

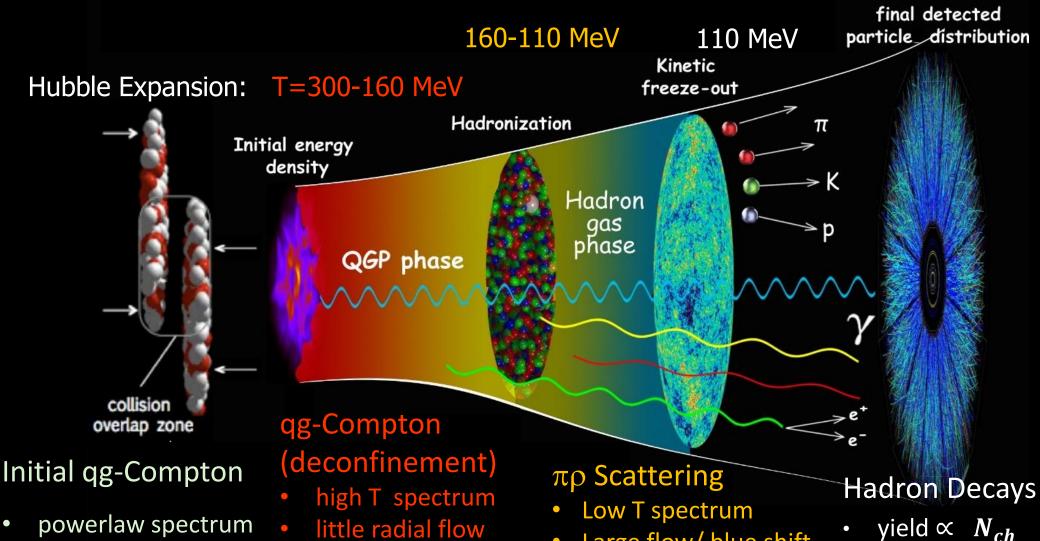


# The big picture



# **Electromagnetic Radiation in A+A Collisions:**





• yield  $\propto N_{ch}^{\alpha} \alpha \leq 2$ 



- $\propto N_{coll}$
- no collective motion

- Large flow/ blue shift
- yield  $\propto N_{ch}^{\alpha} \alpha > 1$
- Spectra derived from parent particles

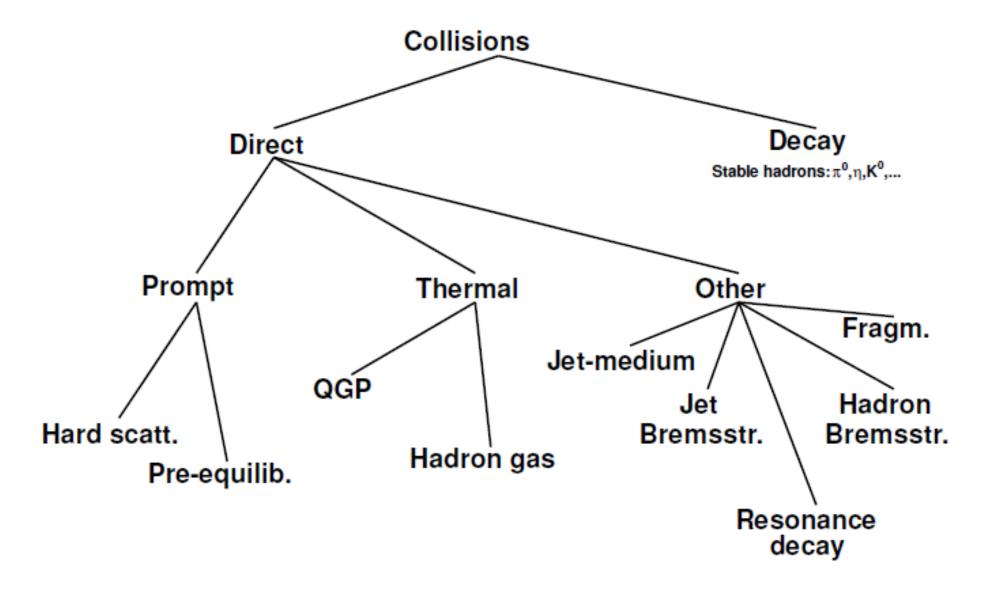


# The terminology Historians of the collision





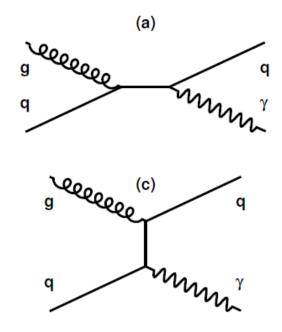
# **Terminology – reflecting different sources**

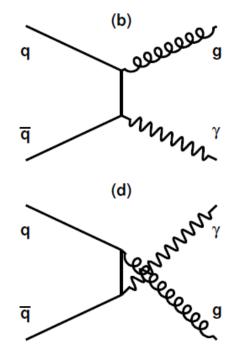


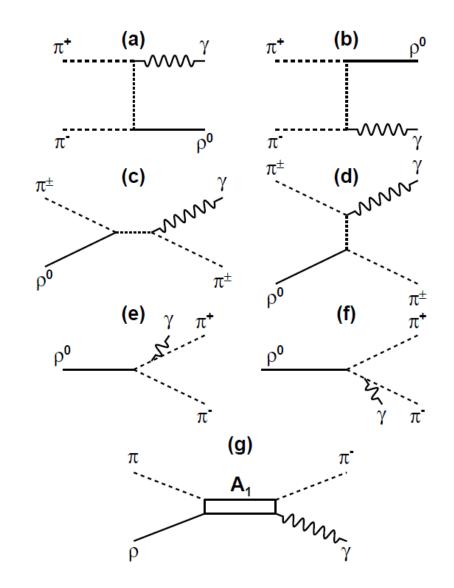




# The basic partonic and hadronic sources







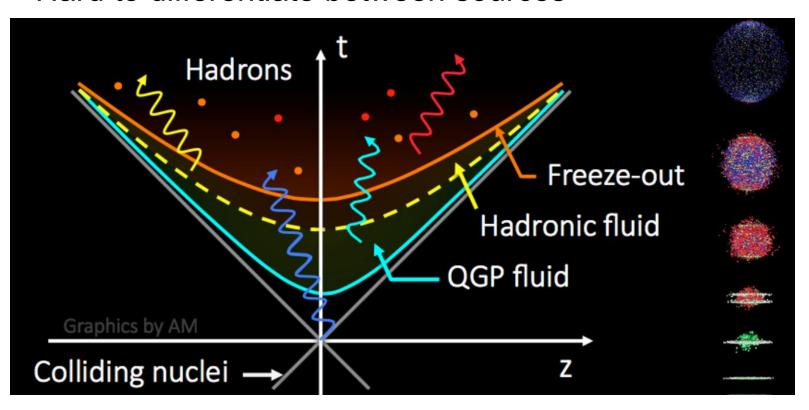


# The blessing – and curse – of photons

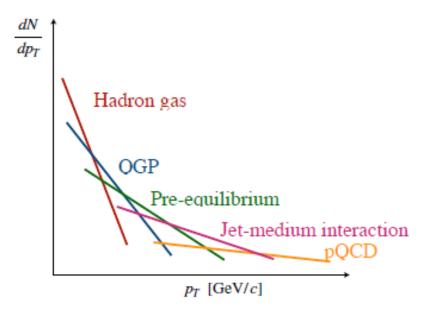


# Penetrating probes, created at all stages

Huge background (FS decays, ~10% S/B) Hard to differentiate between sources





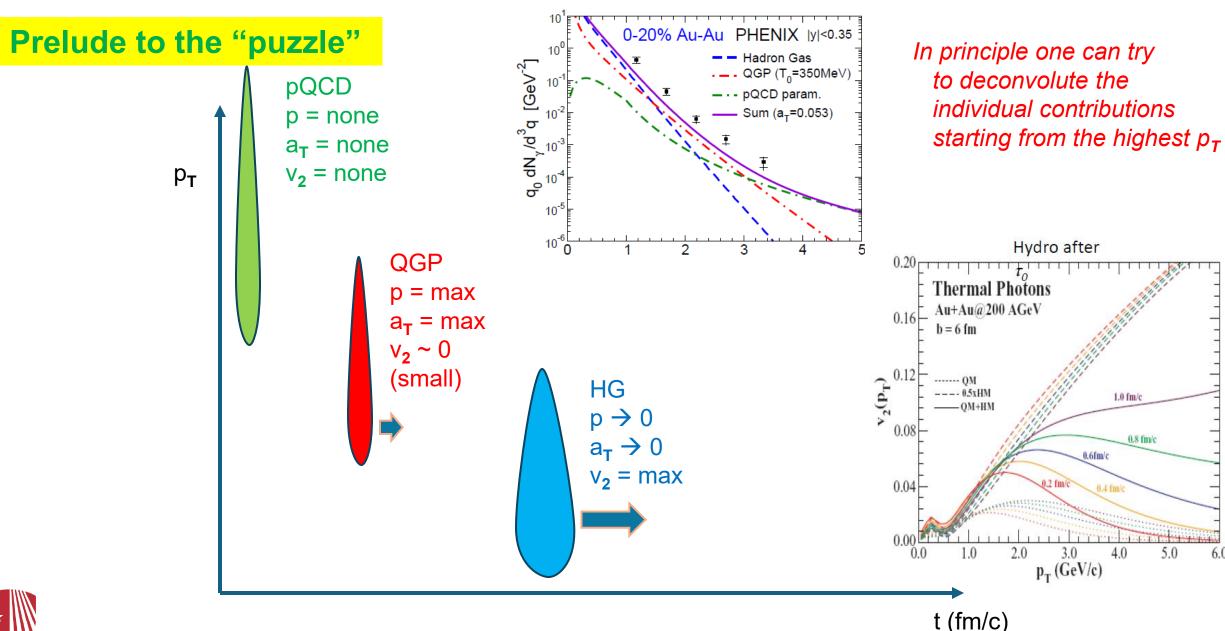




"Historians" of the collision – but convoluted

# **Dominant photon sources (simplified)**





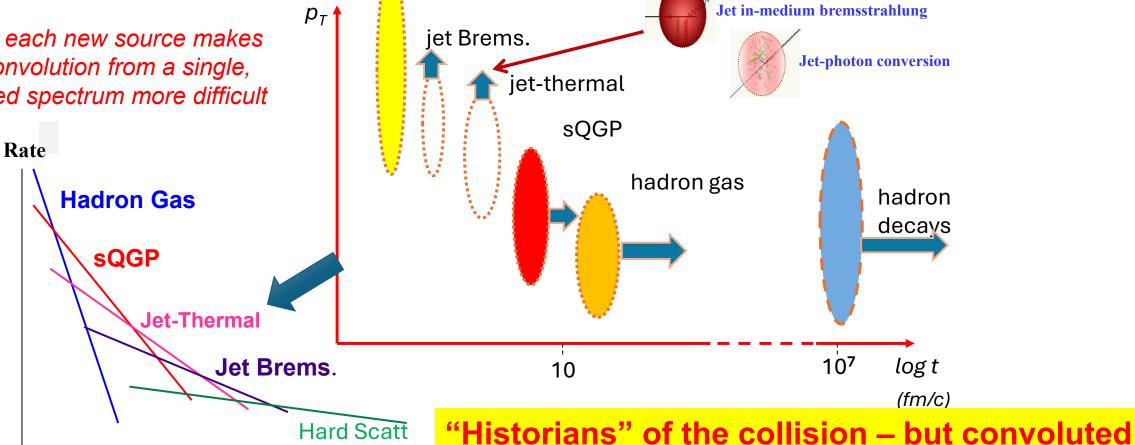
#### More sources...

hard scatt



Jet fragmentation, Jet-thermal interaction (jet-photon conversion) Initial magnetic field Bremsstrahlung (hadron gas) ...???

Obviously each new source makes the deconvolution from a single, integrated spectrum more difficult



 $E_{\gamma}$ 



See e.g., Turbide, Gale, Jeon and

Moore, PRC <u>72</u>, 014906 (2005)

parton-medium interaction



# The reference – or is it? (p+p)

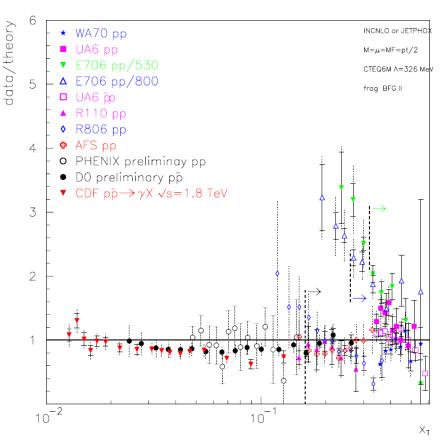


# Direct photons in hadron (p+p) collisions, 22 - 7000 GeV

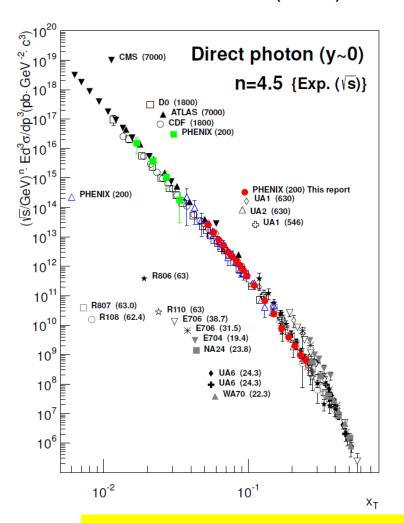


# High $p_T$ , scaling with $x_T$

PRD 73, 094007 (2006)



PRD 86, 072008 (2012)



$$E\frac{d^3\sigma}{dp^3} = \frac{d^3\sigma}{p_T dp_T dy d\phi} = \frac{1}{p_T^{n_{\text{eff}}(x_T,\sqrt{s})}} F\left(\frac{p_T}{\sqrt{s}}\right)$$
$$= \frac{1}{(\sqrt{s})^{n_{\text{eff}}(x_T,\sqrt{s})}} G(x_T),$$

Over 18 orders of magnitude

Without evolution of  $\alpha_s$  and the structure and fragm. functions n=4



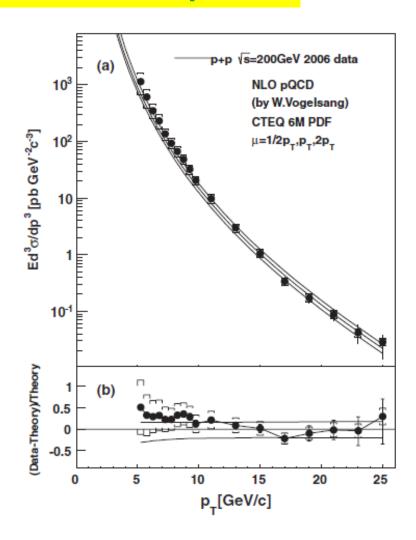
High p<sub>⊤</sub> reasonably well understood

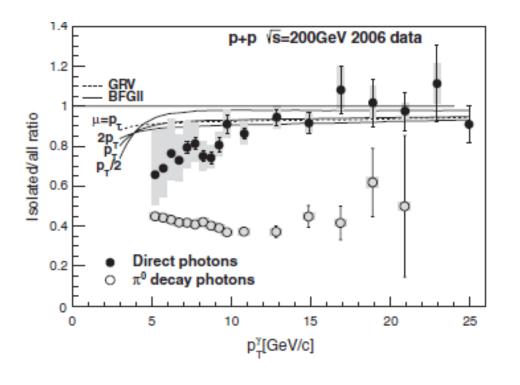
# Direct photons in hadron (200 GeV p+p) collisions, 2012



# **Photons vs NLO pQCD**

PRD 86, 072008 (2012)





Fraction of isolated direct photons

Fragmentation very small above 10 GeV/c

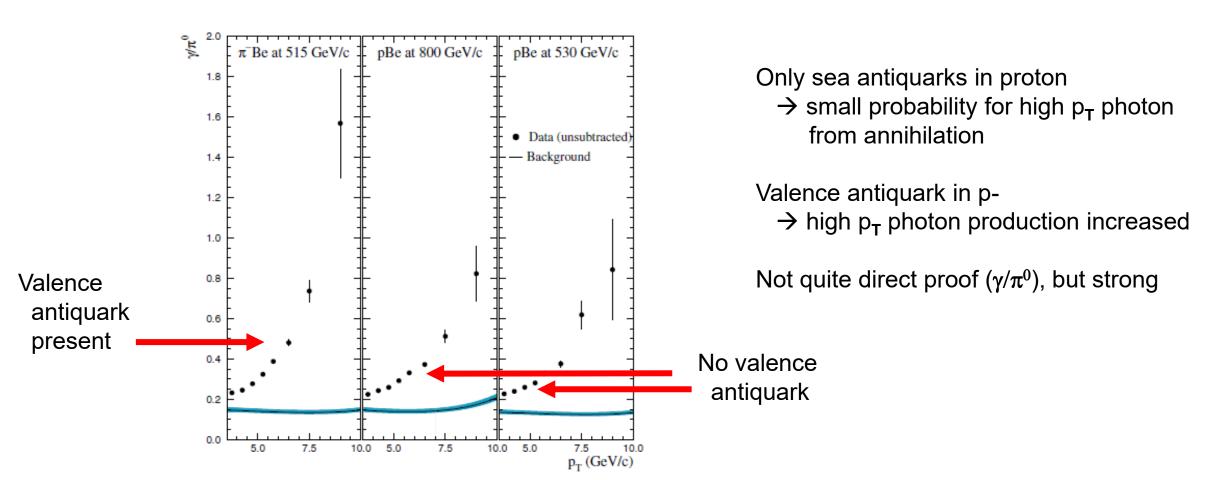


# Interlude: Compton scattering dominates over annihilation



E706, fixed target, p and  $\pi^{-}$  beam

PRD 70, 092009 (2004)



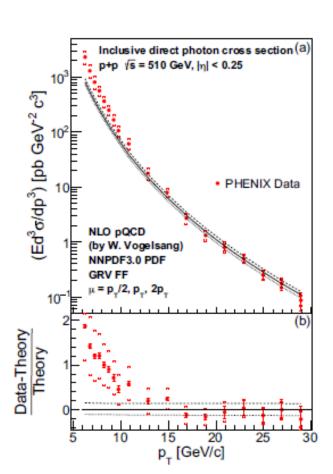


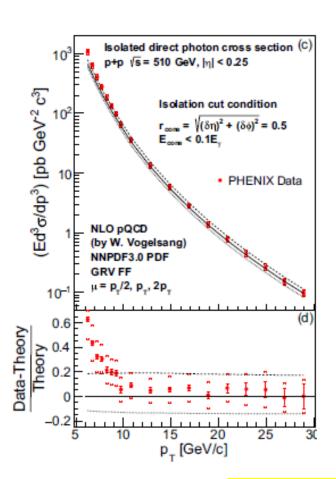
We knew this, but good to see experimentally

# Direct photons in polarized 510 GeV p+p collisions



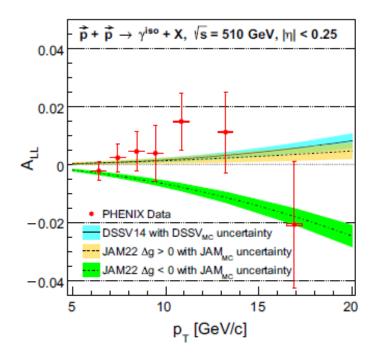
# Polarized gluon contribution to proton spin?





PRL 130, 251901 (2023)

A<sub>LL</sub> – longitudinal double spin asymmetry



Disfavors netagive  $\Delta g$ 



Similar discrepancy below 10 GeV/c, as in 200 GeV collisions – even for isolated photons

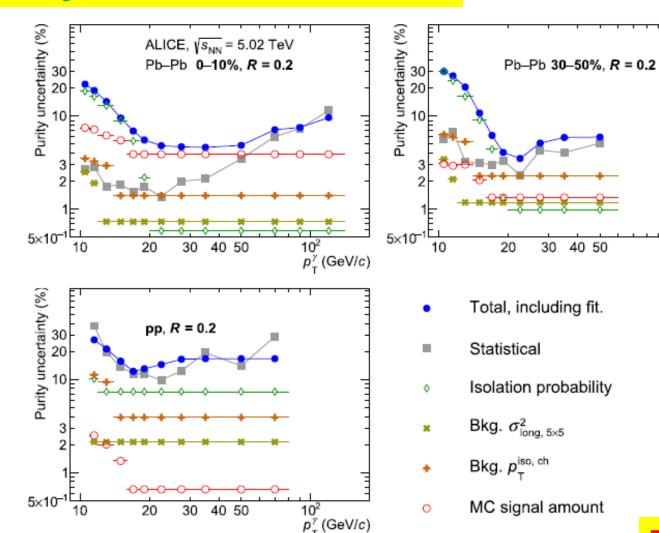
Small, positive contribution from gluon

# **Technical – isolated photon uncertainties (ALICE)**

 $p_{\tau}^{\gamma}$  (GeV/c)



# Can you still isolate in PbPb?



Eur. Phys. J. C (2025) 85:553

→ this is the overwhelming term in the total uncertainty on cross-section

#### Isolation probability

→ ratio of shape distributions of iso and non-iso constant?

#### MC signal amount

varying signal (γ-jet) by 20%w.r.t. bgd (jet-jet)



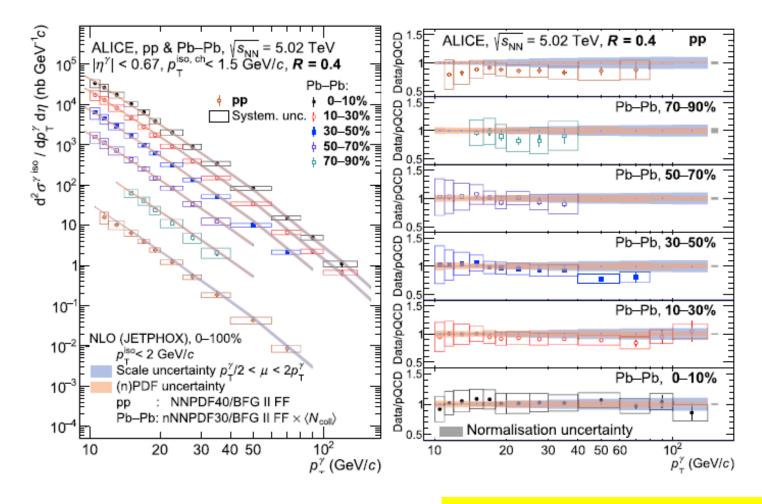
Reverse ordering in PbPb vs pp

# High p<sub>T</sub> isolated photons,, pp and PbPb 5.02 TeV (ALICE)



# Is the high p<sub>T</sub> region understood?

Eur. Phys. J. C (2025) 85:553



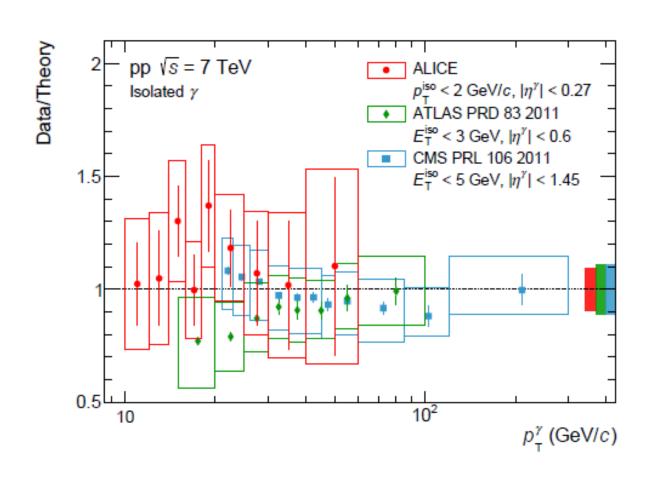


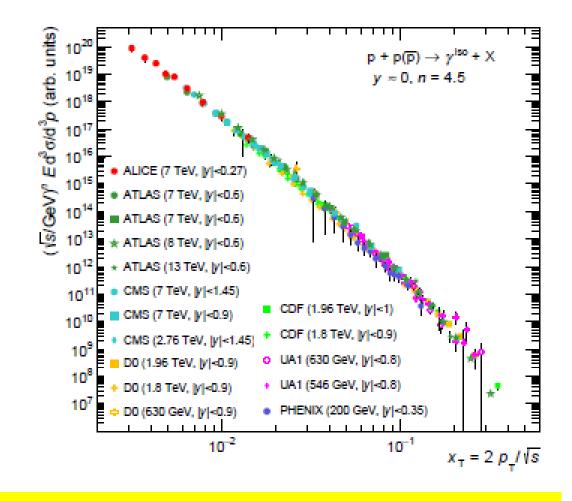
# High p<sub>T</sub> isolated photons,, pp 7 TeV (ALICE, ATLAS, CMS)



# Is the high p<sub>T</sub> region understood?

Eur. Phys. J. C (2019) 79:896







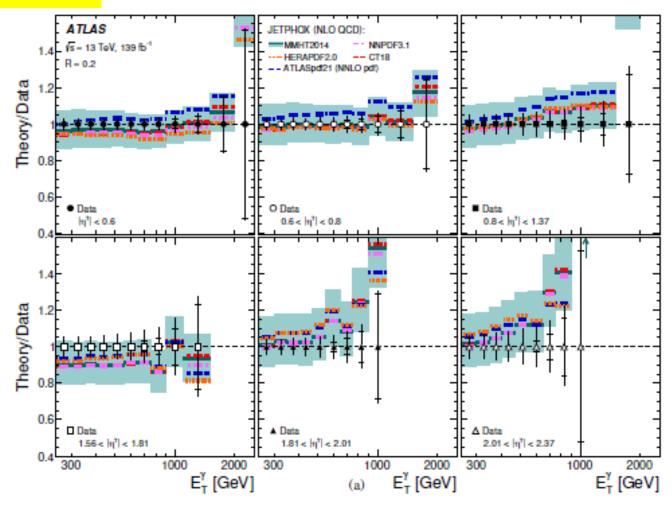
# All consistent with JETPHOX pQCD

# High $p_T$ isolated photons,, pp 13 TeV (ATLAS)



# Rapidity dependence?

JHEP07 (2023) 086





Not well described by JETPHOX pQCD

# Digression: low p<sub>T</sub> photons,, pp, 13 TeV (ALICE)

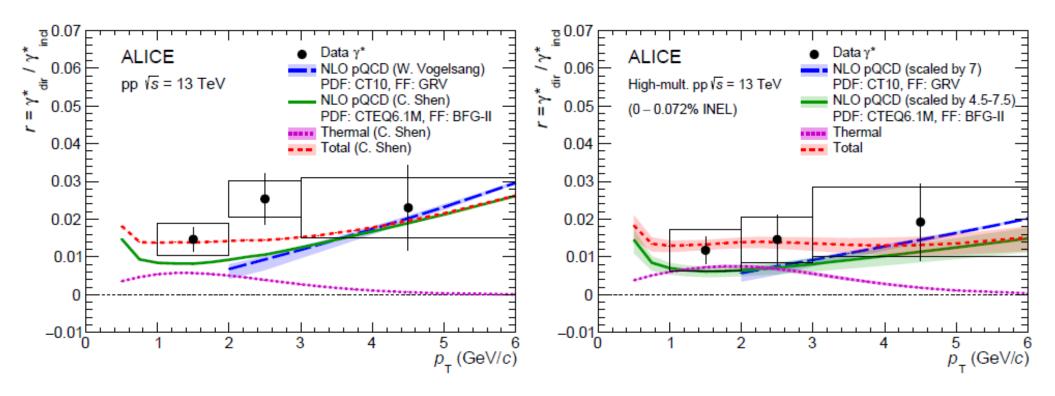


# QGP even in high multiplicity pp?

Something like QGP formed already in the highest multiplicity pp events, too? Internal conversion  $\rightarrow \gamma^*$ 

PLB 868 (2025) 139645

$$r = \gamma^*_{dir} / \gamma^*_{had}$$





No sign of additional "thermal" radiation

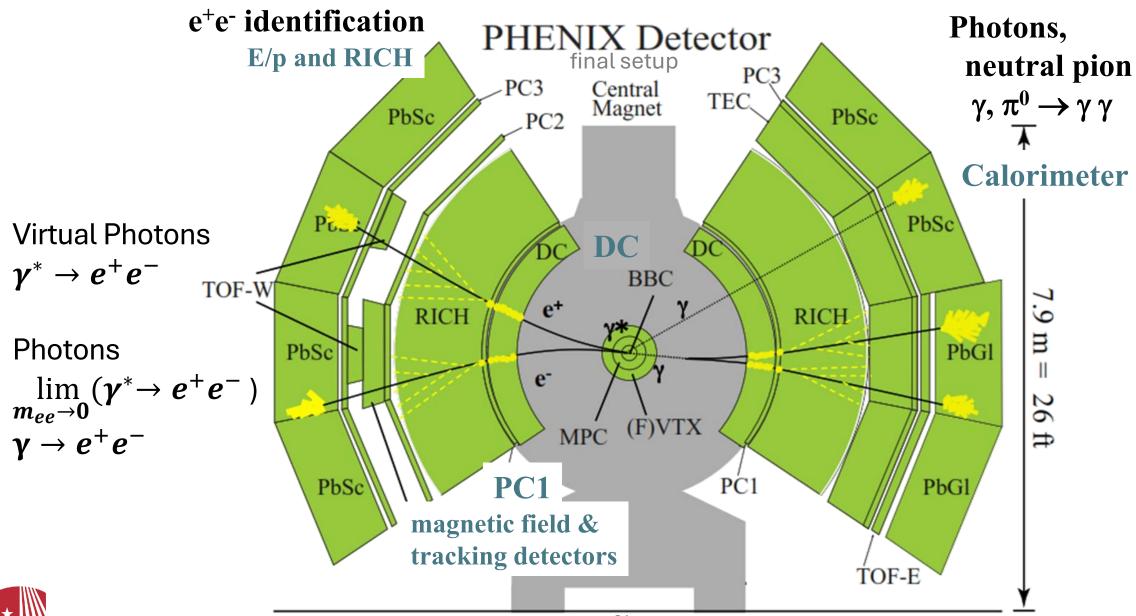


# **Experimental techniques**



### **Photon measurements in PHENIX**







# **Basic techniques of photon measurements:**



# Real photon captured in an electromagnetic calorimeter

#### Electromagnetic calorimeter

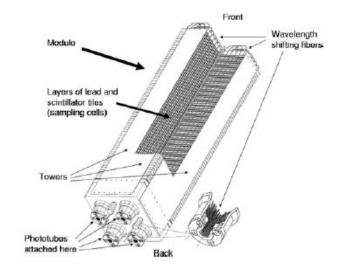
- relatively cheap, large acceptance possible (sampling)
- resolution improves with energy (ideal at high p<sub>T</sub>)
- "thin" for hadrons
- → high energy cluster (almost) always electromagnetic deposited energy counts, not particle momentum!
- measurement can be almost standalone

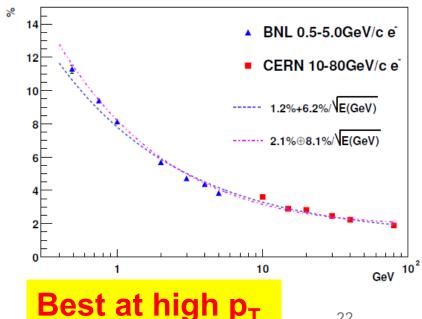
#### But

- issues with resolution and hadron background at low p<sub>T</sub>
- issues with merged  $\pi^0$  decay photons at very high p<sub>T</sub>

These can be mitigated: shower maximum detectors (SMD), small Moliere-radius crystal calorimeters, digital calorimeters

→ but at substantial cost increase



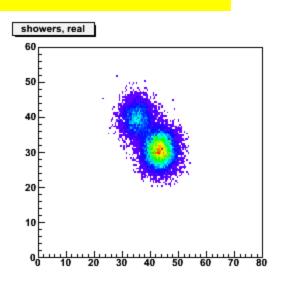


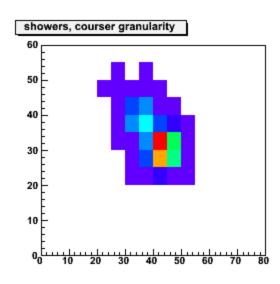


# Close-by showers – real and as seen

# 0

### Nature meets detector...





Very high granularity:
well separated,
good measurement
for both energy and
position

Medium granularity:
well separated,
decent measurement
for energy, some
shift in position,
pair p<sub>T</sub> still correct

Low granularity:

poor separation.

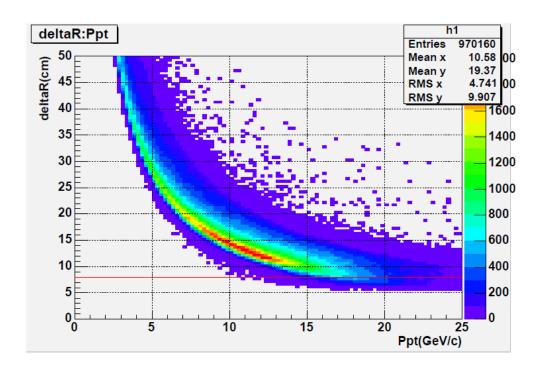
questionable energy
sharing, position
shift,
pair p<sub>T</sub> still ~correct



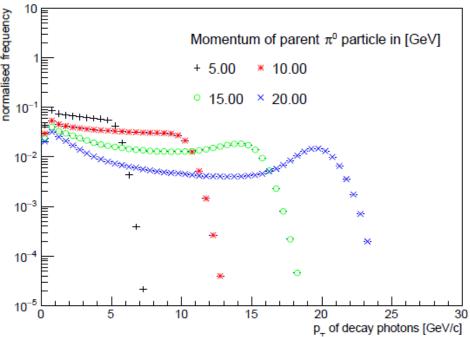
# Zooming in on the $\pi^0$ problem (at high $p_T$ )



Do you think  $\pi^0$  decays to two ordinary photons? WRONG! It decays to two correlated photons



PHENIX – distance between two clusters photons from a  $\pi^0$  decay



The distribution of measured decay photons should be flat up to the kinematic limit. Instead, you see depletion at medium, enhancement at high  $p_T$ . Unless you fully understand this, your direct photons at high  $p_T$  will be strongly overestimated!

Possible clue to E706?



# Basic techniques of photon measurements:

external conversion



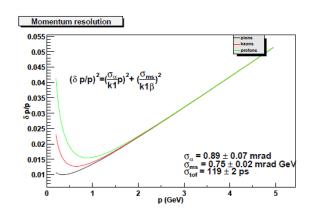
# Real photon converts to e<sup>+</sup>e<sup>-</sup> on detector material

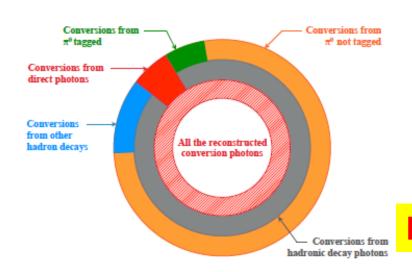
#### Pro:

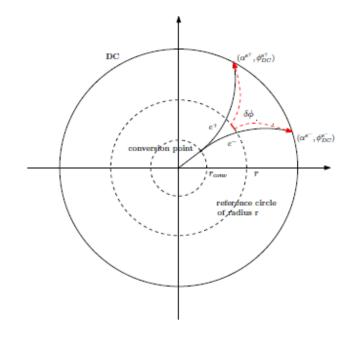
- tracking: higher resolution
- viable at low p<sub>T</sub>
- not sensitive to Dalitz (origin!)
- feasible in "crowded" detectors

#### Con

- electron ID difficult
- small acceptance\*efficiency
- sensitive to material location
- cocktail (discuss later)







# Method of choice at low p<sub>T</sub>



## **External conversion, PHENIX**

BG/FG = 3.0%

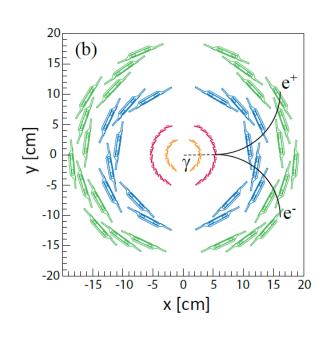
BG/FG = 0.4%

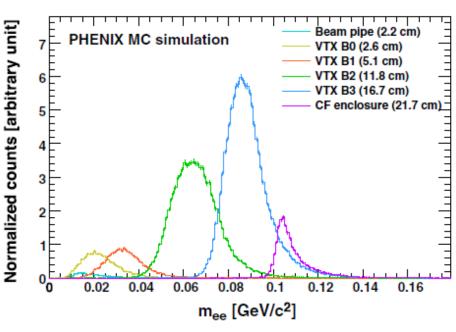
0.15

PRC 109, 044912 (2024)

 $1.0 < p_T < 1.2 \text{ GeV/c}$ 

# Real photon converts to e<sup>+</sup>e<sup>-</sup> on detector material

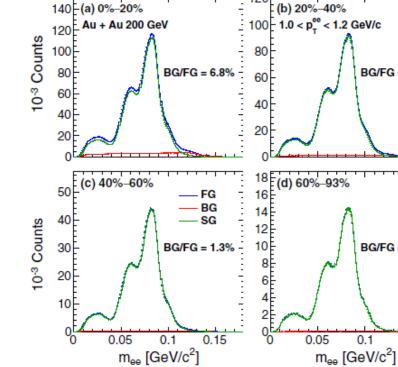


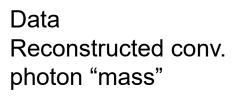


Simulation Reconstructed conv. photon "mass" (DC tracking assumes collision vtx!)









Viable at low p<sub>T</sub> − complements calorimetry

0.1

# Tagging photons from hadron decay



# **Majority of photon are from hadron decays**

#### On the face of it simple:

- pair up conversion photons with real photons in the calorimeter
- tag those that reconstruct a  $\pi^0$

#### But:

- tremendous combinatorial at low p<sub>T</sub>
- limited acceptance, efficiency
- ignores double conversions

#### Delicate balance:

- increasing material increases conversion (i.e. our signal)
- too much material promotes double conversions

# $M_{ee\gamma}$

Conv + real photon invariant mass

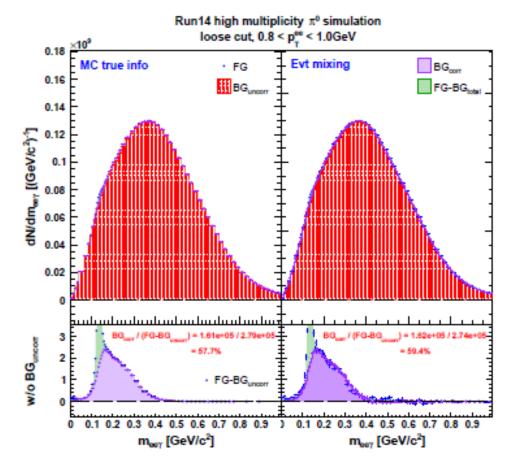


Figure 3.26: Invariant mass distributions of  $e^+e^-\gamma$  pairs.

Large combinatorial



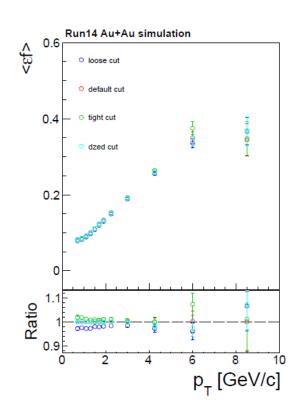
# Conditional probability of tagging, Ry



# How often is the partner $\pi^0$ photon found in the calorimeter?

$$\left(\frac{N_{\gamma}^{\text{incl}}}{N_{\gamma}^{\pi^{0},\text{tag}}}\right)_{\text{Data}} = \frac{\gamma^{\text{incl}} \cdot p_{\text{eonv}} \cdot a_{e^{+}e^{-}} \cdot \epsilon_{e^{+}e^{-}}}{\gamma^{\pi^{0}} \cdot p_{\text{eonv}} \cdot a_{e^{+}e^{-}} \cdot \epsilon_{e^{+}e^{-}} \cdot \langle \epsilon_{\gamma} f \rangle} = \frac{\gamma^{\text{incl}}}{\gamma^{\pi^{0}} \cdot \langle \epsilon_{\gamma} f \rangle}$$

Many systematics cancel



After accounting for other hadrons (cocktail, to be discussed later)

$$R_{\gamma} = rac{N_{\gamma}^{incl}}{N_{\gamma}^{hadr}} = rac{\left\langle \mathcal{E}f \right\rangle imes \left( rac{1 \cdot \gamma}{N_{\gamma}^{\pi^{0}tag}} \right)}{\left( rac{N_{\gamma}^{hadr}}{N_{\gamma}^{\pi^{0}}} 
ight)^{MC}}$$



Powerful for R<sub>y</sub> in terms of uncertainties

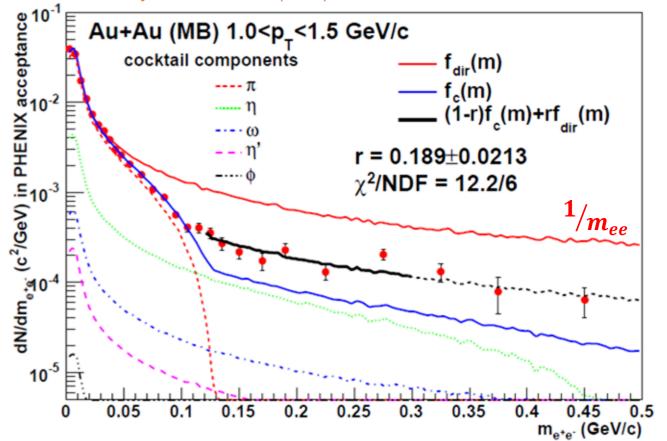
# Virtual photon ("internal conversion") method





$$\frac{d^2 n_{ee}}{dm_{ee}} = \frac{2\alpha}{3\pi} \frac{1}{m_{ee}} \sqrt{1 - \frac{4m_e^2}{m_{ee}^2}} \left(1 + \frac{2m_e^2}{m_{ee}^2}\right) S dn_{\gamma}$$

PHENIX: Phys. Rev. Lett. 104 (2010) 132301



 $\gamma_{dir}$ :  $S \sim 1$  for  $m_{ee} \ll p_T$ 

$$\gamma_{had}$$
:  $S = \left(1 - \frac{m_{ee}^2}{M_{had}^2}\right)^3 |F(m_{ee}^2)|^2$ 

- Using virtual photons  $\gamma^* 
  ightarrow e^+ e^-$ :
  - any process that radiates  $\gamma$  will also radiate  $\gamma^* o e^+ e^-$
  - ullet for  $m_{ee} \ll p_{T}$  extrapolate  $oldsymbol{\gamma}^{*}$  to  $m_{ee} = oldsymbol{0}$
  - $m_{ee} > m_{\pi}$  cut improves S/B by factor 10
  - sys. uncertainty cancelation in ratio  $\gamma^*_{dir}/\gamma^*_{incl}$

Works above 1 GeV/c Hadronic "cocktail" important



Direct  $\gamma^*$  yield fitted in range 120 to 300 MeV Insensitive to  $\pi^0$  yield



# Direct photons in heavy ion collisions



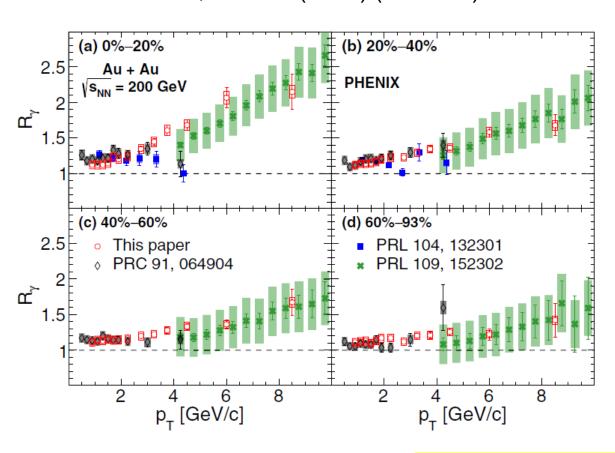
# Direct photons with multiple methods (PHENIX, ALICE)

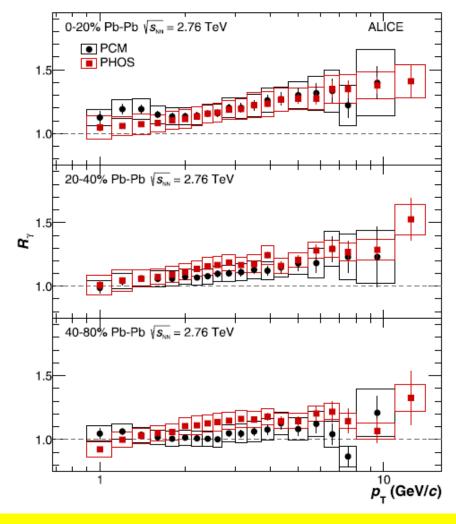


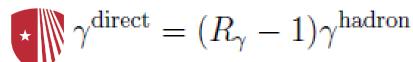
# R<sub>y</sub> with different methods

PLB 754 (2016) 235-248 (ALICE)

PRC 109, 044912 (2024) (PHENIX)







Remember this, when interpreting spectra

# **Invariant yields – 200 GeV AuAu – PHENIX**



# **Direct photons and non-prompt photons**

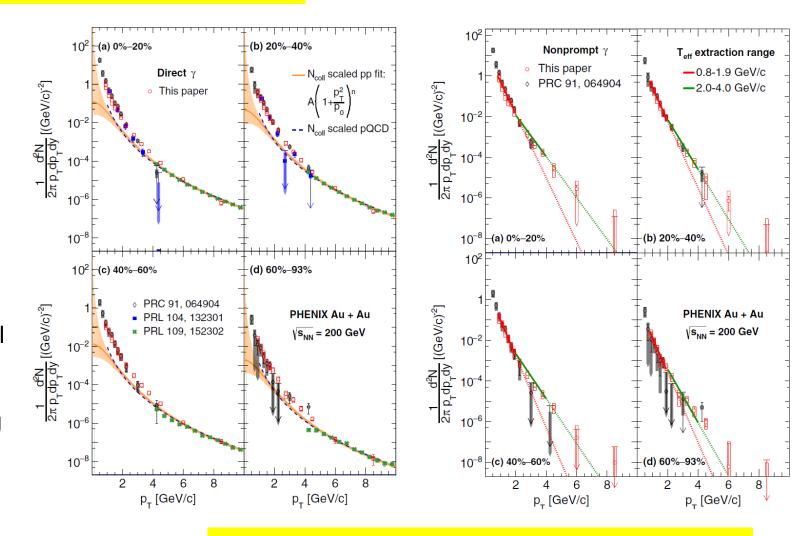
PRC 109, 044912 (2024)

Non-prompt: properly scaled p+p yields subtracted

N<sub>coll</sub> → number of binary nucleon-nucleon collisions (estimated with Glauber MC)

At lower p<sub>T</sub> varying exponential slopes (T<sub>eff</sub>)

Above 4 GeV/c hard scattering dominates





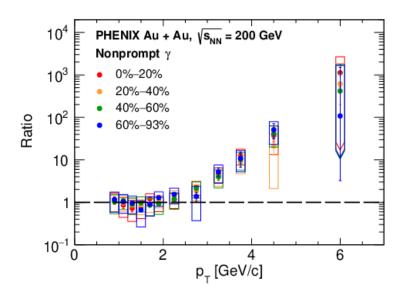
Three methods, consistent results

# T<sub>eff</sub> vs the p<sub>T</sub> range and centrality

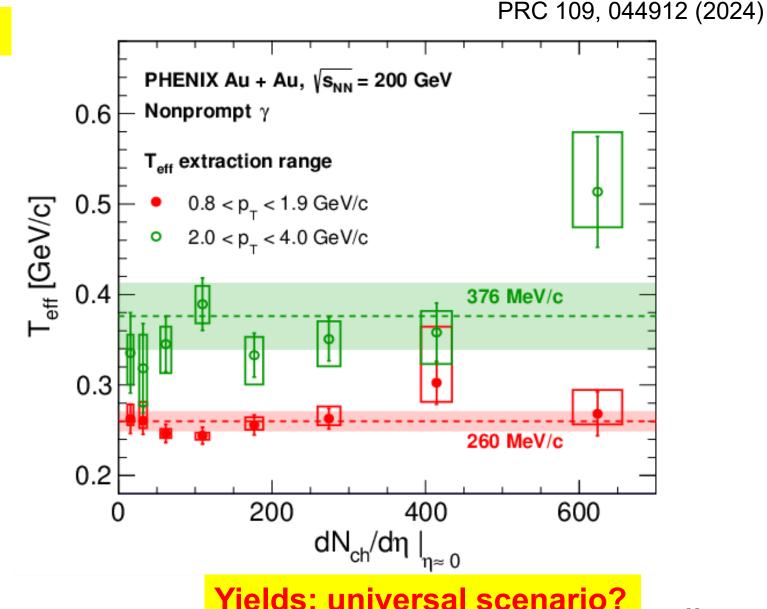


# No dependence on centrality

Strong dependence on p<sub>T</sub> Composition and relative weight of different sources always the same?



Ratio of the yield to a fixed exponential (T<sub>eff</sub> = 0.26 GeV)



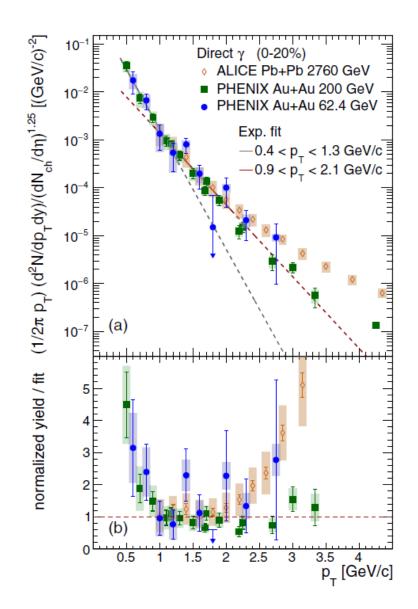


# Slope vs p<sub>T</sub>

# 0

# Dependence on collision energy

PRC 107, 024914 (2023)



Data vary from 62 to 2760 GeV!

Slopes vary with pT

→ imprint of T<sub>eff</sub> evolution?

(different times, different sources, different p<sub>T</sub>)

Normalized yields very similar at low p<sub>T</sub>

Possible interpretation:
no matter where the QGP started, the endgame
will always be the same
(no surprise, but good to see)

# By freeze-out everything is similar

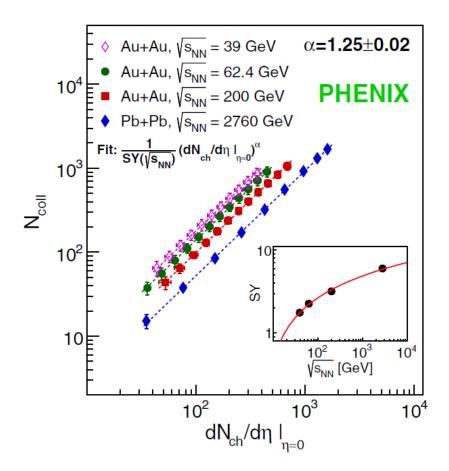


# Scaling with $N_{ch}$ ( $N_{coll}$ )?

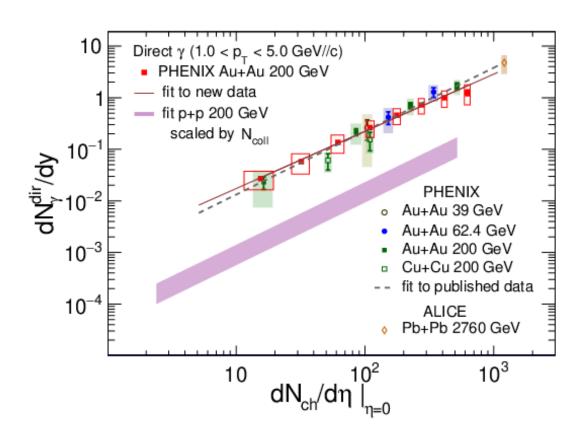


# In larger systems yields scale with N<sub>ch</sub>?

PRL 123, 022301 (2019)



PRC109, 044912 (2024)



Yes, but  $\alpha$  (slope) controversial



#### STAR 200 GeV Au+Au

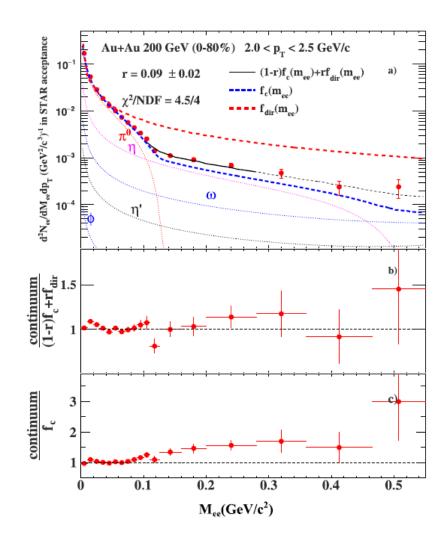


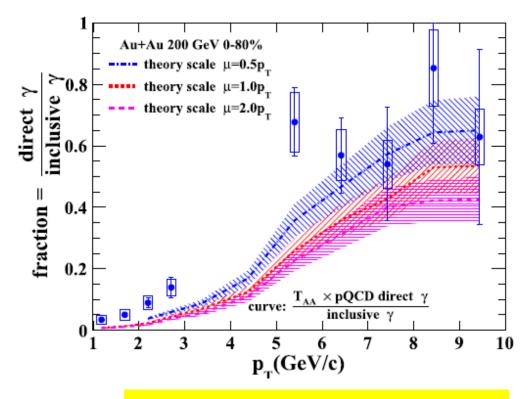
### Two methods: internal conversion / calorimeter

PLB 770 (2017) 451-458

(same as the first PHENIX measurement)

fraction  $\rightarrow$  r = (R $\gamma$  – 1) / R $\gamma$ 





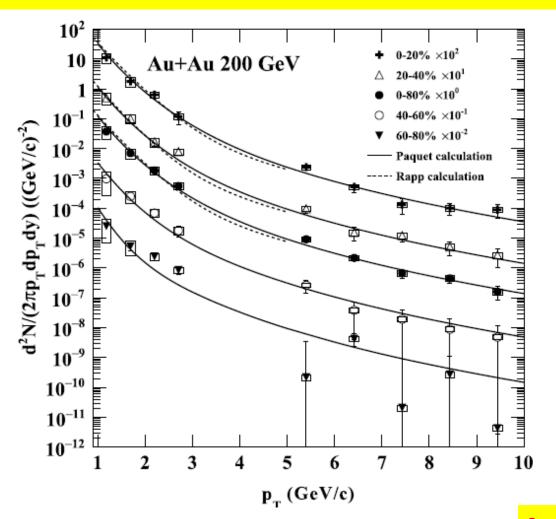
"Tension" with PHENIX, will discuss later



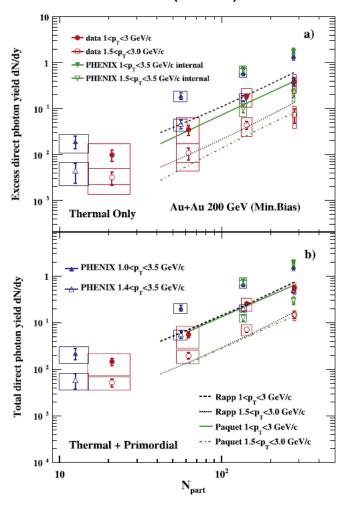
# STAR 200 GeV Au+Au



# Two methods: internal conversion / calorimeter



# PLB 770 (2017) 451-458







# **ALICE 2.76 TeV PbPb -- updated**



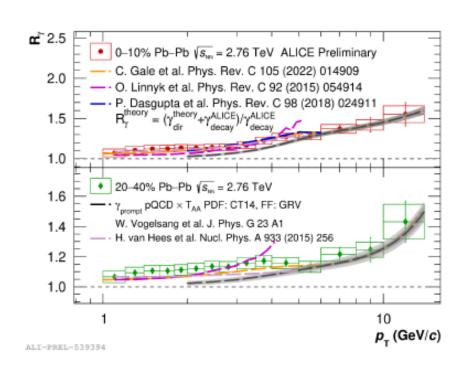
# **Improved significance**

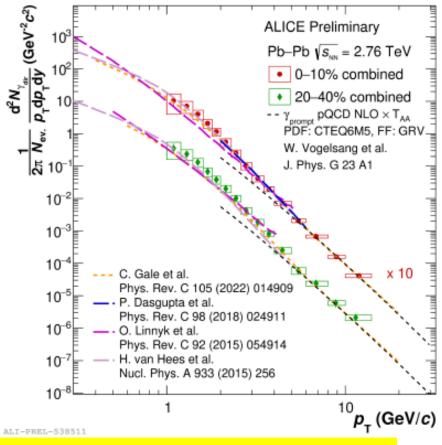
PoS (HardProbes2023) 061

First publication: PLB 754 (2016) 235-248

Direct photon signal down to 1 GeV/c but quite weak

T<sub>eff</sub> ~ 340 MeV for both centralities (1.1-2.1 GeV/c region)





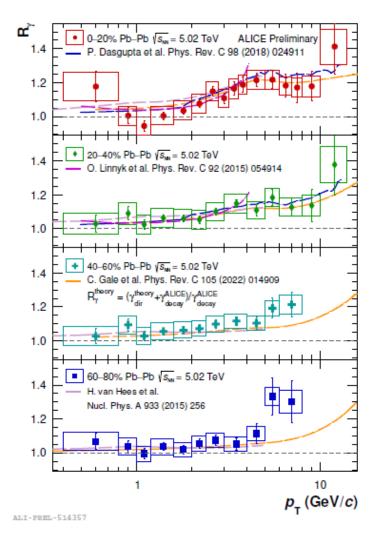


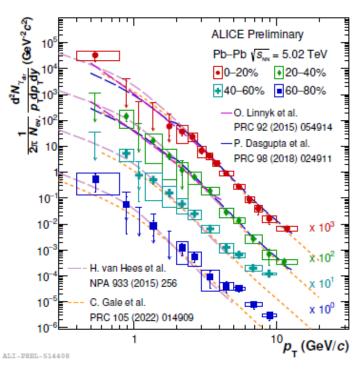
"Thermal" radiation not exceeding PHENIX

# $\otimes$

# **Increase energy**

No significant direct photon signal at low p<sub>T</sub>



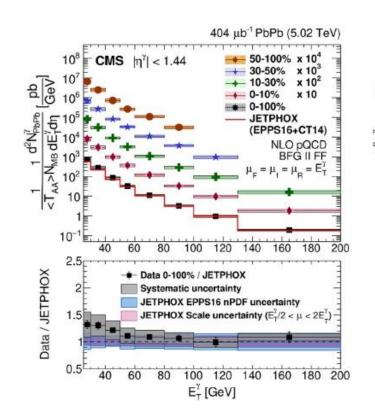


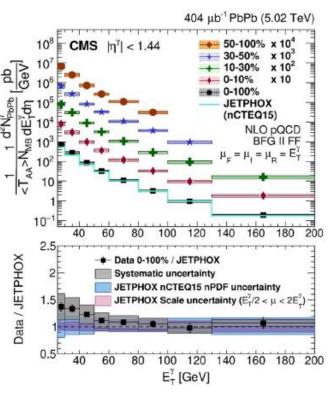


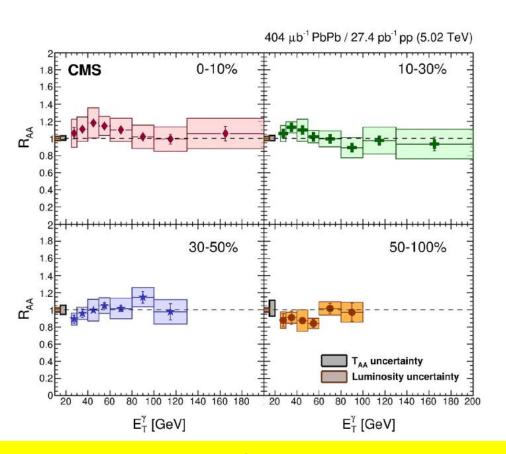
Thermal signal fading out???



# Does high p<sub>T</sub> still look as expected?









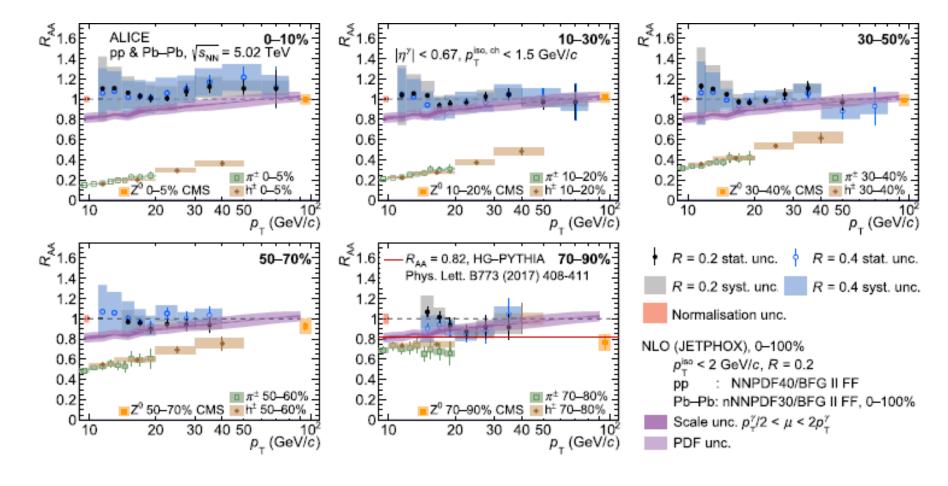


# R<sub>AA</sub> of isolated photons, ALICE PbPb 5.02 TeV



# **Any nuclear modification?**

Eur. Phys. J. C (2025) 85:553







# The centrality, R<sub>AA</sub> and photon saga



# What's the big deal about $R_{AA} = 1$ ?



# **Connecting impact parameter to observables**

Ann.Rev.Nucl.Part.Sci.57:205-243,2007

3.1 Methodology

**N**<sub>ch</sub> charged mult (large η gap)

N<sub>part</sub>
participating
(wounded) nucleons

N<sub>coll</sub> binary (NN) collisions

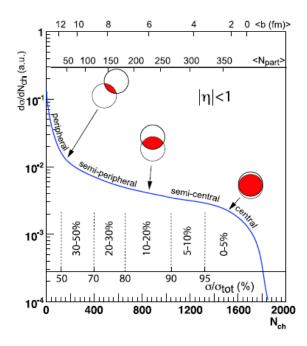
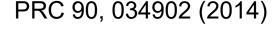
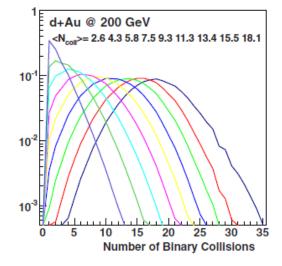
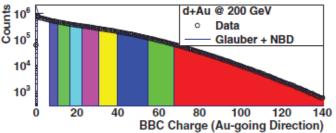


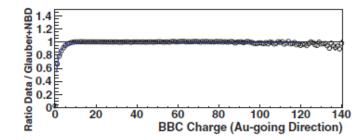
Figure 8: A cartoon example of the correlation of the final state observable  $N_{\rm ch}$  with Glauber calculated quantities  $(b, N_{\rm part})$ . The plotted distribution and various values are illustrative and not actual measurements (T. Ullrich, private communication).

"In heavy ion collisions, we manipulate the fact that the majority of the initial state nucleon-nucleon collisions will be analogous to minimum bias p+p collisions..."









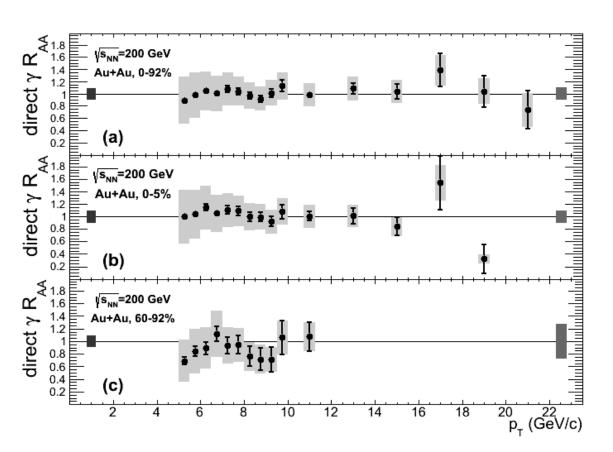


# Glauber MC works in A+A – experimental proof

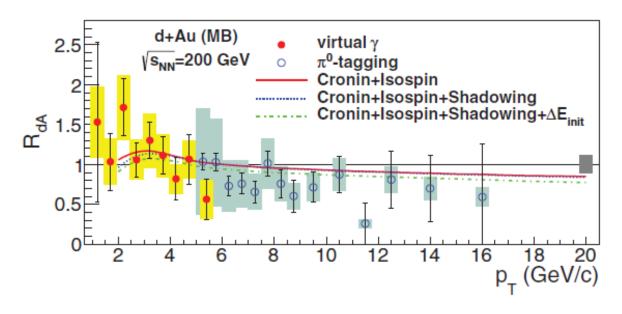


# Hard e.m. probes (mostly) immune to FS effects





PHYSICAL REVIEW C 87, 054907 (2013)



Large uncertainties, but hints of isospin effect



PHENIX, PRL 109, 152302 (2012)

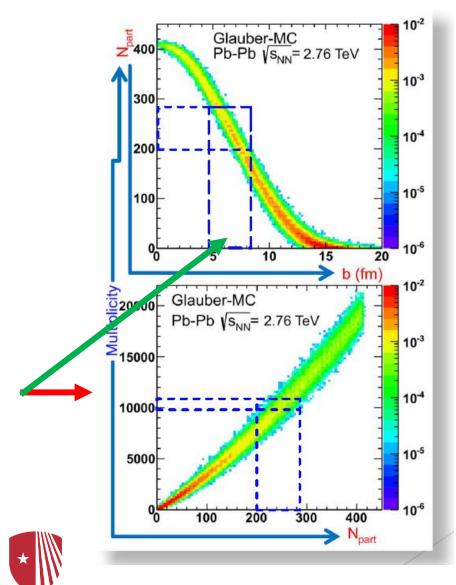
Some caveats (like isospin effect)

# (Un)ambiguity of centrality determination – PbPb, pPb

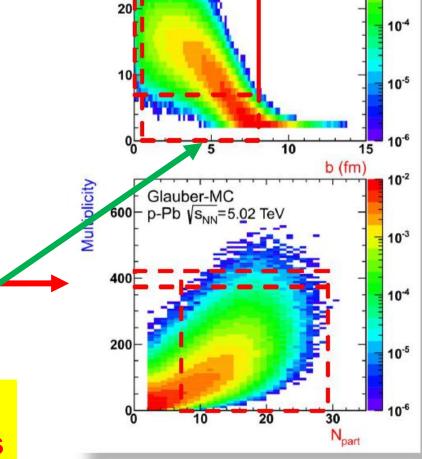


10<sup>-3</sup>





N<sub>ch</sub>, N<sub>part</sub>, **b** correlation tight in PbPb, very loose in pPb



Glauber-MC

n-Ph vs\_=5.02 TeV

Ambiguity in small systems

# Also: centrality from bulk observables – but where?



# Signal: $\eta \sim 0$ , "bulk": $|\eta| >> 0$

ALICE PRC 91 (2015) 064905

CL1  $\rightarrow$   $|\eta|$  < 0.9

V0A  $\rightarrow$  2.8 <  $\eta$  < 5.1 (Pb-going side)

V0C  $\rightarrow$  -3.7 <  $\eta$  < -1.7 (p-going side)

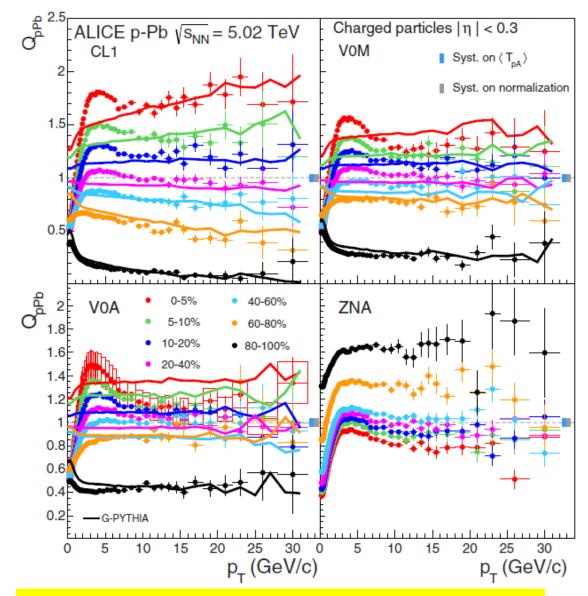
 $VOM \rightarrow VOA + VOC$ 

ZNA → ZDC on Pb-going side

Watch strong auto-correlation in CL1 central, jet veto bias in peripherals

Smaller fluctuations in V0A, mostly around unity, except vastly displaced peripheral due to multiplicity bias (?)

Reverse ordering for ZNA, as expected (as expected???)





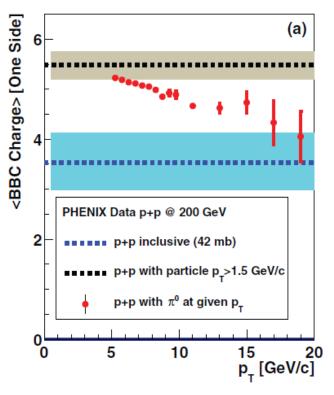
# Color transparency or energy conservation in $R_{xA}$ ?



# Max $p_T$ at $\eta \sim 0$ , $N_{ch}$ at $|\eta| >> 0$ in p+p

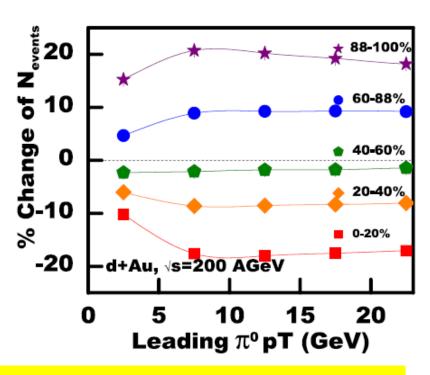
Kordell, Majumder PRC 97, 054907 (2018)

"...the puzzling enhancement in peripheral events ... as well as the suppression seen in central events... are possibly due to *mis*-binning of central and semicentral events, containing a jet, as peripheral events... due to suppression of soft particle production away from the jet, caused by the depletion of energy available in a nucleon of the deuteron in d-Au or proton in p-Pb after the production of a hard jet... "



PRC 90, 034902 (2014)

Initial enhancement of N<sub>ch</sub> forward with mid-rapidity p<sub>T</sub>, then depletion (p+p data)



Anticorrelation at sufficiently high p<sub>T</sub>

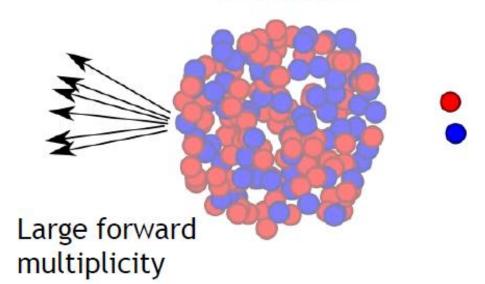
# Very large $p_T$ at $\eta = 0$ – mis-binning of centrality?

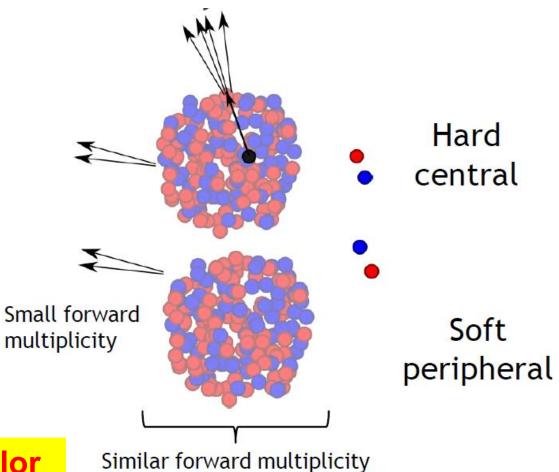
 $\otimes$ 

The same (few) projectile nucleons have to produce both hard scattering and "bulk" forward

This is essentially an energy conservation argument

# Soft central





You can give it another name (like "color fluctuation") but it is still the same thing

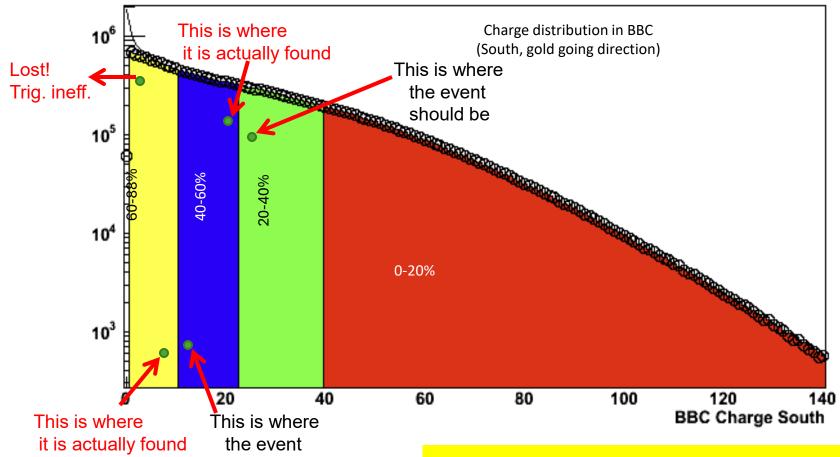


# Illustration: shift between multiplicity classes



# Mapping N<sub>ch</sub> to b using average events...

should be



...breaks down for special (high  $p_T$ ) events



If (experimental) centrality is determined with fixed (forward) multiplicity thresholds, irrespective of what happened at  $\eta$ ~0, events may end up in the wrong centrality class – and attributed an incorrect  $\langle N_{coll} \rangle$ 

# Does bulk observable-based centrality fix N<sub>coll</sub> once and for all?

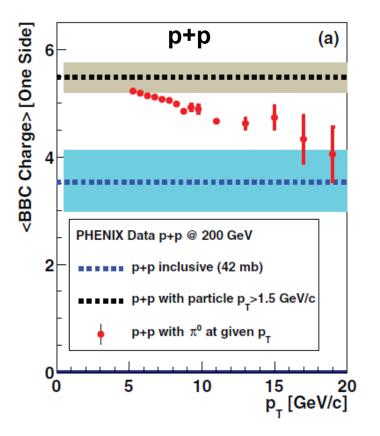


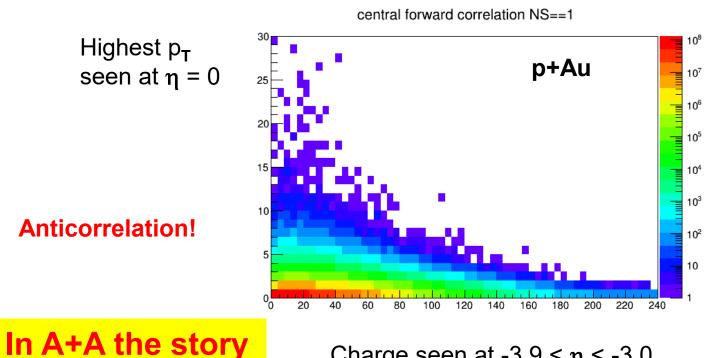
# One more time...

"In heavy ion collisions, we manipulate the fact that the majority of the initial state nucleon-nucleon collisions will be analogous to minimum bias p+p collisions..."

No, it is biased, and the bias changes as a function of the hardest scattering seen at mid-rapidity!

is very different!





PRC 90, 034902 (2014)

Charge seen at  $-3.9 < \eta < -3.0$ 

# A way out: actually measure N<sub>coll</sub>



# You still categorize events with N<sub>ch</sub>, but override Glauber MC

Is it possible? Yes, at least you can get close, and at the very least get rid of fake final state effects in  $R_{xA}$ .

Remember, photons don't care about FS → mostly true, at high p<sub>T</sub> most of them are from initial hard scattering and have 200+ *fm* mean free path in QGP (e.g. *Rept.Prog.Phys.* 83 (2020) 4, 046301)

For an arbitrary "centrality" classification just take the ratio of the direct photon and hadron spectra

- → pure centrality bias (even if p<sub>T</sub>-dependent) will affect both similarly
- if the ratios change with centrality, there's a genuine final state effect on hadrons

Same idea, different realization: you can get  $N_{coll}$  experimentally from the  $Y^g(AB,p_T)/Y^g(pp,p_T)$  direct photon yield ratios

$$N_{coll}^{EXP} = \frac{\left(\frac{d^2 N_{\gamma}}{dp_T d\eta}\right)_{AB}}{\left(\frac{d^2 N_{\gamma}}{dp_T d\eta}\right)_{pp}} = \frac{Y_{AB}^{\gamma}}{Y_{pp}^{\gamma}}$$

Some caveats...

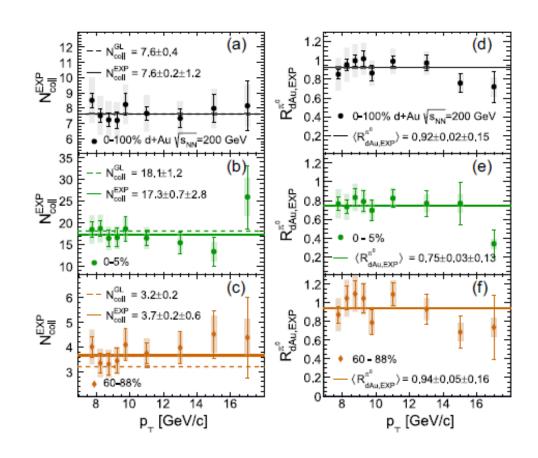


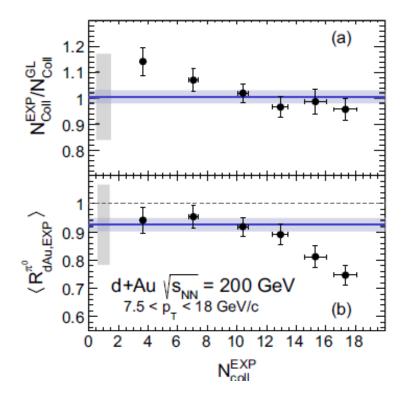
# Nuclear modification of $\pi^0$ in dAu



# Earlier unphysical enhancement in peripherals gone

PRL 134, 022302 (2025)





Ultimate test: pAu, dAu, <sup>3</sup>HeAu "excitation" function (in the works)





# PHENIX / STAR discrepancy Main culprit: cocktail?



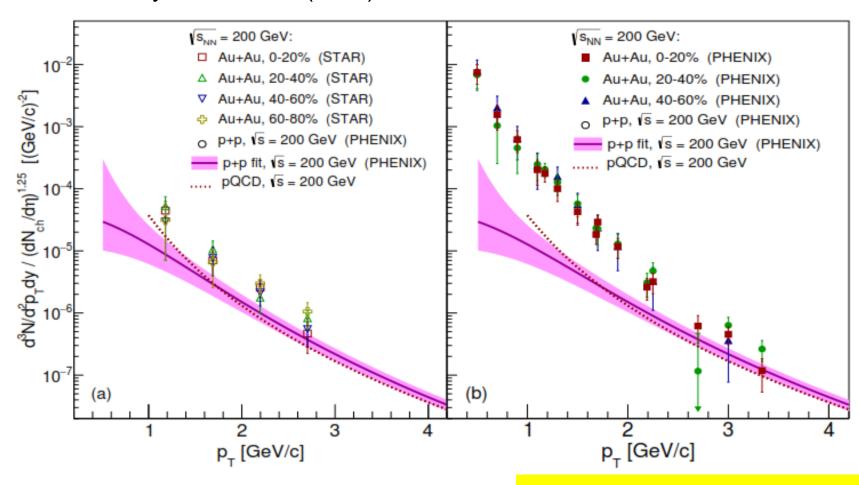
# PHENIX / STAR discrepancy (at low $p_T$ )



A sore point since 2016

STAR: Phys. Lett. B770 (2017) 451

PHENIX  $\gamma \rightarrow e^+e^-$ : Phys. Rev. C91 (2015) 064904  $\gamma^*$ : Phys. Rev. Lett. 104 (2010) 132301



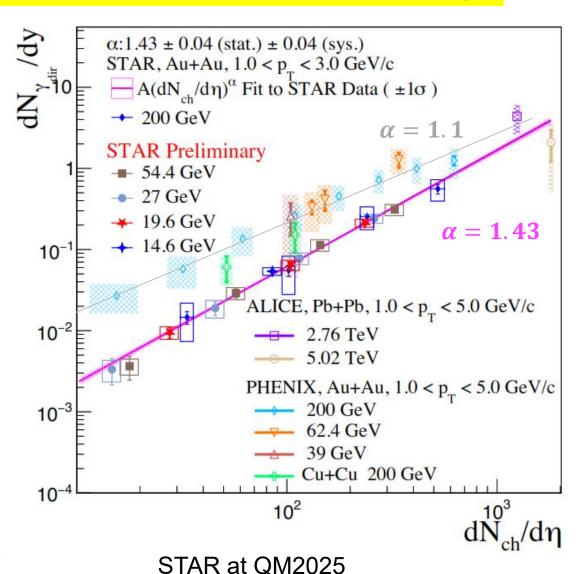


Renewed efforts since QM2025

# Integrated yields vs N<sub>ch</sub> (and slopes)



# **Different evolution with centrality**



### STAR results:

- Virtual photon analysis  $\gamma^* \rightarrow e^+e^-$  for Au+Au at 14.6, 19.6, 27, 54.4, 200 GeV
- Self consistent analyses & results across energies
- Direct photon yield scales with  $N_{\gamma}^{dir} \sim \left(\frac{dN_{ch}}{d\eta}\right)^{\alpha}$  with  $\alpha \sim 1.43$
- Compared to PHENIX results:
  - STAR yield a factor of 3-5 below PHENIX
  - $\alpha \sim 1.43$  versus  $\alpha \sim 1.1$  for PHENIX

Credit: Axel Drees, SBU

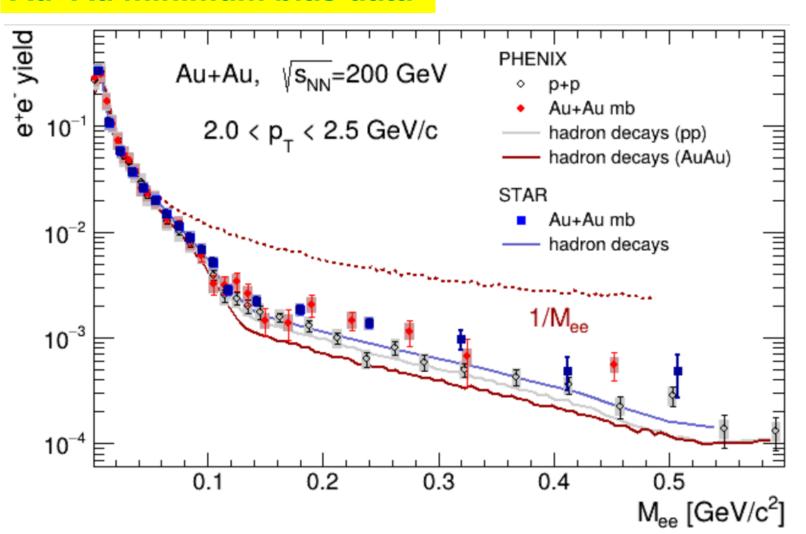
Very different physics message!



# Compare in a particular p<sub>⊤</sub> range

# 0

# Au+Au minimum bias data



Credit: Axel Drees, SBU

Same pt selection

All data is normalized for m<30 MeV

STAR and PHENIX data agree well in  $\pi^0$  and  $\eta$  Dalitz region

 $\eta/\pi^0$  at low p<sub>T</sub> higher in STAR

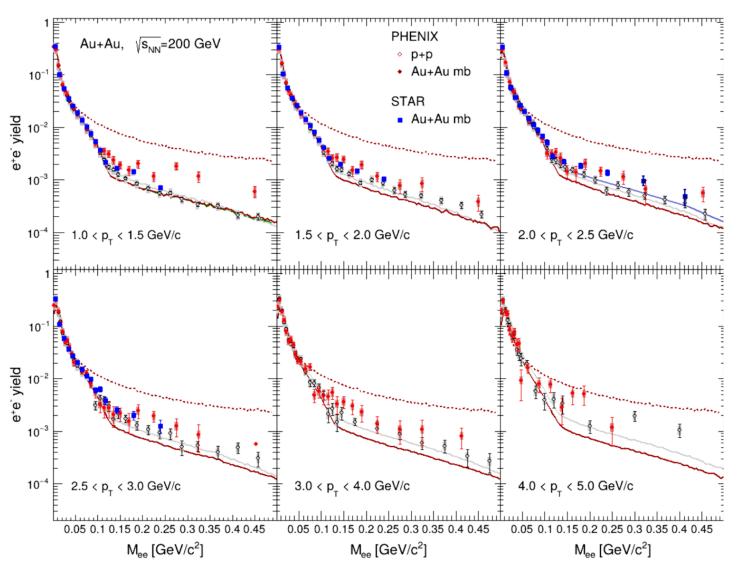
Cocktails do not agree!



# **Evolution with p**<sub>T</sub>

# How does it evolve with $p_T$ ?

Credit: Axel Drees, SBU



Data seem to be consistent for  $p_T > 2 \text{ GeV/c}$ 

Difference at lower p<sub>T</sub>

Converges



# What next?



## Collaborate with STAR

Exchange AuAu virtual gamma data (not all in public domain or HEP data)

Provide standalone software that reproduces experimental data

Independently analyse data from both experiments with same cocktail

Compare & discuss

## On the PHENIX side

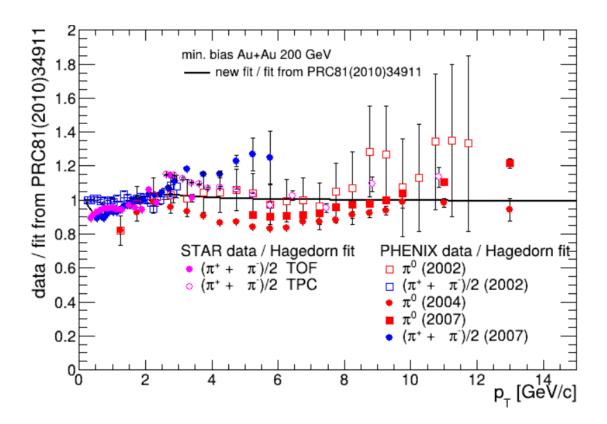
- Collect data in sharable form
- Develop standalone PHENIX simulator
- Create new cocktail based on today's knowledge

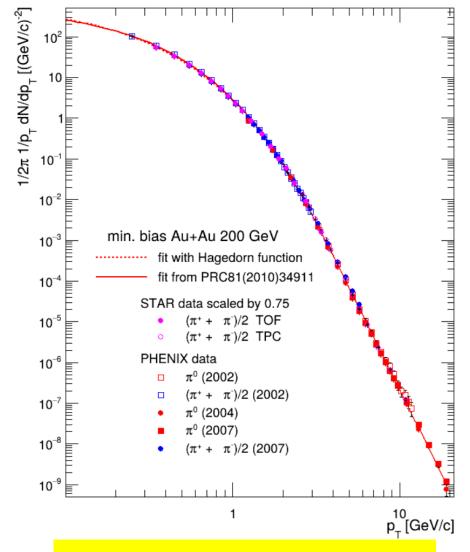
**Ongoing work with STAR** 





# **Compare to charged pions, STAR**





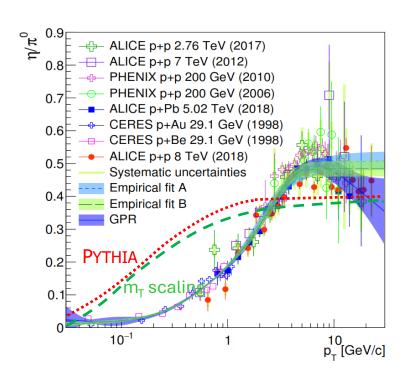
# Reasonable agreement





# The second highest contributor: η

# **Recent study of world data**



 $\eta/\pi^0$  PRC 104, 054902 (2021)

# $\eta/\pi^0$ vs collision energy

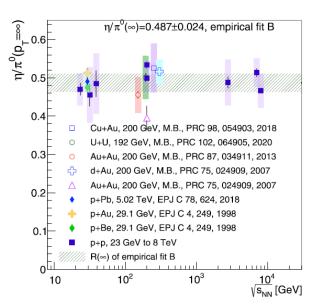


FIG. 4. Values of  $R^{\infty} = \eta/\pi^0(p_T \to \infty)$  as a function of  $\sqrt{s_{NN}}$  for the minimum bias p+p, p+A, and A+B data sets. Statistical errors are shown as bars, and systematic uncertainties are shown as bands. Also shown is a band representing  $0.487 \pm 0.024$ , the result of the empirical fit B to the combined p+p and p+A data. Note that the A+B data at 200 GeV are offset in  $\sqrt{s_{NN}}$  to avoid overlap of data sets.

# $\eta/\pi^0$ vs N <sub>ch</sub> (centrality)

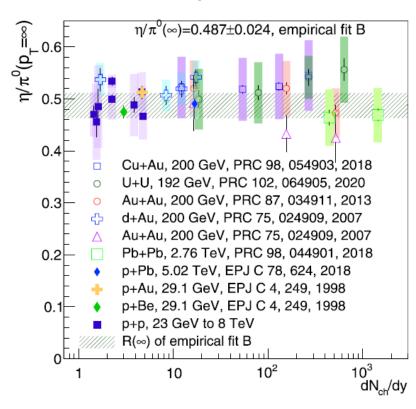


FIG. 5. Values of  $R^{\infty} = \eta/\pi^0(p_T \to \infty)$  as a function of  $dN_{\rm ch}/d\eta$ . The presentation is identical to Fig. 4; however, for A+B collisions results from different centrality classes are shown rather than results for the minimum bias sample.

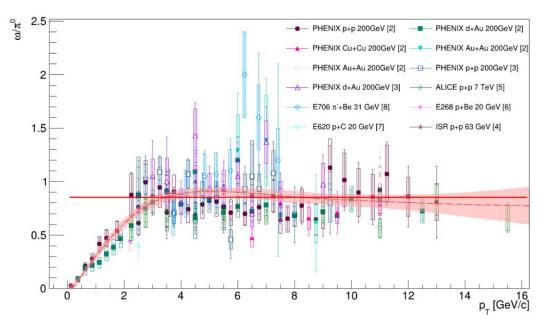
# Is it truly constant at high $p_T$ ?

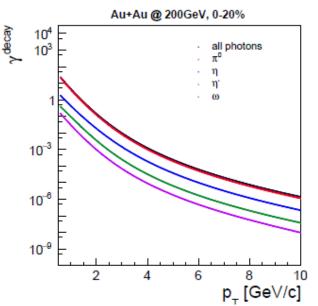


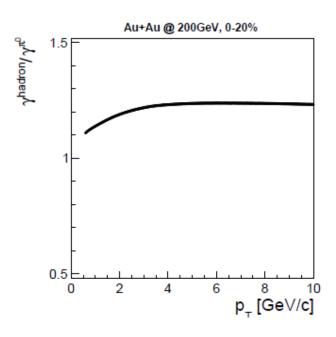


# Other contributors – $\omega$ , $\eta$ ' ...

# Small contribution, but we need a precision measurement







Axel Drees, Konstantin Bauer, SBU 2023 (unpublished)

Wenqing Fan, SBU, PhD thesis

Few measurements available Changes with collision energy





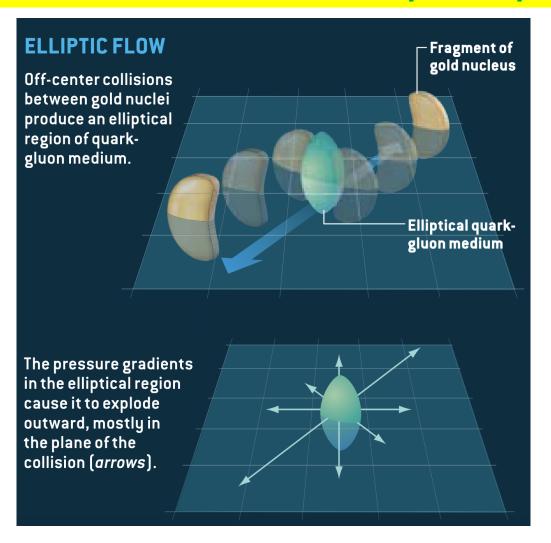
# "Direct photon puzzle"



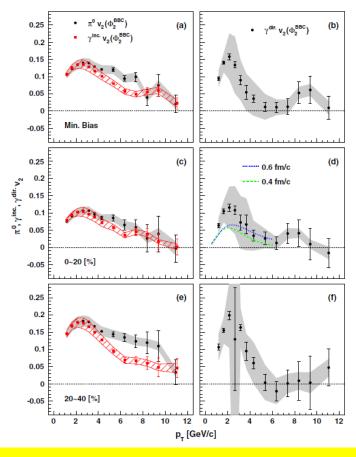
# Azimuthal anisotropy in photon emission



# I hate the word "flow"... it implies a particular dynamics



PRL 109, 122302 (2012)



At low  $p_T \pi^0$  and hadron  $v_2$  similar!



# The direct photon puzzle – Au+Au, 200 GeV



# What is it?

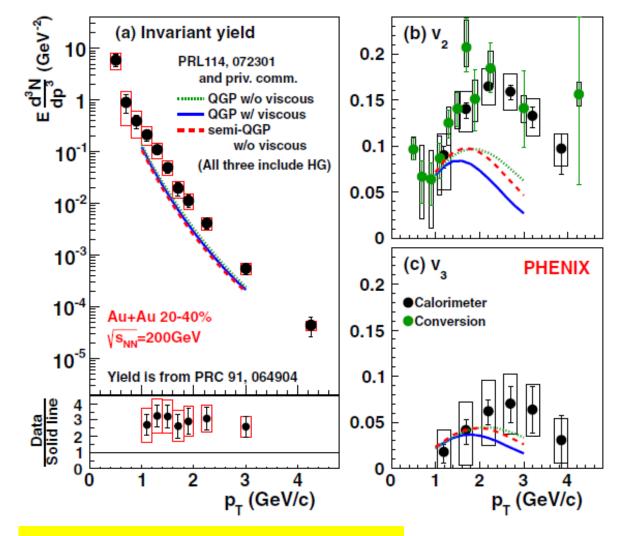
Long-time paradigm: mostly thermal radiation from QGP and HG

Large yield: early emission, high T

Large azimuthal anisotropies: late emission, low T

Models could not explain large yield and v<sub>2</sub> simultaneously

PHENIX PRC 94, 064901 (2016)



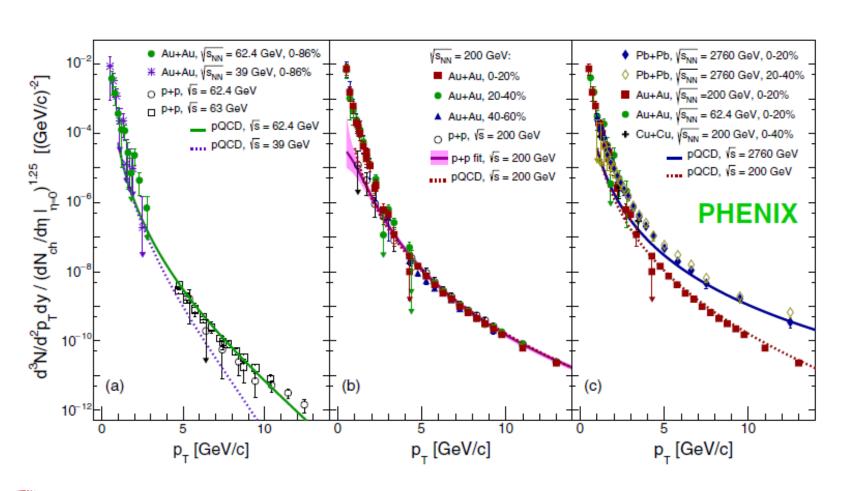


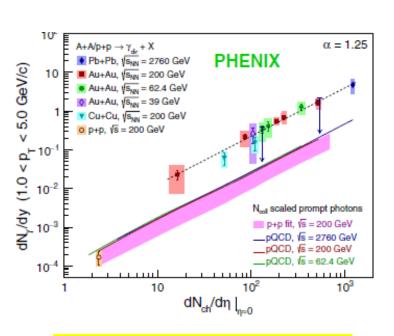
# Scaling: part of the puzzle?



# Universality over system size, collision energy?

$$\frac{dN_{\gamma}}{dy} = \int_{p_{T,\text{min}}}^{p_{T,\text{max}}} \frac{dN_{\gamma}^{\text{dir}}}{dp_{T}dy} dp_{T} = A \times \left(\frac{dN_{\text{ch}}}{d\eta}\right)^{\alpha}$$





Surprisingly small power ( $\alpha$ =1.25)

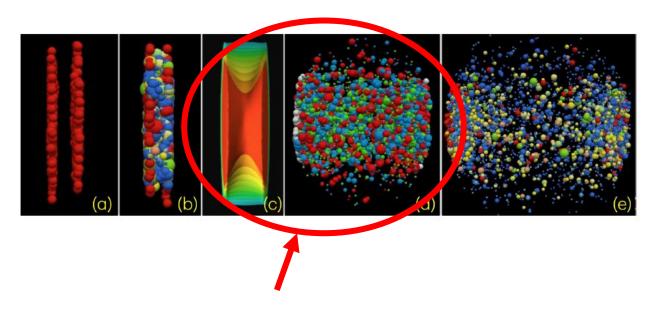


PRL 123, 022301 (2019)

# Is there a "puzzle" at all?



Thermal/hydro paradigm: yield → "early", flow → "late"



Logical – and tractable! –
assumption
but why would thermal source
be super-dominant?
No a priori reason
"Unconventional" sources?

Concentrating on this region and as a (locally) thermalized hydro system is too narrow-minded?



Is the puzzle just due to some tunnel-vision?

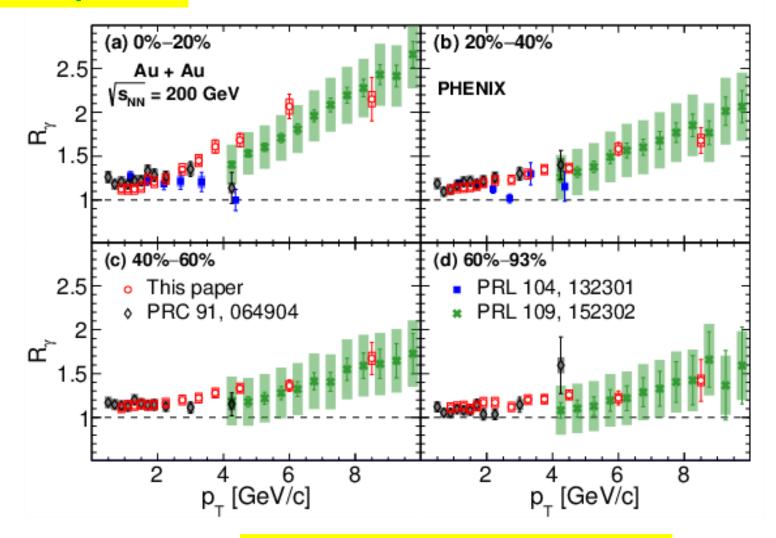
# Consistency



# Four different analysis techniques...

$$R_{\gamma} = \frac{\gamma^{inclusive}}{\gamma^{decay}}$$

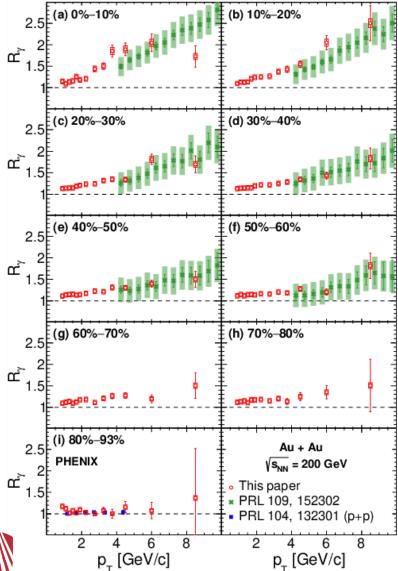
Clear excess in the "thermal" region, depending on centrality



PRC 109, 044912 (2024)



# **External conversion in VTX**

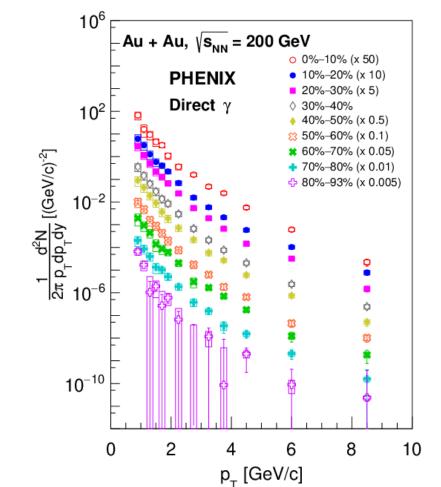


# A closer look at R $\gamma$ and direct $\gamma$



Rγ – and its uncertainty – is crucial in the analysis of azimuthal asymmetries

PRC 109, 044912 (2024)



Limited range→ high precision(Conversion)

Consistent, solid overlap with calorimeter data

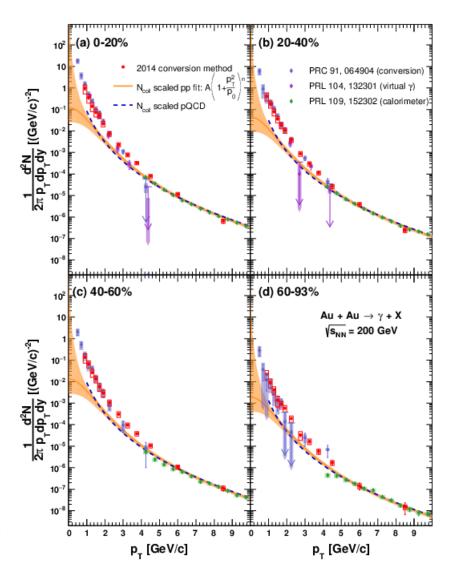


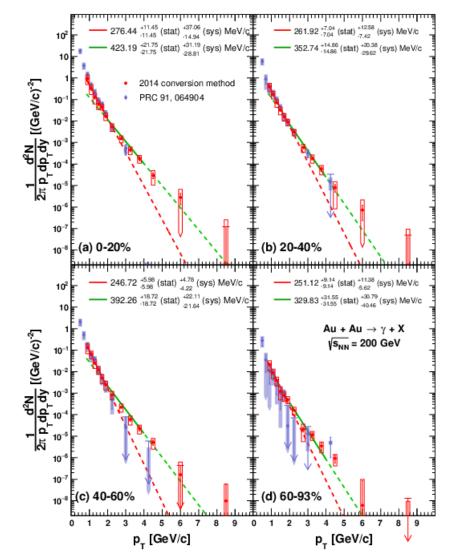
# **Direct and non-prompt photons**



# Non-prompt: subtract p+p (extrapolated)

PRC 109, 044912 (2024)





Non-prompt: not a single exponential ("Thermal" is frowned upon ©)

T<sub>eff</sub> increases with p<sub>T</sub>



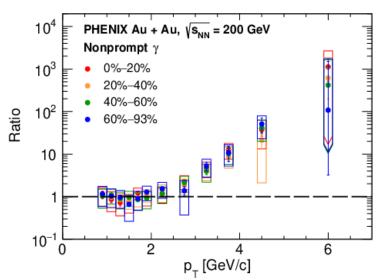
69

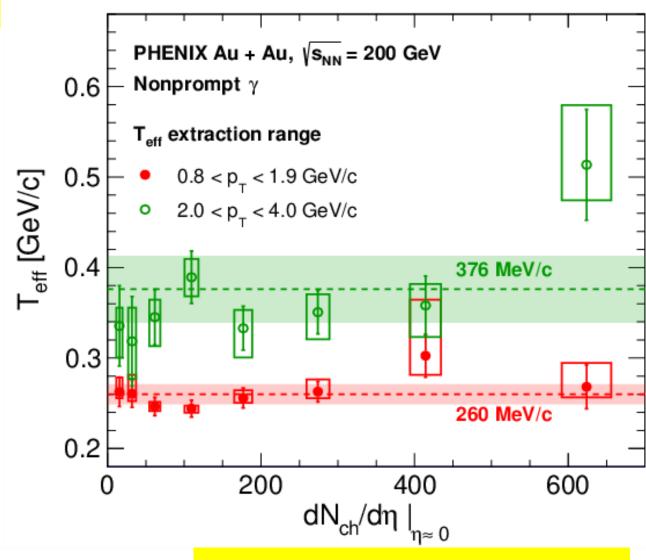
# PRC 109, 044912 (2024)

# **T**<sub>eff</sub> vs the p<sub>T</sub> range and centrality

# No dependence on centrality

Strong dependence on p<sub>T</sub> Composition and relative weight of different sources always the same?











# Easing a tension on $\alpha$ being too small?



# Issue: photon yield vs N<sub>ch</sub>

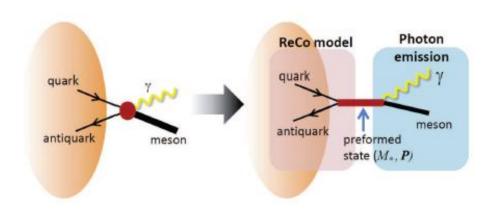
 $dN\gamma/dy \sim (dN_{ch}/d\eta)^{\alpha}$ 

 $\alpha$  from QGP ~ 1.85

 $\alpha$  from HG ~ 1.25

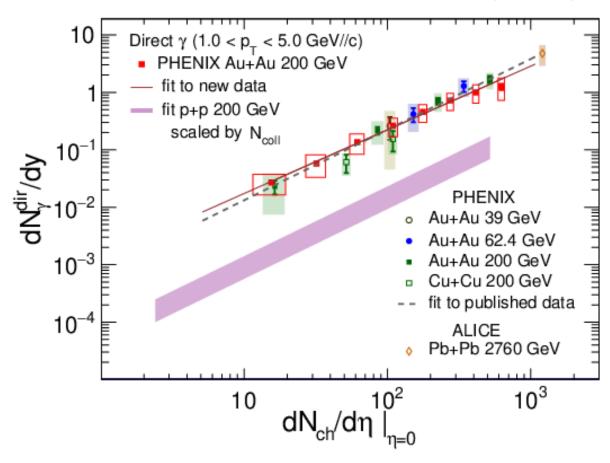
 $\alpha$  measured ~ 1.11

But radiative hadronization:  $N\gamma \sim N_{ch}$ 



PRC 106, 034906 (2022)

# PRC109, 044912 (2024)



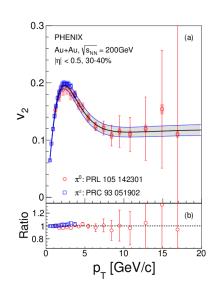
# May help to explain small $\alpha$ observed



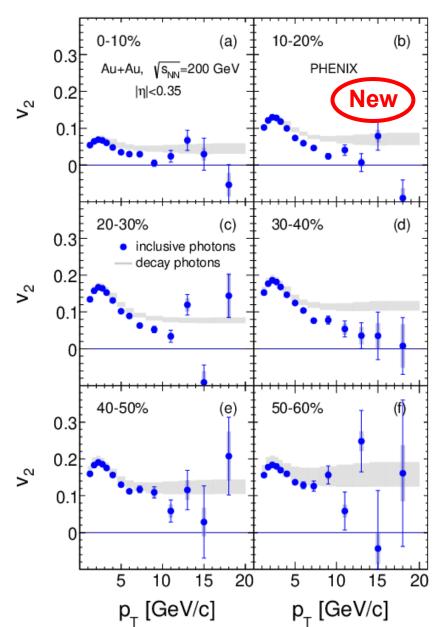
# Direct photon $v_2$ – starts with inclusive and decay $v_2$



Inclusive γ
Decay γ
(cocktail)



Pion v<sub>2</sub> and fit (used for decay v<sub>2</sub>)



arXiv:2504.02955

At low p<sub>T</sub> inclusive (blue) and decay (gray band) v<sub>2</sub> almost identical

At high  $p_T$  strong separation (prompt photon  $v_2 \sim 0$ )

$$v_2^{dir} = \frac{R_{\gamma}v_2^{inc} - v_2^{dec}}{R_{\gamma} - 1}$$

v<sub>2</sub> hyper-sensitive to Rγ

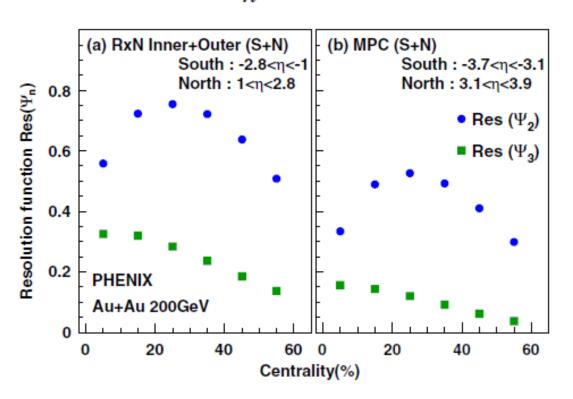


### Reaction plane resolution, uncertainties on v<sub>2</sub>

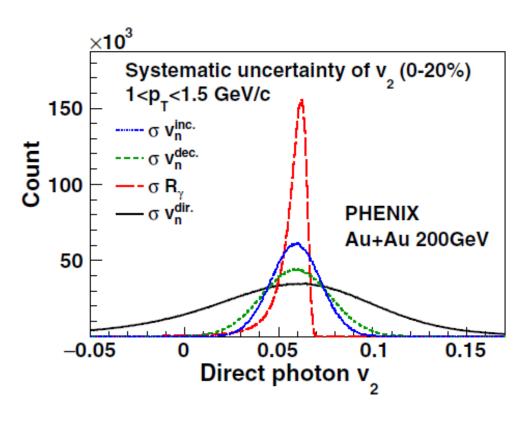


#### **Main sources of uncertainties**

$$v_n = v'_n / \text{Res}(\Psi_n)$$
.



PRC 94, 064901 (2016)



$$v_2^{dir} = \frac{R_\gamma v_2^{inc} - v_2^{dec}}{R_\gamma - 1}$$

v<sub>2</sub> hyper-sensitive to Rγ

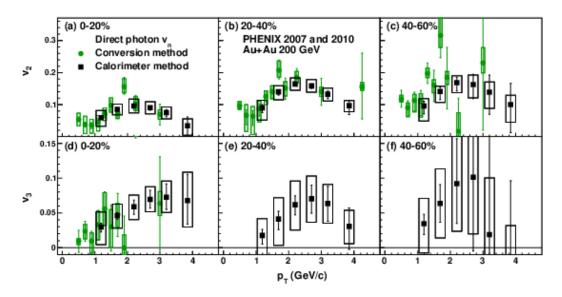


## Direct photon $v_2$ – the new result

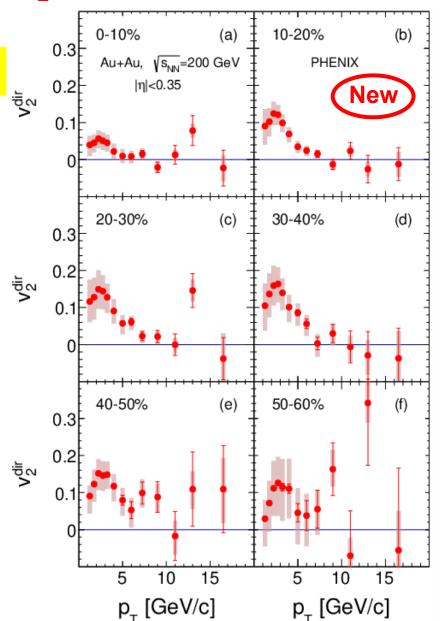


#### **Issue:** magnitude, p<sub>T</sub> dependence

Previous: PRC 94, 064901 (2016)



Conversion, calorimeter Limited p<sub>T</sub> range Significant v<sub>2</sub>



arXiv:2504.02955

10% centrality

Wide p<sub>⊤</sub> range

Calorimeter

Large at low p<sub>T</sub> vanishes at high p<sub>T</sub>

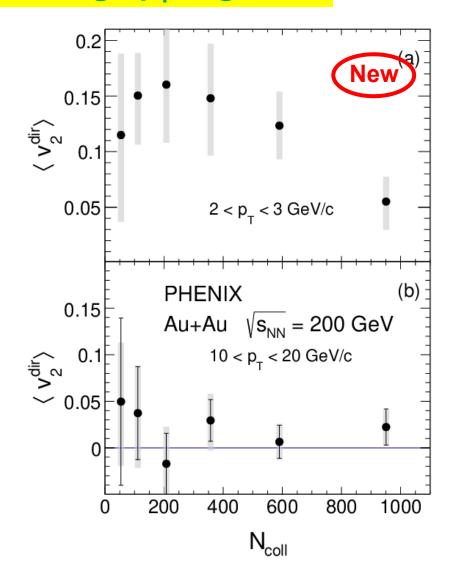




#### Average $v_2$ vs $p_T$ and centrality

arXiv:2504.02955

#### Low and high p<sub>T</sub> regions



Low p<sub>T</sub> (non-prompt)

→ large v<sub>2</sub>, centrality dependence

High p<sub>T</sub> (prompt)

 $\rightarrow$   $v_2 \sim 0$ , no centrality dependence

Hard scattering –  $v_2 \sim 0$ 

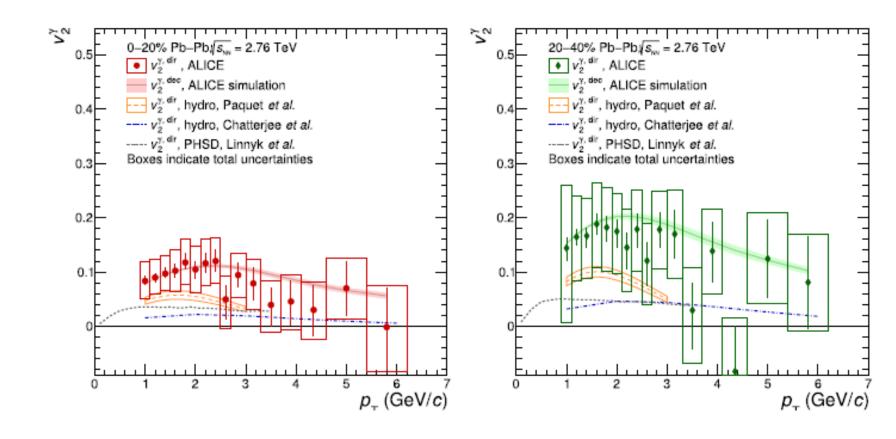




#### Direct photon v<sub>2</sub> in 2.76 PbPb (ALICE)

PLB 789 (2019) 308-322

## Low and high p<sub>T</sub> regions





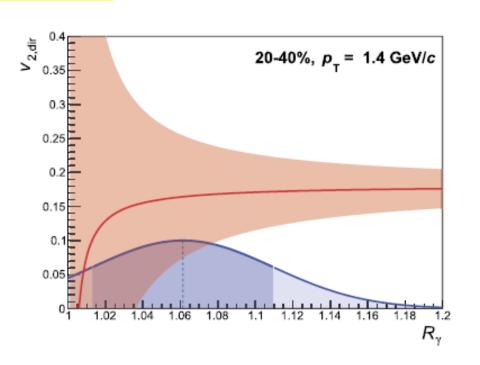
### Direct photon v<sub>2</sub> uncertainties in 2.76 PbPb (ALICE)

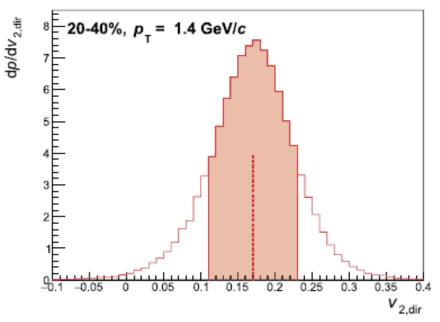


## Precision on R, crucial

PLB 789 (2019) 308-322

$$v_2^{dir} = \frac{R_\gamma v_2^{inc} - v_2^{dec}}{R_\gamma - 1}$$





**Fig. 4.** Left: Central value (solid red line) and uncertainty of the direct-photon  $v_2$  for a selected  $p_T$  interval, The upper and lower edges of the red shaded area correspond to the total uncertainty of  $v_2^{\gamma, \text{dir}}$  as obtained from linear Gaussian propagation of the uncertainties  $\sigma(v_2^{\gamma, \text{inc}})$  and  $\sigma(v_2^{\gamma, \text{dec}})$ . The Gaussian (with arbitrary normalization) reflects the measured value of  $R_{\gamma}$  in this  $p_T$  interval (blue dashed line) and its  $\pm 1\sigma$  uncertainty (dark-blue shaded interval). Right: Posterior distribution of the true value of  $v_2^{\gamma, \text{dir}}$  for the same interval in the Bayesian approach, Note that the distribution has a non-Gaussian shape, implying that the  $\pm 2\sigma$  interval typically corresponds to a probability of less than 95.45% as would be the case for a Gaussian.

Asymmetric, and v<sub>2</sub> can even go negative

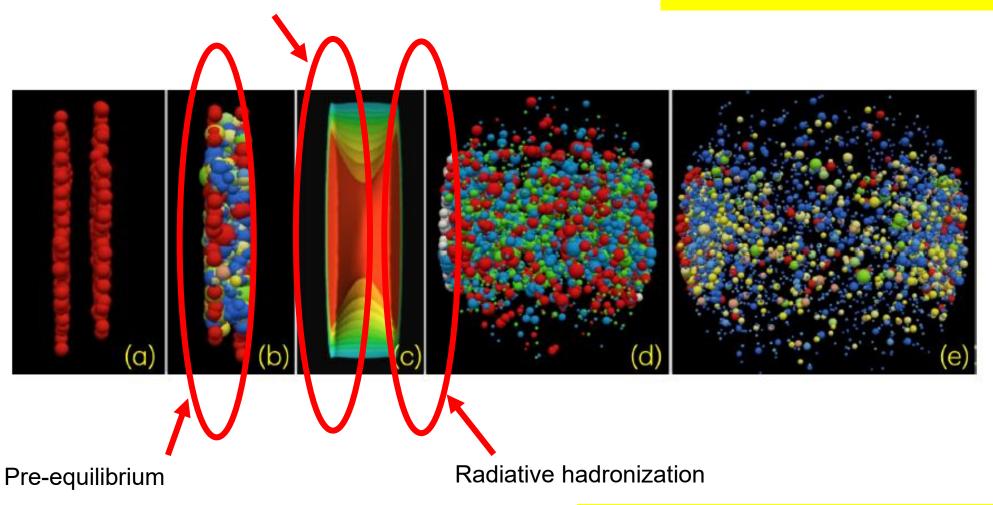


#### New photon sources considered



Weak magnetic γ from QGP

Add "unconventional" sources





**Stepping past the thermal paradigm** 

#### Comparisons of photon v<sub>2</sub> to recent models

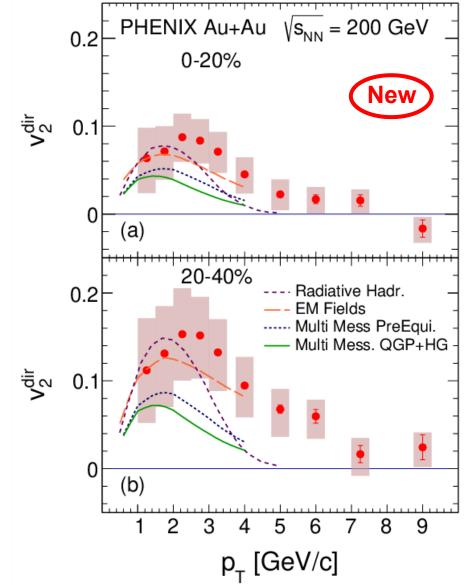


Data from arXiv:2504.02955

### **Different approaches**

Multi-messenger PRC 105, 014909 (2022) Radiative hadronization PRC 106, 034906 (2022) Magnetic emission Nucl. Phys. Rev 41, 1 (2024)

source	multim.	rad. hadr.	magnetic
prompt	X	X	X
magnetic			X
pre-eq	X		X
QGP th	X	X	X
HG th	X	X	X
rad. hadr.		X	





Could (should???) all those sources be combined?



# A few recent models to resolve the puzzle



### Strongly coupled plasma with constant magnetic field



### An early attempt

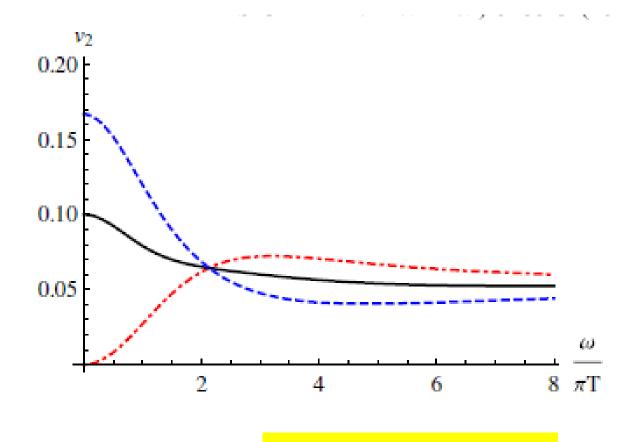
Mueller PRD 89, 026013 (2014)

Generated solely by magnetic field in the strong coupling scenario

Photon v<sub>2</sub> does not vanish at low p<sub>T</sub>

Two polarizations: in-plane (blue) and out-of-plane (red)

Only a fraction of all v<sub>2</sub>







#### Magnetic field – synchrotron radiation



#### Non-prompt photon yields

Tuchin PRC 91, 014902 (2015)

#### PHENIX AuAu data

#### **Authors conclusions:**

- a significant fraction of photon excess in the region  $k\perp$  = 1–3 GeV can be attributed to the synchrotron radiation
- this source alone would predict  $v_2 = 4/7$  and  $v_4 = 1/10$  independent of photon momentum and centrality
- odd terms (v<sub>3</sub>, ...) should vanish
- can contribute a significant fraction of non-prompt photons at 2-3 GeV/c
- temperatures should be below T = 400 MeV
   (since then synchrotron photons would account for all observed photons)

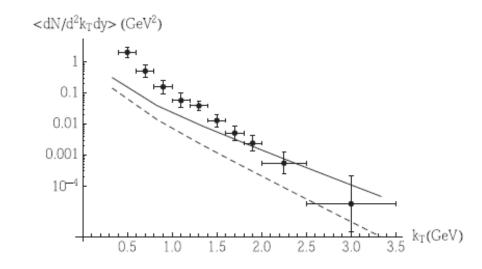


FIG. 2. Spectrum of synchrotron photons averaged over the azimuthal angle versus photon transverse momentum  $k_{\perp}$  at rapidity y = 0 and centrality 40%–60% (b = 10.2 fm [27]). Solid line:  $T = 400 \,\text{MeV}$ ; dashed line:  $T = 200 \,\text{MeV}$ . Data are from [1]; they represent the direct photon  $k_T$  spectra after subtraction of the Ncoll scaled p + p contribution (Fig. 8 there).

### Early attempt, very incomplete

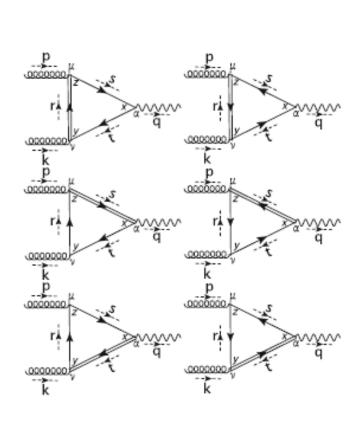


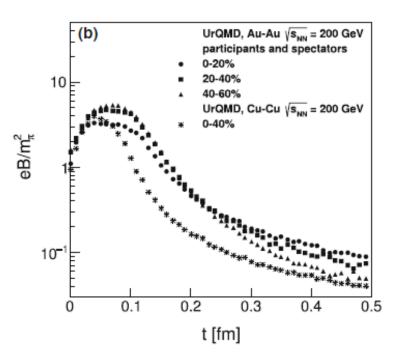
#### Magnetic field induced gluon fusion and splitting



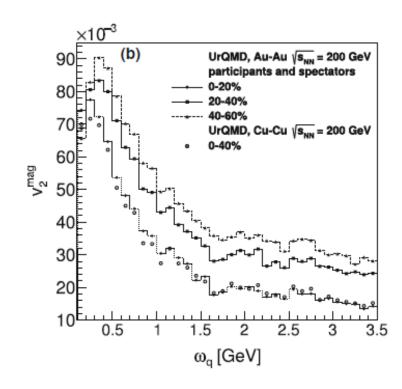
### **Pre-hydro source**

Ayala+ Eur. Phys. J. A (2020) 56:53





Gluon fusion / splitting
Issue: short lifetime of B
Thermal effects (hydro) start
after magnetic pulse
Similar magnitude as Mueller
Non-zero flow at low p<sub>T</sub>!



"Magnetic flow" only – needs to be weighted

Short lifetime,  $v_2(p_T=0)>0$ 

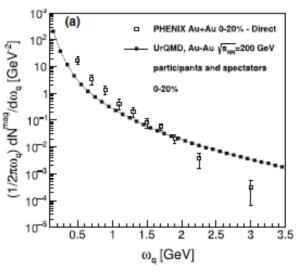


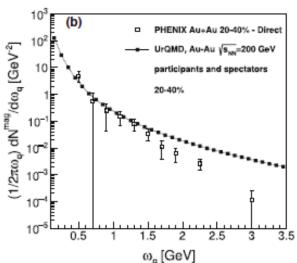
#### Participants, spectators



#### Ayala+ Eur. Phys. J. A (2020) 56:53

### Yields and v<sub>2</sub> have to be explained simultaneously

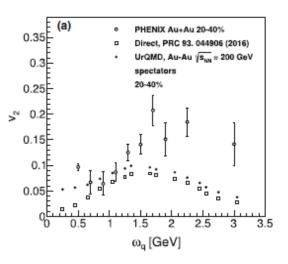


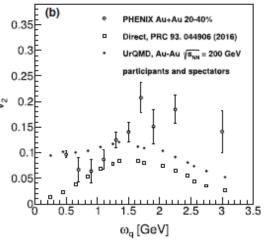


Moving in the right direction but significant shape difference

Overpredict yield, underpredict v<sub>2</sub> above 1.5 GeV/c

Enhancement at  $p_T = 0$  (what is hard to measure)





Challenge to experimentalists: get  $v_2(p_T=0)$ 

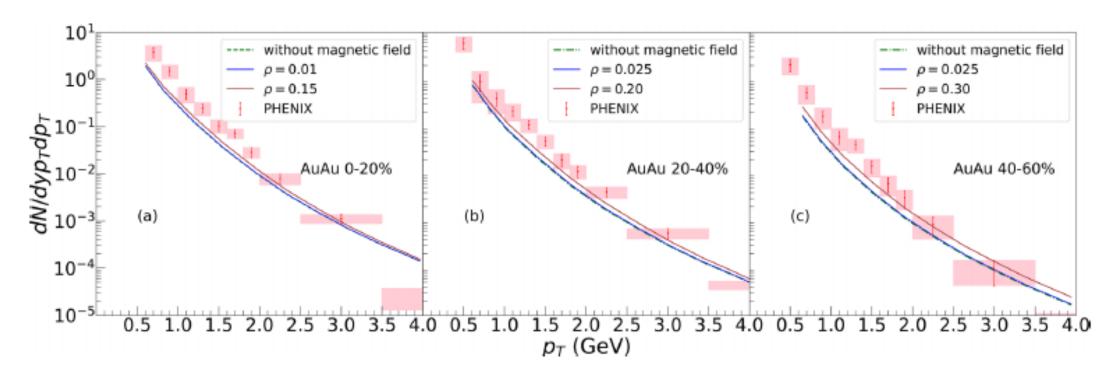




### "Weak" field in QGP, hydro, RHIC

#### Induced radiation, e-b-e viscous hydro

Sun, Yan PRC 109, 034917 (2024)



Effects of the weak magnetic field during QGP evolution Interplay of magnetic field and longitudinal dynamics of medium

$$\rho \equiv \frac{\sigma_{\rm el}}{T} \frac{\overline{eB_{\rm y}}}{m_{\pi}^2},$$

PHENIX data, PRC 91, 064904 (2015)

**Small correction to yield** 

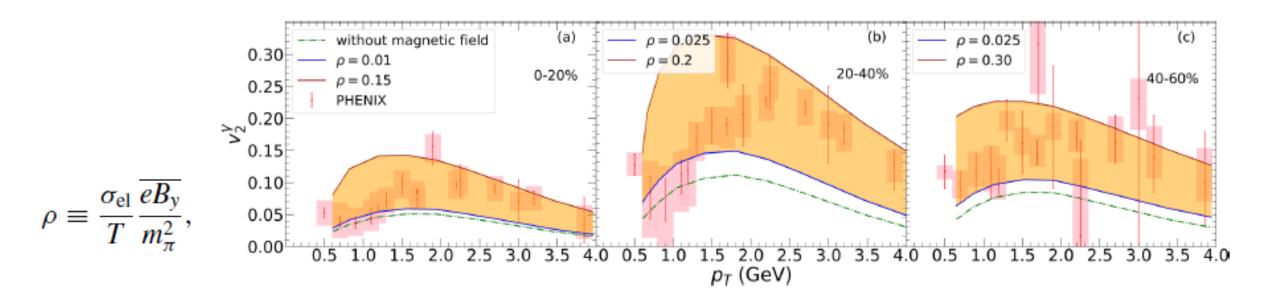




### "Weak" field in QGP, hydro, RHIC

#### Induced radiation, e-b-e viscous hydro

Sun, Yan PRC 109, 034917 (2024)



v<sub>3</sub> emerges (novel) Yield increase marginal flow increase large

Field  $0.1 \text{m}^2_{\pi} \sim 10^{16} \text{ G}$ 

PHENIX data PRC 94, 064901 (2016)

Get  $\rho$  (field strength) from flow

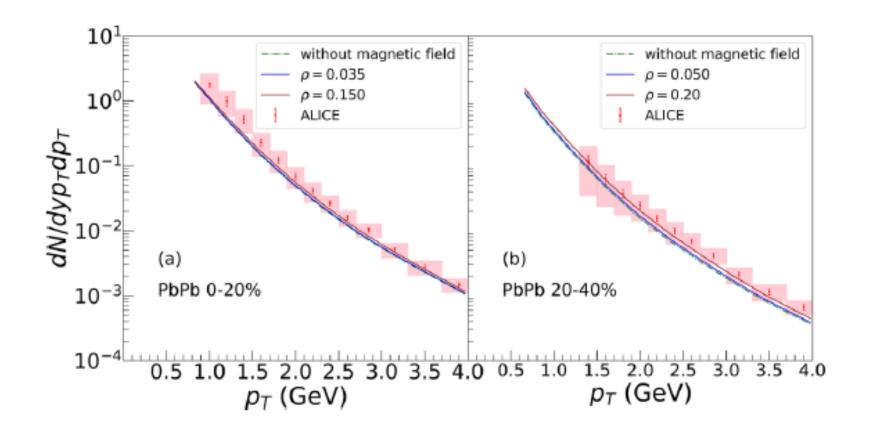




### "Weak" field in QGP, hydro, LHC

#### Induced radiation, e-b-e viscous hydro

Sun, Yan PRC 109, 034917 (2024)



ALICE, PbPb, 2.76 TeV (data PLB 789, 308 (2019))

**Small correction to yield** 

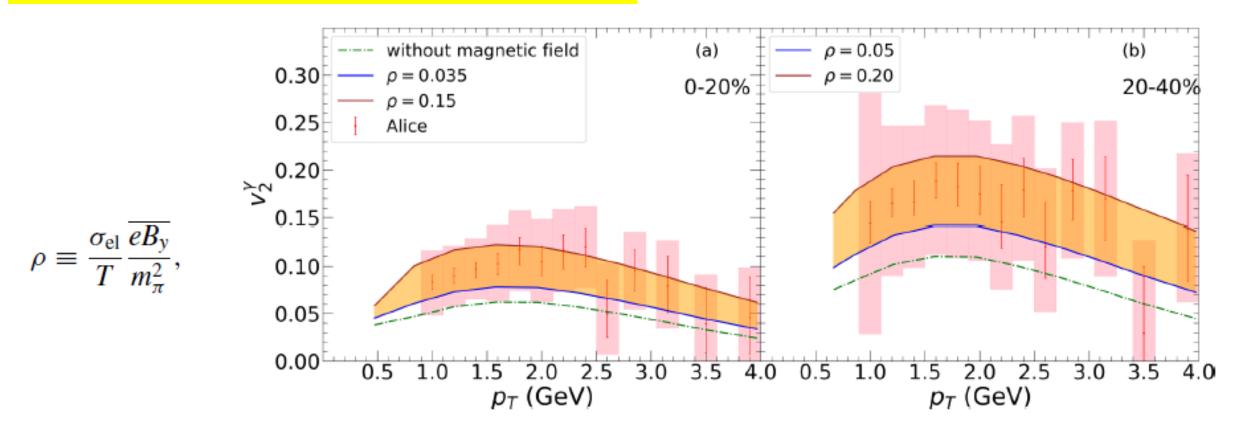




#### "Weak" field in QGP, hydro, LHC

#### Induced radiation, e-b-e viscous hydro

Sun, Yan PRC 109, 034917 (2024)



ALICE, PbPb, 2.76 TeV (data PLB 789, 308 (2019))

 $\rho$  (field strength) somewhat higher than at RHIC



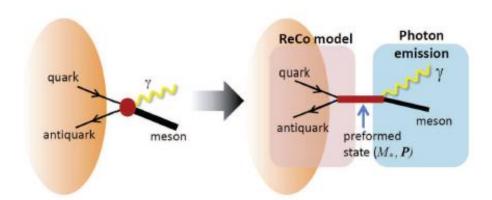
#### Radiative hadronization – RHIC



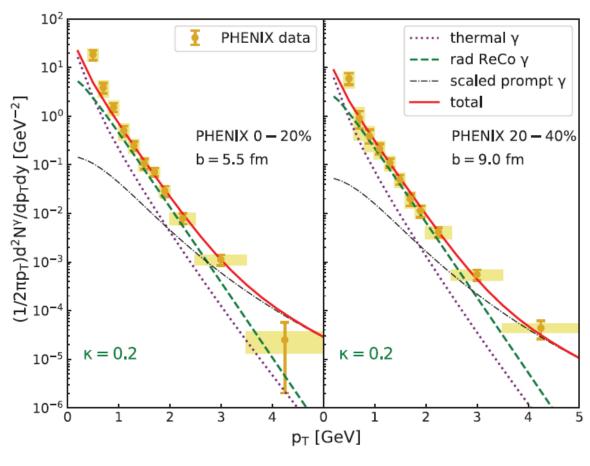
### Circumvents the "old paradigm"

Old paradigm: yield early, high T, flow late Here: new source at the time of hadronization Recombination off-shell, then come on-shell by radiating a photon Will inherit v<sub>2</sub> at the time of phase transition

κ chosen so that thermal, prompt and radiative photons describe the data



Fujii et al, PRC 106, 034906 (2022)



PHENIX, PRC 91, 064904 (2015)

κ is adjusted to describe data



#### Radiative hadronization – RHIC



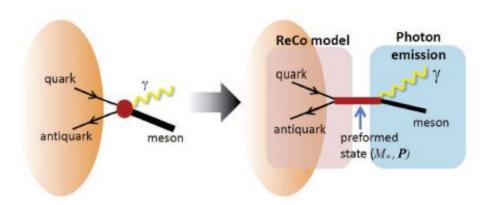
## What about v<sub>2</sub>?

Old paradigm: yield early, high T, flow late

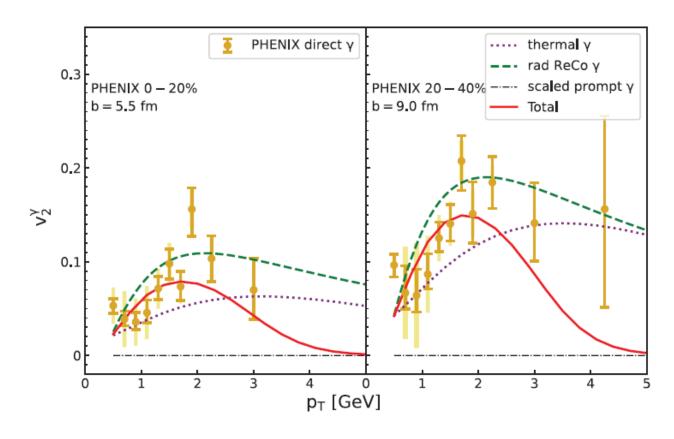
Here: new source at the time of hadronization

Recombination off-shell, then come on-shell by radiating a photon

Will inherit v<sub>2</sub> at the time of phase transition



Fujii et al, PRC 106, 034906 (2022)



PHENIX, PRC 94, 064901 (2016)

Starts to fail above 2 GeV/c

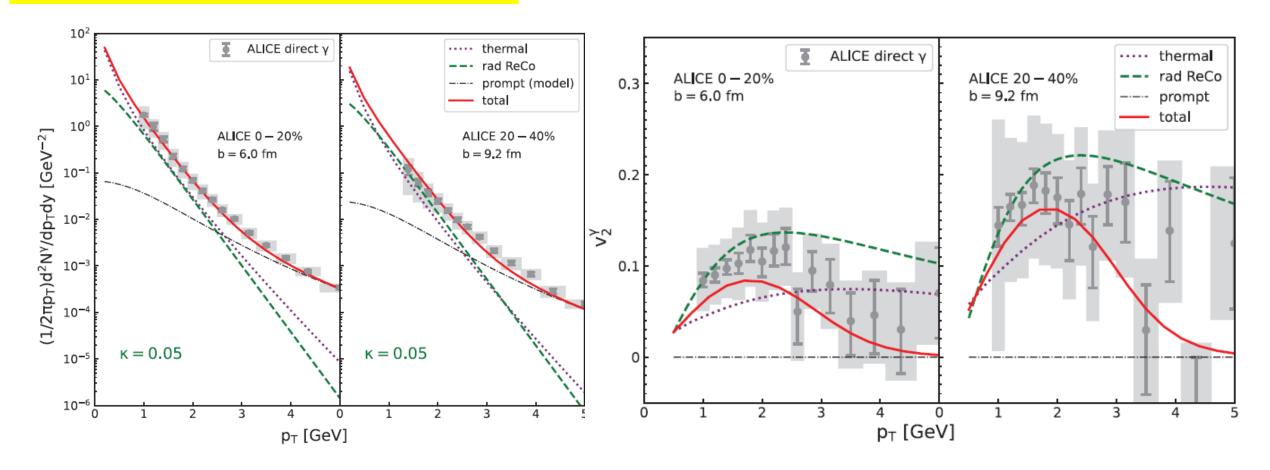


#### Radiative hadronization – LHC



#### Does it work at LHC? (should!)

Fujii et al, PRC 106, 034906 (2022)



Remember: the transition itself at RHIC and LHC appeared to be similar (low p<sub>T</sub> slopes, scaling)

κ very different at RHIC and LHC

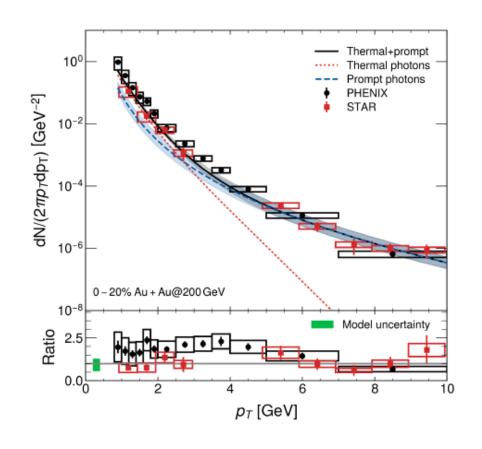


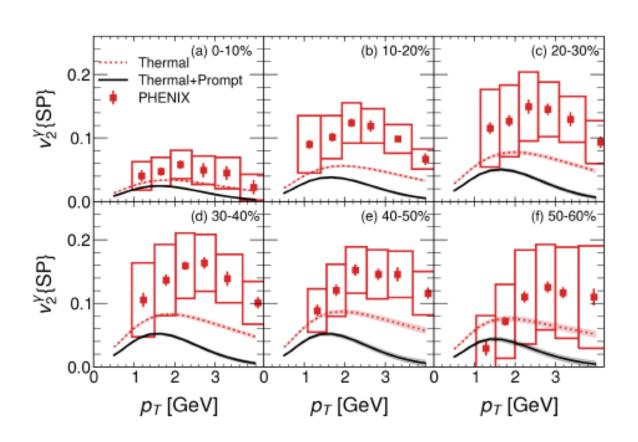
### Multistage (multi-messenger) model



Pre-equilibrium, MUSIC, UrQMD afterburner

Gale et al. arXiv 2511.08773





Add rad. hadr., magnetic?





## Where does this leave us?







Photon at high p<sub>T</sub> well understood

At low p<sub>T</sub> the direct photon puzzle still not resolved

Interesting and plausible new sources suggested

Experimental uncertainties still too large

We never had an honest-to-God direct photon experiment, built without compromises to other physics (and we are paying the price)

We should have one!



Further argument in GD Rept. Prog. Phys. 83 (2020) 4, 046301







