One-loop contributions to *WWZ* from a seesaw variant with radiatively induced light neutrino masses

H. Novales, Mónica Salinas, H. Vázquez,





Calculation, estimation and analysis of the one loop contributions from Majorana neutrinos

> Lorentz-covariant WWZ parametrization

#### BSM description characterized by $\mathcal{L}_{BSM}$ [1]

#### Two phases of spontaneous symmetry breaking

#### Symmetry breaking at some highenergy scale w

Brout-Englert-Higgs mechanism  $SU(2)_L \otimes U(1)_Y \rightarrow U(1)_e$ 

at v = 246 GeV

[1] Pilaftsis, A. (1992). Radiatively induced neutrino masses and large Higgs-neutrino couplings in the Standard Model with Majorana fields. Zeitschrift für Physik C Particles and Fields, **55**(2), 275-282.

 $\mathcal{L}_{\rm BSM} = \mathcal{L}_{\rm mass}^{\nu} + \mathcal{L}_{\rm CC}^{\rm SM} + \mathcal{L}_{\rm NC}^{\rm SM} + \mathcal{L}_{\rm HC}^{\rm SM} + \cdots$ 

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

$$\mathcal{L}_{\text{mass}}^{\nu} = -\sum_{j=1}^{3} \sum_{k=1}^{3} \left( \overline{\nu_{j,L}^{0}}(m_{\text{D}})_{jk} \,\nu_{k,R}^{0} + \frac{1}{2} \overline{\nu_{j,R}^{0\,\text{c}}}(m_{\text{M}})_{jk} \,\nu_{k,R}^{0} \right) + \text{H.c.}$$
(1)

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

 $\mathcal{L}_{\text{mass}}^{\nu} = -\sum_{k=1}^{3} \sum_{k=1}^{3} \left( \overline{\nu_{j,L}^{0}}(m_{\text{D}})_{jk} \nu_{k,R}^{0} + \frac{1}{2} \overline{\nu_{j,R}^{0 \text{ c}}}(m_{\text{M}})_{jk} \nu_{k,R}^{0} \right) + \text{H.c.}$ (1) i = 1 k = 1

 $\nu_k^{\rm c} = C \overline{\nu}_k^{\rm T}$ 

4

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

 $\mathcal{L}_{\text{mass}}^{\nu} = -\sum_{k=1}^{3} \sum_{k=1}^{3} \left( \overline{\nu_{j,L}^{0}(m_{\text{D}})}_{jk} \nu_{k,R}^{0} + \frac{1}{2} \overline{\nu_{j,R}^{0 \text{ c}}(m_{\text{M}})}_{jk} \nu_{k,R}^{0} \right) + \text{H.c.}$ (1)  $i = 1 \ k = 1$ 

3 × 3 matrix assumed to emerge from EW symmetry breaking

3 × 3 matrix assumed to emerge from symmetry breaking at w

4

Defining

$$f_{L} = \begin{pmatrix} \nu_{1,L}^{0} \\ \nu_{2,L}^{0} \\ \nu_{3,L}^{0} \end{pmatrix}$$

$$F_{L} = \begin{pmatrix} \nu_{1,R}^{0 \, c} \\ \nu_{2,R}^{0 \, c} \\ \nu_{3,R}^{0 \, c} \end{pmatrix}$$

 $f_R = C\overline{f_L}^{\mathrm{T}}$ 

 $F_R = C\overline{F_L}^{\mathrm{T}}$ 

$$\Rightarrow \mathcal{L}_{\text{mass}}^{\nu} = -\frac{1}{2} (\overline{f_L} \ \overline{F_L}) \mathcal{M} \left( \begin{array}{c} f_R \\ F_R \end{array} \right) + \text{H.c.}$$
(2)

Defining

 $f_{L} = \begin{pmatrix} \nu_{1,L}^{0} \\ \nu_{2,L}^{0} \\ \nu_{3,L}^{0} \end{pmatrix}$ 

 $F_L = \begin{pmatrix} \nu_{1,R}^{0 \text{ c}} \\ \nu_{2,R}^{0 \text{ c}} \\ \nu_{3,R}^{0 \text{ c}} \end{pmatrix}$ 

 $f_R = C\overline{f_L}^{\mathrm{T}}$ 

 $F_R = C\overline{F_L}^{\mathrm{T}}$ 

symmetric

Unitary diagonalization transformation

Define the mass eigenspinors base

Assume  $6 \times 6$  unitary matrix

 $\mathcal{U}_{
u} = egin{pmatrix} \mathcal{U}_{11} & \mathcal{U}_{12} \ \mathcal{U}_{21} & \mathcal{U}_{22} \end{pmatrix}$ 

 $3 \times 3$  matrix blocks  $\mathcal{U}_{kj}$ 

Unitary transformation

$$\mathcal{U}_{
u}^{ ext{T}}\mathcal{M} \; \mathcal{U}_{
u} = \left( egin{array}{cc} m_{
u} & 0 \ 0 & m_N \end{array} 
ight)$$

 $3 \times 3$  real-valued diagonal matrix

$$(m_{
u})_{jk} = m_{
u_j} \delta_{jk} \ (m_N)_{jk} = m_{N_j} \delta_{jk}$$

6

Unitary diagonalization transformation

Define the mass eigenspinors base

$$\mathcal{L}_{\text{mass}}^{\nu} = \sum_{j=1}^{3} \left( -\frac{1}{2} m_{\nu_j} \overline{\nu_j} \nu_j - \frac{1}{2} m_{N_j} \overline{N_j} N_j \right) \tag{3}$$

#### Majorana fields

$$u_j^{
m c} = 
u_j \qquad N_j^{
m c} = N_j$$

Light neutrinos

Heavy neutrinos

6

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

 $\mathcal{L}_{\rm CC}^{\rm SM} = \sum_{\alpha} \sum_{j=1}^{9} \left( \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha\nu_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L \nu_j + \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha N_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L N_j \right) + \text{H.c.}$ (4)

$$\mathcal{B}_{\alpha\nu_j} = \sum_{k=1}^3 V_{\alpha k}^{\ell} (\mathcal{U}_{11}^*)_{kj}$$
$$\mathcal{B}_{\alpha N_j} = \sum_{k=1}^3 V_{\alpha k}^{\ell} (\mathcal{U}_{12}^*)_{kj}$$

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

 $\mathcal{L}_{\rm CC}^{\rm SM} = \sum_{\alpha} \sum_{j=1}^{9} \left( \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha\nu_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L \nu_j + \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha N_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L N_j \right) + \text{H.c.}$ (4)

$$\mathcal{B}_{\alpha\nu_j} = \sum_{k=1}^3 V_{\alpha k}^{\ell} (\mathcal{U}_{11}^*)_{kj}$$
$$\mathcal{B}_{\alpha N_j} = \sum_{k=1}^3 V_{\alpha k}^{\ell} (\mathcal{U}_{12}^*)_{kj}$$



3 × 3 matrix Lepton mixing

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

0

$$\mathcal{L}_{\rm CC}^{\rm SM} = \sum_{\alpha} \sum_{j=1}^{3} \left( \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha\nu_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L \nu_j + \frac{g}{\sqrt{2}} \mathcal{B}_{\alpha N_j} W_{\rho}^{-} \overline{l_{\alpha}} \gamma^{\rho} P_L N_j \right) + \text{H.c.}$$
(4)

Unitary block parametrization [2,3]

7

[2] J. G. Korner, A. Pilaftsis, and K. Schilcher, Leptonic CP asymmetries in flavor-changing H0 decays, Phys. Rev. D 47, 1080 (1993).
 [3] P. S. B. Dev and A. Pilaftsis, Minimal radiative neutrino mass mechanism for inverse seesaw models, Phys. Rev. D 86, 113001 (2012).

 $\mathcal{L}_{\text{BSM}} = \mathcal{L}_{\text{mass}}^{\nu} + \mathcal{L}_{\text{CC}}^{\text{SM}} + \mathcal{L}_{\text{NC}}^{\text{SM}} + \mathcal{L}_{\text{HC}}^{\text{SM}} + \cdots$ 

 $\mathcal{L}_{\mathrm{NC}}^{\mathrm{SM}} = \sum_{k=1}^{3} \sum_{j=1}^{3} \frac{g}{4c_{\mathrm{W}}} \Big( Z_{\rho} \overline{\nu_{k}} \gamma^{\rho} \big( i \mathcal{C}_{\nu_{k} \nu_{j}}^{\mathrm{Im}} - \mathcal{C}_{\nu_{k} \nu_{j}}^{\mathrm{Re}} \gamma_{5} \big) \nu_{j}$ + $\left(Z_{\rho}\overline{\nu_{k}}\gamma^{\rho}\left(i\mathcal{C}_{\nu_{k}N_{j}}^{\mathrm{Im}}-\mathcal{C}_{\nu_{k}N_{j}}^{\mathrm{Re}}\gamma_{5}\right)N_{j}+\mathrm{H.c.}\right)$  $+Z_{\rho}\overline{N_{k}}\gamma^{\rho}\left(i\mathcal{C}_{N_{k}N_{j}}^{\mathrm{Im}}-\mathcal{C}_{N_{k}N_{j}}^{\mathrm{Re}}\gamma_{5}\right)N_{j}\right)$ 

8

(6)

 $\mathcal{L}_{\rm BSM} = \mathcal{L}_{\rm mass}^{\nu} + \mathcal{L}_{\rm CC}^{\rm SM} + \mathcal{L}_{\rm NC}^{\rm SM} + \mathcal{L}_{\rm HC}^{\rm SM} + \cdots$ 

$$\mathcal{L}_{\rm HC}^{\rm SM} = \sum_{k=1}^{3} \sum_{j=1}^{3} \frac{g}{4m_W} \Big( h \overline{\nu_k} \big( (m_{\nu_k} + m_{\nu_j}) \mathcal{C}_{\nu_k \nu_j}^{\rm Re} - i \gamma_5 (m_{\nu_k} - m_{\nu_j}) \mathcal{C}_{\nu_k \nu_j}^{\rm Im} \big) \nu_j \\ + h \overline{\nu_k} \big( (m_{\nu_k} + m_{N_j}) \mathcal{C}_{\nu_k N_j}^{\rm Re} - i \gamma_5 (m_{\nu_k} - m_{N_j}) \mathcal{C}_{\nu_k N_j}^{\rm Im} \big) N_j \\ + h \overline{N_k} \big( (m_{N_k} + m_{N_j}) \mathcal{C}_{N_k N_j}^{\rm Re} - i \gamma_5 (m_{N_k} - m_{N_j}) \mathcal{C}_{N_k N_j}^{\rm Im} \big) N_j \Big)$$

(8)

Consider the neutrino mass model [1]



$$(\mathcal{M}\mathcal{U}_{\nu})_{jk} = 0$$
  $j = 1, 2, 3, 4, 5, 6$ 

(9)

11

#### the k - th neutrino mass vanishes at tree level.

Heavy neutrinos:

Tree-level mass are left untouched

$$m_N \simeq m_{\rm M} \left( 1 + \frac{1}{2} m_{\rm M}^{-1} (\xi^{\dagger} m_{\rm D} + m_{\rm D}^{\rm T} \xi^*) \right)$$

Smallness of light-neutrino masses comes about if heavy neutrino masses are nearly-degenerate

[1] Pilaftsis, A. (1992). Radiatively induced neutrino masses and large Higgs-neutrino couplings in the Standard Model with Majorana fields. Zeitschrift für Physik C Particles and Fields, **55**(2), 275-282.

### One loop effects from Majorana neutrinos on the vertex WWZ

#### The WWZ vertex

Lagrangian

Where

$$\mathcal{L}_{WWZ}^{\text{even}} = -ig_{Z} \Big( g_{1} \Big( W_{\mu\nu}^{+} W^{-\mu} - W_{\mu\nu}^{-} W^{+\mu} \Big) Z^{\nu} \\
+ \kappa W_{\mu}^{+} W_{\nu}^{-} Z^{\mu\nu} + \frac{\lambda}{m_{W}^{2}} W_{\mu\nu}^{+} W^{-\nu}{}_{\rho} Z^{\rho\mu} - ig_{2} \epsilon^{\rho\mu\lambda\nu} Z_{\rho} \Big( W_{\nu}^{+} \partial_{\lambda} W_{\mu}^{-} - W_{\mu}^{-} \partial_{\lambda} W_{\nu}^{+} \Big) \Big) (10)$$

$$\mathcal{L}_{WWZ}^{\text{odd}} = -ig_{Z} \Big( \tilde{\kappa} W_{\mu}^{+} W_{\nu}^{-} \tilde{Z}^{\mu\nu} + \frac{\tilde{\lambda}}{m_{W}^{2}} W_{\mu\nu}^{+} W^{-\nu}{}_{\rho} \tilde{Z}^{\rho\mu} - i\tilde{g}_{1} W_{\mu}^{+} W_{\nu}^{-} \Big( \partial^{\mu} Z^{\nu} + \partial^{\nu} Z^{\mu} \Big) \Big) \Big)$$

$$(11)$$



[4] W. A. Bardeen, R. Gastmans, and B. Lautrup, Static quantities in Weinberg's model of weak and electromagnetic interactions, Nucl. Phys. B **46**, 319 (1972).

CP-even WWZ vertex function

$$Z_{\mu}(2q)$$

$$\Gamma_{\sigma\rho\mu}^{\text{even}} = ig_{Z} \left( g_{1} \left( 2p_{\mu}g_{\sigma\rho} + 4(q_{\rho}g_{\sigma\mu} - q_{\sigma}g_{\rho\mu}) \right) + \frac{4\Delta Q}{m_{W}^{2}} p_{\mu} \left( q_{\sigma}q_{\rho} - \frac{q^{2}}{2}g_{\sigma\rho} \right) + 2\Delta\kappa(q_{\rho}g_{\sigma\mu} - q_{\sigma}g_{\rho\mu}) + if_{1}\epsilon_{\sigma\rho\mu\alpha}p^{\alpha} \right)$$
(12)
$$\Delta\kappa = -g_{1} + \kappa + \lambda \qquad \Delta Q = -2\lambda$$

 $f_1 = -g_2$ 

 $W^+_{\sigma}$ 

Anomalous couplings

CP-odd WWZ vertex function

$$\begin{aligned}
Z_{\mu}(2q) \\
& \int \int_{\sigma\rho\mu}^{\text{odd}} = ig_Z \left( 2\Delta \tilde{\kappa} \epsilon_{\sigma\rho\mu\alpha} q^{\alpha} + \frac{4\Delta \tilde{Q}}{m_W^2} q_{\rho} \epsilon_{\sigma\mu\alpha\beta} p^{\alpha} q^{\beta} \\
& + i \tilde{f}_1 (q_{\rho} g_{\sigma\mu} + q_{\sigma} g_{\rho\mu}) \\
& + \tilde{f}_2 p^{\lambda} \epsilon_{\sigma\rho\lambda\alpha} (q^2 \delta^{\alpha}{}_{\mu} - q^{\alpha} q_{\mu}) \right),
\end{aligned}$$
(13)

 $f_2 =$ 

$$\Delta \tilde{\kappa} = \tilde{\kappa} + \frac{m_W^2 - 2q^2}{m_W^2} \tilde{\lambda} \qquad \Delta \tilde{Q} = -2\tilde{\lambda}$$
$$\tilde{f}_1 = 2\tilde{g}_1 \qquad \tilde{f}_2 = -\frac{4\tilde{\lambda}}{m_W^2}$$

Anomalous couplings

#### One-loop contributions to WWZ

Vertex function  $\Gamma^{WWZ}_{\sigma\rho\mu} = \Gamma^{lln}_{\sigma\rho\mu} + \Gamma^{nnl}_{\sigma\rho\mu} + \Gamma^{nnh}_{\sigma\rho\mu}$  (14)

 $W_{\rho}^{-}$ 

 $Z_{\mu}$ 

$$\Gamma^{lln}_{\sigma\rho\mu} = \sum_{k=1}^{6} \sum_{\alpha} I_{\alpha} I_{\alpha}$$







#### One-loop contributions to WWZ

Vertex function  $\Gamma^{WWZ}_{\sigma\rho\mu} = \Gamma^{lln}_{\sigma\rho\mu} + \Gamma^{nnl}_{\sigma\rho\mu} + \Gamma^{nnh}_{\sigma\rho\mu}$  (14)



 $n_k$  represents both  $\nu_k$   $N_k$ 



#### One-loop contributions to WWZ



**Do not contribute** to  $\Delta Q$  and  $\Delta \kappa$ 

### **Estimations and discussion**

$$\xi = \hat{\rho} X$$
Considering the matrix texture
$$X = e^{i\varphi} \cdot 1_3$$

$$V^{\ell} = U_{\nu}$$

 $m_{\nu_i} \approx 0.8 \text{eV}$  [5]

#### Nearly degenerated Neutrino spectrum

 The International Linear Collider technical design report estimated sensitivity [6]

 $W^{-}$ 

e+

 $e^{-}$ 

Coupling	$\sqrt{s} = 500 \mathrm{GeV}$	$\sqrt{s} = 800 \mathrm{GeV}$
$ \Delta \kappa_{ m ILC} $	$3.20 \times 10^{-4}$	$1.90 \times 10^{-4}$
$ \Delta Q_{\rm ILC} $	$1.34 \times 10^{-3}$	$6.00 \times 10^{-4}$
$ \Delta \tilde{\kappa}_{ m ILC} $	$5.33 \times 10^{-2}$	$5.77 \times 10^{-2}$
$ \Delta  ilde{Q}_{ ext{ILC}} $	$1.5 \times 10^{-3}$	$6.00 \times 10^{-4}$

[5] M. Aker et al. (The KATRIN Collaboration), Direct neutrino-mass measurement with sub-electronvolt sensitivity, Nature Phys. 18, 160 (2022)

[6] H. Baer et al. (ILC Collaboration), The international linear collider technical design report-Volume 2: Physics, arXiv:1306.6352.

### **CP-Even** contributions

$$\Delta \kappa = \sum_{\alpha} \left( \frac{1}{1+\hat{\rho}^2} \left( \Delta \kappa_{\alpha\nu}^{(1)} + \hat{\rho}^2 \Delta \kappa_{\alpha N}^{(1)} \right) + \frac{1}{(1+\hat{\rho}^2)^2} \left( \Delta \kappa_{\alpha\nu\nu}^{(2)} + \Delta \kappa_{\alpha\nu\nu}^{(3)} \right) + \frac{\hat{\rho}^4}{(1+\hat{\rho}^2)^2} \left( \Delta \kappa_{\alpha NN}^{(2)} + \Delta \kappa_{\alpha NN}^{(3)} \right) + \frac{2\hat{\rho}^2}{(1+\hat{\rho}^2)^2} \left( \Delta \kappa_{\alpha\nu N}^{(2)} + \cos(2\varphi) \Delta \kappa_{\alpha\nu N}^{(3)} \right)$$

$$(16)$$



Accessible values to the

25

## Contributions to the light + heavy neutrinos to the CP-even form factor $\Delta \kappa$

 $|\Delta \kappa|$  CP-even



## Contributions to the light + heavy neutrinos to the CP-even form factor $\Delta \kappa$

 $|\Delta\kappa|$  CP-even for  $\hat{\rho}$ =0.65  $\sqrt{s}$  = 800 GeV

 $Log_{10}{|\Delta\kappa|}$ 



### **CP-Odd contributions**

Implementing the X-matrix texture

$$\Delta \tilde{\kappa} = 3 \left( \tilde{\Omega} + \tilde{\mathcal{J}} \right) + \frac{2i\hat{\rho}^2 \sin 2\varphi}{\left( 1 + \hat{\rho}^2 \right)^2} \sum_{\alpha} \Delta \tilde{\kappa}^{(3)}_{\alpha\nu N} \quad (17)$$

 $|\Delta \tilde{\kappa}| = 1.19 \times 10^{-3}$ 

### Summary and conclusions

 BSM physics effects on the WWZ with virtual Majorana neutrinos at one-loop level.

• SM *W* bosons on-shell and Z boson off-shell.

$$\Delta \kappa \longrightarrow \text{CP-even} + \text{CP-odd effects} \longrightarrow \Delta \tilde{\kappa}$$

- Ultraviolet finite.
- Analytical expressions in terms of three parameters:  $m_N$ , s,  $\hat{\rho}$ .



• $\Delta \kappa$  — Contributions  $\mathcal{O}(10^{-3})$ We could reach ILC sensitivity at CME of 800 GeV due to the effects from virtual heavy and light neutrinos.

•  $\Delta Q \rightarrow$  At least one order of magnitude below the estimated sensitivity for ILC.

•  $\Delta \tilde{\kappa}$  Contributions  $\mathcal{O}(10^{-3})$ . One order of magnitude beyond expected ILC sensitivity.

## Thanks!

[1] Pilaftsis, A. (1992). Radiatively induced neutrino masses and large Higgs-neutrino couplings in the Standard Model with Majorana fields. Zeitschrift für Physik C Particles and Fields, **55**(2), 275-282.

[2] J. G. Korner, A. Pilaftsis, and K. Schilcher, Leptonic CP asymmetries in flavor-changing H0 decays, Phys. Rev. D 47, 1080 (1993).

[3] P. S. B. Dev and A. Pilaftsis, Minimal radiative neutrino mass mechanism for inverse seesaw models, Phys. Rev. D **86**, 113001 (2012).

[4] W. A. Bardeen, R. Gastmans, and B. Lautrup, Static quantities in Weinberg's model of weak and electromagnetic interactions, Nucl. Phys. B **46**, 319 (1972).

[5] H. Baer et al. (ILC Collaboration), The international linear collider technical design report-Volume 2: Physics, arXiv:1306.6352.

[6] A. M. Sirunyan et al. (CMS Collaboration), Search for heavy neutral leptons in events with three charged leptons in proton-proton collisions at ps = 13TeV, Phys. Rev. Lett. **120**, 221801 (2018).

[7] G. 't Hooft and M. Veltman, Regularization and renormalization of gauge fields, Nucl. Phys. B **44**, 189 (1972).

[8] F. Jegerlehner, Facts of life with 5, Eur. Phys. J. C 18, 67 (2001).

[9] G. Weiglein et al., Physics interplay of the LHC and the ILC, Phys. Rept. 426, 47 (2006).

[10] L. Bian, J. Shu, and Y. Zhang, Prospects for triple gauge coupling measurements at future lepton colliders and the 14 TeV LHC, JHEP **09**, 206 (2015).

# Thanks!

