

In-medium Finite Energy Sum Rules: advances and perspectives

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Motivation:

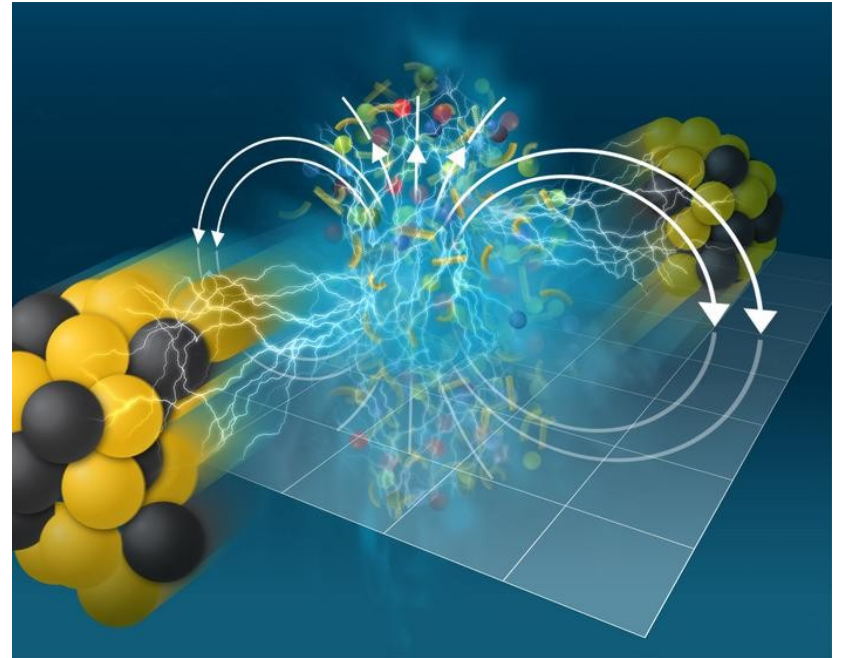
Extreme density,
Extreme temperature
Extreme magnetic field

ρ_B up to $\sim 10\rho_0$ (compact stars)

T up to $\sim 0.5 \text{ GeV} \sim 10^{12} \text{ K}$ (LHC)

B up to $\sim 0.3 \text{ GeV}^2/e \sim 10^{19} \text{ G}$ (LHC)

Some alternative model to obtain medium
modification of parameters



Outline

- Introduction to finite energy sum rules (FESR)
- In-medium sum rules
- Advantages and disadvantages
- some properties of FESR - magnetic field and chemical potential
- Summary and outlook

The spectral function

Two current correlator

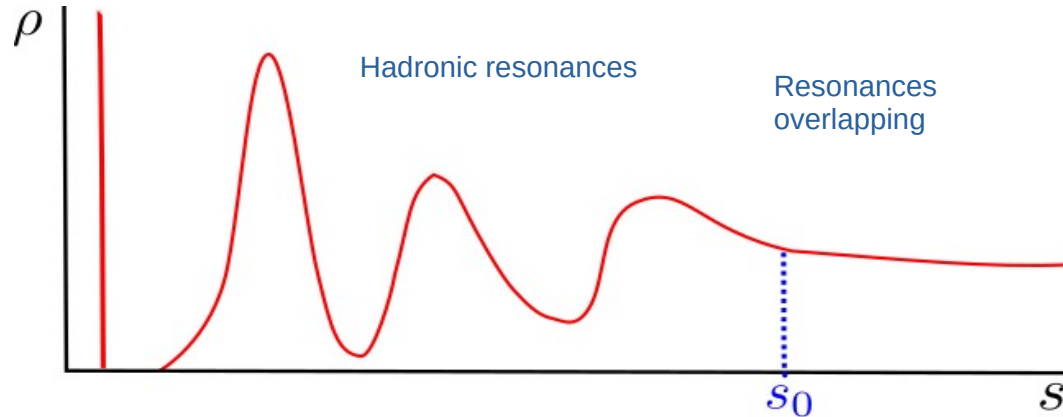
$$\Pi_{\mu\nu}(x - y) = i\langle 0|T J_\mu(x) J_\nu^\dagger(y)|0\rangle$$

Fourier transformation

$$\Pi_{\mu\nu}(q) = q_\mu q_\nu \Pi_L(q^2) + (g_{\mu\nu} q^2 - q_\mu q_\nu) \Pi_T(q^2)$$

Spectral function

$$\rho(s) = \frac{1}{\pi} \text{Im} \Pi(s + i\epsilon)$$



$s_0 \rightarrow$ Hadronic continuum threshold

FESR Quark-hadron duality $\Pi^{\text{Had}} \leftrightarrow \Pi^{\text{QCD}}$

Cauchy's theorem

$$\frac{1}{\pi} \int_0^{s_0} ds s^N \text{Im} \Pi^{\text{had}}(s + i\epsilon) = \frac{-1}{2\pi i} \oint_{|s|=s_0} ds s^N \Pi^{\text{QCD}}(s)$$

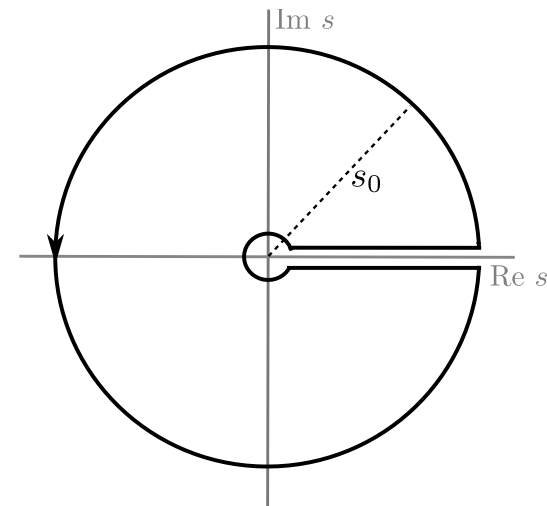
Operator Product Expansion (OPE)

$$\Pi^{\text{QCD}}(x, y) = \Pi^{\text{pQCD}}(x - y) + \sum_{n>0} \text{Wilson coefficients } C_n(x - y) \langle \Omega | : \mathcal{O}_n(x + y) : | \Omega \rangle$$

condensates

In-medium FESR

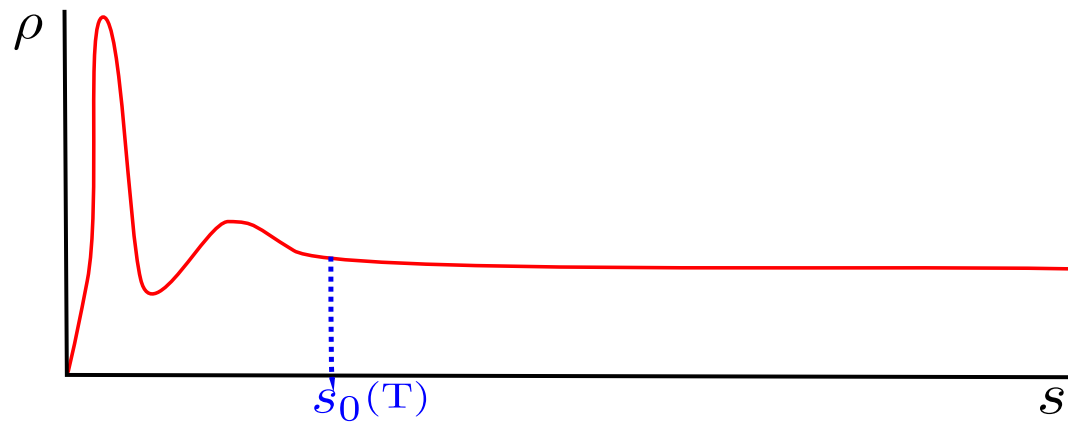
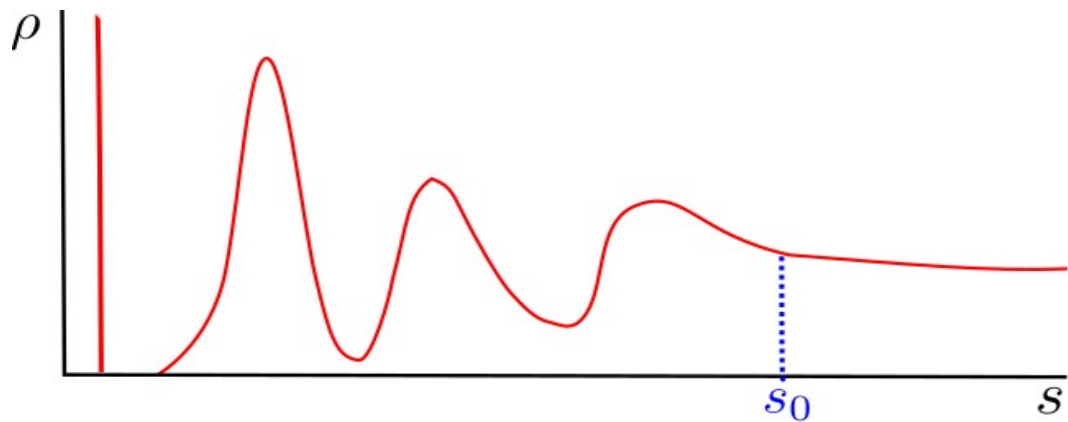
$$\Pi(p; T, B, \mu_q) = \Pi_{\text{pQCD}}(p, \mu; T, B, \mu_q) + \sum_{n>0} C_n(p, \mu) \langle : \mathcal{O}_n(T, B, \mu_q) : \rangle$$



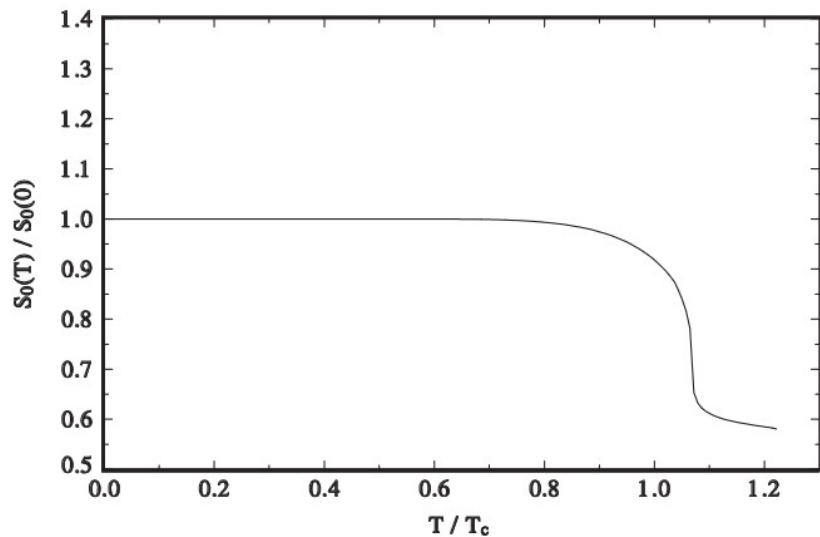
The good of FESR

- Hadronic threshold acts as a deconfinement order parameter
- The kernel s^N cuts the OPE series (in-vacuum, no radiative corrections, low current masses)
- Contour integrals can be performed before loop integration → skip full form factor calculation
- More equations → more parameters to find → more information
- No unphysical Borel mass parameter → no need of window of applicability

Hadronic threshold acts as a deconfinement order parameter



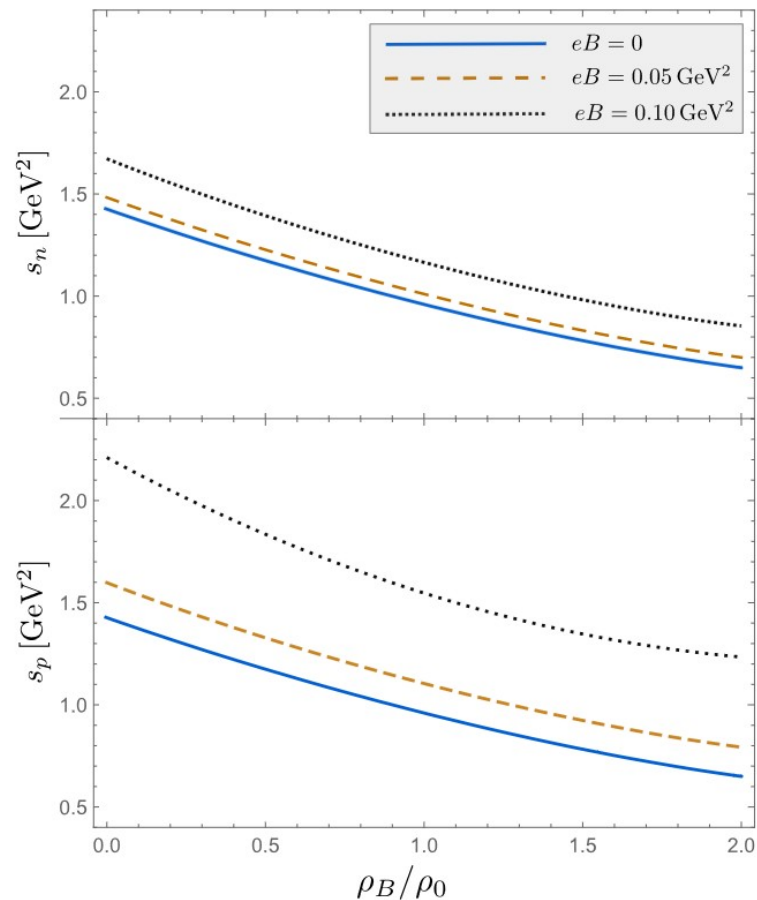
In-medium hadronic thresholds



Dominguez, Loewe, Rojas, Zhang
PRD **81**, 014007 (2010)

- hadronic threshold related with Polyakov loop

Carlomagno, Loewe,
PRD **95**, 036003 (2017)



Dominguez, Loewe, C.V., Zamora
PRD **108**, 074024 (2023)

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In momentum space, high s expansion

$$\Pi^{\text{QCD}}(s) = c_0 \ln(-s) + \sum_{n>0} c_{2n} \frac{\langle \mathcal{O}_{2n} \rangle}{s^n} + \text{regular terms}$$

$$\frac{-1}{2\pi i} \oint_{s_0} ds s^N \Pi^{\text{QCD}}(s) = \frac{c_0 s_0^{N+1}}{N+1} + c_{2(N+1)} \langle \mathcal{O}_{2(N+1)} \rangle$$

Integrated Wilson coefficient in general

$$\sim C \left(1 - \frac{4m^2}{s_0} \right) \frac{\langle \mathcal{O}_{2n} \rangle}{s_0^{n-N-1}}$$

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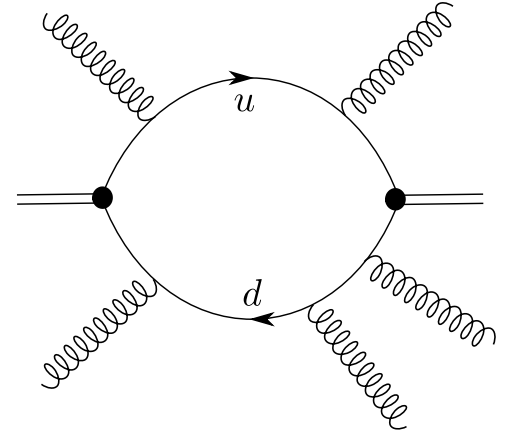
example: one-loop multi-gluon condensate using Feynman parameters

$$I_n^N(s_0) = \frac{-1}{2\pi i} \oint_{C(s_0)} ds s^N \int_0^1 dx \frac{f(x)}{[s - M^2(x)]^n}$$

$$M^2(x) \equiv \frac{m_u^2}{1-x} + \frac{m_d^2}{x}$$

$$I_{N+1}^N(s_0) = - \int_0^1 dx f(x) \theta(s_0 - M^2(x)) = - \int_{x_-}^{x_+} dx f(x)$$

$$x_{\pm} = \frac{1}{2} \left[1 + \frac{m_d^2}{s_0} - \frac{m_u^2}{s_0} \pm \sqrt{1 - 2 \left(\frac{m_d^2}{s_0} + \frac{m_u^2}{s_0} \right) + \left(\frac{m_d^2}{s_0} - \frac{m_u^2}{s_0} \right)^2} \right].$$



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The bad of FESR

- Strong dependence on Hadronic continuum threshold \rightarrow scale dependence
- Some care when Hadronic threshold diminishes too much \rightarrow OPE “convergence”

The ugly of FESR

- Too many equations! \rightarrow to extract all the info is an art

Example: charmonium using AA, VV, AP, channels

$$\mu = 3 \text{ GeV} \quad \Rightarrow \quad m_c = 1 \text{ GeV}$$

$$\sqrt{s_V} = 2.70 \text{ GeV}, \quad \sqrt{s_A} = 2.62 \text{ GeV}, \quad \sqrt{s_P} = 2.55 \text{ GeV}$$

$$\sqrt{s_0} > 3.1 \text{ GeV}$$

$$f_{J/\Psi} = 0.30 \text{ GeV}, \quad f_{\eta_c} = 0.25 \text{ GeV}, \quad \left\langle \frac{\alpha_s}{\pi} GG \right\rangle = 0.035 \text{ GeV}^4$$

$$\mu = m_c, \quad \Rightarrow \quad m_c = 1.27 \text{ GeV}$$

$$\sqrt{s_V} = 3.32 \text{ GeV}, \quad \sqrt{s_A} = 3.13 \text{ GeV}, \quad \sqrt{s_P} = 3.20 \text{ GeV}$$

$$f_{J/\Psi} = 0.28 \text{ GeV}, \quad f_{\eta_c} = 0.23 \text{ GeV}, \quad \left\langle \frac{\alpha_s}{\pi} GG \right\rangle = 0.022 \text{ GeV}^4$$

$$\mu = 1 \text{ GeV} \quad \Rightarrow \quad m_c = 1.42 \text{ GeV}$$

$$\sqrt{s_V} = 3.72 \text{ GeV}, \quad \sqrt{s_A} = 3.50 \text{ GeV}, \quad \sqrt{s_P} = 2.57 \text{ GeV}$$

$$\sqrt{s_0} \lesssim 3.6 \text{ GeV}$$

$$f_{J/\Psi} = 0.28 \text{ GeV}, \quad f_{\eta_c} = 0.23 \text{ GeV}, \quad \left\langle \frac{\alpha_s}{\pi} GG \right\rangle = 0.022 \text{ GeV}^4$$

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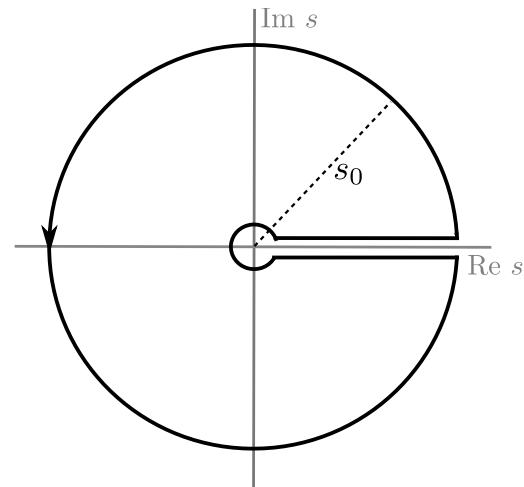
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The “scattering term”

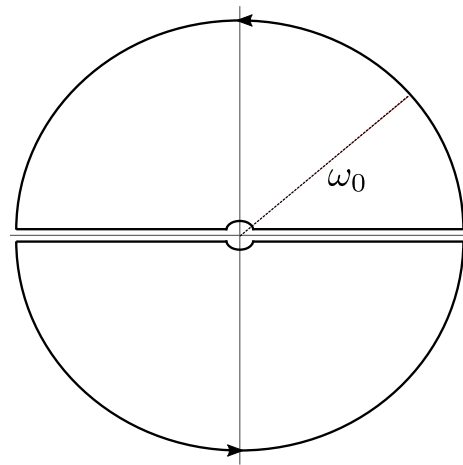
$$\Pi(\omega, \mathbf{p}^2) = \Pi^{\text{even}}(\omega^2, \mathbf{p}^2) + \omega \Pi^{\text{odd}}(\omega^2, \mathbf{p}^2) \quad \omega^2 = s$$

$$\begin{aligned} \int_0^{s_0} \frac{ds}{\pi} s^{N-1} \text{Im} \Pi^{\text{had}}(s + i\epsilon, \mathbf{p}^2) \\ = - \oint_{|s|=s_0} \frac{ds}{2\pi i} s^{N-1} \Pi^{\text{QCD}}(s, \mathbf{p}^2) + \text{Res}_{s \rightarrow 0} \Pi^{\text{QCD}}(s, \mathbf{p}^2) \end{aligned}$$



Alternative integration path $p_0 = \omega$ instead of $p^2 = s$

$$\begin{aligned} \int_{-\omega_0}^{\omega_0} \frac{d\omega}{\pi} \omega^{n+1} \text{Im} \Pi^{\text{had}}(\omega + i\epsilon, \mathbf{p}^2) \\ = - \oint_{\omega_0} \frac{d\omega}{2\pi i} \omega^{n+1} \Pi^{\text{QCD}}(\omega, \mathbf{p}^2) + \text{Res}_{\omega=0} [\omega^{n+1} \Pi^{\text{QCD}}(\omega, \mathbf{p}^2)] \end{aligned}$$



The “scattering term”

Bochkarev, Shaposhnikov, Nucl. Phys. B **268**, 220–252 (1986).

Dominguez, Loewe, Phys. Lett. B **233**, 201–204 (1989).

Furnstahl, Hatsuda, Lee, Phys. Rev. D **42**, 1744–1756 (1990)

appears at finite temperature and/or chemical potential

careful with these limits:

$$m_u = m_d, \quad \mathbf{p} = 0$$

scattering term can be obtained in rest frame keeping different masses

Operator mixing

$$\Pi(s) = \Pi_{\text{pQCD}}(s, \mu) + \sum_{n>0} C_n(s) \langle : \mathcal{O}_n : \rangle \quad \text{normal ordered condensates}$$

Problems with the chiral limit: $\ln(-s/m_q^2)$ terms

limits $m_u \rightarrow 0, m_d \rightarrow 0 \neq m_d \rightarrow 0, m_u \rightarrow 0 \neq m_u = m_d \rightarrow 0$

Solution:

$$\Pi(s) = \tilde{\Pi}_{\text{pQCD}}(s, \mu) + \sum_{n>0} \tilde{C}_n(s, \mu) \langle \mathcal{O}_n \rangle(\mu) \quad \text{non-normal ordered condensates}$$

$$\langle : \bar{q} \Gamma q : \rangle = \langle \bar{q} \Gamma q \rangle + \text{tr} \int \frac{d^d k}{(2\pi)^d} \Gamma S_q(k, \mu; G) \quad \text{Gluon background field}$$

→ This replaces $\ln(-s/m_q^2) \rightarrow \ln(-s/\mu^2)$

In-medium operator mixing

$$\langle : \bar{q} \Gamma q : \rangle = \langle \bar{q} \Gamma q \rangle + \text{tr} \int \frac{d^d k}{(2\pi)^d} \Gamma S_q(k, \mu; \textcolor{blue}{G}, \textcolor{red}{T}, \textcolor{red}{B}, \mu_q)$$

At finite magnetic field: removes log terms

$$eB^2 \ln(-s/m_q^2) \rightarrow eB^2 \ln(-s/\mu^2)$$

Gubler, Hattori, Lee, Oka, Ozaki, Suzuki, Phys. Rev. D **93**, 054026 (2016).

At finite chemical potential:

- Silver Blaze behavior in pion channels
- Predicted condensate values will be incorrect if not considered

Raya, Villavicencio, Eur. Phys. J. C **85**, 13 (2025)

At finite temperature: probably must be considered

Some properties of in-medium FESR

Finite magnetic field

QCD sector \rightarrow B-expansion $\left(\frac{eB}{s_0(B)} \right)^n$ must be treated as the gluon background

eB for light quarks can reach values $\gtrsim 0.1 \text{ GeV}^2$

Hadronic threshold increases with B

Hadron sector \rightarrow Sum of Landau levels with Euler-Maclaurin formula

Some properties of in-medium FESR

Finite chemical potential \rightarrow phase transition in pion channels

$$\mu_q > \mu_c = \sqrt{s_A}/2$$

$$\text{In vacuum} \quad \mu_c \sim 0.4 \text{ GeV}$$

$$f_\pi^2 = \frac{\mu^2}{2\pi^2}(1+a)$$

$$v_\pi^2 = \frac{\frac{2}{5} - a}{1+a}$$

$$\langle \bar{q}q \rangle \sim m \quad m_\pi^2 \sim m^2$$

$$a = \frac{4\pi^2}{9} \frac{\langle \theta_g \rangle}{s_A \mu^2}$$

Raya, Villavicencio, Eur. Phys. J. C **85**, 13 (2025)

Apparently describes Color-Flavor-Locked phase $\rightarrow SU(3)_{c+L+R}$ models for high μ

$$f_\pi \sim 0.2\mu$$

$$v_\pi^2 \sim \frac{1}{3}$$

$$\langle \bar{q}q \rangle \sim m$$

$$m_\pi^2 \sim m m_s$$

Some results

Axial-Axial, Axial-Pseudoscalar, Pseudoscalar Pseudoscalar channels

nucleon – nucleon correlator

masses, decay constants, hadronic thresholds, condensates of pion at finite B

nucleon – Axial current - nucleon correlator

nucleon axial coupling constant at finite B and ρ_B

Summary

- All the in-medium and vacuum parameters obtained are consistent with other approaches
- Contour integration before loop integration → skips full correlator calculation
→ low mass expansion not necessary
- Scattering term and in-medium operator mixed **must** be incorporated
- Magnetic field expansion can reach high values → considered as operator expanded
- For $\mu > \mu_c$ the model (apparently) describes $SU(3)_{c+L+R}$ models

Outlook

A lot of incoming projects

References

A.Raya and C.Villavicencio,
Finite-energy sum rules at finite chemical potential and zero temperature,
Eur. Phys. J. C **85**, 13 (2025)

C.A. Dominguez, M. Loewe, C. Villavicencio and R.Zamora,
Nucleon axial-vector coupling constant in magnetar environments,
Phys. Rev. D **108**, 074024 (2023)

C. Villavicencio,
Axial coupling constant in a magnetic background,
Phys. Rev. D **107**, 076009 (2023)

C.A. Dominguez, L.A. Hernández, M. Loewe, C. Villavicencio and R. Zamora,
Magnetic field dependence of nucleon parameters from QCD sum rules,
Phys. Rev. D **102**, 094007 (2020)

C.A. Dominguez, M. Loewe and C. Villavicencio,
QCD determination of the magnetic field dependence of QCD and hadronic parameters,
Phys. Rev. D **98**, 034015 (2018)

A. Ayala, C.A. Dominguez, L.A. Hernandez, M. Loewe, J.C. Rojas and C. Villavicencio,
Quark deconfinement and gluon condensate in a weak magnetic field from QCD sum rules,
Phys. Rev. D **92**, 016006 (2015)