \delta < \infty$
- A special case happens when  $\beta = 0$ ,  $\gamma = R/2^{1/\alpha}$  and  $\delta = x_0$ , then
- $\tilde{f}(q) = \exp (i q x_0 + |q R|^\alpha) \approx 1 + i q x_0 - \frac{1}{2} |q R|^\alpha$
- Therefore,  $C_2(q; \alpha) = 1 + \exp (-|q R|^\alpha)$

# Correlation functions

## Relationship between the Lévy index and QCD critical exponents

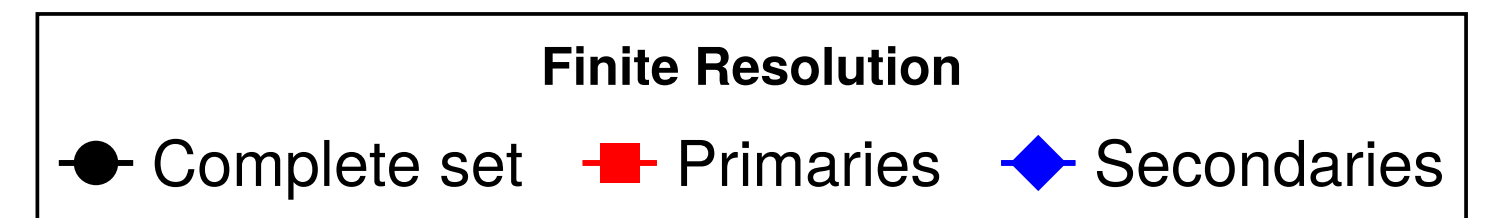
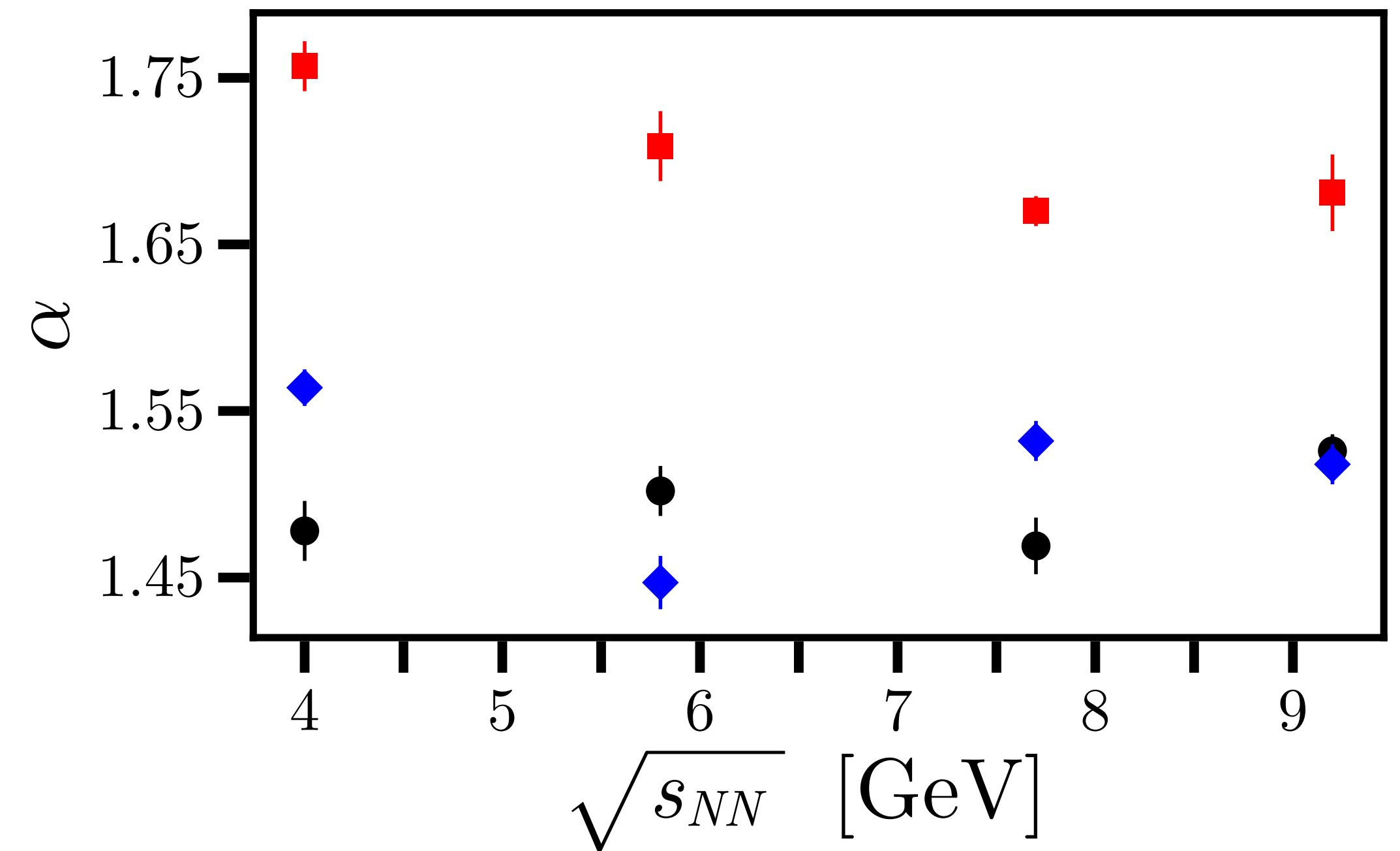
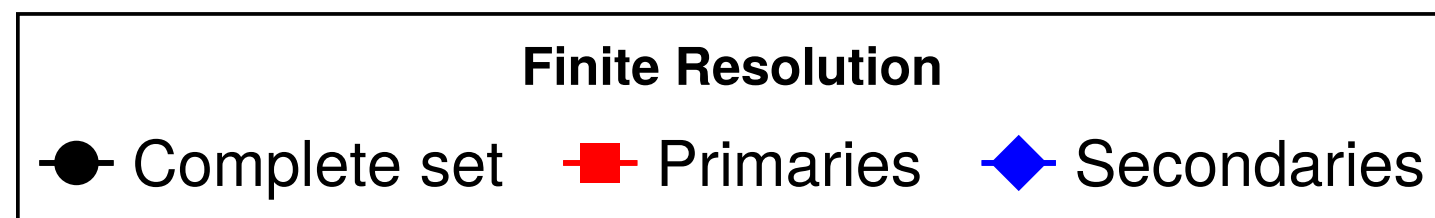
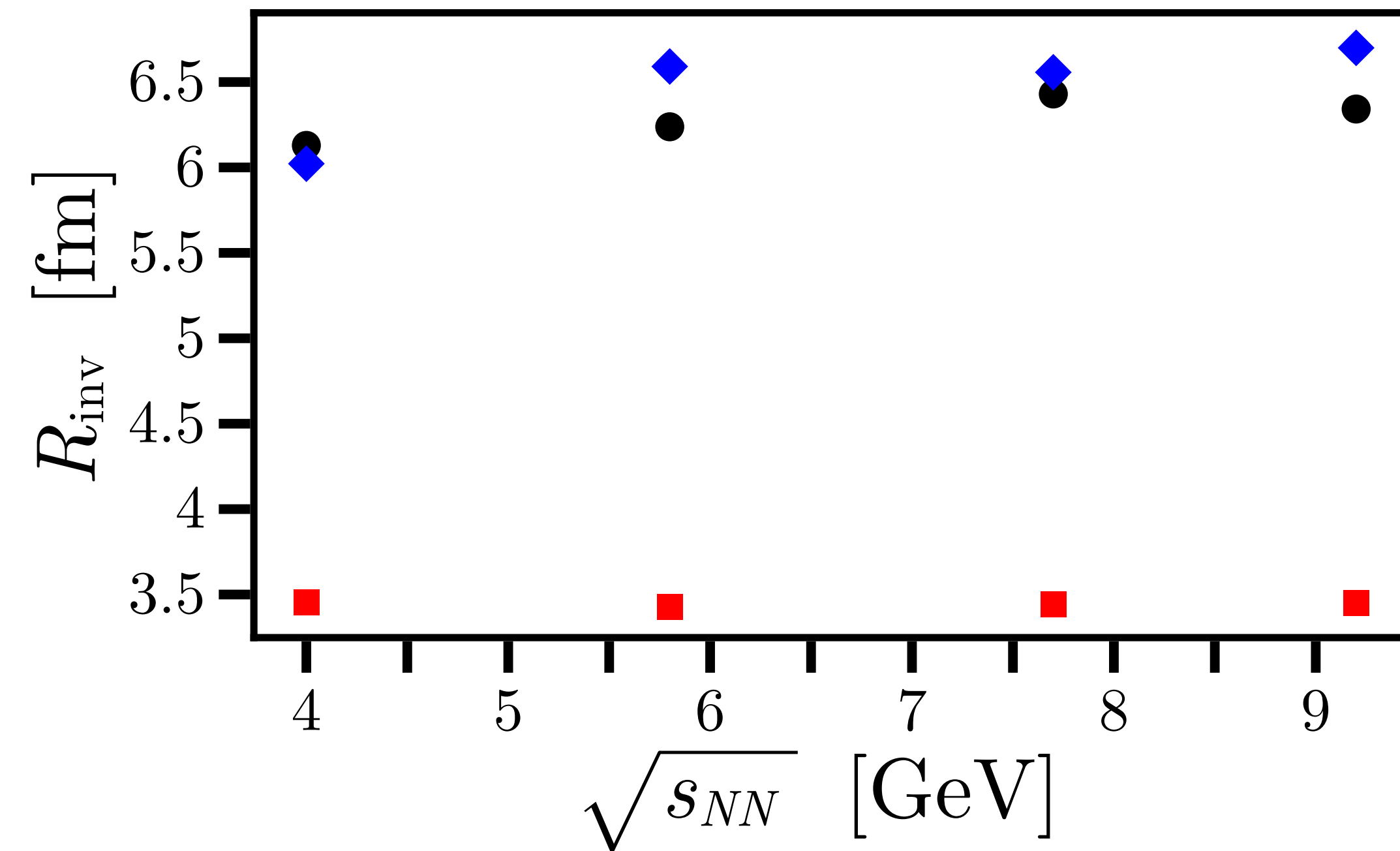
- For QCD, the correlation function of the order parameter decays as a power-law in the vicinity of the critical point  
 $\rho \propto r^{-(d-2+\eta)}$
- For Lévy-type sources, the correlation between initial and final positions decays also as a power-law  $\rho \propto r^{-(1+\alpha_{Levy})}$
- Para  $d = 3$ ,  $\eta = \alpha_{Lévy}$
- In this model,  $\eta = 0.5 \pm 0.05$





# Two-pion correlation functions

With finite resolution at different energies



EPJ A **60**, 135 (2024)

# Summary

## And future work

- Two-pion correlation functions are a prime tool to study the interaction region of heavy-ion collisions
- Radii grow as the energy increases, and the Lévy index of stability decreases
- At low collision energies, the core has a large contribution from secondary pions
- This type of studies including EoS effects could help us determine if it can be used to signal critical phenomena
- Stay tuned for new results!

# Thank you!

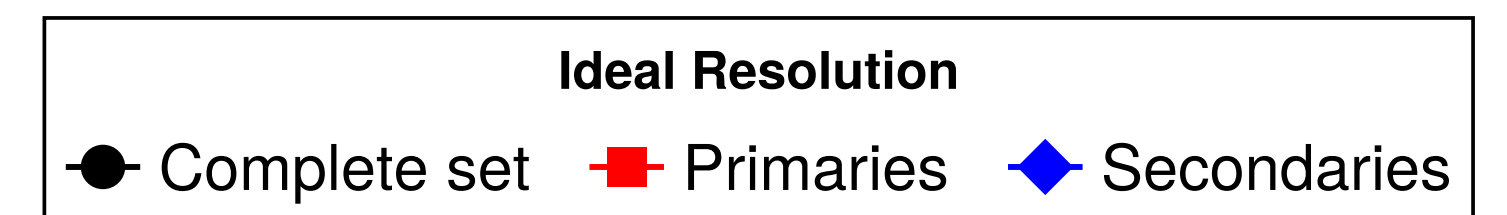
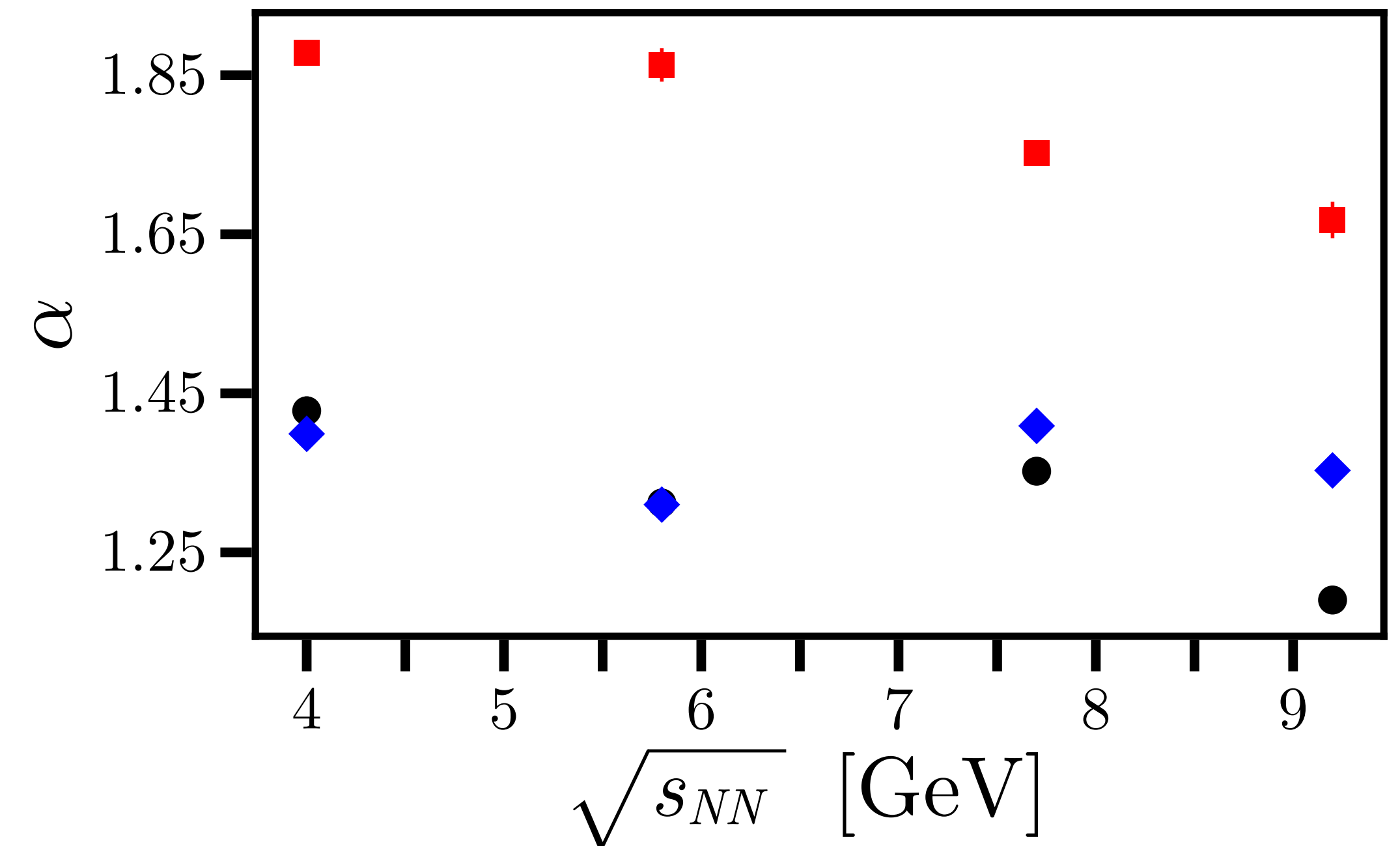
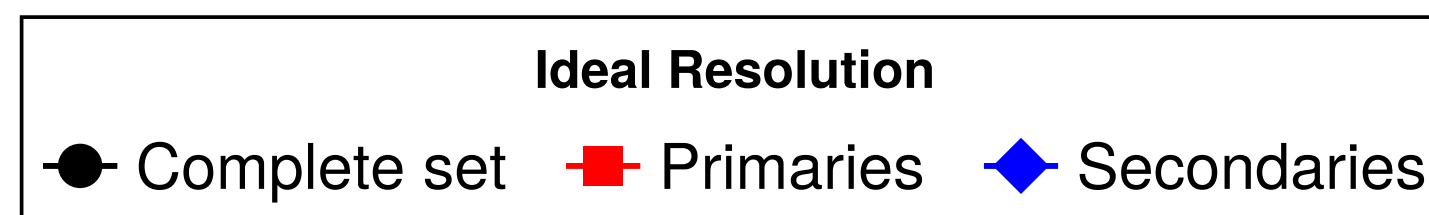
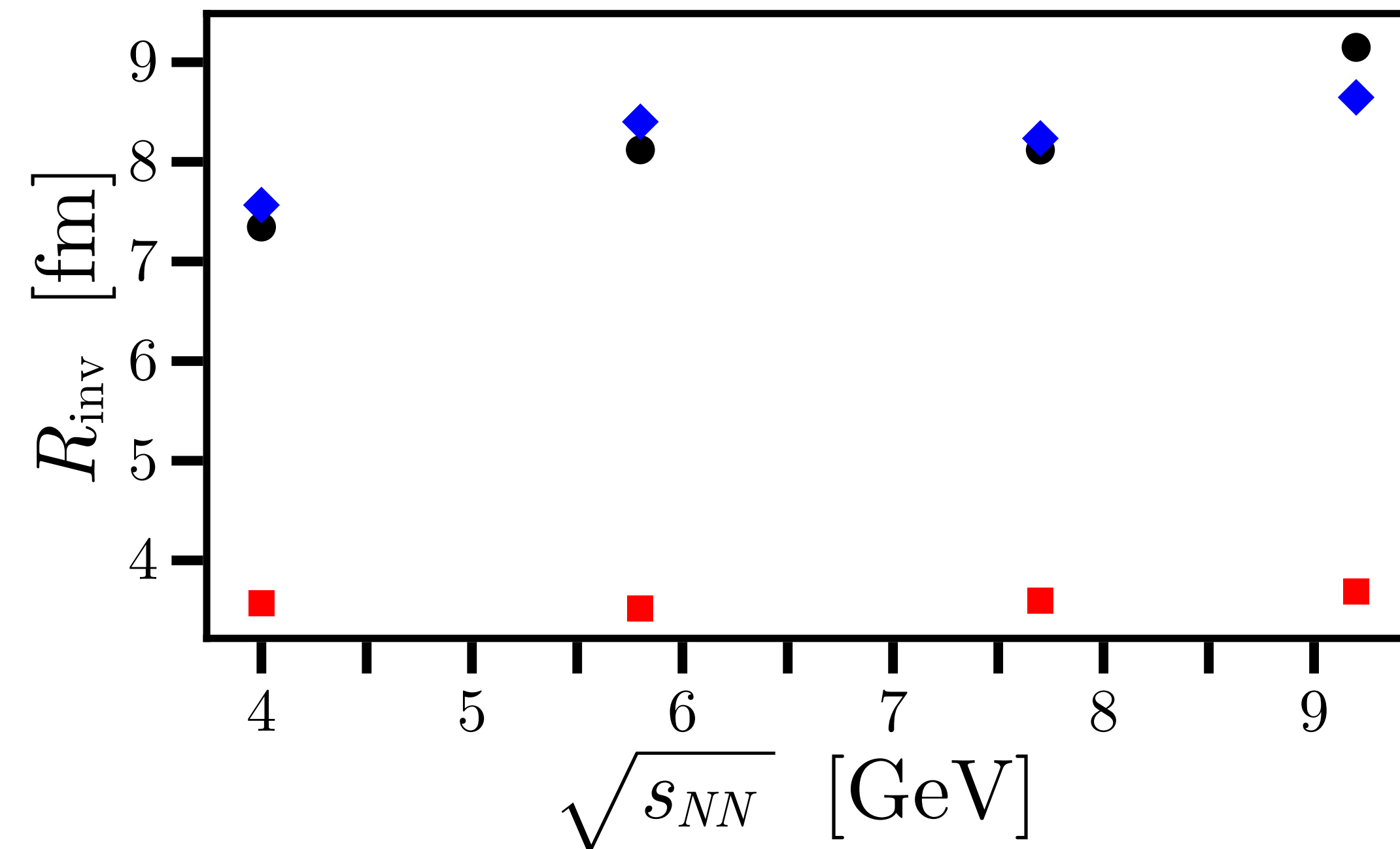
[santiago.bernal@correo.nucleares.unam.mx](mailto:santiago.bernal@correo.nucleares.unam.mx)



# Backup

# Two-pion correlation functions

With ideal resolution at different energies



EPJ A **60**, 135 (2024)

# Two-pion correlation functions

## And core - halo model

- In the core - halo model

- $$\lambda = \left( \frac{N_{\text{core}}}{N_{\text{core}} + N_{\text{halo}}} \right)^2$$

- This means that between 82 and 77 % of the pions come from the core
- Between 6 y 13 % of the pions are produced from primary processes, hence the core has a large fraction of secondary pions
- Since  $\Delta_q \cdot R_{\text{inv}} \sim 1$ , then  $R_{\text{inv}} \lesssim 20$  fm

$\sqrt{s_{\text{NN}}}$	$\lambda_{\text{all}}$	$\lambda_{\text{primary}}$	$\lambda_{\text{secondary}}$
4.0	$0.677 \pm 0.003$	$0.907 \pm 0.002$	$0.651 \pm 0.004$
5.8	$0.632 \pm 0.004$	$0.905 \pm 0.003$	$0.647 \pm 0.005$
7.7	$0.625 \pm 0.004$	$0.9 \pm 0.003$	$0.608 \pm 0.003$
9.2	$0.595 \pm 0.007$	$0.887 \pm 0.005$	$0.602 \pm 0.003$