# Nambu-Jona—Lasinio model in the presence of intense magnetic fields

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# First Latin American Workshop on Electromagnetics Effects in QCD



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## Outline for the lectures

- (i) NJL model at finite T and B basics
- (ii) Issues related to regularizing thermo and magnetic contributions within nonrenormalizable theories and **applications**
- (iii) hot quark matter and hot bosonic matter with a strong electric field

# Outline

- Motivation
- Schrödinger Equation and Dirac Equation
- Particle in the Presence of an Electromagnetic Field
- NJL model in MFA
- NJL model at finite eB
- Magnetic Catalysis
- Thermodynamical Quantities
- Magnetic Field Independent Regularization MFIR

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### **SIMEE** 10 **Strongly Interacting Matter under Extreme Environments**



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### Quarks and gluons in extreme conditions

- heavy ion collisions  $T \lesssim 10^{12} \, {}^\circ C = 200$  MeV,  $n \lesssim 0.12$  fm<sup>-3</sup>  $B \lesssim 10^{19}$  G = 0.3 GeV<sup>2</sup>/e
- neutron stars  $T \lesssim 1$  keV,  $n \lesssim 2$  fm<sup>-3</sup> magnetars  $B \lesssim 10^{15}$  G
- ▶ neutron star mergers  $T \lesssim 50$  MeV
- eary universe, QCD epoch T  $\leq 200$  MeV standard scenario:  $n \approx 0$  also allowed:  $n_Q = 0$ ,  $n_\ell/s \leq 0.01$



## Strengths of magnetic fields

• Strong magnetic fields are also present in magnetars: C. Kouveliotou et al., Nature 393, 235 (1998). magnetars: at surface  $B \lesssim 10^{15} \,\mathrm{G}$ Duncan, Thompson, Astrophys.J. 392, L9 (1992) larger in the interior,  $B \sim 10^{18-20} \,\mathrm{G?}$ Lai, Shapiro, Astrophys.J. 383, 745 (1991) E. J. Ferrer et al., PRC 82, 065802 (2010)



A. K. Harding, D. Lai, Rept. Prog. Phys. 69, 2631 (2006)

• and might have played an important role in the physics of the early universe. T. Vaschapati, Phys. Lett. B 265, 258 (1991).

D. Grasso and H.R. Rubinstein, Phys. Rep. 348, 163 (2001).

# B Effects on QCD phase transitions?

 $\Lambda_{\rm QCD}^2 \sim (200 \,{\rm MeV})^2 \sim 2 \times 10^{18} \,{\rm G}$ 



IMC: Bali, Bruckmann, Endrodi, Fodor, Katz et al. JHEP 02 (2012) 044 Phys.Rev.D 86 (2012) 071502

### **Phase diagram**

▶ control parameters: T,  $n \leftrightarrow \mu$ , B

$$\mu_{\{u,d,s\}} \ / \ \mu_{\{B,Q,S\}} \ / \ \mu_{\{B,I,S\}}$$

well-known famous phase diagram

▶ well-known, less famous phase diagram: T - B



\* G. Endrodi slide - SQM 2024

# B Effects on QCD phase transitions?



M. D'Elia , L. Maio, F. Sanfilippo, A. Stanzione, Phys. Rev. D 105 , 034511 (2022).

### Strength of the magnetic fields

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	B [Gauss]	eB [MeV <sup>2</sup> ]
Earth surface	0.5	(0.05x10 <sup>-6</sup> MeV) <sup>2</sup>
Magnetic Ressonance	1.5x10 <sup>4</sup>	(8.6x10 <sup>-6</sup> MeV) <sup>2</sup>
magnet - CERN	8.4x10 <sup>4</sup>	(20.5x10 <sup>-6</sup> MeV) <sup>2</sup>
frog levitation *	10 <sup>5</sup>	(25x10 <sup>-6</sup> MeV) <sup>2</sup>
Critical quantum field of the electron	4.4 x 10 <sup>13</sup>	(0.5 MeV) <sup>2</sup> = <b>m</b> <sub>e</sub> <sup>2</sup>
Magnetares (field on the surface)	5.0x10 <sup>15</sup>	(5 MeV) <sup>2</sup> = (10 <b>m</b> <sub>e</sub> ) <sup>2</sup>
(Au+Au) Heavy ion collisions	10 <sup>19</sup>	$(400 \text{ MeV})^2 = (3\mathbf{m}_{\pi})^2$

(1 Tesla = 10<sup>4</sup> Gauss)

\* Andre Geim - Ig Nobel-2000 and Nobel-2010 (graphene)

Let's consider relativistic particles and, therefore, we will start by discussing the appropriate equation of motion for this case, namely the Dirac equation. In the non-relativistic case, we heuristically obtain the Schrödinger equation from the energy

$$E = \frac{\vec{p}^2}{2m} = \frac{p_x^2 + p_y^2 + p_z^2}{2m}$$

using the prescription:

$$E 
ightarrow i rac{\partial}{\partial t} \ , \ p_x 
ightarrow rac{\hbar}{i} rac{\partial}{\partial x} \ , \ p_y 
ightarrow rac{\hbar}{i} rac{\partial}{\partial y}, \ \ p_z 
ightarrow rac{\hbar}{i} rac{\partial}{\partial z}$$

we obtain

$$i\hbar \frac{\partial}{\partial t}\psi(\vec{r},t) = -\frac{\hbar^2}{2m}\nabla^2\psi(\vec{r},t)$$

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In relativistic case the energy is given by:

$$\mathsf{E} = \sqrt{\vec{p}^2 c^2 + m^2 c^4}$$

using the prescription:

$$E 
ightarrow i rac{\partial}{\partial t} \ , \ p 
ightarrow rac{\hbar}{i} ec
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we obtain

$$i\hbar \frac{\partial}{\partial t}\psi(\vec{r},t) = \sqrt{-c^2\hbar^2\vec{
abla}^2 + m^2c^4} \ \psi(\vec{r},t)$$

Extremely complex equation (nature is simpler!)

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Dirac's idea was to "take the square root" of:

$$E = \sqrt{\vec{p}^2 c^2 + m^2 c^4} , \quad E \to i \frac{\partial}{\partial t} , \quad p \to \frac{\hbar}{i} \vec{\nabla}$$
$$E = i\hbar \frac{\partial}{\partial t} \psi(\vec{r}, t) = H_D \psi(\vec{r}, t) = \left( c\vec{\alpha} \cdot \vec{p} + \beta m c^2 \right) \psi(\vec{r}, t)$$

Requiring that in the operatorial form  $E^2 = H_D^2 = \rho^2 c^2 + m^2 c^4$ ( $\rightarrow$  relativistic dispersion relation) we can determine  $\vec{\alpha}$  and  $\beta$ .

$$\left(c\vec{\alpha}\cdot\vec{p}+\beta mc^{2}\right) \left(c\vec{\alpha}\cdot\vec{p}+\beta mc^{2}\right)\psi(\vec{r},t)=(c^{2}\vec{p}^{2}+m^{2}c^{4})\psi(\vec{r},t)$$

For the last equation to have a solution  $\vec{\alpha}$  and  $\beta$  must be matrices.

The minimum dimension of the matrices  $\alpha_i$ ,  $i = x, y, z \in \beta$  that satisfy the desired conditions is 4. A standard representation is the following:

$$\alpha_i = \begin{pmatrix} \mathbf{0}_{2\times 2} & \sigma_i \\ \sigma_i & \mathbf{0}_{2\times 2} \end{pmatrix} , \ i = \mathbf{x}, \mathbf{y}, \mathbf{z}, \ , \ \beta = \begin{pmatrix} \mathbf{1}_{2\times 2} & \mathbf{0}_{2\times 2} \\ \mathbf{0}_{2\times 2} & -\mathbf{1}_{2\times 2} \end{pmatrix} ,$$

where  $\sigma_i$  are the Pauli matrices:

$$\sigma_{X} = \left(\begin{array}{cc} 0 & 1 \\ 1 & 0 \end{array}\right) , \ \sigma_{Y} = \left(\begin{array}{cc} 0 & -i \\ i & 0 \end{array}\right) , \ \sigma_{Z} = \left(\begin{array}{cc} 1 & 0 \\ 0 & -1 \end{array}\right) ,$$

as an example:

$$\alpha_{\mathbf{X}} = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & 1 & 0 & 0 \\ 1 & 0 & 0 & 0 \end{pmatrix} , \ \beta = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} , \ \psi(\vec{r},t) = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}$$

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The Dirac equation is given by:

$$i\hbar\frac{\partial}{\partial t}\psi(\vec{r},t) = H_D \,\psi(\vec{r},t) = \left(c\vec{\alpha}\cdot\vec{p} + \beta mc^2\right) \,\psi(\vec{r},t)$$

Let's rewrite the Dirac equation in a more compact form using the  $\gamma$  Dirac matrices:

$$\gamma_0 = \beta$$
 ,  $\gamma_i = \beta \alpha_i$  ,  $i = x, y, z$ 

multiplying the Dirac equation by  $\beta$ 

$$i\hbar\beta \frac{\partial}{\partial t}\psi(\vec{r},t) = \left(c\beta\vec{\alpha}\cdot\vec{p} + (\beta)^2mc^2\right)\psi(\vec{r},t)$$

or even

$$\left(i\hbar\gamma_0\frac{\partial}{\partial ct}+\vec{\gamma}\cdot i\hbar\vec{\nabla}\right)\psi(\vec{r},t)=\textit{mc}\;\psi(\vec{r},t)$$

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Using natural units: h=1 e c=1

$$\begin{pmatrix} \boldsymbol{p} - \boldsymbol{m} \end{pmatrix} \Psi(t, \vec{r}) = 0 , \quad \boldsymbol{p} = \boldsymbol{p}^{\mu} \gamma_{\mu} , \\ i \partial_{t} \Psi(t, \vec{r}) = H_{D} \Psi(t, \vec{r}) = \left( \vec{\alpha} \cdot \hat{\vec{p}} + \beta \boldsymbol{m} \right) \Psi(t, \vec{r}) \\ i \frac{\partial}{\partial t} \psi(\vec{r}, t) = \left[ \begin{pmatrix} 0 & \vec{\sigma} \cdot \vec{p} \\ \vec{\sigma} \cdot \vec{p} & 0 \end{pmatrix} + \begin{pmatrix} \boldsymbol{m} & 0 \\ 0 & -\boldsymbol{m} \end{pmatrix} \right] \psi(\vec{r}, t)$$

Ansatz to find the positive energy solution:

$$\Psi(\vec{r},t) = \Psi(\vec{p})e^{-ip^{\mu}x_{\mu}} = \begin{bmatrix} \chi \\ \phi \end{bmatrix} e^{-i(Et-\vec{p}\cdot\vec{r})}$$

Substituting into the Dirac Equation:

$$E\left[\begin{array}{c} \chi\\ \phi \end{array}\right] = \left(\begin{array}{c} m & \vec{\sigma} \cdot \vec{p}\\ \vec{\sigma} \cdot \vec{p} & -m \end{array}\right) \left[\begin{array}{c} \chi\\ \phi \end{array}\right]$$

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Which results in the following  $2 \times 2$  matrix equations

$$E \chi = \vec{\sigma} \cdot \vec{p} \phi + m \chi \ , \ E \phi = \vec{\sigma} \cdot \vec{p} \chi - m \phi$$

Which, isolating  $\phi$  on the right-hand side, results in :

$$\phi = \frac{\vec{\sigma} \cdot \vec{p}}{\mathbf{E} + \mathbf{m}} \, \chi$$

The positive and negative energy solutions are:

$$\begin{split} \Psi^{(+)}(\vec{r},t) &= N \begin{bmatrix} \chi_s \\ \frac{\vec{\sigma} \cdot \vec{p}}{\vec{E} + m} \chi_s \end{bmatrix} e^{-ip^{\mu} x_{\mu}} &= u_s e^{-ip^{\mu} x_{\mu}} \\ \Psi^{(-)}(\vec{r},t) &= N \begin{bmatrix} \frac{\vec{\sigma} \cdot \vec{p}}{\vec{E} + m} \chi_s \\ \chi_s \end{bmatrix} e^{ip^{\mu} x_{\mu}} &= v_s e^{ip^{\mu} x_{\mu}} \end{split}$$

*N* corresponds to the normalization constant, and  $\chi$  to the Pauli spinor:

$$N = \sqrt{rac{E+m}{2E}}$$
,  $\chi_+ = \begin{bmatrix} 1\\0 \end{bmatrix}$ ,  $\chi_- \begin{bmatrix} 0\\1 \end{bmatrix}$ 

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Considering the case where the particle's momentum is zero:

$$\vec{p} = 0 \rightarrow i\partial_t \Psi(t, \vec{r}) = H_D \Psi(t, \vec{r}) = \left(\vec{\alpha} \cdot \hat{\vec{p}} + \beta m\right) \Psi(t, \vec{r}) = \beta m \Psi(t, \vec{r})$$

And the ansatz for the positive and negative energy solutions:

$$\Psi^{(+)}(\vec{r},t) = u_s e^{-ip^{\mu}x_{\mu}} = u_s e^{-iEt}$$
,  $\Psi^{(-)}(\vec{r},t) = v_s e^{ip^{\mu}x_{\mu}} = v_s e^{iEt}$ 

$$E \begin{bmatrix} u_1 \\ u_2 \\ u_3 \\ u_4 \end{bmatrix} e^{-iEt} = m \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{bmatrix} u_1 \\ u_2 \\ u_3 \\ u_4 \end{bmatrix} e^{-iEt}$$

$$i\partial_t \Psi^{(-)}(t,\vec{r}) = \beta m \Psi^{(-)}(t,\vec{r}) \Longrightarrow$$
$$-E \begin{bmatrix} v_1 \\ v_2 \\ v_3 \\ v_4 \end{bmatrix} e^{iEt} = m \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{bmatrix} v_1 \\ v_2 \\ v_3 \\ v_4 \end{bmatrix} e^{iEt}$$

The four independent solutions are:

$$\begin{split} \Psi_{\uparrow}^{(+)}(\vec{r},t) &= \begin{bmatrix} \chi_{+} \\ 0 \end{bmatrix} e^{-iEt} , \ \Psi_{\downarrow}^{(+)}(\vec{r},t) = \begin{bmatrix} \chi_{-} \\ 0 \end{bmatrix} e^{-iEt} \\ \Psi_{\uparrow}^{(-)}(\vec{r},t) &= \begin{bmatrix} 0 \\ \chi_{+} \end{bmatrix} e^{iEt} , \ \Psi_{\downarrow}^{(-)}(\vec{r},t) = \begin{bmatrix} 0 \\ \chi_{-} \end{bmatrix} e^{iEt} \\ \mathcal{H}_{D}\Psi_{s}^{(+)}(t,\vec{r}) &= E\Psi_{s}^{(+)}(t,\vec{r}) = m\Psi_{s}^{(+)}(t,\vec{r}) , \ s = \{\uparrow,\downarrow\} \text{ positive energy} \end{split}$$

 $H_D \Psi_s^{(-)}(t, \vec{r}) = E \Psi_s^{(-)}(t, \vec{r}) = -m \Psi_s^{(-)}(t, \vec{r}) , \ s = \{\uparrow, \downarrow\}$  negative energy



**Dirac Sea**  $\rightarrow$  set of negative energy states

$$E=\pm\sqrt{p^2c^2+m^2c^4}$$

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Particle-Hole Pair Creation ( $e^-e^+$ ) (Electron-Positron) Dirac Sea Hole ightarrow Positron

### Particle in the Presence of an Electromagnetic Field

The free Dirac equation:

$$i\hbar\frac{\partial}{\partial t}\psi(\vec{r},t) = H_D \,\psi(\vec{r},t) = \left(\vec{\alpha}\cdot\hat{\vec{p}} + \beta m\right) \,\psi(\vec{r},t)$$

**Transforms in the Presence of an Electromagnetic Field**  $A^{\mu}(\vec{x}, t) = (\phi(\vec{x}, t), \vec{A}(\vec{x}, t))$  in:

$$i\hbar\frac{\partial}{\partial t}\psi(\vec{r},t) = H_{\mathcal{D}}\,\psi(\vec{r},t) = \left(\vec{\alpha}\cdot(\vec{\hat{p}}-q\vec{A})+\beta m\right)\,\psi(\vec{r},t) + q\phi\psi(\vec{r},t)$$

Rearranging and multiplying by  $\beta = \gamma_0$  to rewrite the equation in terms of matrices,  $\gamma^{\mu} = (\gamma_0, \vec{\gamma}) = (\beta, \beta \alpha)$ 

$$\left(\gamma_0(\hat{p}_0-q\phi)-\vec{\gamma}\cdot(\hat{\vec{p}}-q\vec{A})-m\right)\,\psi(\vec{r},t)=0$$

### Particle in the Presence of an Electromagnetic Field

Recalling the 4-vector notation, which the prescription corresponds to

$$\begin{aligned} x^{\mu} &= (t, \vec{x}) \ , \ x_{\mu} &= (t, -\vec{x}) \ , \ \partial^{\mu} \equiv \frac{\partial}{\partial x_{\mu}} = (\frac{\partial}{\partial t}, -\nabla) \ , \\ p^{\mu} &= (\hat{p_0}, \hat{\vec{p}}) = i\hbar(\frac{\partial}{\partial t}, -\nabla) \ , \ \ \boldsymbol{A}^{\mu} = (\phi, \vec{A}) \end{aligned}$$

$$\left(\gamma_0(\hat{p}_0 - q\phi) - \vec{\gamma} \cdot (\hat{\vec{p}} - q\vec{A}) - m\right) \psi(\vec{r}, t) = \left(\gamma_\mu(p^\mu - qA^\mu) - m\right) \psi(\vec{r}, t) = 0$$

Therefore, to describe a particle in the presence of an external electromagnetic field, we use the prescription: (Mininal coupling) :

$$p^{\mu} \rightarrow p^{\mu} - qA^{\mu} \Rightarrow i\hbar \frac{\partial}{\partial t} \rightarrow i\hbar \frac{\partial}{\partial t} - q\phi \ , \ -i\hbar \nabla \rightarrow -i\hbar \nabla - q\overline{A}$$

### Particle in a Magnetic Field

Let's Introduce the External Magnetic Field  $\vec{B}$  via Minimal Coupling:

$$p\!\!\!/ p \equiv \hat{p}^{\mu} \gamma_{\mu} 
ightarrow (\hat{p}^{\mu} - q A^{\mu}) \gamma_{\mu} \; ,$$

q= particle charge  $A^{\mu}=~(\phi,ec{A})~=~(0,0,Bx,0)$  (Landau gauge)

$$\Rightarrow \vec{B} = \nabla \times \vec{A} = B\hat{z} , \quad \nabla \cdot \vec{A} = 0 \quad , \vec{E} = 0 \quad , \phi = 0$$

$$(\not p - q \vec{A} - m) \Psi(t, \vec{r}) = 0 ,$$
  
$$i \partial_t \Psi(t, \vec{r}) = H(A^{\mu}(\vec{r})) \Psi(t, \vec{r}) = \left( \vec{\alpha} \cdot \left[ \hat{\vec{p}} - q \vec{A}(x^{\mu}) \right] + \beta m \right) \Psi(t, \vec{r})$$

### Particle in a Magnetic Field

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Let's consider an electron (q=-e) , and Landau gauge ( $A^{\mu}=(0,0,Bx,0)$ ), e= proton charge > 0

The Dirac equation assumes the following expression:

$$i\partial_t \Psi(t,\vec{r}) = H(x)\Psi(t,\vec{r}) = \left(\vec{\alpha} \cdot \left[\hat{\vec{p}} + eBx\hat{j}\right] + \beta m\right)\Psi(t,\vec{r})$$
$$i\frac{\partial}{\partial t}\psi(\vec{r},t) = \left[ \begin{pmatrix} 0 & \vec{\sigma} \cdot \left[\hat{\vec{p}} + eBx\hat{j}\right] \\ \vec{\sigma} \cdot \left[\hat{\vec{p}} + eBx\hat{j}\right] & 0 \end{pmatrix} + \begin{pmatrix} m & 0 \\ 0 & -m \end{pmatrix} \right]\psi(\vec{r},t)$$

Analogously to what we did when B=0, we will use an ansatz for the solution (positive energy)

$$\Psi(t,\vec{r}) = f(x)e^{-iEt+ip_yy+ip_zz}$$
,  $f(x) \rightarrow 4$ -spinor

#### Johnson-Lippmann solution

 $\Rightarrow$  ansatz for the solution (positive energy):

$$\Psi(t,\vec{r}) = \begin{pmatrix} C_1 v_{n-1}(\xi) \\ C_2 v_n(\xi) \\ C_3 v_{n-1}(\xi) \\ C_4 v_n(\xi) \end{pmatrix} e^{-iEt+ip_y y+ip_z z}$$

A given choice of  $C_1$ ,  $C_2$ ,  $C_3$ ,  $C_4 \Rightarrow$  four independent solutions:

$$\Psi^{\epsilon}(\vec{r}) = \begin{bmatrix} \left(\epsilon E_{n} + m\right)v_{n-1}(\xi)\\ 0\\ \epsilon \rho_{z}v_{n-1}(\xi)\\ i\rho_{n}v_{n}(\xi) \end{bmatrix} + \frac{1-s}{2} \begin{bmatrix} 0\\ (\epsilon E_{n} + m)v_{n}(\xi)\\ -i\rho_{n}v_{n-1}(\xi)\\ -\epsilon \rho_{z}v_{n}(\xi) \end{bmatrix} \end{bmatrix}$$

$$\Psi^{\epsilon}(t,\vec{r}) = \frac{(eB)^{1/4}}{(2\pi)} \frac{1}{\sqrt{2\epsilon E_n(\epsilon E_n + m)}} \Psi^{\epsilon}(\vec{r}) e^{-i\epsilon(Et+p_y y+p_z z)}$$

 $\epsilon = +1(-1) \rightarrow \text{positive (negative) state of energy s=+1(-1) <math>\rightarrow \text{spin states up (down)}$   $p_n = \sqrt{2eBn} \quad \xi = (eB)^{1/2}(x + \epsilon \frac{p_y}{eB})$ Convenient notation:

$$\psi^{\epsilon}(\vec{x},t) = \phi^{(\epsilon)}_{n,s,p_{y},p_{z}}(\vec{x})e^{-i\epsilon Et}$$

### The Positive Energy Solution for an Electron in the Presence of a Magnetic Field $\vec{B}$ :

$$\Psi^{(+)}(t,\vec{r}) = \begin{pmatrix} C_1 v_{n-1}(\xi) \\ C_2 v_n(\xi) \\ C_3 v_{n-1}(\xi) \\ C_4 v_n(\xi) \end{pmatrix} e^{-iEt+ip_y y+ip_z z}$$
$$= (eB)^{1/2}(x + \frac{p_y}{eB}) , \quad v_n(\xi) = \frac{1}{(\pi^{1/2} 2^n n!)^{1/2}} H_n(\xi) e^{-\frac{1}{2}\xi^2}$$

ξ

 $\frac{p_{v}}{eB}$  Determine the position where the oscillator wave functions are centered. If our system is contained in a box of side *L*:

$$0 \leq \frac{p_y}{eB} \leq L ,$$

$$\sum_{p_x} \to \sum_{n=0}^{\infty} g_n , \quad \sum_{p_y} \to \frac{L}{2\pi} \int dp_y = \frac{L}{2\pi} L eB , \quad \sum_{p_z} \to \frac{L}{2\pi} \int_{-\infty}^{\infty} dp_z$$

$$\frac{2}{V} \sum_{p_x, p_y, p_z} \equiv \frac{2}{(2\pi)^3} \int d^3p \to \sum_{n=0}^{\infty} g_n \frac{eB}{(2\pi)^2} \int_{-\infty}^{\infty} dp_z , \quad g_n = 2 - \delta_{n,0}$$

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### 3. IS THE DIMENSIONAL REDUCTION $3+1 \rightarrow 1+1$ ( $2+1 \rightarrow 0+1$ ) CONSISTENT WITH SPONTANEOUS CHIRAL SYMMETRY BREAKING?

V.P. Gusynin, V.A. Miransky, I.A. Shovkovy, Nucl. Phys. B462, 249 (1996)

In this section we consider the question whether the dimensional reduction  $3+1 \rightarrow 1+1$   $(2+1 \rightarrow 0+1)$  in the dynamics of the fermion pairing in a magnetic field is consistent with spontaneous chiral symmetry breaking. This question occurs naturally since, due to the Mermin-Wagner-Coleman (MWC) theorem [12], there cannot be spontaneous breakdown of continuous symmetries at D = 1 + 1 and D = 0+1. The MWC theorem is based on the fact that gapless Nambu-Goldstone (NG) bosons cannot exist in dimensions less than 2+1. This is in particular reflected in that the (1 + 1)-dimensional propagator of would be NG bosons would lead to infrared divergences in perturbation theory (as indeed happens in the  $1/N_c$  expansion in the (1 + 1)-dimensional Gross-Neveu model with a continuous symmetry [13]).

However, the MWC theorem is not applicable to the present problem. The central point is that the condensate  $\langle 0|\bar{\psi}\psi|0\rangle$  and the NG modes are **neutral** in this problem and the dimensional reduction in a magnetic field does not affect the dynamics of the center of mass of **neutral** excitations. Indeed, the dimensional reduction  $D \to D-2$  in the fermion propagator, in the infrared region, reflects the fact that the motion of **charged** particles is restricted in the directions perpendicular to the magnetic field. Since there is no such restriction for the motion of

### Particle in a Magnetic Field

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- Landau Levels with n=1,2,3... are doubly degenerate (spin s = ±1)
- Ground state, n = 0, is not degenerate and has spin s=-1 (for the electron) (In the figure we take  $p_z = 0$ ,  $\frac{eB}{M} = 1$ )

### **DIRAC FERMIONS AT B**≠0

• Dirac equation for charged fermions:

 $(i\gamma^{\mu}D_{\mu}-m)\psi=0$ 

where  $A_{\mu} = (A_0, -\vec{A})$  and the Landau gauge  $\vec{A} = (-By, 0, 0)$  is used.

• Look for a solution in the form:  $\psi = (i\gamma^{\mu}D_{\mu} + m)\phi$ . Then,

$$\left[-\partial_0^2 + (\partial_x + ieBy)^2 + \partial_y^2 + \partial_z^2 + i\gamma^1\gamma^2eB - m^2\right]\phi = 0$$

• Normalized solutions for  $\phi$  have the form

$$\phi_{k,\pm} \propto \frac{1 \pm i \operatorname{sgn}(eB) \gamma^1 \gamma^2}{2} \varphi_k(y) e^{-i\omega t + i p_x x + i p_z z}$$

where  $\varphi_k$  are harmonic oscillator wave functions, i.e.,

$$\varphi_k \propto H_k(\xi) e^{-\frac{\xi^2}{2}}, \quad \xi = \frac{y}{l} + p_x l \operatorname{sgn}(eB) \quad \text{and} \quad l = \frac{1}{\sqrt{|eB|}}$$

• The dispersion relation is given by

$$\omega = E_n^{\pm} = \pm \sqrt{2n|eB| + p_z^2 + m^2}$$

where 
$$n = k + \frac{1}{2} + \text{sgn}(eB)s_z$$
 and  $s_z = \pm \frac{1}{2}$  is an eigenvalue of  $\frac{i}{2}\gamma^1\gamma^2$   
orbital spin

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### **DEGENERACY OF LANDAU LEVELS**

• The Landau level energies are independent of  $p_x$ 

 $E_n^{\pm} = \pm \sqrt{2n|eB| + p_z^2 + m^2}$ 

- This means that each level is highly degenerate
- Let's calculate the degeneracy by confining the  $L_x$ system in a finite box of size  $L_x \times L_y$  with periodic boundary conditions
- The wave function is a plane wave in the x direction:  $\psi(x) \propto e^{ip_x x}$

$$\psi(0) = \psi(L_x) \implies e^{ip_x L_x} = 1 \implies p_x = \frac{2\pi n}{L_x}, n = 1, 2, ..., N_{\text{max}}$$

• The value of  $p_x$  sets the center of the Landau orbit in *y*-direction:

$$y_c \approx p_x l^2 \implies p_{x,\max} l^2 \lesssim L_y \implies \frac{2\pi N_{\max}}{L_x} \frac{1}{|eB|} \approx L_y \implies \frac{N_{\max}}{L_x L_y} \approx \frac{|eB|}{2\pi}$$

• The degeneracy is proportional to the field strength and the size (area) of the system in the spatial directions perpendicular to  $\vec{B}$  $N_{\max} \approx \frac{|eB|}{2\pi} L_x L_y$ 

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### LANDAU ENERGY SPECTRUM



### Example: Dirac Lagrangian

 $\mathcal{L} = \mathcal{L} \left( \bar{\Psi}(t,\vec{r}), \Psi(t,\vec{r}), \partial^{\mu} \Psi(t,\vec{r}) \right) = \bar{\Psi}(t,\vec{r}) \left( i \gamma_{\mu} \partial^{\mu} - m \right) \Psi(t,\vec{r}) , \ \bar{\Psi} \equiv \Psi^{\dagger} \gamma_{0}$ Equation of motion for  $\bar{\Psi}$ :

$$\frac{\partial \mathcal{L}}{\partial \bar{\Psi}} - \partial_{\mu} \frac{\partial \mathcal{L}}{\partial \partial_{\mu} \bar{\Psi}} = 0 \ , \ \frac{\partial \mathcal{L}}{\partial \bar{\Psi}} = (i \gamma_{\mu} \partial^{\mu} - m) \Psi \ , \ \frac{\partial \mathcal{L}}{\partial \partial_{\mu} \bar{\Psi}} = 0$$

that results in the Dirac equation:

$$(i\gamma_{\mu}\partial^{\mu}-m)\Psi(t,\vec{r})=0 \rightarrow (\not p-m)\Psi(t,\vec{r})=0$$

Equation of motion for  $\Psi$ :

$$\frac{\partial \mathcal{L}}{\partial \Psi} - \partial_{\mu} \frac{\partial \mathcal{L}}{\partial \partial_{\mu} \Psi} = 0 \ , \ \frac{\partial \mathcal{L}}{\partial \Psi} = -\bar{\Psi}m \ , \ \frac{\partial \mathcal{L}}{\partial \partial_{\mu} \Psi} = \bar{\Psi}i\gamma^{\mu}$$

that results in the Dirac equation:

$$-\bar{\Psi}m - \partial_{\mu}\bar{\Psi}i\gamma^{\mu} = 0 \rightarrow \bar{\Psi}\left(\overleftarrow{\partial_{\mu}}i\gamma^{\mu} + m\right) = 0 \rightarrow \bar{\Psi}\left(\overleftarrow{\not{p}} + m\right) = 0$$

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### Example: Dirac Hamiltonian

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The conjugate momentum to the fields  $\Psi \in \overline{\Psi}$  are given by:

$$\Pi_{\Psi} = \frac{\partial \mathcal{L}}{\partial \frac{\partial \Psi}{\partial t}} = \frac{\partial \mathcal{L}}{\partial \partial_0 \Psi} , \quad \Pi_{\bar{\Psi}} = \frac{\partial \mathcal{L}}{\partial \frac{\partial \bar{\Psi}}{\partial t}} = \frac{\partial \mathcal{L}}{\partial \partial_0 \bar{\Psi}}$$
$$\mathcal{L} = \bar{\Psi}(t, \vec{r}) (i\gamma_{\mu} \partial^{\mu} - m) \Psi(t, \vec{r}) \Longrightarrow \Pi_{\Psi} = \bar{\Psi} i\gamma_0 = i\Psi^{\dagger} , \quad \Pi_{\bar{\Psi}} = 0$$

Therefore the Hamiltonian density is given by:

$$\mathcal{H} = \Pi_{\Psi} \dot{\Psi} + \Pi_{\bar{\Psi}} \dot{\bar{\Psi}} - \mathcal{L} = i \Psi^{\dagger} \partial_{0} \Psi - \bar{\Psi} (i \gamma_{\mu} \partial^{\mu} - m) \Psi$$

simplifying the expression:

$$\mathcal{H} = i\Psi^{\dagger}\partial_{0}\Psi - \Psi^{\dagger}\gamma_{0}\left(i\gamma_{\mu}\partial^{\mu} - m\right)\Psi = i\Psi^{\dagger}\partial_{0}\Psi - \Psi^{\dagger}\gamma_{0}\left(i\gamma_{0}\partial_{0} + i\vec{\gamma}\cdot\nabla - m\right)\Psi$$

we obtain:

$$\mathcal{H} = \Psi^{\dagger} (-i\vec{\alpha} \cdot \nabla + \beta m) \Psi \Longrightarrow H = \int d^3 r \mathcal{H} = \int d^3 r \Psi^{\dagger} (-i\vec{\alpha} \cdot \nabla + \beta m) \Psi$$

#### Free Fields Quantization

Let's consider the **canonical quantization** in Quantum Field Theory. As an example we will take the scalar field:

$$\operatorname{QM} \left\{ \begin{array}{l} \left[ q_i, p_j \right] = i\hbar\delta_{i \ j} \\ \left[ q_i, q_j \right] = \left[ p_i, p_j \right] = 0 \end{array} \right. , \operatorname{QFT} \left\{ \begin{array}{l} \left\{ \Psi_{\alpha}(\vec{r}, t), \Pi_{\beta}(\vec{r}', t) \right\} = i\delta(\vec{r} - \vec{r}')\delta_{\alpha\beta} \\ \left\{ \Psi_{\alpha}(\vec{r}, t), \Psi_{\beta}(\vec{r}', t) \right\} = \left\{ \Pi_{\alpha}(\vec{r}, t), \Pi_{\beta}(\vec{r}', t) \right\} = 0 \end{array} \right.$$

we obtain

$$\Pi_{\alpha} = \frac{\partial \mathcal{L}}{\partial \partial^0 \Psi_{\alpha}} = i \Psi^{\dagger}(\vec{r}, t)$$

and, therefore the commutators of the scalar fields need to satisfy the canonical quantization relations:

$$\begin{split} \left\{ \Psi_{\alpha}(\vec{r},t),\Psi_{\beta}(\vec{r}\,',t) \right\} &= \left\{ \Psi_{\alpha}^{\dagger}(\vec{r},t),\Psi_{\beta}^{\dagger}(\vec{r}\,',t) \right\} = 0, \quad \left\{ \Psi_{\alpha}(\vec{r},t),\Psi_{\beta}^{\dagger}(\vec{r}\,',t) \right\} = i\delta(\vec{r}-\vec{r}\,')\delta_{\alpha\beta} \\ \hat{\Psi}(x) &= \sum_{r} \left( \hat{a}_{r}\phi_{r}^{(+)}(\vec{x})e^{-iE_{r}t} + \hat{b}_{r}^{\dagger}\phi_{r}^{(-)}(\vec{x})e^{iE_{r}t} \right) \\ \hat{\Psi}^{\dagger}(x) &= \sum_{r} \left( \hat{a}_{r}^{\dagger}\phi_{r}^{(+)}(\vec{x})^{\dagger}e^{iE_{r}t} + \hat{b}_{r}\phi_{r}^{(-)}(\vec{x})^{\dagger}e^{-iE_{r}t} \right) , \end{split}$$

We need to do the interpretation of  $\hat{a}_r$  and  $\hat{a}_r^\dagger$  as creation and annihilation operators of fermionic particles

and for  $\hat{b}_r$  and  $\hat{b}_r^{\dagger}$  as creation and annihilation operators for fermionic anti-particles (electron-positron or quark-antiquark)

$$\{\hat{a}_{r}, \hat{a}_{r'}^{\dagger}\} = \{\hat{b}_{r}, \hat{b}_{r'}^{\dagger}\} = \delta_{r \ r'} \quad , \quad \{\hat{a}_{r}, \hat{a}_{r'}\} = \{\hat{b}_{r}, \hat{b}_{r'}\} = 0 \; .$$

### SU(2) Nambu-Jona-Lasinio model (NJL)

The Lagrangian of the NJL model with two flavors (u and d quarks):

$$\mathcal{L} = \overline{\psi} \left( i \partial \!\!\!/ - \tilde{m} \right) \psi + G \left[ (\overline{\psi} \psi)^2 + (\overline{\psi} i \gamma_5 \vec{\tau} \psi)^2 \right]$$

interaction terms : scalar-isoscalar + pseudoscalar-isovector

 $ec{ au}$  are the isospin Pauli matrices

 $\psi$  is the Dirac fields of quaks u and d,

$$\psi = \begin{pmatrix} \psi_u \\ \psi_d \end{pmatrix}, \quad \tilde{m} = \begin{pmatrix} m_u & 0 \\ 0 & m_d \end{pmatrix}, \quad Q = \begin{pmatrix} q_u = \frac{2}{3}e & 0 \\ 0 & q_d = -\frac{1}{3}e \end{pmatrix}$$

We consider  $m_u = m_d = m$ 

### SU(2) Nambu-Jona-Lasinio model (NJL)

The Lagrangian of the NJL model to be suitable as an effective model for QCD (Quantum Chromodynamics)

 $\rightarrow$  It must reflect the symmetries (properties) of the strong interaction! Positive points:

- Invariant under global phase transformations  $\rightarrow$  Baryon number conservation
- The Lagrangian has chiral symmetry( in the limit  $m_u = m_d = 0$  )
- It has the spontaneous symmetry breaking mechanism (dynamic mass generation)
- The entire QCD phase diagram can be described by a single effective model (a single equation of state)

Negative points:

- The model is non-renormalizable (requires regularization, Λ-cutoff )
- The interaction does not have confinement (there are no gluons or color charge)

### NJL model in the mean field approximation (MFA)

$$\mathcal{L} = \overline{\psi} \left( i \not\!\!\!\partial - \tilde{m} \right) \psi + G \left[ (\overline{\psi} \psi)^2 + (\overline{\psi} i \gamma_5 \vec{\tau} \psi)^2 \right]$$

 $\text{MFA} \rightarrow \text{Linearization}$  of the interaction terms of  $\mathcal L$  neglecting quadratic fluctuations:

$$\begin{split} \hat{O} &\equiv \langle \hat{O} \rangle + (\hat{O} - \langle \hat{O} \rangle) &= \langle \hat{O} \rangle + \Delta \hat{O} \quad , \quad \hat{O} &= (\overline{\psi}\psi) \text{ or } (\overline{\psi}i\gamma_5\tilde{\tau}\psi) \\ \\ \mathsf{MFA} \to (\Delta \hat{O})^2 &\cong 0 ; \langle \overline{\psi}i\gamma_5\vec{\tau}\psi \rangle = 0 \text{ (symmetry)} \\ \hat{O}_1\hat{O}_2 &= (\langle \hat{O}_1 \rangle + \Delta \hat{O}_1)(\langle \hat{O}_2 \rangle + \Delta \hat{O}_2) \approx \langle \hat{O}_1 \rangle \langle \hat{O}_2 \rangle + \langle \hat{O}_1 \rangle \Delta \hat{O}_2 + \langle \hat{O}_2 \rangle \Delta \hat{O}_1 \\ \\ &= \langle \hat{O}_1 \rangle \langle \hat{O}_2 \rangle + \langle \hat{O}_1 \rangle (\hat{O}_2 - \langle \hat{O}_2 \rangle) + \langle \hat{O}_2 \rangle (\hat{O}_1 - \langle \hat{O}_1 \rangle) = \langle \hat{O}_1 \rangle \hat{O}_2 + \langle \hat{O}_2 \rangle \hat{O}_1 - \langle \hat{O}_1 \rangle \langle \hat{O}_2 \rangle \end{split}$$

therefore:

$$(\overline{\psi}\psi)^2 \approx 2 \langle \overline{\psi}\psi \rangle \overline{\psi}\psi - \langle \overline{\psi}\psi \rangle^2$$

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### NJL model in the mean field approximation (MFA)

$$\mathcal{L} \to \mathcal{L}_{MFA} = \overline{\psi} \left( i\partial \!\!\!/ - \tilde{m} \right) \psi + G \left[ 2 \langle \overline{\psi} \psi \rangle \overline{\psi} \psi - \langle \overline{\psi} \psi \rangle^2 \right]$$

defining the constituent mass

$$M = m - 2G\left\langle \overline{\psi}\psi \right\rangle$$

we obtain

$$\mathcal{L}_{MFA} = \overline{\psi} \left( i \partial \!\!\!/ - M 
ight) \psi - G \left\langle \overline{\psi} \psi \right\rangle^2 \,,$$

As we have seen, the Hamiltonian is easily obtained from the above Lagrangian:

$$\hat{H}_{MFA} = \int d^{3}r \mathcal{H} = \int d^{3}r \left[ \Psi^{\dagger} \left( -i\vec{\alpha} \cdot \nabla + \beta M \right) \Psi + G \left\langle \overline{\psi}\psi \right\rangle^{2} \right]$$

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### NJL model in the mean field approximation (MFA)

From the Hamiltonian operator, we obtain the energy of the system., *E*, calculating its statistical average value at T=0:

$$E = \langle \hat{H}_{MFA} \rangle = \int d^{3}r \mathcal{H} = \int d^{3}r \left[ \langle \Psi^{\dagger} \left( -i\vec{\alpha} \cdot \nabla + \beta M \right) \Psi \rangle + G \left\langle \overline{\psi}\psi \right\rangle^{2} \right]$$

noting that  $H_{Dirac} = -i\vec{\alpha} \cdot \nabla + \beta M$  and that Dirac field is expanded in a basis of  $H_{Dirac}$ :

$$\Psi(\vec{r},t) = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{i\vec{k}\cdot\vec{r}} \left( v_{s} e^{iE(k)t} b^{\dagger}_{-\vec{k},-s} + u_{s} e^{-iE(k)t} a_{\vec{k},s} \right)$$

where  $a_{\vec{k},s}^{\dagger}$  is the fermion creation operator(quark) with linear momentum  $\vec{k}$  and spin s and  $b_{\vec{k},s}^{\dagger}$  is the anti-fermion creation operator(antiquark) corresponding to the linear momentum  $\vec{k}$  and spin s and  $E(k) = \sqrt{k^2 + M^2}$ . The operators  $a_{\vec{k},s}$  and  $b_{\vec{k},s}$  are the corresponding annihilation operators.

### NJL in MFA - quark gas (fermions)

Substituting the expression for the field into the Dirac Hamiltonian operator, we can show that:

$$\frac{1}{V}\int d^{3}r\Psi^{\dagger}\left(-i\vec{\alpha}\cdot\nabla+\beta M\right)\Psi=\sum_{\xi}\int\frac{d^{3}p}{2\pi^{3}}\left(b_{\vec{p},\xi}^{\dagger}b_{\vec{p},\xi}+a_{\vec{p},\xi}^{\dagger}a_{\vec{p},\xi}-1\right)$$

The vacuum energy density can be calculated using the expression above for a quark gas at T = 0:

$$\begin{split} \epsilon = &< 0|\sum_{\xi} \int \frac{d^3p}{2\pi^3} \left( b^{\dagger}_{\vec{p},\xi} b_{\vec{p},\xi} + a^{\dagger}_{\vec{p},\xi} a_{\vec{p},\xi} - 1 \right) |0> + G\langle \overline{\psi}\psi \rangle^2 \\ \epsilon = &-\frac{2N_c N_f}{(2\pi)^3} \int_{|\vec{p}| < \Lambda} d^3p \sqrt{p^2 + M^2} + G\langle \overline{\psi}\psi \rangle^2 \\ = &- \frac{N_c N_f}{8\pi^2} \left( 2\Lambda E_{\Lambda}^3 - M^2 \Lambda E_{\Lambda} - M^4 \ln\left[\frac{\Lambda + E_{\Lambda}}{M}\right] \right) + G\langle \overline{\psi}\psi \rangle^2 \end{split}$$

where  $N_f=2$ ,  $N_c=3$  and  $E_{\Lambda}=\sqrt{\Lambda^2+M^2}$  and we introduce the cutoff  $\Lambda$  to regularize the integral.

### NJL in MFA - quark gas (fermions)

usando que

$$M = m - 2G\left\langle \overline{\psi}\psi \right\rangle \rightarrow \left\langle \overline{\psi}\psi \right\rangle = -\frac{M - m}{2G}$$

Therefore, we can rewrite the energy density,  $\epsilon$ , as:

$$\epsilon = -\frac{N_c N_f}{8\pi^2} \left(2\Lambda E_{\Lambda}^3 - M^2 \Lambda E_{\Lambda} - M^4 \ln\left[\frac{\Lambda + E_{\Lambda}}{M}\right]\right) + \frac{(M-m)^2}{4G}$$

#### **Gap Equation**

To obtain the Gap equation, we need to calculate

$$\left\langle \overline{\psi}\psi\right\rangle = \left\langle \psi^{\dagger}\gamma_{0}\psi\right\rangle$$

where

$$\psi = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{i\vec{k}\cdot\vec{r}} \left( v_{s}(-\vec{k})e^{iE(k)t}b^{\dagger}_{-\vec{k},-s} + u_{s}(\vec{k})e^{-iE(k)t}a_{\vec{k},s} \right)$$
  
$$\psi^{\dagger} = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{-i\vec{k}\cdot\vec{r}} \left( v_{s}^{\dagger}(-\vec{k})e^{-iE(k)t}b_{-\vec{k},-s} + u_{s}^{\dagger}(\vec{k})e^{iE(k)t}a^{\dagger}_{\vec{k},s} \right)$$

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### NJL in MFA - calculation of the gap equation

$$\begin{split} \left\langle \overline{\psi}\psi\right\rangle &= \left\langle \psi^{\dagger}\gamma_{0}\psi\right\rangle = \sum_{s}\int\frac{d^{3}k}{(2\pi)^{3/2}}e^{-i\vec{k}\cdot\vec{r}}\sum_{s'}\int\frac{d^{3}k'}{(2\pi)^{3/2}}e^{i\vec{k}'\cdot\vec{r}}\times\\ &< 0|\left(v_{s}^{\dagger}e^{-iEt}b_{-\vec{k},-s} + u_{s}^{\dagger}e^{iEt}a_{\vec{k},s}^{\dagger}\right)\gamma_{0}\left(v_{s'}e^{iEt}b_{-\vec{k}',-s'}^{\dagger} + u_{s'}e^{-iEt}a_{\vec{k}',s'}\right)|0> \end{split}$$

simplifying

$$\left\langle \overline{\psi}\psi \right\rangle = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{-i\vec{k}\cdot\vec{r}} \sum_{s'} \int \frac{d^{3}k'}{(2\pi)^{3/2}} e^{i\vec{k}'\cdot\vec{r}} v_{s}^{\dagger} \gamma_{0} v_{s'} < 0 |b_{-\vec{k},-s} b_{-\vec{k}',-s'}^{\dagger}|0>$$

but,

$$<0|b_{-\vec{k},-s}b^{\dagger}_{-\vec{k}',-s'}|0>=<0|b_{-\vec{k},-s}b^{\dagger}_{-\vec{k}',-s'}+b^{\dagger}_{-\vec{k}',-s'}b_{-\vec{k},-s}|0>=\\<0|\{b_{-\vec{k},-s}, b^{\dagger}_{-\vec{k}',-s'}\}|0>=\delta(\vec{k}-\vec{k}')\delta_{s\,s'}$$

$$\left\langle \overline{\psi}\psi \right\rangle = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{-i\vec{k}\cdot\vec{r}} \sum_{s'} \int \frac{d^{3}k'}{(2\pi)^{3/2}} e^{i\vec{k}'\cdot\vec{r}} v_{s}^{\dagger}(-\vec{k})\gamma_{0}v_{s'}(-\vec{k}')\delta(\vec{k}-\vec{k}')\delta_{s\,s'}$$

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### NJL in MFA - calculation of the gap equation

$$\begin{split} \left\langle \overline{\psi}\psi \right\rangle &= \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3/2}} e^{-i\vec{k}\cdot\vec{r}} \sum_{s'} \int \frac{d^{3}k'}{(2\pi)^{3/2}} e^{i\vec{k}'\cdot\vec{r}} v_{s}^{\dagger} \gamma_{0} v_{s'} \delta(\vec{k}-\vec{k}') \delta_{s\,s'}} \\ \left\langle \overline{\psi}\psi \right\rangle &= \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} v_{s}^{\dagger}(-\vec{k}) \gamma_{0} v_{s}(-\vec{k}) = \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \\ &\times \left[ \chi_{s}^{\dagger} \quad \chi_{s}^{\dagger} \frac{-\vec{\sigma}\cdot\vec{k}}{E+M} \right] \quad \begin{bmatrix} I & 0\\ 0 & -I \end{bmatrix} \quad \begin{bmatrix} \chi_{s} \\ -\vec{\sigma}\cdot\vec{k} \\ E+M \chi_{s} \end{bmatrix} = \\ &= \sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \quad \left[ \chi_{s}^{\dagger} \quad \chi_{s}^{\dagger} \frac{-\vec{\sigma}\cdot\vec{k}}{E+M} \right] \quad \begin{bmatrix} \chi_{s} \\ \vec{\sigma}\cdot\vec{k} \\ E+M \chi_{s} \end{bmatrix} \\ &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \quad \left[ -\chi_{s}^{\dagger}\chi_{s} + \quad \chi_{s}^{\dagger} \frac{\vec{\sigma}\cdot\vec{k}}{E+M} \frac{\vec{\sigma}\cdot\vec{k}}{E+M} \chi_{s} \right] \\ &\text{using that } \vec{\sigma} \cdot \vec{a} = \vec{a} \cdot \vec{b} + \vec{\sigma} \cdot \vec{a} \times \vec{b} \rightarrow \frac{\vec{\sigma}\cdot\vec{k}}{E+M} \frac{\vec{\sigma}\cdot\vec{k}}{E+M} = \frac{k^{2}}{(E+M)^{2}} \end{split}$$

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### NJL in MFA - calculation of the gap equation

$$\begin{split} \left\langle \overline{\psi}\psi \right\rangle &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \chi_{s}^{\dagger} \chi_{s} (1 - \frac{k^{2}}{(E+M)^{2}}) \\ \left\langle \overline{\psi}\psi \right\rangle &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \frac{(E+M)^{2} - k^{2}}{(E+M)^{2}} \\ \left\langle \overline{\psi}\psi \right\rangle &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \frac{E^{2} + 2EM + M^{2} - k^{2}}{(E+M)^{2}} \\ \left\langle \overline{\psi}\psi \right\rangle &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \frac{k^{2} + M^{2} + 2EM + M^{2} - k^{2}}{(E+M)^{2}} \\ \left\langle \overline{\psi}\psi \right\rangle &= -\sum_{s} \int \frac{d^{3}k}{(2\pi)^{3}} \frac{E+M}{2E} \frac{2M(E+M)}{(E+M)^{2}} = -2 \int \frac{d^{3}k}{(2\pi)^{3}} \frac{M}{\sqrt{k^{2} + M^{2}}} \end{split}$$

therefore, we obtain the Gap Equation:

$$\left\langle \overline{\psi}\psi \right\rangle = -rac{M-m}{2G} 
ightarrow rac{M-m}{2G} = 2N_f N_c \int rac{d^3k}{(2\pi)^3} rac{M}{\sqrt{k^2+M^2}}$$

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### SU(2)-NJL model in the presence of a B field

NJL Lagrangian with two flavors:

$$\mathcal{L} = \overline{\psi} \left( i \not\!\!D - \tilde{m} \right) \psi + G \left[ (\overline{\psi} \psi)^2 + (\overline{\psi} i \gamma_5 \vec{\tau} \psi)^2 \right] - \frac{1}{4} F^{\mu\nu} F_{\mu\nu}$$

 $F^{\mu
u}=\partial^{\mu}A^{
u}-\partial^{
u}A^{\mu}$  - electromagnetic field tensor

 $D^{\mu} = (i\partial^{\mu} - QA^{\mu})$  - covariant derivative (minimal coupling) we work in Landau gauge  $\rightarrow \vec{B} = B\hat{z}$ . Using the prescription:

$$\frac{2}{(2\pi)^3} \int d^3 p = \to \sum_{n=0}^{\infty} g_n \frac{eB}{(2\pi)^2} \int_{-\infty}^{\infty} dp_z$$

Thus, the Gap equation transforms into:

$$\frac{M-m}{2G} = N_c \sum_{q=u,d} \sum_{n=0}^{\infty} g_n \frac{|e_q|B}{(2\pi)^2} \int_{-\infty}^{\infty} dp_z \frac{M}{\sqrt{p_z^2 + M^2 + 2eBn}}$$

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### Gap equation - NJL

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### Effective mass increase with $B \rightarrow$ magnetic catalysis effect (MC)

Refs: parameters NJL : M. Buballa, Physics Reports 407 (2005)205

su(2)-NJL EOS: D. P. Menezes, M. Benghi Pinto, S. S. Avancini, A. Pérez Martinez and C. Providência, Phys. Rev. C 79, 035807 (2009).

### NJL equation of state with two flavors

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Equation of state using NJL model with two flavors.  $B_0 = 1 \times 10^{19}$  Gauss

### Mass-Radius diagram of a neutron star



Mass-radius diagram of a neutron star using the NJL model with two flavors.  $B_0 = 1 \times 10^{19}$  Gauss  $\beta$ -equilibrium is imposed  $\rightarrow$  chemical equilibrium for the reaction:  $n \rightleftharpoons p + e^-$ 

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#### Thermodynamical properties of the model

The mean-field Hamiltonian for the quarks in second quantization is given by:

$$H^{MFA} = \sum_{q=u,d} \sum_{n=0}^{\infty} \sum_{s=\pm 1} \sum_{p_2} \sum_{p_3} \sqrt{M^2 + p_3^2 + 2|Q_q|Bn} \left( \hat{a}_{nsp_2p_3}^{q\dagger} \hat{a}_{nsp_2p_3}^{q} + \hat{b}_{nsp_2p_3}^{q\dagger} \hat{b}_{nsp_2p_3}^{q} - 1 \right)$$

Grand canonical partition function:

$$Z = Tr[e^{-eta(H^{MFA}-\sum_q \mu_q \hat{N}_q)}] \quad , \quad \Omega = -rac{1}{eta}\ln Z \quad ,$$

where  $\beta$ =1/T. Thermodynamic quantities are related to  $\Omega$  through the following relations:

$$\Omega = \Omega(T, V, \mu_q, \mu_l) = E - TS - \sum_q \mu_q \bar{N}_q - \sum_l \mu_l \bar{N}_l ,$$
  

$$\Omega(T, V, \mu_q) = -PV , \ F = \Omega(T, V, \mu_q, \mu_l) + \sum_q \mu_q \bar{N}_q , \qquad (1)$$

where F = E - TS is the Helmholtz free energy and the average number of particles is obtained from the expression:

$$\bar{N}_{\alpha} = \frac{1}{\beta} \frac{\partial \ln Z}{\partial \mu_{\alpha}} = -\frac{\partial \Omega}{\partial \mu_{\alpha}} \,.$$

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Due to the particular form of the mean-field Hamiltonian, we have:

$$Z = Tr[e^{-\beta(H^{MFA} - \sum_{q} \mu_q \hat{N}_q)}] = e^{-\beta V(G\sigma^2 + \frac{1}{2}B^2)} Tr[e^{-\beta(H^{MFA} - \sum_{q} \mu_q \hat{N}_q)}]$$

where  $\bar{H}^{MFA}$  It corresponds to the NJL model Hamiltonian without  $VG\sigma^2$  and  $V\frac{1}{2}B^2$ . The representation of the occupation numbers in terms of the quark  $(n_{q_r})$  and antiquark  $(\bar{n}_{q_r})$  occupation numbers can be written as:

$$|\tilde{\alpha}\rangle = |n_{q_1}, n_{q_2}...; \bar{n}_{q_1}, \bar{n}_{q_2}, ...\rangle$$
 onde  $n_{q_r}, \bar{n}_{q_r} = 0, 1, ..., \infty$  e r = 1, 2, ...  $\infty$ 

We order, for example, the set of independent quark particle states according to the rule:

$$\{n_{q_r}\} = \{n_{q_{nsp_2p_3}}\} = (n_{q_1}, n_{q_2}, ..., n_{l_{\infty}}),$$

$$Tr[e^{-\beta(\tilde{H}^{MFA}-\sum_{q}\mu_{q}\hat{N}_{q})}] = \sum_{\tilde{\alpha}} \langle \tilde{\alpha} | e^{-\beta(\tilde{H}^{MFA}-\sum_{q}\mu_{q}\hat{N}_{q})} | \tilde{\alpha} \rangle$$
$$= e^{\beta \sum_{q,r} E_{r}^{q}} \times e^{-\beta \sum_{q,r,nqr} (E_{r}^{q}-\mu_{q})n_{qr}} e^{-\beta \sum_{q,r,\bar{n}qr} (E_{r}^{q}+\mu_{q})\bar{n}_{qr}}$$

Fermion occupation numbers can only take the values 0 or 1, and therefore we can write it using products:

$$Z = e^{-\beta V(G\sigma^{2} + \frac{1}{2}B^{2})} e^{\beta \sum_{q,r} E_{r}^{q}} \times \prod_{q,r} \left( 1 + e^{-\beta (E_{r}^{q} - \mu_{q})} \right) \prod_{q,r} \left( 1 + e^{-\beta (E_{r}^{q} + \mu_{q})} \right)$$

From the partition function, we can obtain the grand canonical thermodynamic potential:

$$\begin{split} \Omega_{Q} &= -\frac{1}{\beta} \ln Z = V(G\sigma^{2} + \frac{1}{2}B^{2}) - \sum_{q,r} E_{r}^{q} \\ &- \frac{1}{\beta} \sum_{q,r} \ln \left( 1 + e^{-\beta(E_{r}^{q} - \mu_{q})} \right) - \frac{1}{\beta} \sum_{q,r} \frac{1}{\beta} \ln \left( 1 + e^{-\beta(E_{r}^{q} + \mu_{q})} \right) \,. \end{split}$$

We have already shown that:

$$\sum_{r} = \sum_{n,s,p_2,p_3} \Rightarrow V \sum_{n} g_n \frac{|Q|B}{(2\pi)^2} \int_{-\infty}^{\infty} dp_3 .$$

The grand canonical potential can be written as:

$$\begin{split} \omega_{Q} &= \frac{\Omega_{Q}}{V} = \omega_{Q}(0,B) + \frac{1}{2}B^{2} \\ &- \frac{1}{\beta}\sum_{q,n}g_{n}\frac{N_{c}|Q_{q}|B}{(2\pi)^{2}}\int_{-\infty}^{\infty}dp_{3}\left(\ln\left(1 + e^{-\beta(E^{q} - \mu_{q})}\right) + \ln\left(1 + e^{-\beta(E^{q} + \mu_{q})}\right)\right) \,. \end{split}$$

$$\begin{split} \omega_Q(0,B) &= G\sigma^2 - \sum_{q,n} g_n \frac{N_c |Q_q| B}{(2\pi)^2} \int_{-\infty}^{\infty} dp_3 \sqrt{M^2 + p_3^2 + 2|Q_q| Bn} \\ &= \frac{(M - m_c)^2}{4G} + N_c \sum_{q=u,d} l_1^q(B) = \frac{(M - m_c)^2}{4G} + \Omega_{T=0}^{(1-Loop)} , \quad N_c = 3. \end{split}$$

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### Interesting expression

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$$I_1^q(B) = -\sum_{n=0}^{\infty} (2-\delta_{n0}) \frac{|Q_q|B}{(2\pi)^2} \int_{-\infty}^{\infty} dp_3 \sqrt{M^2 + p_3^2 + 2|Q_q|Bn}$$

This contribution essentially corresponds to the vacuum energy, that is, to the expectation value of the quark Hamiltonian in the vacuum state:

$$\begin{split} & I_{1}^{q}(B) = \left\langle 0 \left| \frac{H_{q}}{V} \right| 0 \right\rangle \\ & = \left\langle 0 \left| \frac{1}{V} \sum_{n=0}^{\infty} \sum_{s=\pm 1} \sum_{P_{2}} \sum_{\rho_{3}} \sqrt{M^{2} + \rho_{3}^{2} + 2|Q_{q}|Bn} \left( \hat{a}_{nsp_{2}p_{3}}^{q} \hat{a}_{nsp_{2}p_{3}}^{q} + \hat{b}_{nsp_{2}p_{3}}^{q} \hat{b}_{nsp_{2}p_{3}}^{q} - 1 \right) \right| 0 \right\rangle \end{split}$$

The contributions in  $l_1^q$  are clearly divergent and need to be regularized. We will rewrite them in a more convenient form using the generalized Riemann zeta function or the Hurwitz-Riemann zeta function:

$$\zeta(z,x) = \sum_{n=0}^{\infty} \frac{1}{(x+n)^z}$$

We can rewrite  $I_1^q$  as the following:

$$\begin{split} I_{1}^{q}(B) &= -\frac{(2|Q_{q}|B)^{\frac{3}{2}}}{(2\pi)^{2}} \int_{-\infty}^{\infty} dp_{3} \sum_{n=0}^{\infty} \sqrt{\left(\frac{M^{2}+p_{3}^{2}}{2|Q_{q}|B}+n\right)} + \frac{|Q_{q}|B}{(2\pi)^{2}} \int_{-\infty}^{\infty} dp_{3} \sqrt{M^{2}+p_{3}^{2}} \\ &= -\frac{(2|Q_{q}|B)^{\frac{3}{2}}}{(2\pi)^{2}} \int_{-\infty}^{\infty} dp_{3} \, \zeta(-\frac{1}{2},\frac{M^{2}+p_{3}^{2}}{2|Q_{q}|B}) + \frac{|Q_{q}|B}{(2\pi)^{2}} \int_{-\infty}^{\infty} dp_{3} \sqrt{M^{2}+p_{3}^{2}} \, . \end{split}$$

Using the integral representation of the zeta function:

$$\int_0^\infty dy y^{z-1} \exp[-\beta y] \coth(\alpha y) = \Gamma[z] \left\{ 2^{1-z} \alpha^{-z} \zeta(z, \frac{\beta}{2\alpha}) - \beta^{-z} \right\} , \qquad (2)$$

Making the identification:

$$\alpha = |Q_q| B, \ \beta = M^2 + p_3^2, \ z = -\frac{1}{2},$$

we obtain:

$$\begin{split} I_1^q(B) &= -\frac{(2|Q_q|B)^{\frac{3}{2}}}{(2\pi)^2} \int_{-\infty}^{\infty} dp_3 \; \frac{1}{2^{3/2}(|Q_q|B)^{1/2}} \\ &\times \left\{ \frac{1}{\Gamma(-1/2)} \int_0^{\infty} dy y^{-3/2} \exp[-(M^2 + p_3^2)y] \coth(|Q_q|By) + \sqrt{M^2 + p_3^2} \right\} \\ &+ \frac{|Q_q|B}{(2\pi)^2} \int_{-\infty}^{\infty} dp_3 \sqrt{M^2 + p_3^2} \;, \end{split}$$

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using that  $\Gamma(-1/2) = -2\pi^{1/2}$  we can rewrite the last expression as the following:

$$M_1^q(B) = \frac{|Q_q|B}{(2\pi)^2 2\pi^{1/2}} \int_{-\infty}^{\infty} dp_3 \int_0^{\infty} dy y^{-3/2} \exp[-(M^2 + p_3^2)y] \coth(|Q_q|By).$$

The  $p_3$  integration can be easily performed:

$$\int_{-\infty}^{\infty} dp_3 \exp[-p_3^2 y] = \frac{1}{y^{1/2}} \int_{-\infty}^{\infty} dp \exp[-p^2] = \frac{\pi^{1/2}}{y^{1/2}} .$$

The final results for  $I_1^q$  is the following:

$$I_1^q(B) = \frac{|Q_q|B}{8\pi^2} \int_0^\infty dy \frac{e^{-M^2y}}{y^2} \coth(|Q_q|By) = \frac{B_q}{8\pi^2} \int_0^\infty dy \frac{e^{-M^2y}}{y^2} \coth(B_q y), B_q = |Q_q|B.$$

The integration  $I_1^q(B)$  is clearly divergent and needs to be regularized.

$$\Omega_{T=0}^{(1-Loop)} \equiv N_c \sum_{f=u,d} l_1^q(B) = \frac{N_c}{8\pi^2} \sum_{f=u,d} \int_0^\infty \frac{dy}{y^3} e^{-yM^2} B_q y \coth(B_q y) \quad \leftarrow (\text{divergent if } y \to 0) \quad .$$

The origin of the divergences can be understood by using the Taylor series expansion of the function:

$$B_{q}y \coth(B_{q}y) \sim 1 + rac{(B_{q}y)^2}{3} + rac{(B_{q}y)^4}{45} + O[(B_{q}y)^6] ~,$$

 $\Rightarrow To regularize the effective potential, we need to perform two subtractions.$ 

1-Loop efective potential - MFIR regularization

$$\begin{split} \Omega_{T=0}^{(1-Loop)} &\equiv \frac{N_c}{8\pi^2} \sum_{q=u,d} \left\{ \underbrace{\int_0^{\infty} \frac{dy}{y^3} e^{-yM^2} \left[ B_q y \coth(B_q y) - 1 - \frac{(B_q y)^2}{3} \right]}_{finite} \right. \\ &+ \underbrace{\int_0^{\infty} \frac{dy}{y^3} e^{-yM^2}}_{infinity} + \frac{B_q^2}{3} \underbrace{\int_0^{\infty} \frac{dy}{y} e^{-yM^2}}_{infinity} \right\} \\ \Omega_{T=0}^{(mag)} &= \frac{N_c}{8\pi^2} \sum_{q=u,d} \int_0^{\infty} \frac{dy}{y^3} e^{-yM^2} \left[ B_q y \coth(B_q y) - 1 - \frac{(B_q y)^2}{3} \right] \\ \Omega_{T=0}^{(vac)} &= \frac{N_c}{8\pi^2} \sum_{q=u,d} \int_0^{\infty} \frac{dy}{y^3} e^{-yM^2} \rightarrow -\frac{N_c}{\pi^2} \sum_{q=u,d} \int_0^{\Lambda} p^2 \sqrt{M^2 + p^2} , \\ \Omega_{T=0}^{(field)} &= \frac{N_c}{24\pi^2} \sum_{q=u,d} B_q^2 \int_0^{\infty} \frac{dy}{y} e^{-yM^2} \rightarrow \frac{N_c}{24\pi^2} \sum_{q=u,d} B_q^2 \int_{1/\Lambda^2}^{\infty} \frac{dy}{y} e^{-yM^2} \\ &= \frac{N_c}{24\pi^2} \sum_{q=u,d} B_q^2 \Gamma \left[ 0, \frac{M^2}{\Lambda^2} \right] \sim -\frac{N_c}{24\pi^2} \sum_{q=u,d} B_q^2 \left[ \ln \left( \frac{M^2}{\Lambda^2} \right) + \gamma_E \right] \end{split}$$

$$\Omega_{\mathcal{T}=0}^{(1-\textit{Loop})} = \Omega_{\mathcal{T}=0}^{(\textit{mag})} + \Omega_{\mathcal{T}=0}^{(\textit{vac})} + \Omega_{\mathcal{T}=0}^{(\textit{field})}$$

(S. S. Avancini, R. L. S. Farias, M. B. Pinto, T. E. Restrepo and W. Tavares, Phys. Rev. D 103, 056009 (2021)

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## Conferences in Brazil



XXII Escola de Verão Jorge André Swieca de Física Nuclear Teórica

April 28 to May 2 (2025), Niterói, RJ, Brazil

CA1: A modern description of dense matter Palestrante: Veronica Dexheimer (Kent State University, EUA)

CA2: Hot and dense QCD in colliders Palestrante: Carlos Alberto Salgado (Universidade de Santiago de Compostela, Espanha)

CA3: Effective Field Theories Palestrante: Laura Tolos (Institute of Space Sciences, Espanha)

CA4: Nuclear reactions Palestrante: Chloe Hebborn (Michigan State Uni - EUA) Support was received in part by Consejo Nacional de Humanidades, Ciencia y Tecnología (México) grant number CF-2023-G-433.

# Thank you for your attention!