



Production of dark matter in association with a Higgs via exclusive photon fusion

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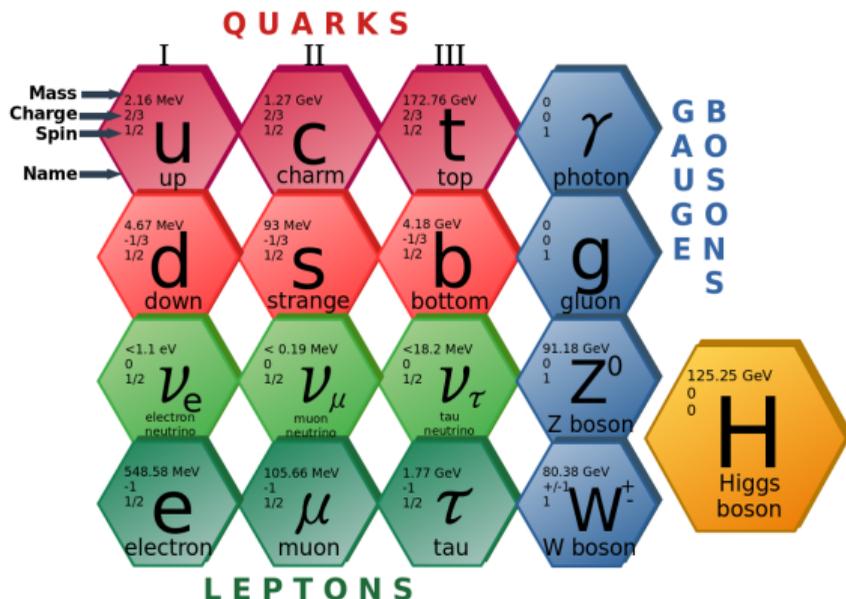
Outline



- 1 Introduction
- 2 Inert Doublet Model plus a complex Singlet (IDMS)
- 3 Central Exclusive Production
- 4 Results
- 5 Conclusions

Standard Model

- Quantum Field Theory: Electromagnetic + Weak + Strong Forces
- Symmetry group:
 $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
- Self-consistent. Free of anomalies.
Consistent perturbative scheme for systems at very high energies (LHC at 13 TeV)
- Higgs: last piece of the theory discovered in 2012, $M_H = 125 \text{ GeV}$.



Dark matter: observational evidence

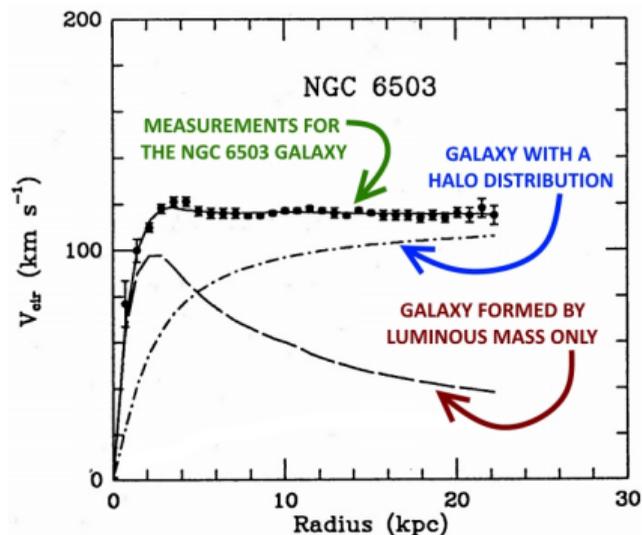


Figure: Rotation curve for the galaxy NGC6503 (Doroshkevich et. al., 2012).

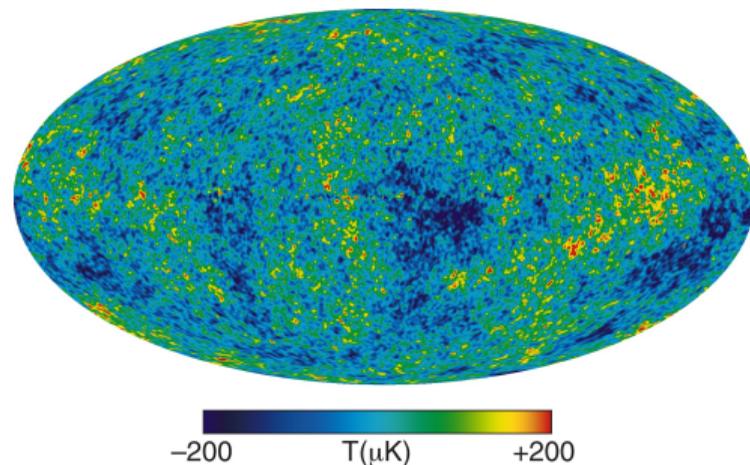


Figure: Internal Linear Combination Map (ILC), which is a linear combination of the WMAP (Wilkinson Microwave Anisotropy Probe) maps, at five different frequencies. This map shows the anisotropy of the CMB. (Bennett et. al., 2013).

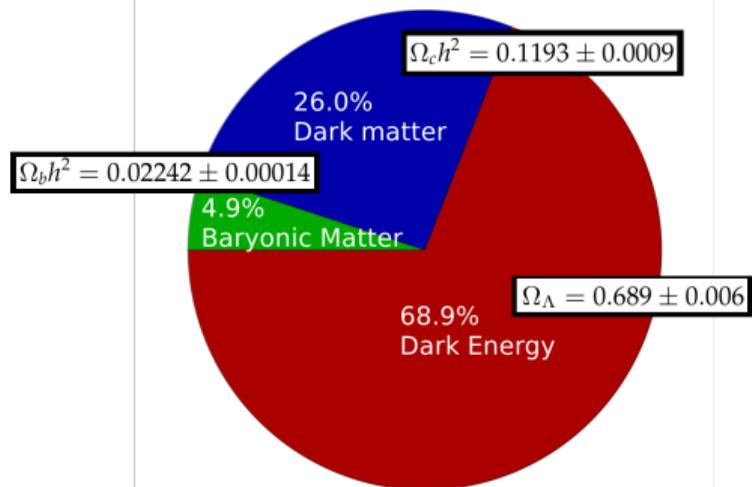
Dark matter: observational evidence

Fixing Planck's data: Λ CDM scheme
+ CMB anisotropies + baryon
acoustic oscillations (BAO) distance
measurements + Planck lensing
(Planck 2018 results):

$$\begin{aligned}\rho_{tot}/\rho_c &\equiv \Omega_{tot} \equiv \sum_i \Omega_i + \Omega_\Lambda \\ &= 0.9993 \pm 0.0019. \quad (1)\end{aligned}$$

Photons: $\Omega_\gamma h^2 = 2.473 \times 10^{-5} \rightarrow \sim 0\%$

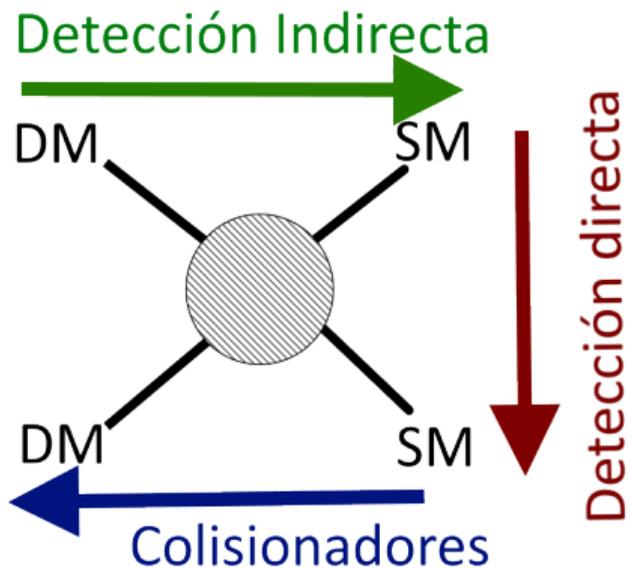
Neutrinos: $\Omega_\nu h^2 = (\sum m_\nu)/93.04 eV \leq 0.00132 \rightarrow \sim 1\%$



Dark Matter Properties



- Relic density observed: $\Omega_c h^2 = 0.1193 \pm 0.0009$.
- Cold: non-relativistic before the matter dominated era to form the cosmological structures we see today. *Hot dark matter* is not sufficient to account for the DM content of the Universe. *Warm dark matter* is a possibility.
- Effectively neutral: interact very weakly with electromagnetic radiation. Effectively a singlet of the $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$ group.
- Leave stellar evolution and Big Bang Nucleosynthesis predictions unchanged. The Tully-Fisher relation is another strong cosmological constraint.
- Consistent with current experimental bounds.



DM search approaches:

- Model based (*WIMPs, axion, ...*)
- Signal based (*fuzzy DM, astrophysics...*)

The Inert Higgs Doublet Model



The Inert Higgs Doublet Model (IDM) [1] is a simple way to introduce a DM candidate via a Lorentz scalar $SU(2)$ doublet (Φ_2), in addition to the Standard Model Higgs doublet (Φ_1).

What distinguishes the IDM to more general models with two Higgs doublets is that its potential has a \mathbb{Z}_2 symmetry unbroken by the vacuum state[2]. The unbroken discrete symmetry implies no Yukawa couplings between Φ_2 and fermions (hence called *inert*).

The most general CP conserving Higgs potential for this case is:

$$\begin{aligned} V = & \mu_1^2 \Phi_1^\dagger \Phi_1 + \mu_2^2 \Phi_2^\dagger \Phi_2 + \lambda_1 \left(\Phi_1^\dagger \Phi_1 \right)^2 + \lambda_2 \left(\Phi_2^\dagger \Phi_2 \right)^2 \\ & + \lambda_3 \left(\Phi_1^\dagger \Phi_1 \right) \left(\Phi_2^\dagger \Phi_2 \right) + \lambda_4 \left| \Phi_1^\dagger \Phi_2 \right|^2 + \lambda_5 \text{Re} \left[\left(\Phi_1^\dagger \Phi_2 \right) \right]^2 \end{aligned} \quad (2)$$

Inert Doublet Model plus a complex Scalar (IDMS)

The IDMS¹ [3] incorporates to the SM gauge symmetry a local $U(1)_X$ gauge symmetry and a second $SU(2)$ scalar doublet, where the DM candidate arises.

Stability of DM

- The second scalar doublet must have a VEV equal to zero.
- Either we impose a discrete \mathcal{Z}_2 symmetry or a $U(1)_X$ symmetry.

The scalar fields and their assignments under the $G_{SM} \otimes U(1)_X$ group are given by:

$$\begin{aligned}
 \Phi_1 &\sim (\mathbf{1}, \mathbf{2}, 1/2, x_1), \\
 \Phi_2 &\sim (\mathbf{1}, \mathbf{2}, 1/2, x_2), \\
 \mathcal{S}_X &\sim (\mathbf{1}, \mathbf{1}, 0, x).
 \end{aligned} \tag{3}$$

¹Eur. Phys. J. C, 80(8):788, 2020

Scalar fields

The scalar fields are written as follows:

$$\begin{aligned}\Phi_1 &= \begin{pmatrix} \phi_1^+ \\ \frac{1}{\sqrt{2}}(v + \phi_1 + i\eta_1) \end{pmatrix}, \\ \Phi_2 &= \begin{pmatrix} \phi_2^+ \\ \frac{1}{\sqrt{2}}(\phi_2 + i\eta_2) \end{pmatrix}, \\ S_X &= \frac{1}{\sqrt{2}}(v_X + s_X + i\eta_X).\end{aligned}\tag{4}$$

The spontaneous symmetry breaking (SSB) occurs as follows

$$G_{\text{SM}} \times U(1)_X \xrightarrow{\langle S_X \rangle} G_{\text{SM}} \xrightarrow{\langle \Phi_1 \rangle} SU(3)_C \times U(1)_{\text{EM}},\tag{5}$$

where $\langle S_X \rangle = v_X/\sqrt{2}$ and $\langle \Phi_1 \rangle^T = (0, v/\sqrt{2})$ with $v = 246.22$ GeV.

Higgs potential

The most general, renormalizable and gauge invariant potential is

$$\begin{aligned} V = & \mu_1^2 \Phi_1^\dagger \Phi_1 + \mu_2^2 \Phi_2^\dagger \Phi_2 + \mu_x^2 \mathcal{S}_X^* \mathcal{S}_X + \left[\mu_{12}^2 \Phi_1^\dagger \Phi_2 + h.c. \right] \\ & + \lambda_x (\mathcal{S}_X^* \mathcal{S}_X)^2 + \lambda_1 (\Phi_1^\dagger \Phi_1)^2 + \lambda_2 (\Phi_2^\dagger \Phi_2)^2 \\ & + \lambda_3 (\Phi_1^\dagger \Phi_1) (\Phi_2^\dagger \Phi_2) + \lambda_4 |\Phi_1^\dagger \Phi_2|^2 + \left[\lambda_5 (\Phi_1^\dagger \Phi_2)^2 \right. \\ & + \lambda_6 (\Phi_1^\dagger \Phi_1) (\Phi_1^\dagger \Phi_2) + \lambda_7 (\Phi_2^\dagger \Phi_2) (\Phi_1^\dagger \Phi_2) + h.c. \left. \right] \\ & + (\mathcal{S}_X^* \mathcal{S}_X) \left[\lambda_{1x} (\Phi_1^\dagger \Phi_1) + \lambda_{2x} (\Phi_2^\dagger \Phi_2) \right] \\ & + \left[\lambda_{12x} (\Phi_1^\dagger \Phi_2) (\mathcal{S}_X^* \mathcal{S}_X) + h.c. \right], \end{aligned} \quad (6)$$

where $\mu_{1,2}^2$, $\lambda_{1,2,3,4,1x,2x}$ are real parameters and μ_{12}^2 , $\lambda_{5,6,7,12x}$ can be complex parameters. Note that $\mu_2^2 > 0$ because the DM candidate arises from Φ_2 , which has $\langle \Phi_2 \rangle = 0$.

Mass matrix

By considering the $\{\phi_1, s_x, \phi_2, \xi_2\}$ basis, and after the SSB, the scalar mass matrix can be written as follows

$$\mathbf{M}_0^2 = \begin{pmatrix} M_{11} & M_{12} & M_{13} & 0 \\ M_{12} & M_{22} & M_{23} & 0 \\ M_{13} & M_{23} & M_{33} & M_{34} \\ 0 & 0 & M_{34} & M_{44} \end{pmatrix}, \quad (7)$$

where

$$\begin{aligned} M_{11} &= 2\lambda_1 v^2, & M_{12} &= \lambda_{1x} v v_x, & M_{13} &= \frac{1}{2} \lambda_6 v^2, \\ M_{22} &= 2\lambda_x v_x^2, & M_{23} &= \frac{1}{2} \lambda_{12x} v v_x, \\ M_{33} &= \mu_2^2 + \frac{1}{2} (\lambda_3 + \lambda_4 + \text{Re}[\lambda_5]) v^2 + \frac{1}{2} \lambda_{2x} v_x^2, \\ M_{34} &= -\text{Im}[\lambda_5] v^2, \\ M_{44} &= \mu_2^2 + \frac{1}{2} (\lambda_3 + \lambda_4 - \text{Re}[\lambda_5]) v^2 + \frac{1}{2} \lambda_{2x} v_x^2. \end{aligned} \quad (8)$$

Mass matrix

The mass matrix for the neutral scalar is diagonalized by

$$\begin{pmatrix} h \\ S \end{pmatrix} = \begin{pmatrix} \cos \alpha_1 & -\sin \alpha_1 \\ \sin \alpha_1 & \cos \alpha_1 \end{pmatrix} \begin{pmatrix} \phi_1 \\ s_x \end{pmatrix} \quad (9)$$

and

$$\begin{pmatrix} \chi \\ A \end{pmatrix} = \begin{pmatrix} \cos \alpha_2 & -\sin \alpha_2 \\ \sin \alpha_2 & \cos \alpha_2 \end{pmatrix} \begin{pmatrix} \phi_2 \\ \xi_2 \end{pmatrix}, \quad (10)$$

where $\tan \alpha_{1,2} = \frac{\rho_{1,2}}{1 + \sqrt{1 + \rho_{1,2}^2}}$ with $\rho_1 = \frac{\lambda_{1x} v v_x}{\lambda_1 v^2 - \lambda_x v_x^2}$ and $\rho_2 = \frac{-\text{Im}[\lambda_5]}{\text{Re}[\lambda_5]}$ [4].

The scalar masses are given by

$$\begin{aligned} M_{S,h}^2 &= \lambda_1 v^2 + \lambda_x v_x^2 \pm (\lambda_1 v^2 + \lambda_x v_x^2) \sqrt{1 + \rho_1^2}, \\ M_{H^\pm}^2 &= \mu_2^2 + \frac{1}{2}(\lambda_3 v^2 + \lambda_{2x} v_x^2), \\ M_{A,\chi}^2 &= M_{H^\pm}^2 + \left(\frac{\lambda_4}{2} \pm |\lambda_5| \right) v^2. \end{aligned} \quad (11)$$

Couplings

The IDMS couplings involved in the calculations of this work are as follows:

Coupling	Expression
$hf_i\bar{f}_i$	$\frac{m_{f_i}}{v} \cos \alpha_1$
hH^-H^+	$(\lambda_3 \cos \alpha_1 + \lambda_{2x}/2)v$
$hW_\mu^-W_\nu^+$	$gm_W \cos \alpha_1 g_{\mu\nu}$
$h\chi\chi$	$(\lambda_{345} \cos \alpha_1 + \lambda_{2x}/2)v$
$Z'_\mu f_i\bar{f}_i$	$\frac{g_x}{2} (1 - \gamma^5) \gamma^\mu$
$Z'_\mu \chi\chi$	$\frac{g_x}{2} (p_{Z'_\mu} - p_\chi)^\mu$
$Sf_i\bar{f}_i$	$\frac{m_{f_i}}{v} \sin \alpha_1$
$SW_\mu^-W_\nu^+$	$gm_W \sin \alpha_1 g_{\mu\nu}$
$S\chi\chi$	$\lambda_{2x}v_x \cos \alpha_1 - \lambda_{345}v \sin \alpha_1$
$\chi(A)H^\pm W_\mu^\mp$	$i\frac{g}{\sqrt{2}}(p_{H^\pm} - p_{\chi(A)})^\mu \cos \alpha_2$
$S_\chi h$	$\frac{\lambda_{12x} \cos \alpha_1}{2} \left(\frac{v_x}{\cos \alpha_2} \right)$

To avoid a decay of DM candidate into $H^\pm W^\mp$ or two neutral scalars, we require that $v_x > \sqrt{\frac{2M_h^2 + 2\mu_2^2}{4\lambda_x - \lambda_{2x}}}$ and $m_\chi < \sqrt{2\lambda_x}v_x$. We define $\lambda_{345} = \lambda_3 + \lambda_4 + 2\lambda_5$.

To enforce dark matter stability (χ), we can take the following approaches:

- Impose a discrete \mathcal{Z}_2 symmetry, then the terms proportional to $\phi_1^\dagger \phi_2$ are eliminated ($\mu_{12} = \lambda_6 = \lambda_7 = \lambda_{12x} = 0$). **No interactions with Higgs!**

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- Impose a discrete \mathcal{Z}_2 symmetry AND allow a lightly parity violating term, $\lambda_{12x} \neq 0$

Central Exclusive Production

In proton-proton collisions, a colorless exchange (γ , color neutralized gluons or a pomeron, \mathbb{P}) could result in a $p + p \rightarrow p + X + p$ process, where the protons remain intact after the collision.

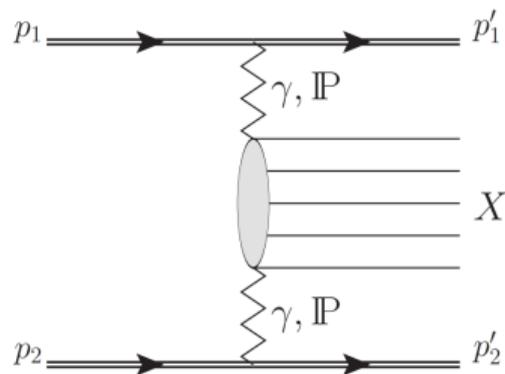


Figure: Central Exclusive Production diagram ($p + p \rightarrow p + X + p$) [5].

Central Exclusive Production

The outgoing protons can be captured by near beam detectors, such as the CMS-TOTEM Precision Proton Spectrometer (PPS) or the ATLAS Forward Proton detectors (AFP).

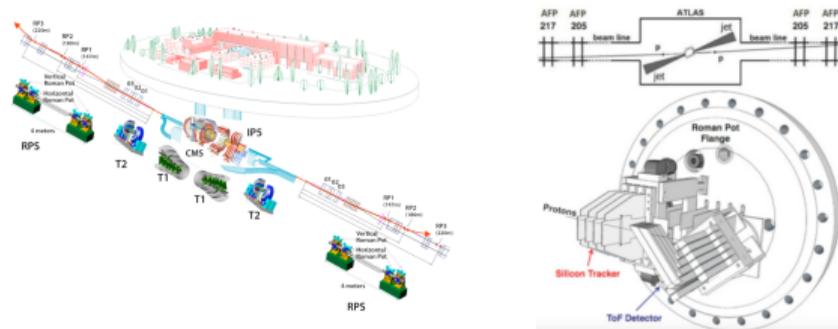


Figure: CMS-TOTEM PPS (left) and AFP (right).

Each proton loses a fraction of its longitudinal momentum, $\xi_{1,2} = \Delta p_{1,2}/p_{1,2}$, and the mass of the X state can be reconstructed, $M(X) = M(pp) \sim \sqrt{\xi_1 \xi_2 s}$.

Dark Matter associated with a Higgs in γ -fusion CEP

The process of interest is $pp \rightarrow h\chi pp$, where the DM particle χ is produced along side a Higgs in photon-fusion induced CEP. The proton kinematics allows for the reconstruction of the scalar S , while measurements of the central objects would allow the reconstructions of the χ kinematics.

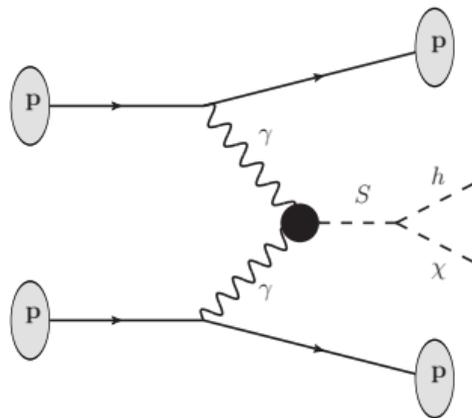


Figure: Diagram of the process $pp \rightarrow h\chi pp$.

Preliminary Results

IDMS model implemented using LanHEP 4.0. We calculate the $pp \rightarrow h\chi pp$ cross-section using MadGraph5_aMC ver. 3.5. The photon-fusion to S vertex is done at an effective level.

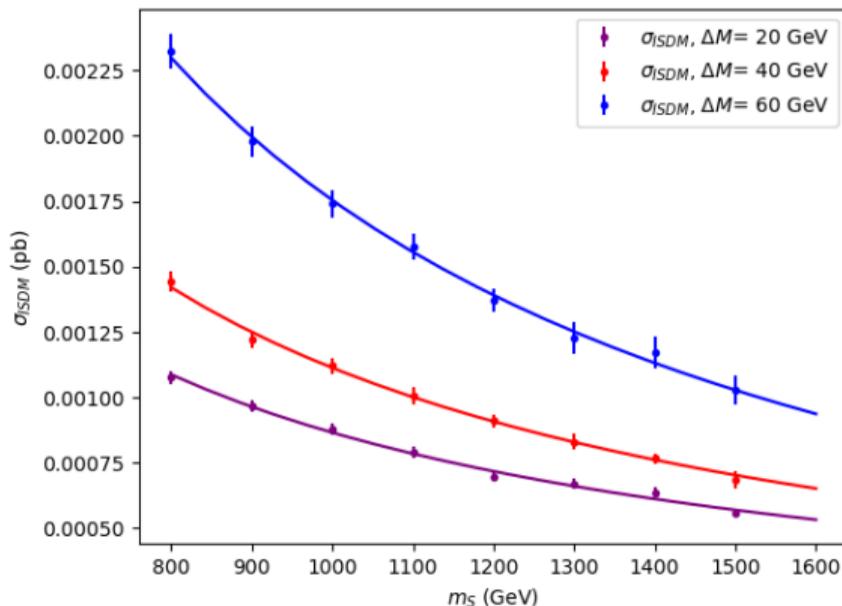


Figure: Cross-section of the $pp \rightarrow h\chi pp$ process for different masses of m_S and for $\Delta M = m_S - m_h - m_\chi = 20, 40$ and 60 GeV.

- The analysis of the Central Exclusive Production opens the door to searches for physics beyond the Standard Model, where processes with a dark matter candidate alongside an SM particle are of particular interest.
- We explore the possibility that a DM candidate can come from an extension of the SM with a local $U(1)_\chi$ gauge symmetry and a second $SU(2)$ scalar doublet.
- We calculate the cross section of the $pp \rightarrow h\chi pp$ process, where the scalar S is produced through a photon-fusion induced CEP and subsequently decays into $h + \chi$, for different values of the m_S and for $\Delta M = m_S - m_h - m_\chi = 20, 40$ and 60 GeV.
- More calculations of the cross-section as well as the missing mass spectrum need to be done. The particular characteristics of the near-beam proton detectors also need to be taken into account to determine the expected signatures of the process.

- [1] Nilendra G. Deshpande and Ernest Ma. Pattern of symmetry breaking with two higgs doublets. *Phys. Rev. D*, 18:2574–2576, Oct 1978. doi: 10.1103/PhysRevD.18.2574. URL <https://link.aps.org/doi/10.1103/PhysRevD.18.2574>.
- [2] Michael Gustafsson. The Inert Doublet Model and Its Phenomenology. *PoS, CHARGED2010:030*, 2010. doi: 10.22323/1.114.0030.
- [3] M. A. Arroyo-Ureña, R. Gaitan, R. Martinez, and J. H. Montes de Oca Yemha. Dark matter in Inert Doublet Model with one scalar singlet and $U(1)_X$ gauge symmetry. *Eur. Phys. J. C*, 80(8):788, 2020. doi: 10.1140/epjc/s10052-020-8316-9.
- [4] L. G. Cabral-Rosetti, R. Gaitán, J. H. Montes de Oca, R. Osorio Galicia, and E. A. Garcés. Scalar dark matter in inert doublet model with scalar singlet. *J. Phys. Conf. Ser.*, 912(1): 012047, 2017. doi: 10.1088/1742-6596/912/1/012047.
- [5] The CMS Precision Proton Spectrometer at the HL-LHC – Expression of Interest. 3 2021.

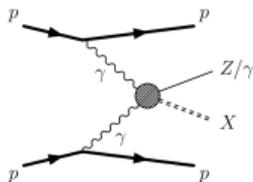


THANK YOU!



Auxiliary Slides

Search for New Physics in CEP



Search for new physics in CEP using the missing mass technique with the CMS-TOTEM PPS detector (Eur. Phys. J. C, 83:827, 2023).

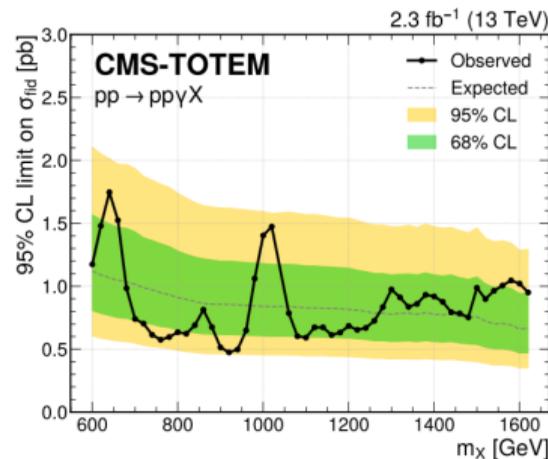
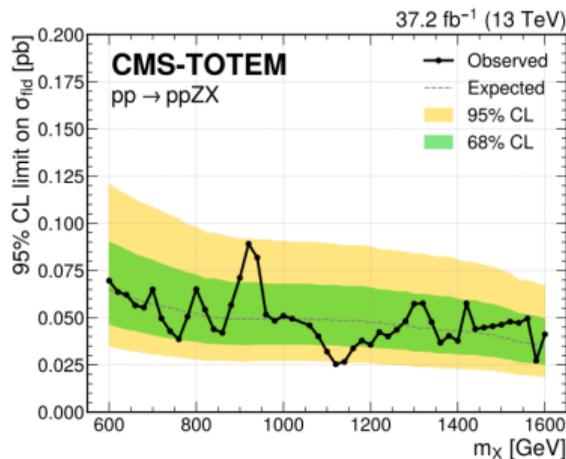


Figure: Upper limits on the $pp \rightarrow ppZ/\gamma + X$ cross section at 95% CL, as a function of m_X found by the CMS-TOTEM study.

Constraints on IDMS free model parameters

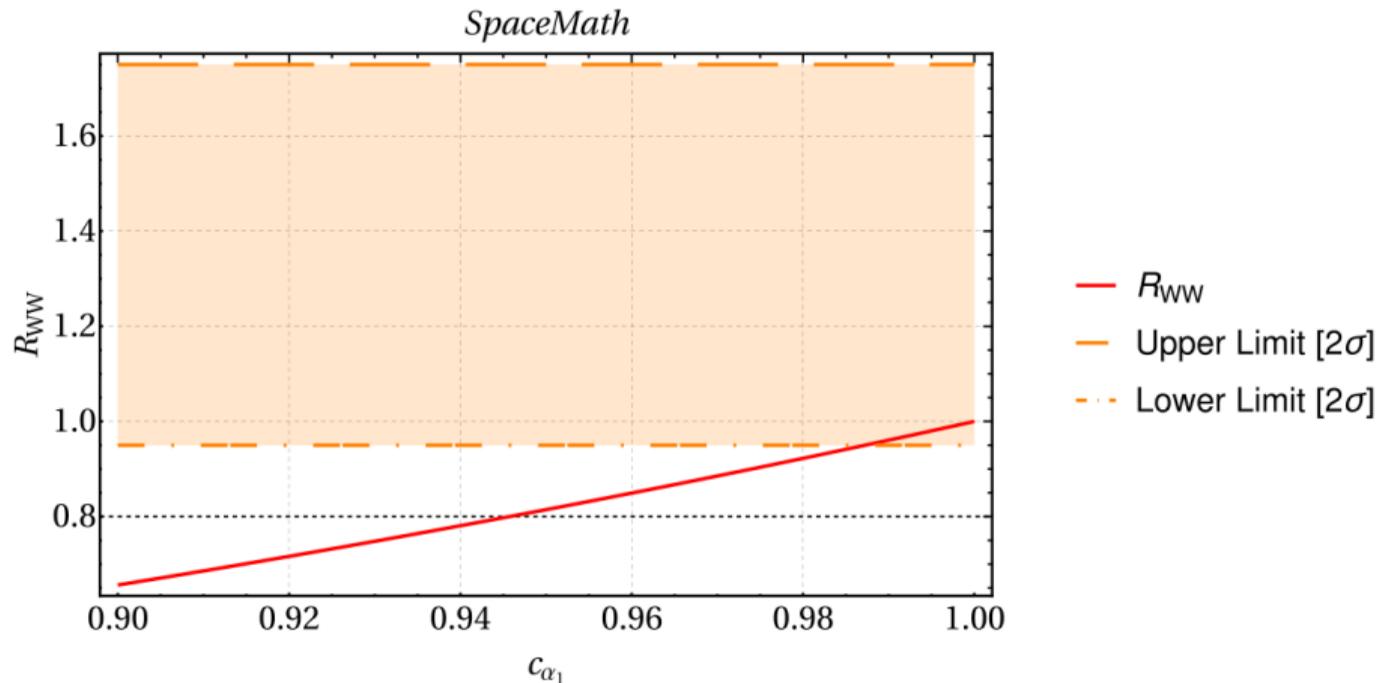


Figure: \mathcal{R}_{WW}^* as a function of $\cos \alpha_1$. The shadowed area represents the allowed region by the signal from collider experiments at 2σ .

Constraints on IDMS free model parameters

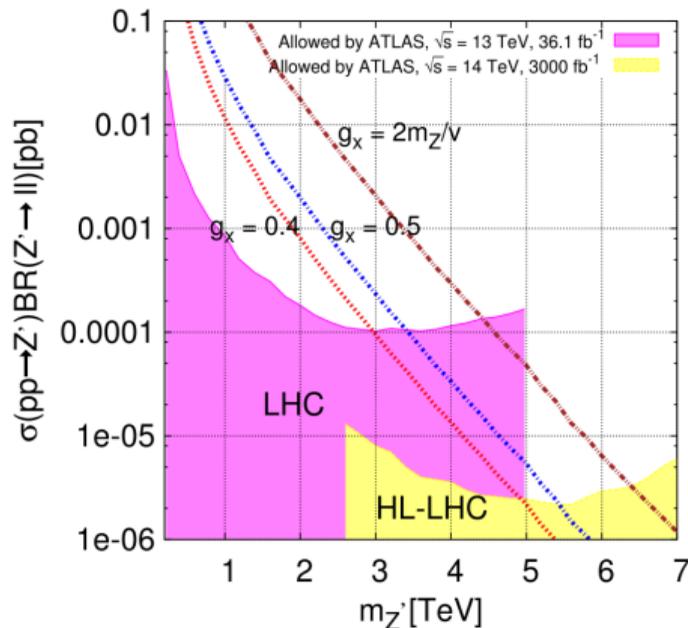


Figure: $\sigma_{Z'} \mathcal{B}_{Z'}$ as a function of the Z' gauge boson mass for $g_x = 0.4, 0.5$ and $2m_{Z'}/v$. Dark areas correspond to allowed regions by ATLAS Collaboration.

Constraints on IDMS free model parameters

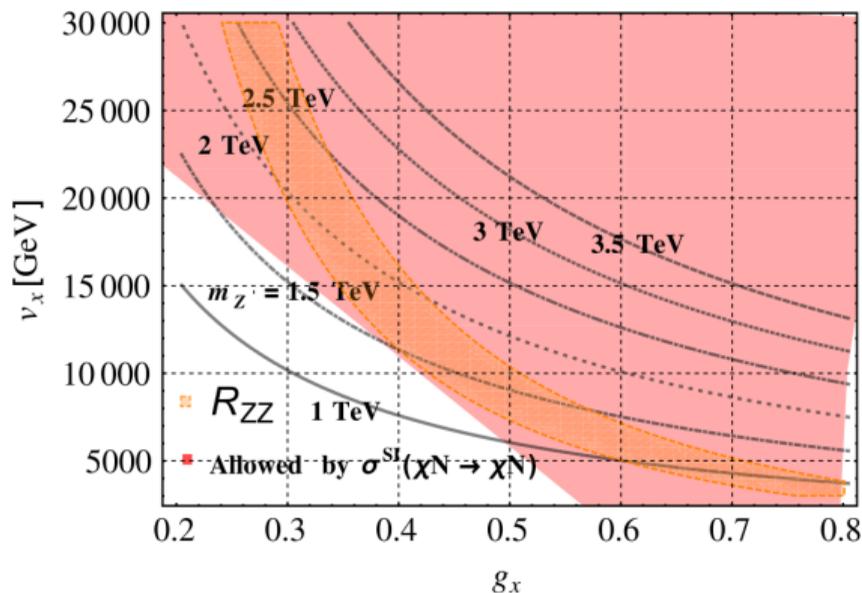


Figure: The inclined and curved orange area represents the consistent region with R_{ZZ}^* while the large area (light red) indicates the allowed values by the upper limit on $\sigma_{SI}(\chi N \rightarrow \chi N)$.