

GENERALIZED NEUTRINO INTERACTIONS IN DARK MATTER and $CE\nu NS$ EXPERIMENTS

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NON STANDARD NEUTRINO INTERACTIONS (NSI)

$$\mathcal{L}_{\text{NSI}} = - 2\sqrt{2}G_F \sum_{\alpha,\beta,f,C} \left\{ \varepsilon_{\alpha\beta}^{fC} (\bar{\nu}_{\alpha}\gamma^{\mu}P_L\nu_{\beta}) (\bar{f}\gamma_{\mu}P_C f) \right\}$$

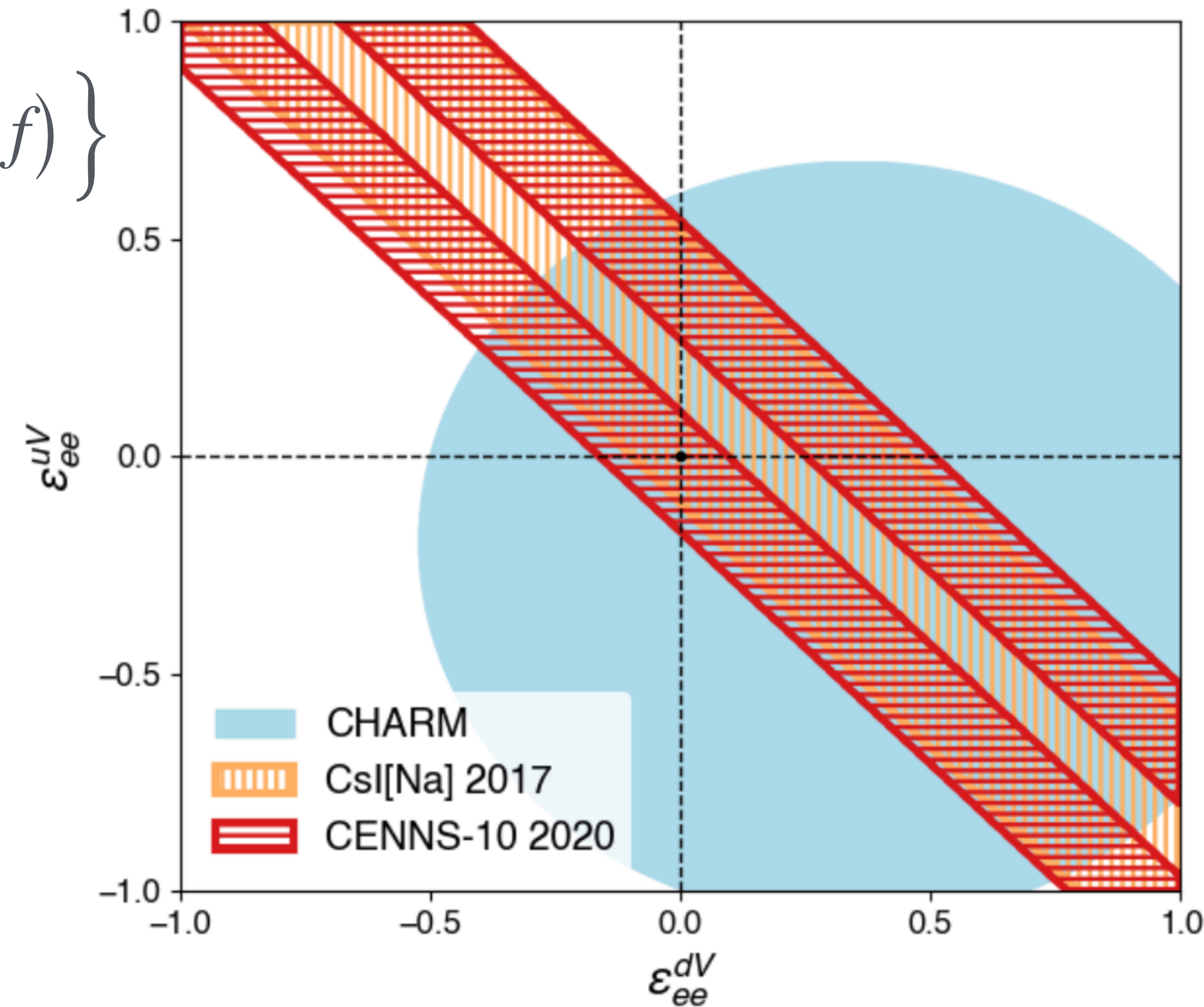
Useful to study new **vector** interactions in a model independent way

Could have sizable effects in:

✱ Production

✱ Detection

✱ Propagation in matter



COHERENT Collaboration • D. Akimov (Moscow, ITEP and Moscow Phys. Eng. Inst.) et al. (Apr 3, 2020)
Published in: *Phys.Rev.Lett.* 126 (2021) 1, 012002 • e-Print: 2003.10630 [nucl-ex]

GENERALIZED NEUTRINO INTERACTIONS (GNI)

- Also referred to as NGI

$$\mathcal{L}_{\text{GNI}} = - \frac{G_F}{\sqrt{2}} \sum_{\alpha, \beta, f} \left\{ \epsilon_{\alpha\beta}^{fS} (\bar{\nu}_\alpha (1 - \gamma^5) \nu_\beta) (\bar{f}f) + \tilde{\epsilon}_{\alpha\beta}^{fS} (\bar{\nu}_\alpha (1 + \gamma^5) \nu_\beta) (\bar{f}f) \right. \\ \left. - \epsilon_{\alpha\beta}^{fP} (\bar{\nu}_\alpha (1 - \gamma^5) \nu_\beta) (\bar{f} \gamma^5 f) - \tilde{\epsilon}_{\alpha\beta}^{fP} (\bar{\nu}_\alpha (1 + \gamma^5) \nu_\beta) (\bar{f} \gamma^5 f) \right. \\ \left. + \epsilon_{\alpha\beta}^{fT} (\bar{\nu}_\alpha \sigma^{\mu\nu} P_L \nu_\beta) (\bar{f} \sigma_{\mu\nu} P_L f) + \tilde{\epsilon}_{\alpha\beta}^{fT} (\bar{\nu}_\alpha \sigma^{\mu\nu} P_R \nu_\beta) (\bar{f} \sigma_{\mu\nu} P_R f) \right\} \\ + \text{h.c.},$$

$$f = e, u, d$$

$$-\mathcal{L}_{\text{GNI}} = \frac{G_F}{\sqrt{2}} \sum_X \bar{\nu} \Gamma^X \nu \bar{q} \Gamma_X \left(C_{\alpha\beta}^{qX} + i \gamma^5 D_{\alpha\beta}^{qX} \right)$$

GNI framework features

- Can describe new interactions in Model-independent framework
- Low-energy Effective Field Theory
- Naturally includes RH neutrinos
- Suitable for low-energy experiments (below the EW scale)
- Not as robust as SMEFT

Interaction with
Electrons

e

CEvNS

COHERENT-CsI
COHERENT-LAr
COHERENT-Ge

DM DD

LUX ZEPLIN
PANDAX-4T
XENONnT

CHARM II

Texono

LEP

DIS

CHARM
CDHS
NuTev

Interactions with
Nuclei

N

to name a few...

DARK MATTER DIRECT DETECTION EXPERIMENTS

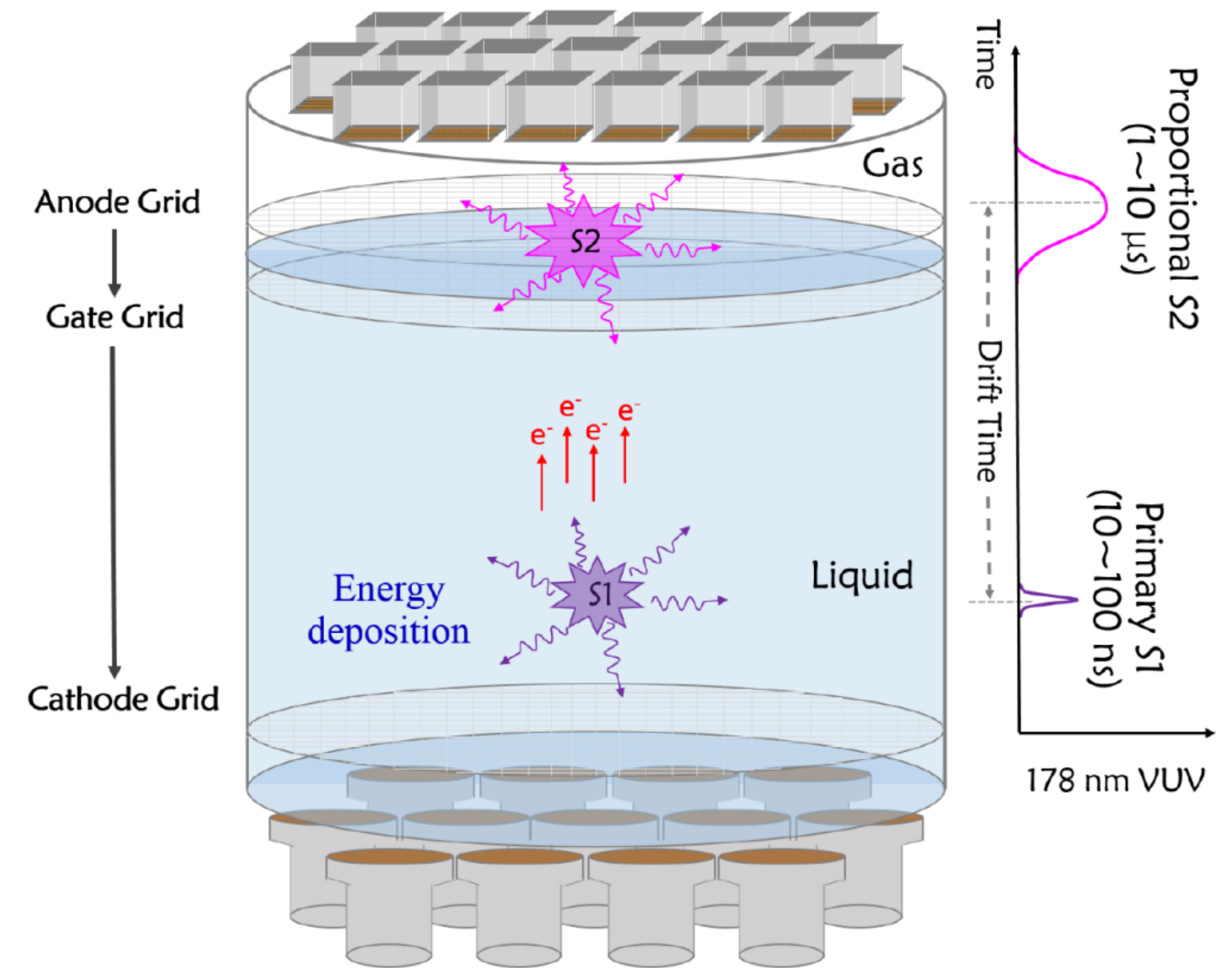


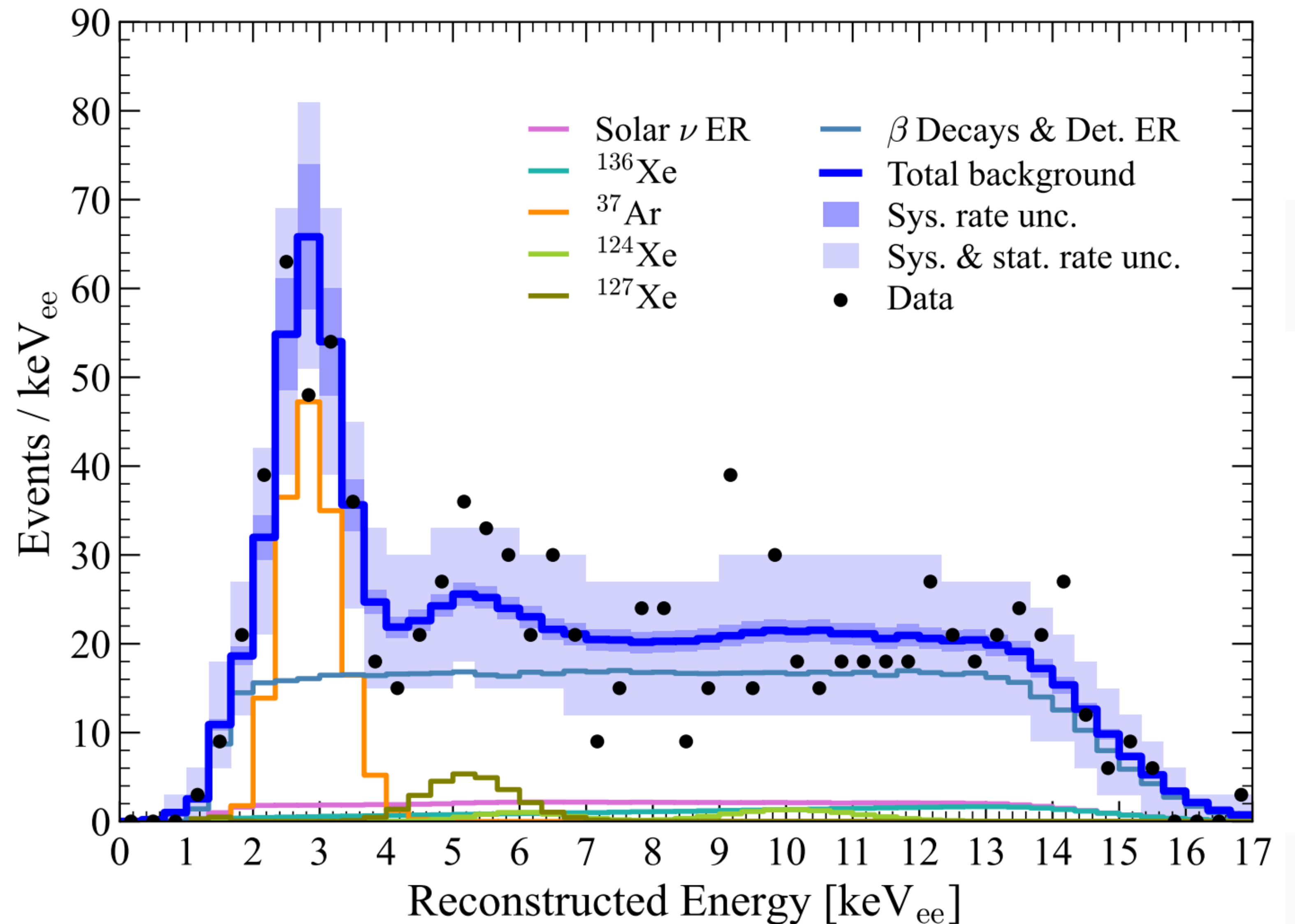
Fig. 1. Working principle of a two-phase TPC.

DARK MATTER DIRECT DETECTION EXPERIMENTS



LUX-ZEPLIN (LZ)
DARK MATTER SEARCH

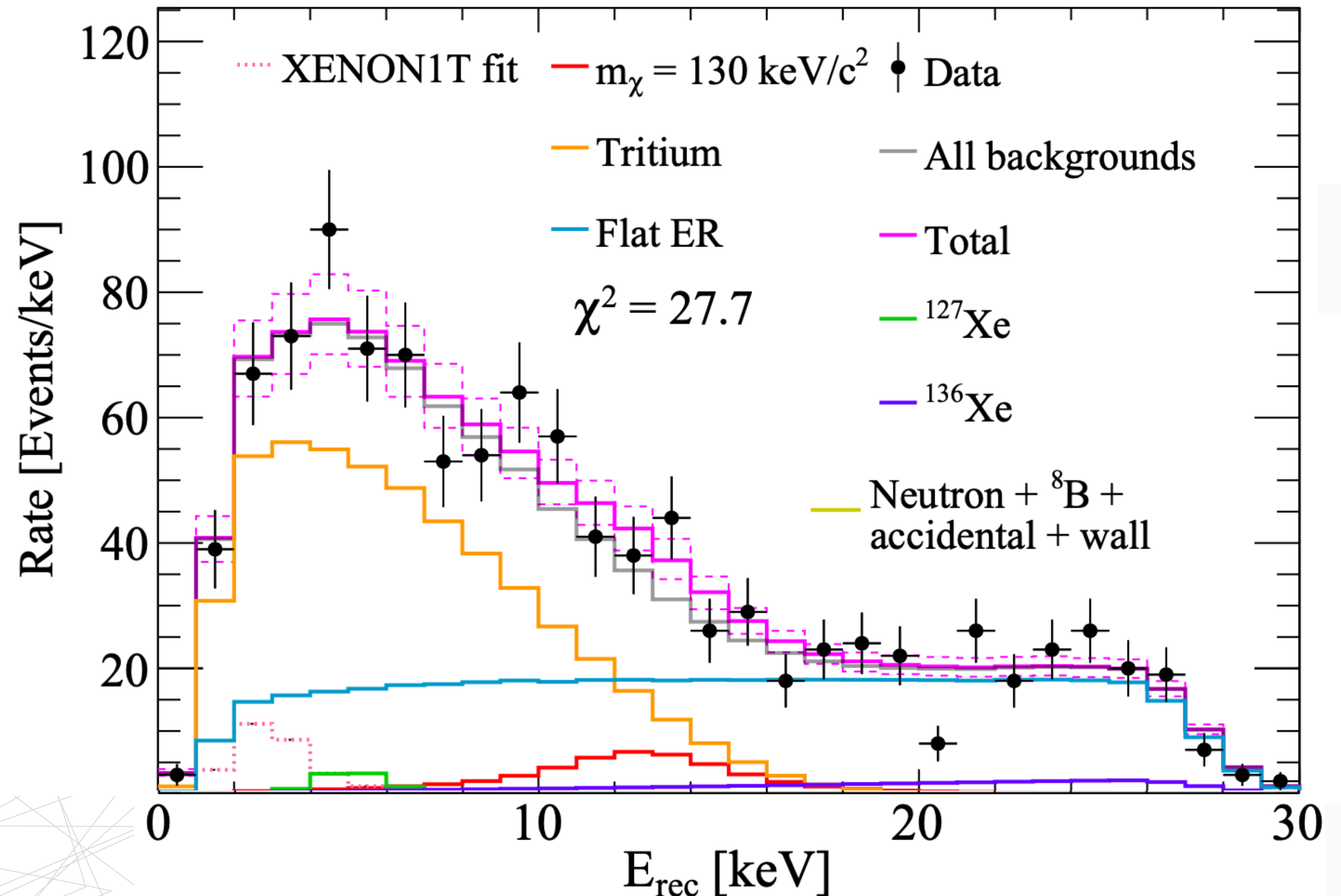
~ 27 solar ν (ER)



DARK MATTER DIRECT DETECTION EXPERIMENTS



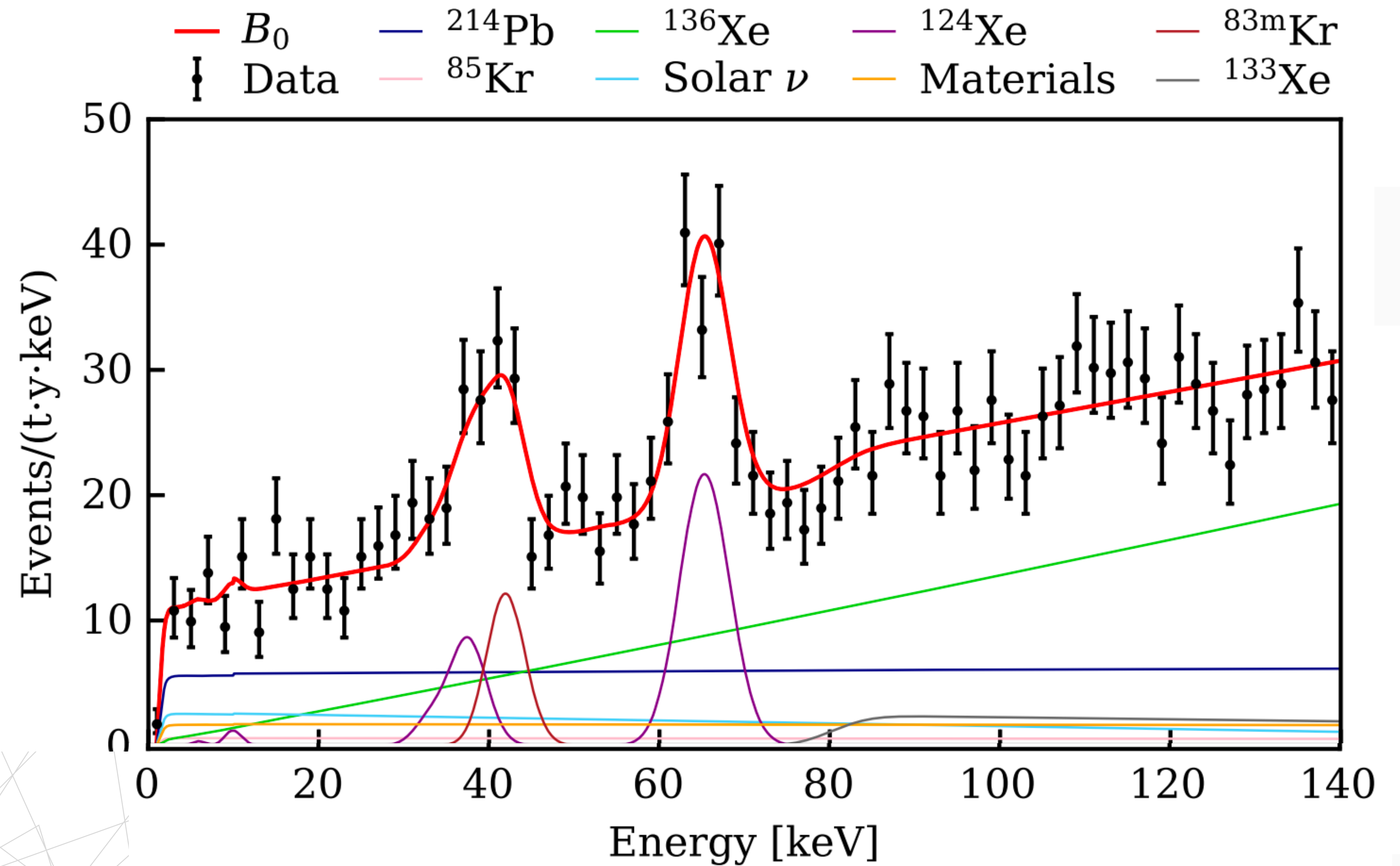
~ 36 solar ν (ER)



DARK MATTER DIRECT DETECTION EXPERIMENTS



~ 300 solar ν (ER)



Neutrino-electron cross section

$$\frac{d\sigma}{dT}(\nu_\alpha + e^- \rightarrow \nu_\beta + e^-) = \frac{G_F^2 m_e}{2\pi} \left[A_{\alpha\beta} + 2B_{\alpha\beta} \left(1 - \frac{T}{E_\nu}\right) + C_{\alpha\beta} \left(1 - \frac{T}{E_\nu}\right)^2 + D_{\alpha\beta} \frac{m_e T}{4E_\nu^2} \right]$$

$$A = \frac{1}{4} |\epsilon^A + \epsilon^V - \tilde{\epsilon}^A - \tilde{\epsilon}^V + 2g^L|^2 + \frac{1}{8} |\epsilon^S + i\tilde{\epsilon}^P|^2 + \frac{1}{8} |\epsilon^P + i\tilde{\epsilon}^S|^2 \\ + |\epsilon^T - i\tilde{\epsilon}^T|^2 + \frac{1}{2} \text{Re} \left[(\epsilon^T - i\tilde{\epsilon}^T)^* (\epsilon^P + i\tilde{\epsilon}^S - \epsilon^S - i\tilde{\epsilon}^P) \right],$$

$$B = -\frac{1}{8} |\epsilon^P + i\tilde{\epsilon}^S|^2 - \frac{1}{8} |\epsilon^S + i\tilde{\epsilon}^P|^2 + |\epsilon^T - i\tilde{\epsilon}^T|^2,$$

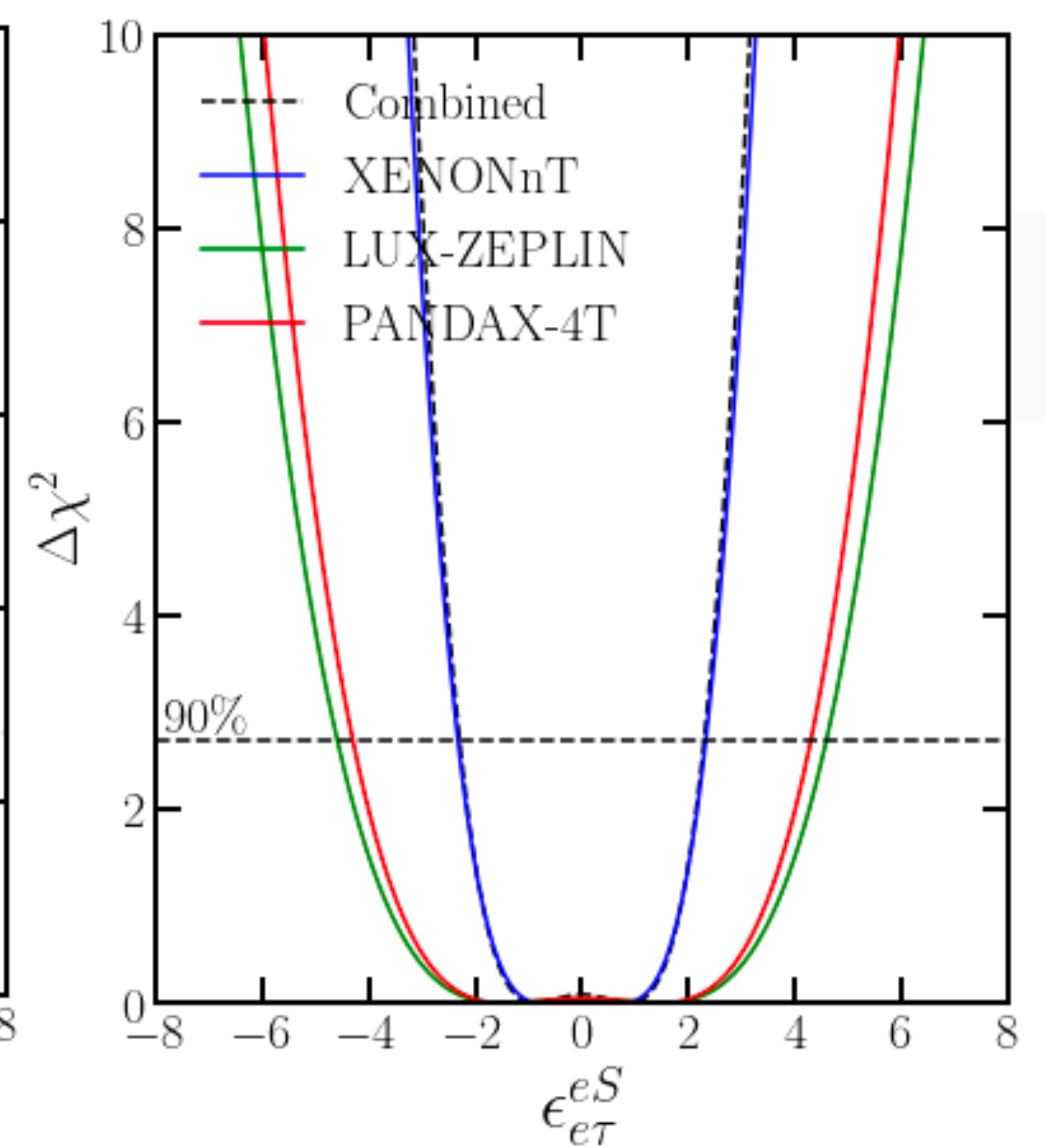
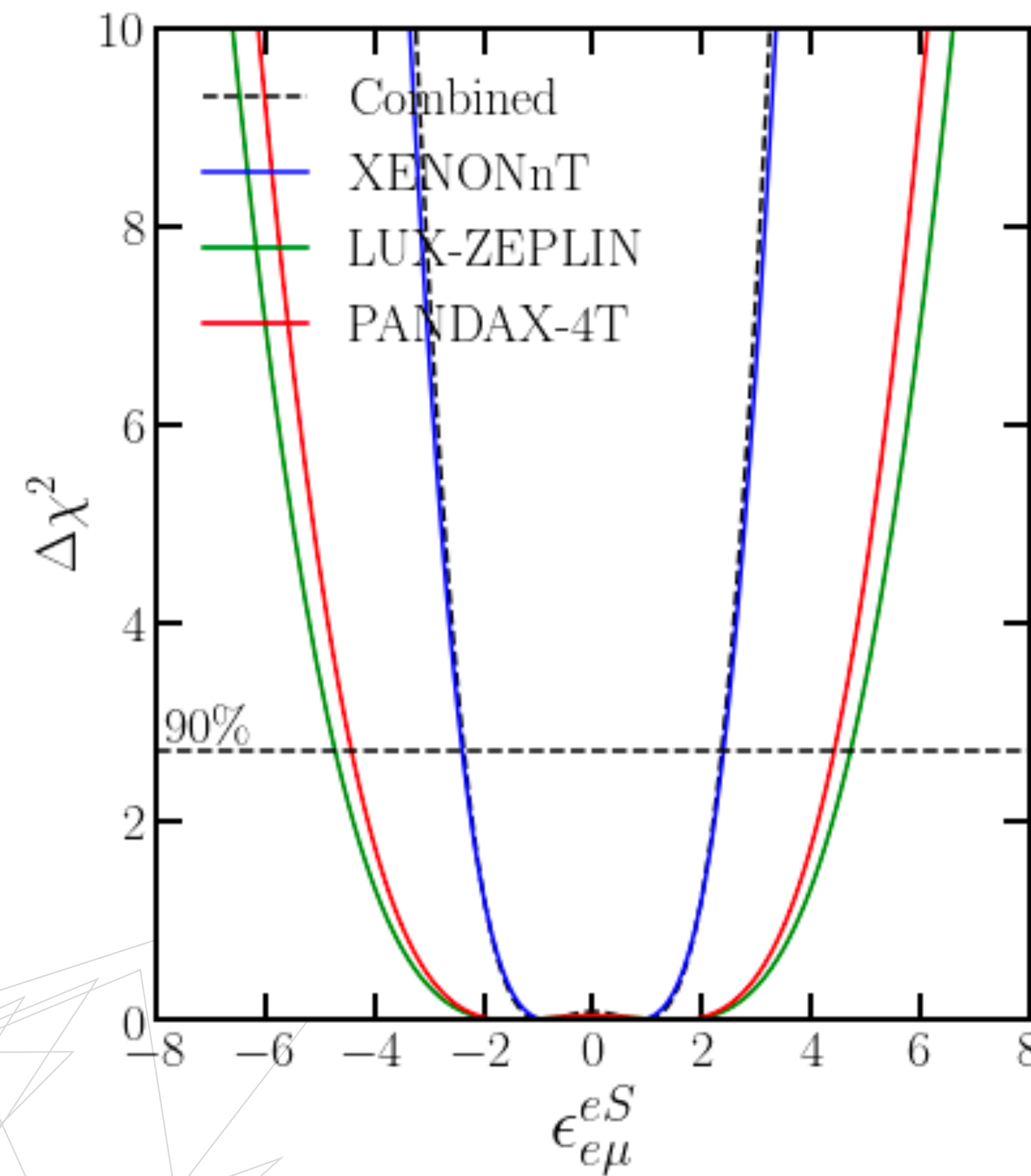
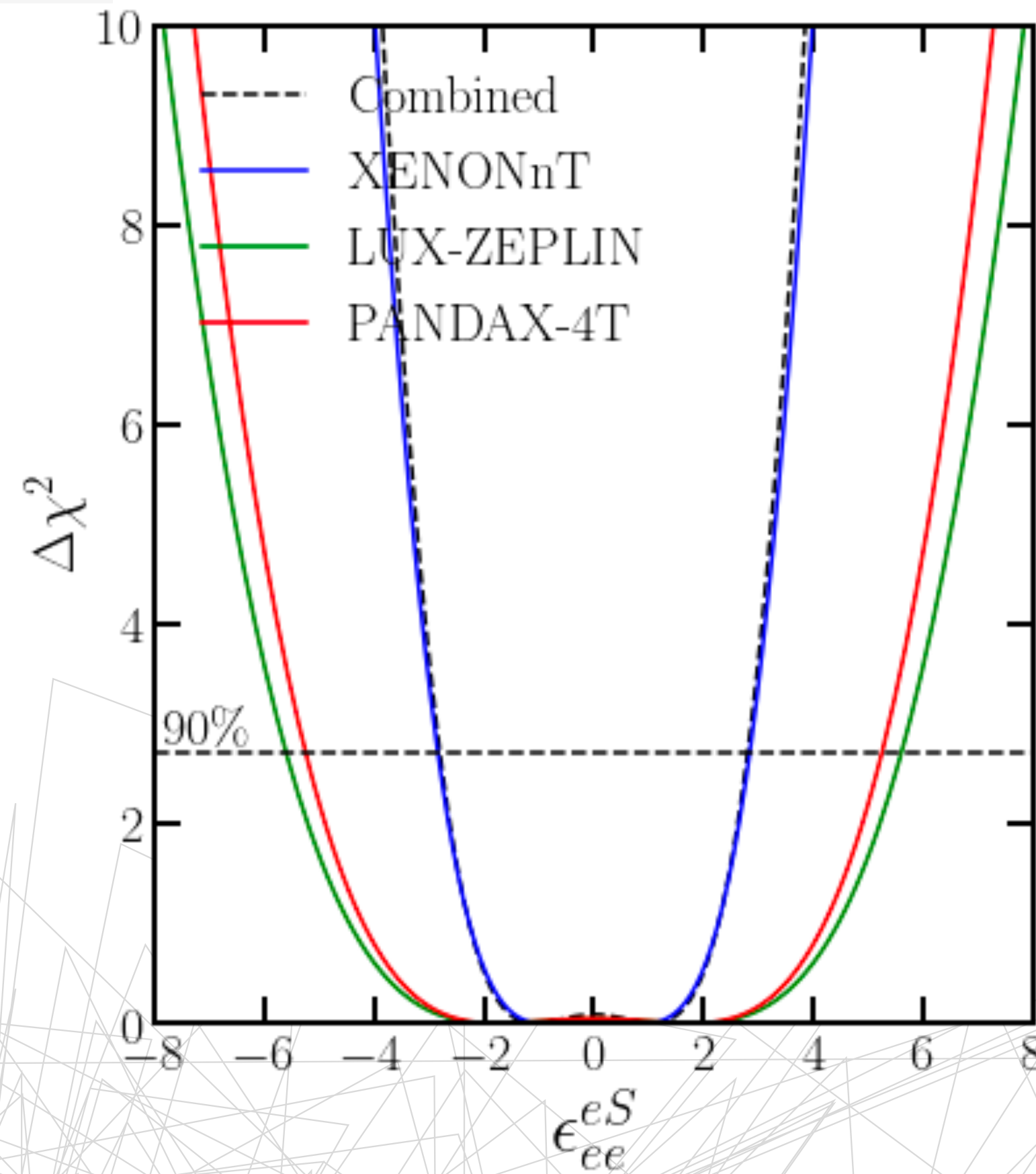
$$C = \frac{1}{4} |\epsilon^A + \tilde{\epsilon}^A - \epsilon^V - \tilde{\epsilon}^V - 2g^R|^2 + \frac{1}{8} |\epsilon^S + i\tilde{\epsilon}^P|^2 + \frac{1}{8} |\epsilon^P + i\tilde{\epsilon}^S|^2 \\ + |\epsilon^T - i\tilde{\epsilon}^T|^2 - \frac{1}{2} \text{Re} \left[(\epsilon^T - i\tilde{\epsilon}^T)^* (\epsilon^P + i\tilde{\epsilon}^S - \epsilon^S - i\tilde{\epsilon}^P) \right],$$

$$D = \text{Re} \left[(\epsilon^A + \epsilon^V - \tilde{\epsilon}^A - \tilde{\epsilon}^V + 2g^L) (\epsilon^A + \tilde{\epsilon}^A - \epsilon^V - \tilde{\epsilon}^V - 2g^R)^* \right] \\ - 4 |\epsilon^T - i\tilde{\epsilon}^T|^2 + |\epsilon^S + i\tilde{\epsilon}^P|^2.$$

RESULTS FOR ϵ^{eS}

In collaboration with:

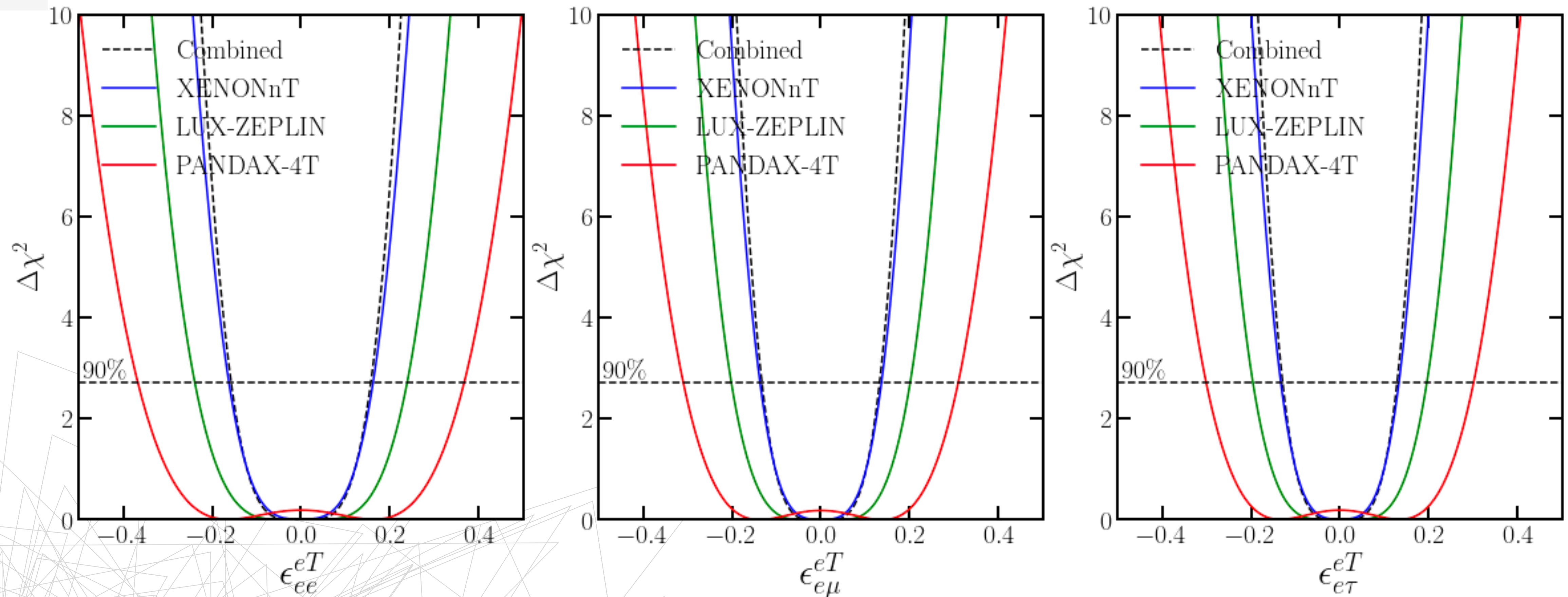
Jesus Celestino, Francisco Escrihuela,
Omar Miranda, Ricardo Sánchez



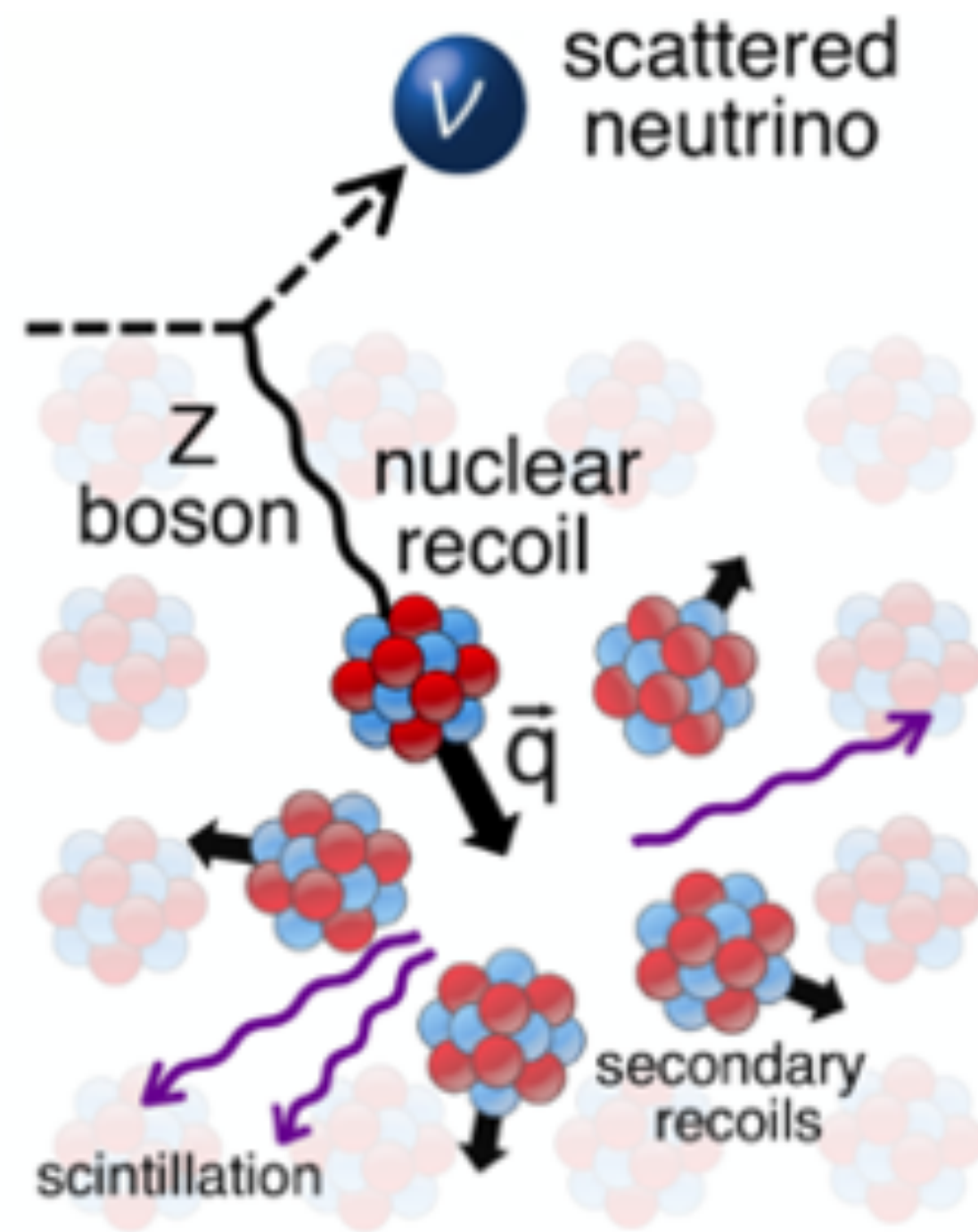
RESULTS FOR ϵ^{eT}

In collaboration with:

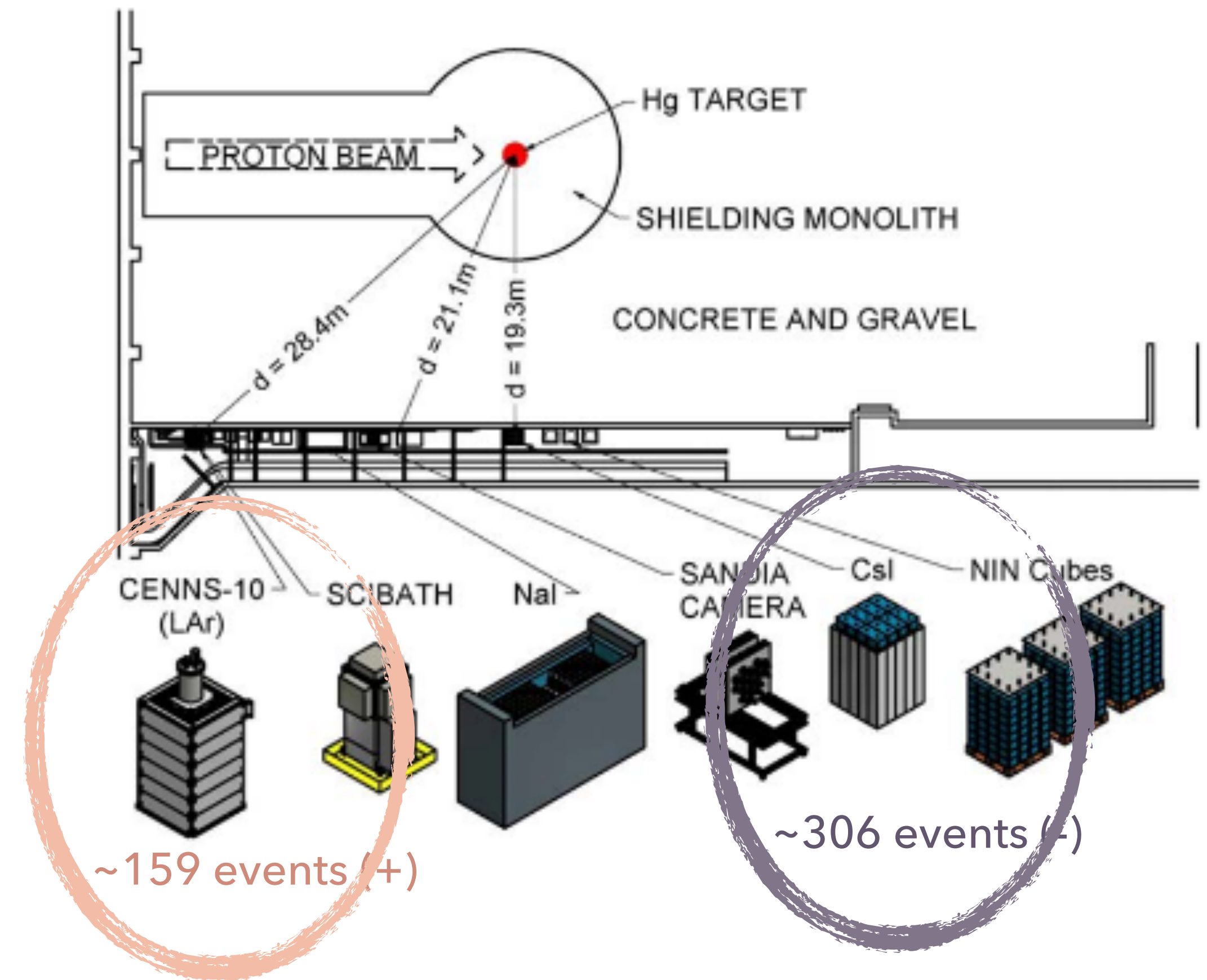
Jesus Celestino, Francisco Escrihuela,
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COHERENT ELASTIC NEUTRINO NUCLEUS SCATTERING EXPERIMENTS

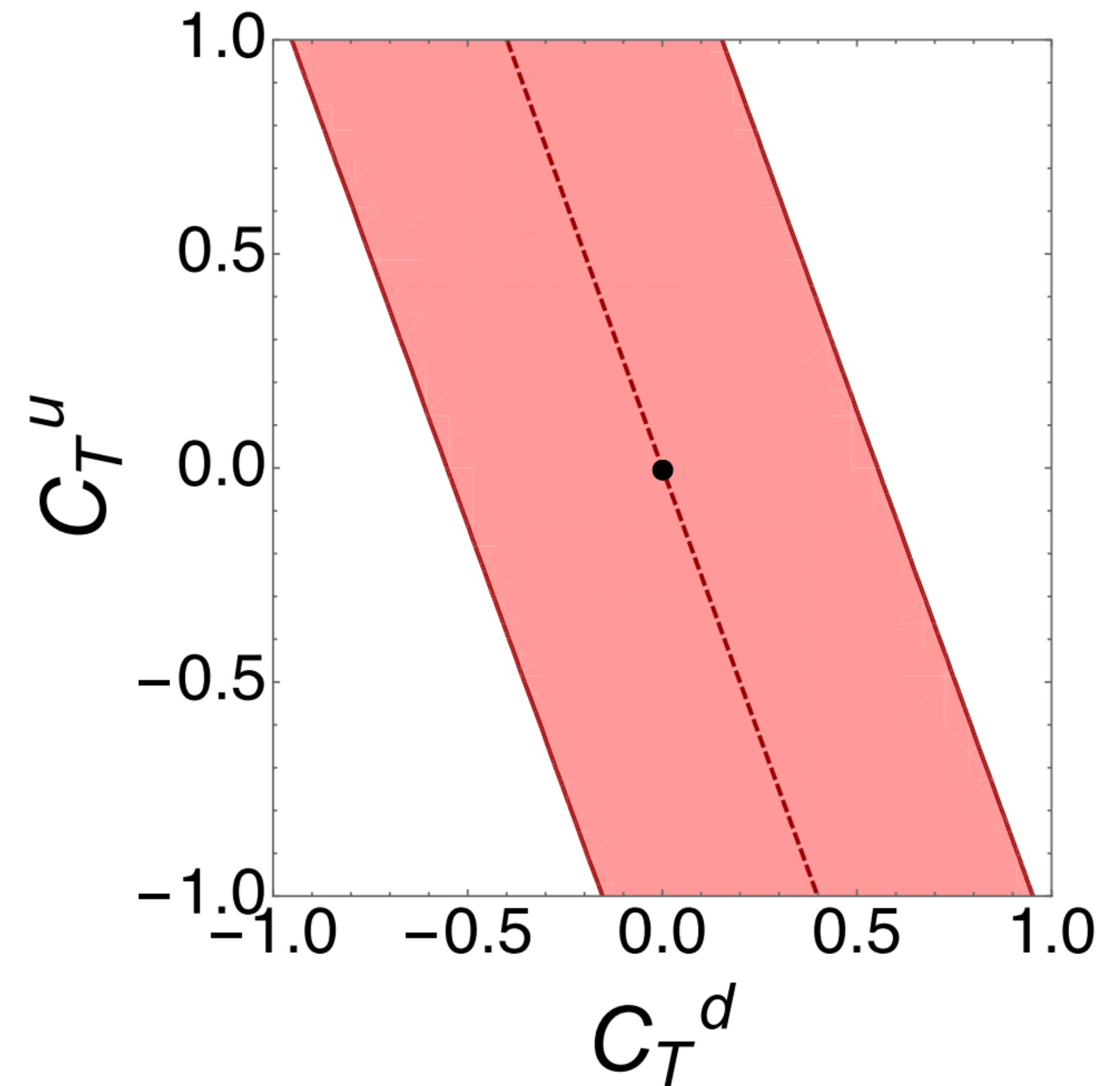
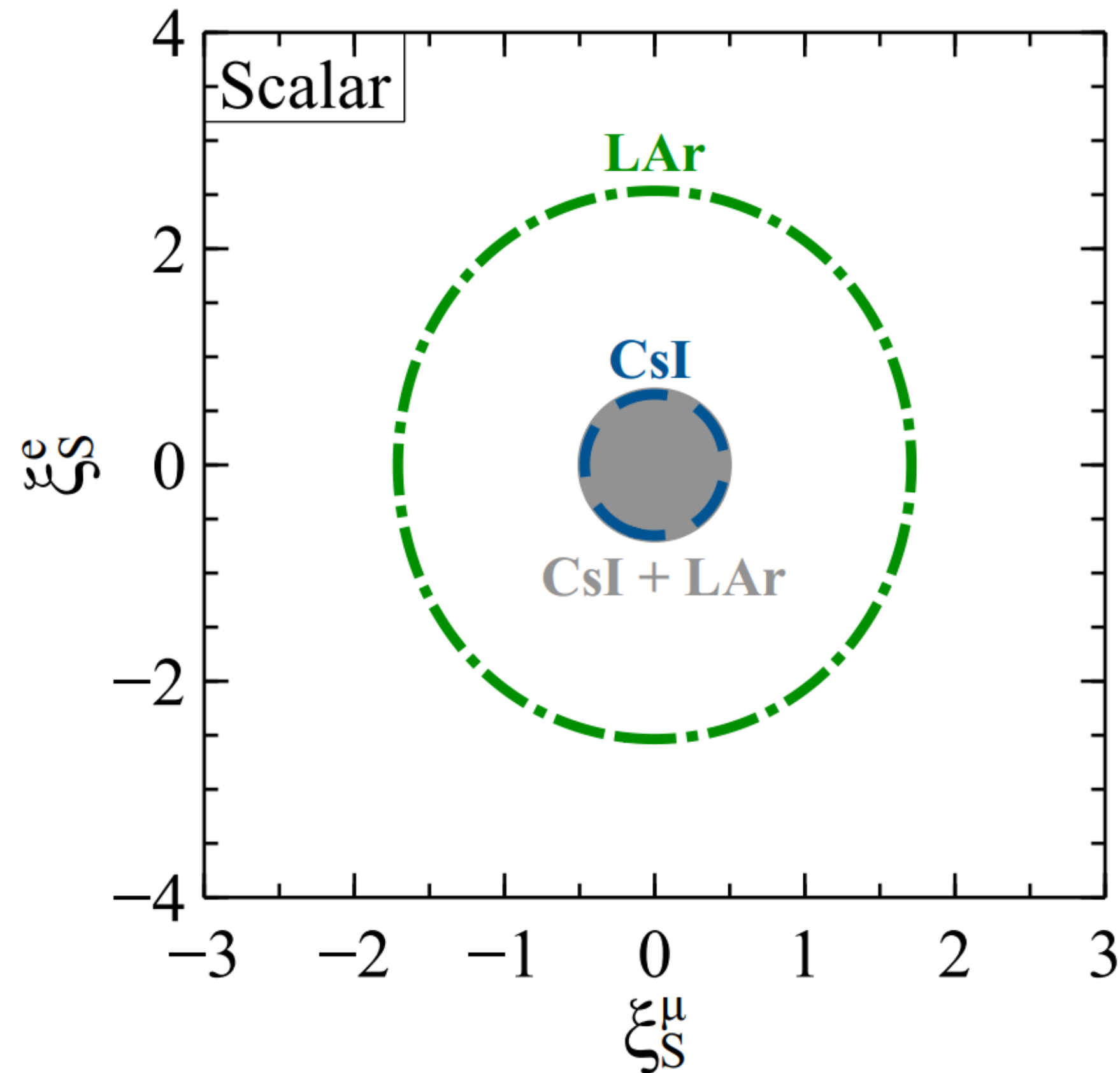


COHERENT SNS



CEvNS cross section

$$\frac{d\sigma}{dT}(E_\nu, T) = \frac{G_F^2 M}{\pi} F^2(|\vec{q}|^2) \sum_{\beta=e, \mu, \tau} \left[\left| \xi_{\alpha\beta}^S \right|^2 \frac{MT}{8E_\nu^2} + \left| \frac{\xi_{\alpha\beta}^V}{2} + Q_{W,\alpha} \delta_{\alpha\beta} \right|^2 \left(1 - \frac{MT}{2E_\nu^2} \right) + 2 \left| \xi_{\alpha\beta}^T \right|^2 \left(1 - \frac{MT}{4E_\nu^2} \right) - \frac{1}{2} R_{\alpha\beta} \frac{T}{E_\nu} \right]$$



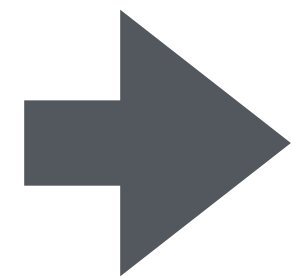
GNI couplings

$$C_{\alpha\beta}^S = \sum_{q=u,d} \left(Z \frac{m_p}{m_q} f_q^p + N \frac{m_n}{m_q} f_q^n \right) C_{\alpha\beta}^{qS} ,$$

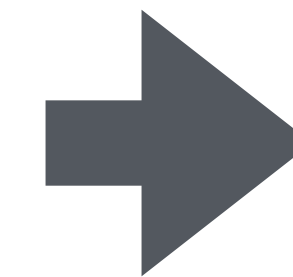
$$\xi_{\alpha\beta}^{S^2} = C_{\alpha\beta}^S{}^2 + D_{\alpha\beta}^P{}^2$$

From Lagrangian

C's and D's



$$C_{\alpha\beta}^V = (2Z + N) C_{\alpha\beta}^{uV} + (Z + 2N) C_{\alpha\beta}^{dV} ,$$



$$\xi_{\alpha\beta}^{V^2} = C_{\alpha\beta}^V{}^2 + D_{\alpha\beta}^A{}^2$$

$$C_{\alpha\beta}^T = \sum_{q=u,d} (Z \delta_q^p + N \delta_q^n) C_{\alpha\beta}^{qT} ,$$

$$\xi_{\alpha\beta}^{T^2} = C_{\alpha\beta}^T{}^2$$

COMMON PARAMETRIZATION

For **Scalar** interactions:

$$\frac{d\sigma}{dT} = \frac{G_F^2 M}{\pi} F^2(|\vec{q}|^2) \left[\left| \sum_{q=u,d} \left(Z \frac{m_p}{m_q} f_q^p + N \frac{m_n}{m_q} f_q^n \right) \left(\epsilon_{\alpha\beta}^{qS} + \tilde{\epsilon}_{\alpha\beta}^{qS} \right) \right|^2 + \left| \sum_{q=u,d} \left(Z \frac{m_p}{m_q} f_q^p + N \frac{m_n}{m_q} f_q^n \right) \left(\epsilon_{\alpha\beta}^{qS} - \tilde{\epsilon}_{\alpha\beta}^{qS} \right) \right|^2 \right] \frac{MT}{8E_\nu^2},$$

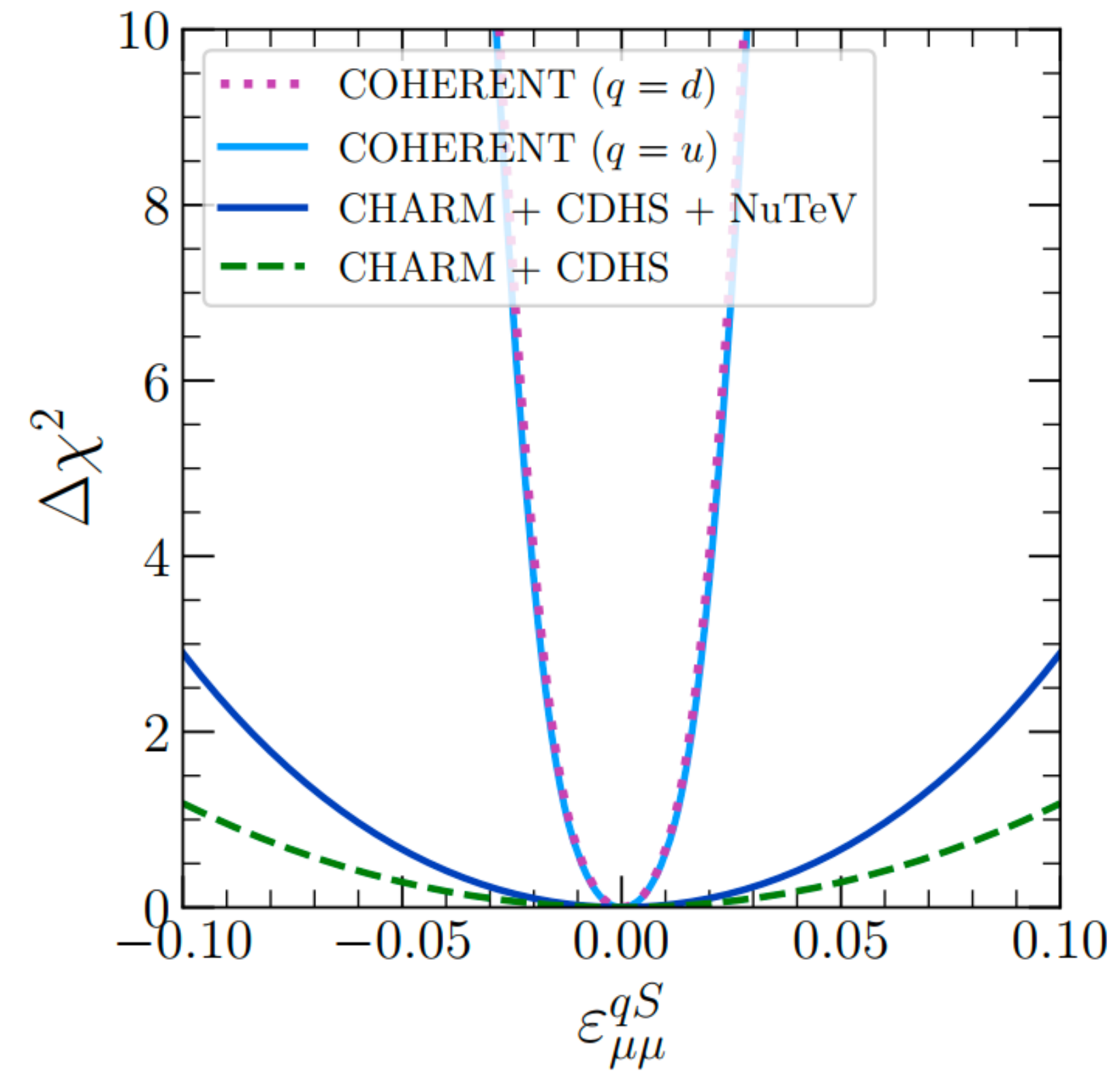
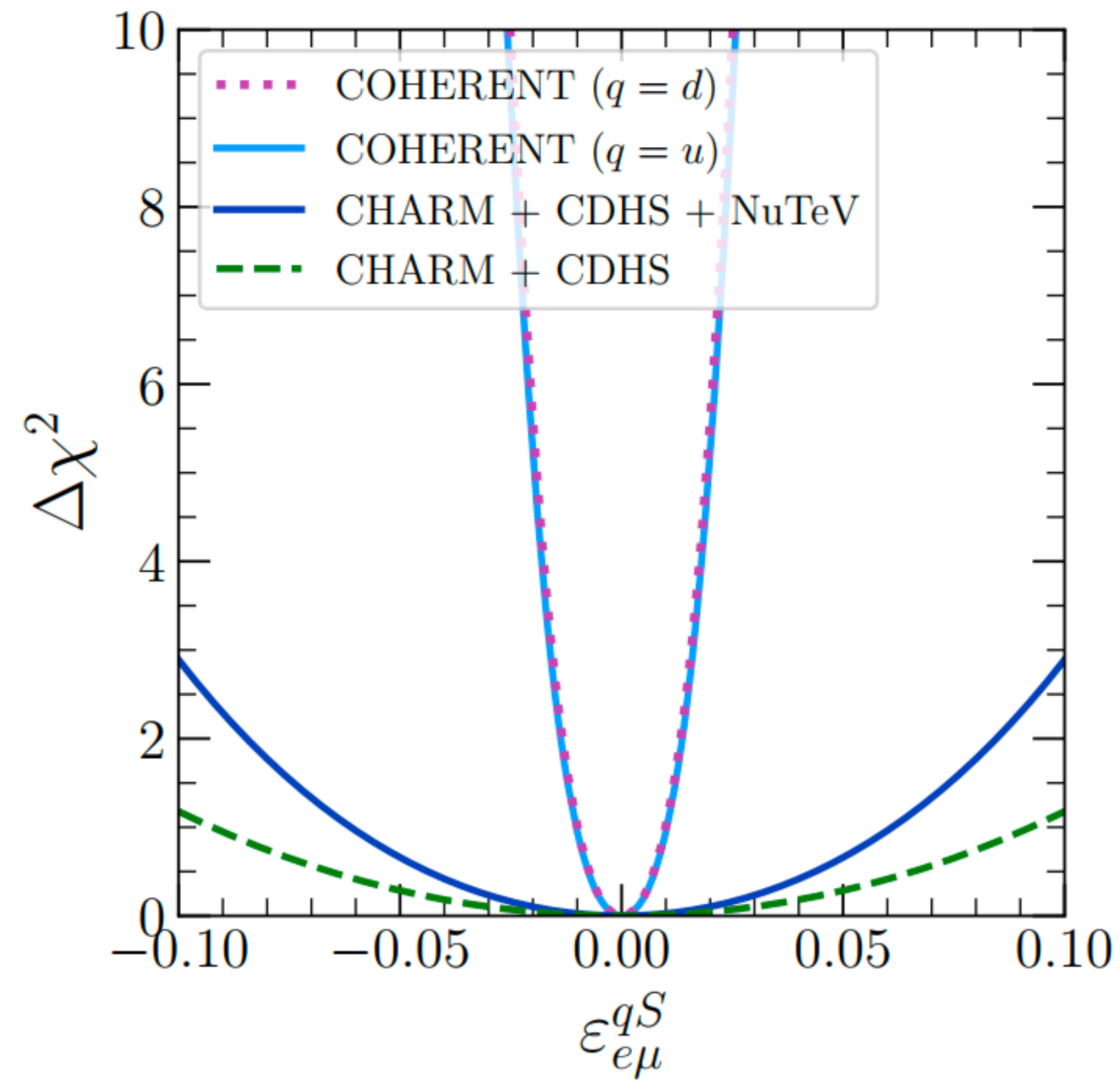
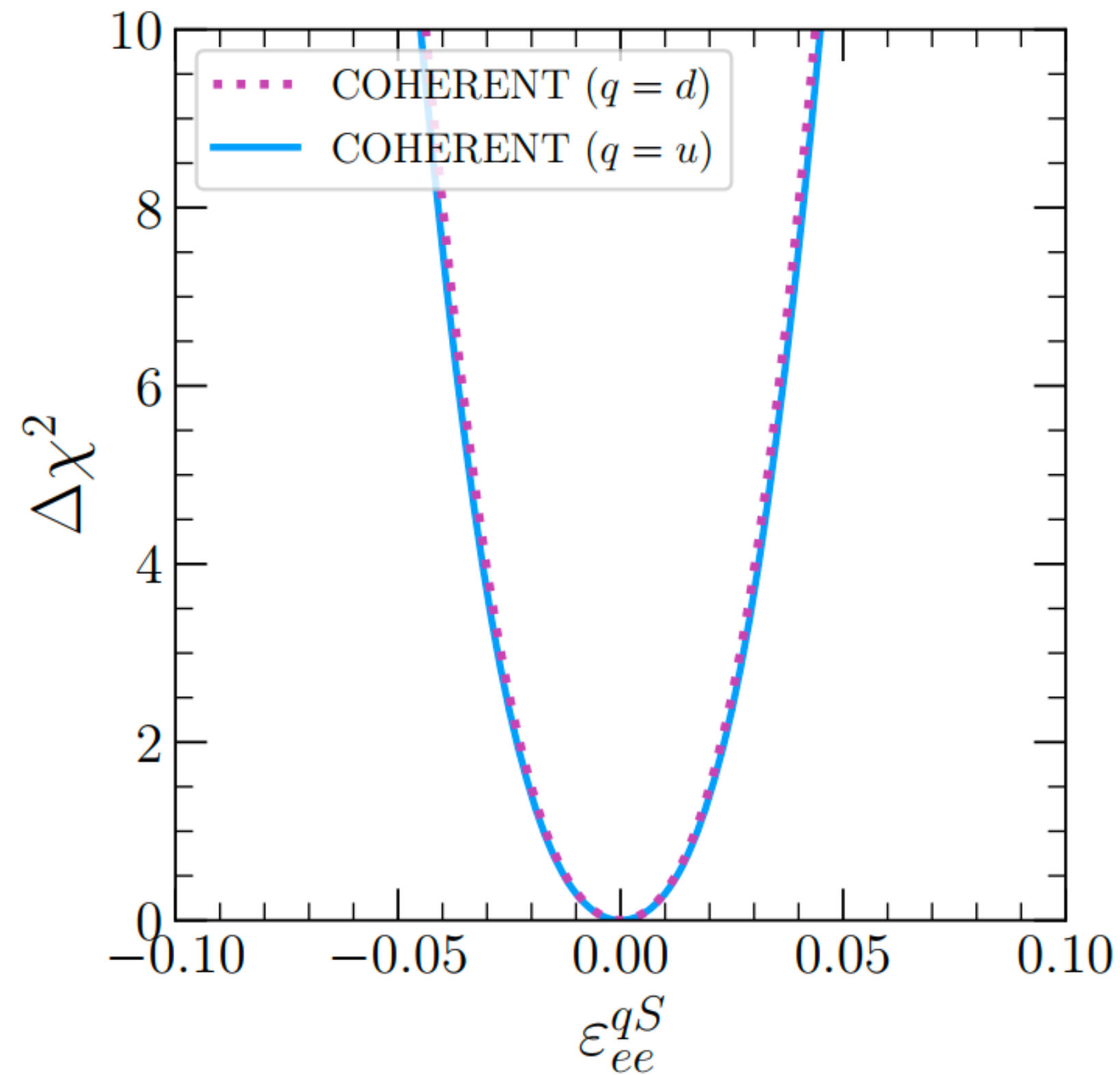
For **Tensor** interactions:

$$\frac{d\sigma}{dT} = \frac{2G_F^2 M}{\pi} F^2(|\vec{q}|^2) \left[4 \left| \sum_{q=u,d} \left(Z \delta_q^p + N \delta_q^n \right) \left(\epsilon_{\alpha\beta}^{qT} + \tilde{\epsilon}_{\alpha\beta}^{qT} \right) \right|^2 + 4 \left| \sum_{q=u,d} \left(Z \delta_q^p + N \delta_q^n \right) \left(\epsilon_{\alpha\beta}^{qT} - \tilde{\epsilon}_{\alpha\beta}^{qT} \right) \right|^2 \right] \left(1 - \frac{MT}{4E_\nu^2} \right),$$

RESULTS FOR ε^{qS}

In collaboration with:

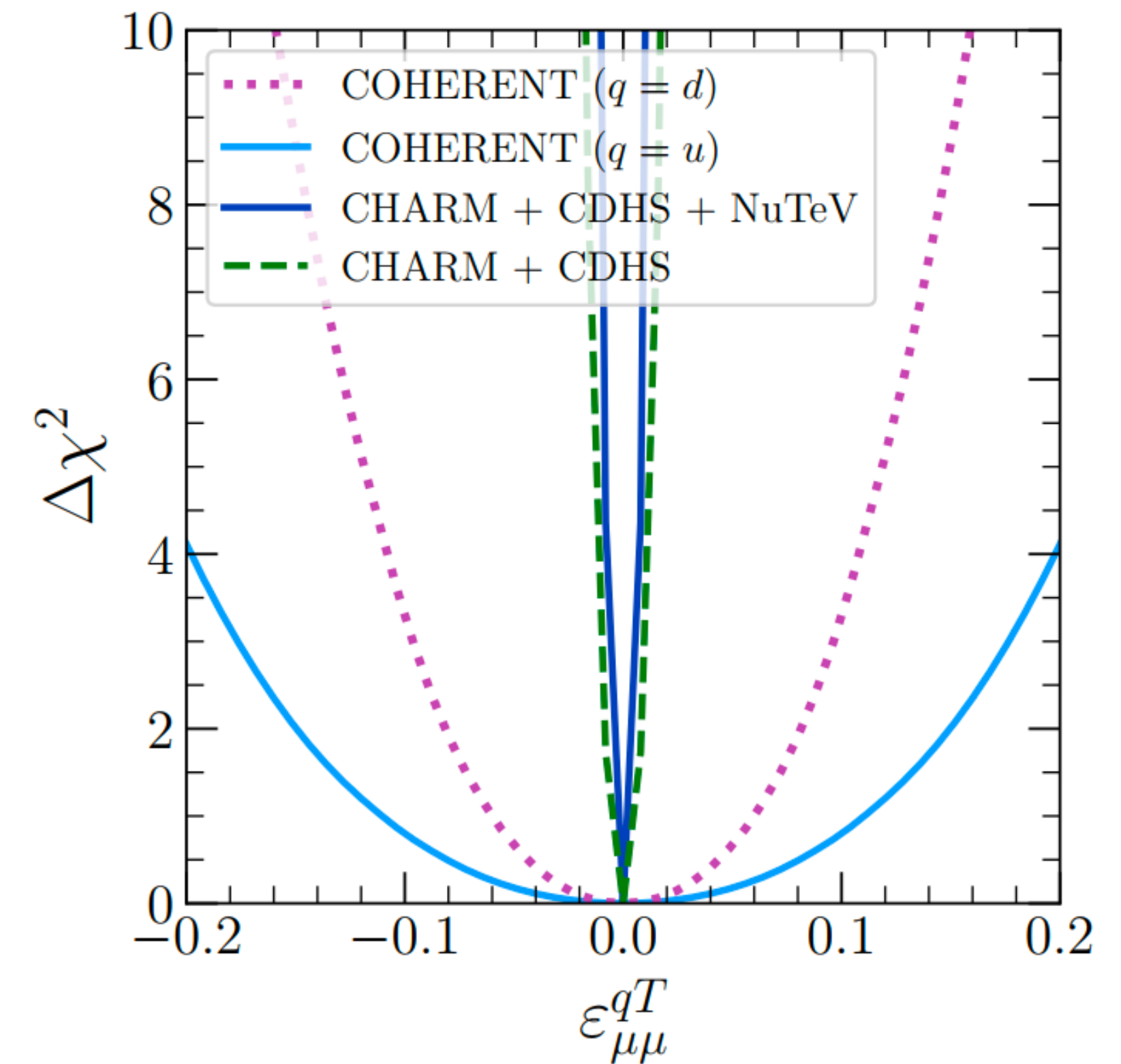
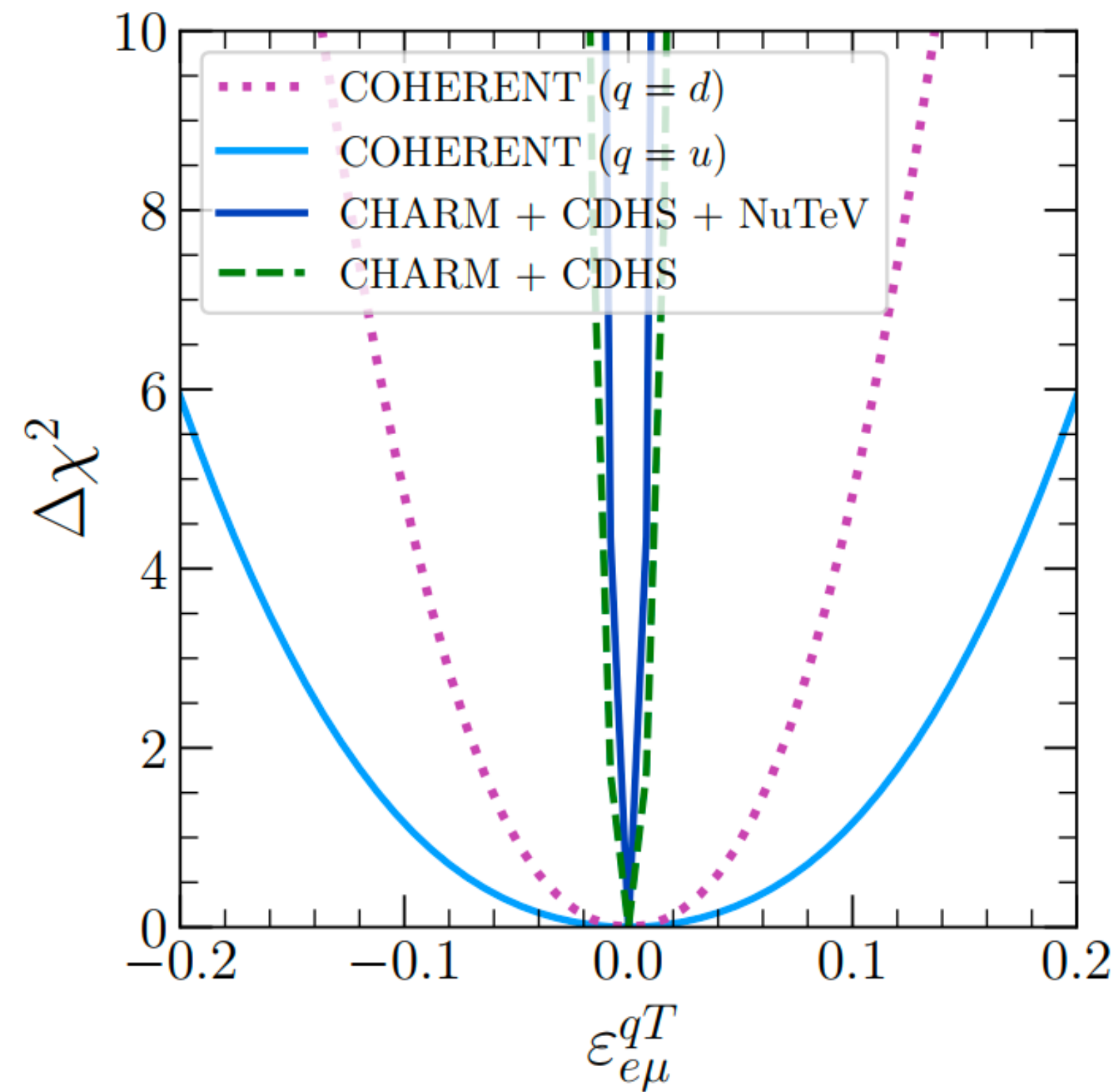
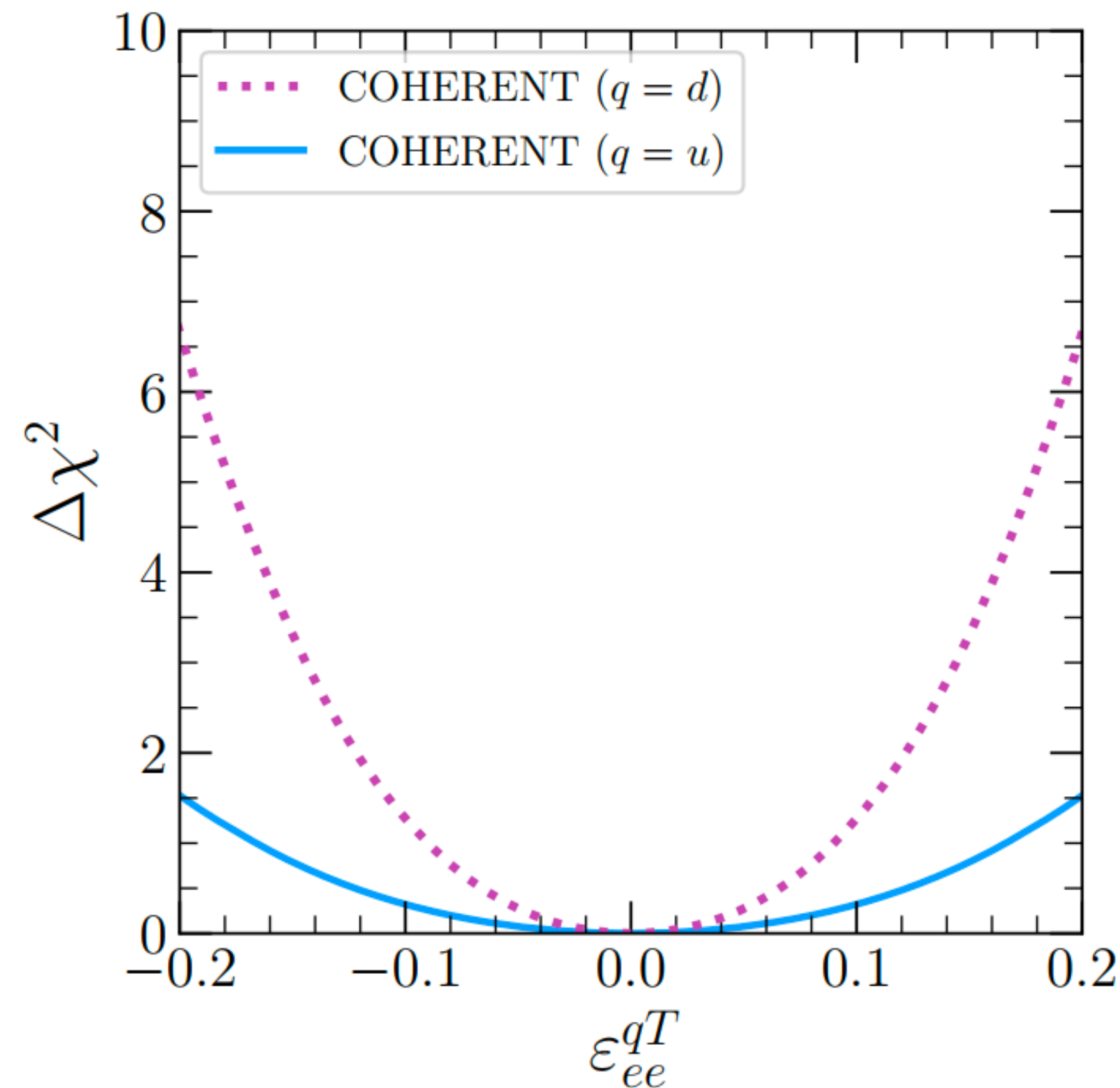
Omar Miranda, Gonzalo Sánchez



RESULTS FOR ε^{qT}

In collaboration with:

Omar Miranda, Gonzalo Sánchez



Comparing with previous limits

Experiments		Scalar	Pseudoscalar	Tensor	
e^-e^+ + TEXONO	$ \varepsilon_{ee}^{eS} < 2.82$	$ \epsilon_{ee}^{e,S} < 0.38$	$ \epsilon_{ee}^{e,P} < 0.40$	$ \varepsilon_{ee}^{eT} < 0.16$	$ \epsilon_{ee}^{e,T} < 0.07$
e^-e^+ + CHARM-II	$ \varepsilon_{\mu\mu}^{eS} < 4.35$	$ \epsilon_{\mu\mu}^{e,X} < 0.31$		$ \varepsilon_{\mu\mu}^{eT} < 0.24$	$ \epsilon_{\mu\mu}^{e,T} < 0.03$
e^-e^+	$ \varepsilon_{\tau\tau}^{eS} < 3.98$	$ \epsilon_{\tau\tau}^{e,X} < 0.40$		$ \varepsilon_{\tau\tau}^{eT} < 0.22$	$ \epsilon_{\tau\tau}^{e,T} < 0.12$
e^-e^+ + TEXONO + CHARM-II	$ \varepsilon_{e\mu}^{eS} < 2.34$	$ \epsilon_{e\mu}^{e,S} < 0.25$	$ \epsilon_{e\mu}^{e,P} < 0.25$	$ \varepsilon_{e\mu}^{eT} < 0.13$	$ \epsilon_{e\mu}^{e,T} < 0.03$
e^-e^+ + TEXONO	$ \varepsilon_{e\tau}^{eS} < 2.28$	$ \epsilon_{e\tau}^{e,S} < 0.28$	$ \epsilon_{e\tau}^{e,P} < 0.29$	$ \varepsilon_{e\tau}^{eT} < 0.13$	$ \epsilon_{e\tau}^{e,T} < 0.07$
e^-e^+ + CHARM-II	$ \varepsilon_{\mu\tau}^{eS} < 2.94$	$ \epsilon_{\mu\tau}^{e,X} < 0.25$		$ \varepsilon_{\mu\tau}^{eT} < 0.16$	$ \epsilon_{\mu\tau}^{e,T} < 0.03$
CHARM- e	$ \varepsilon_{ee}^{qS} < 0.026$	$ \epsilon_{ee}^{q,X} < 1.9$		$ \varepsilon_{ee}^{dT} < 0.13$	$ \epsilon_{ee}^{q,T} < 0.13$
CHARM + CDHS (+ NuTeV)	$ \varepsilon_{\mu\mu}^{qS} < 0.018$	$ \epsilon_{\mu\mu}^{q,X} < 0.15 (0.1)$		$ \varepsilon_{\mu\mu}^{dT} < 0.09$	$ \epsilon_{\mu\mu}^{q,T} < 0.01 (0.006)$
CHARM- e + CHARM + CDHS (+ NuTeV)	$ \varepsilon_{e\mu}^{qS} < 0.015$	$ \epsilon_{e\mu}^{q,X} < 0.15 (0.1)$		$ \varepsilon_{e\mu}^{dT} < 0.08$	$ \epsilon_{e\mu}^{q,T} < 0.01 (0.006)$
CHARM- e		$ \epsilon_{e\tau}^{q,X} < 1.9$			$ \epsilon_{e\tau}^{q,T} < 0.13$
CHARM + CDHS (+ NuTeV)		$ \epsilon_{\mu\tau}^{q,X} < 0.15 (0.1)$			$ \epsilon_{\mu\tau}^{q,T} < 0.01 (0.006)$

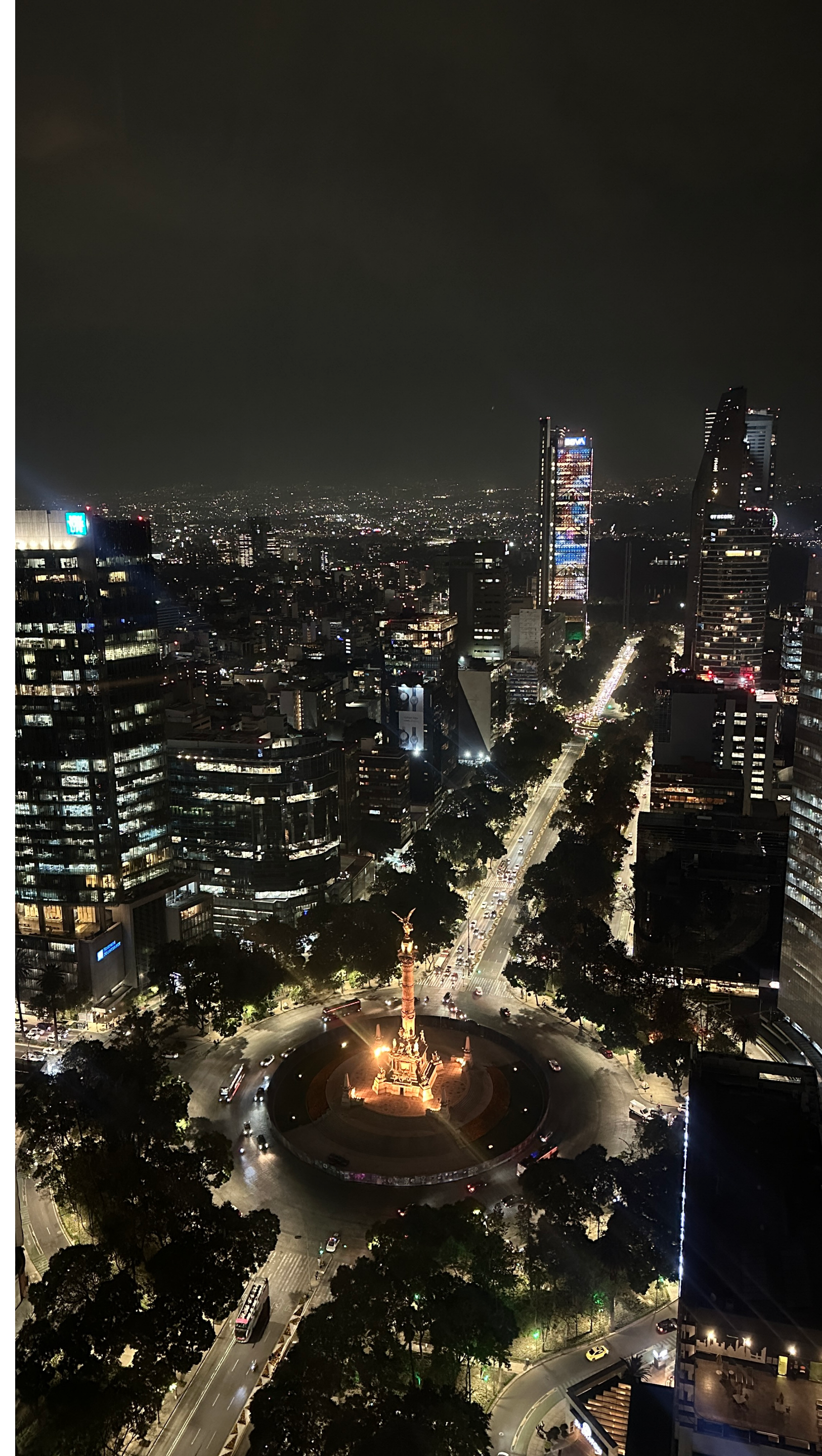
SUMMARY

- So far, no evidence of nonzero GNI couplings
- Dark matter experiments :
 - can improve current TENSOR GNI with electron recoils
 - will soon provide insights of GNI with nuclear recoils
- CEvNS experiments are better at constraining SCALAR GNI

Thank you



XV SILAF AE



BACKUP

$$\mathcal{L} = -\frac{G_F}{\sqrt{2}} \sum_{\alpha,\beta} \sum_{j=1}^{10} \left(\begin{matrix} (\sim)j \\ \epsilon \end{matrix} \right)_{\alpha\beta} (\bar{\nu}^\alpha \mathcal{O}_j \nu^\beta) (\bar{e} \mathcal{O}'_j e) ,$$

j	$\begin{matrix} (\sim) \\ \epsilon_j \end{matrix}$	$\mathcal{O}_j$	$\mathcal{O}'_j$
1	ϵ_L	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
2	$\tilde{\epsilon}_L$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 - \gamma^5)$
3	ϵ_R	$\gamma_\mu(1 - \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
4	$\tilde{\epsilon}_R$	$\gamma_\mu(1 + \gamma^5)$	$\gamma^\mu(1 + \gamma^5)$
5	ϵ_S	$(1 - \gamma^5)$	1
6	$\tilde{\epsilon}_S$	$(1 + \gamma^5)$	1
7	$-\epsilon_P$	$(1 - \gamma^5)$	γ^5
8	$-\tilde{\epsilon}_P$	$(1 + \gamma^5)$	γ^5
9	ϵ_T	$\sigma_{\mu\nu}(1 - \gamma^5)$	$\sigma^{\mu\nu}(1 - \gamma^5)$
10	$\tilde{\epsilon}_T$	$\sigma_{\mu\nu}(1 + \gamma^5)$	$\sigma^{\mu\nu}(1 + \gamma^5)$

RESULTS FOR ϵ^{eP}

In collaboration with:

Jesus Celestino, Francisco Escrihuela,
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