



# Placing Bounds on New Physics with Displaced Photons

**Joel Jones-Pérez**

Pontificia Universidad Católica del Perú (PUCP)

**Based on the following work:**

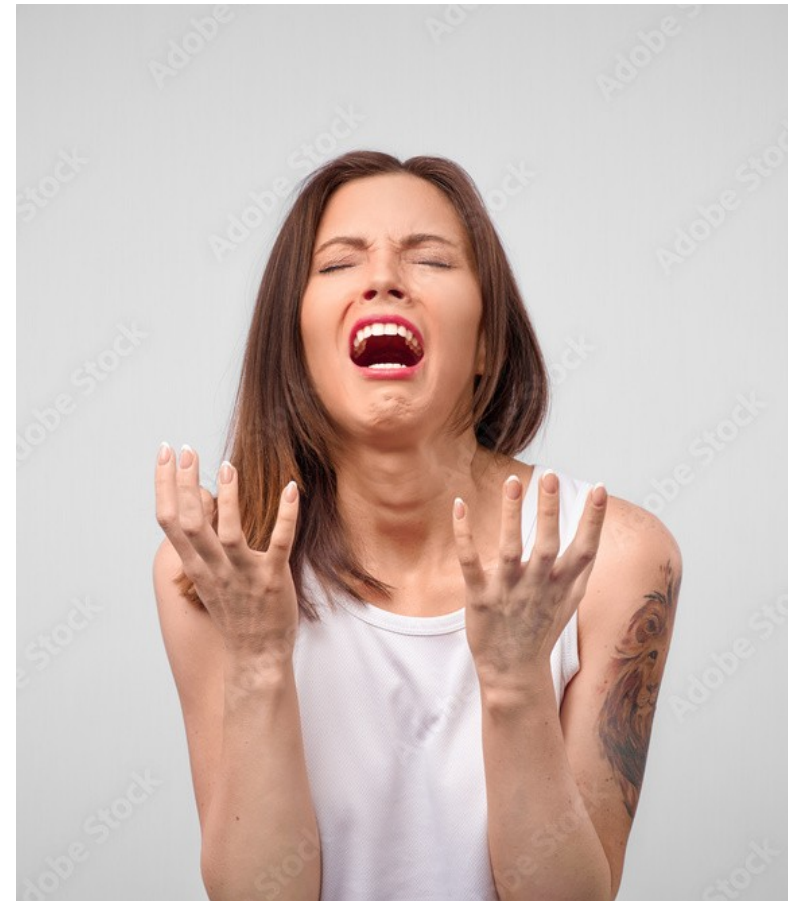
F. Delgado, L. Duarte, JJP, C. Manrique-Chavil, S. Peña (2205.13550)

L. Duarte, JJP, C. Manrique-Chavil (2311.17989)

B. Díaz, L. Duarte, JJP, W. Rodriguez, D. Zegarra (25xx.xxxxx)

# Why don't we find physics beyond the Standard Model?

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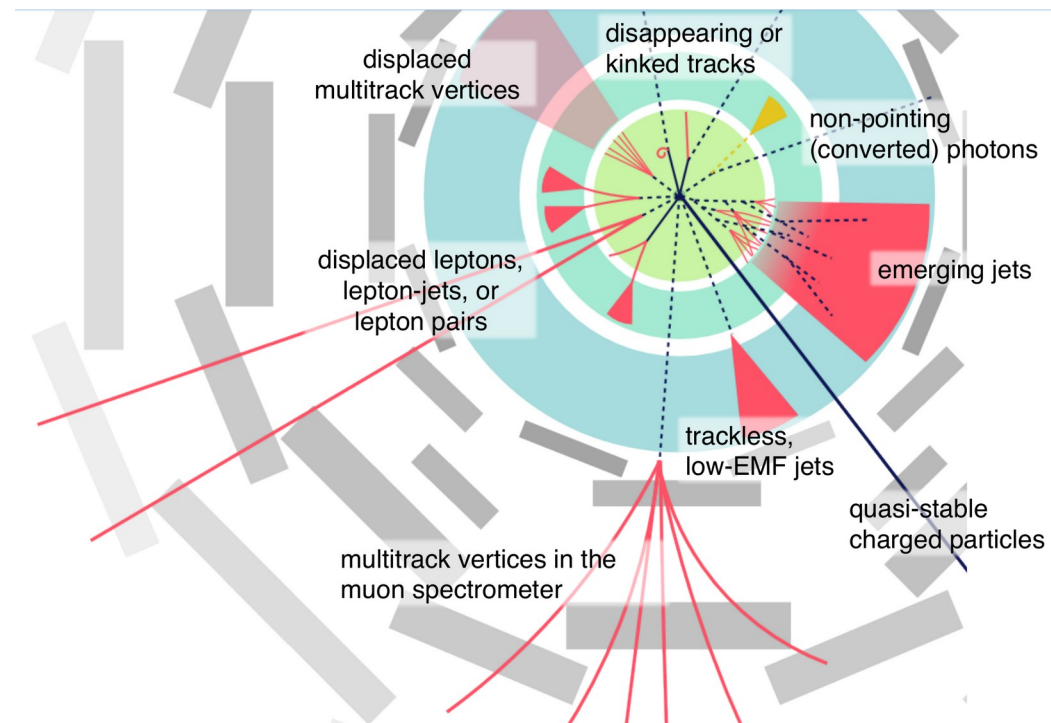
# Why don't we find physics beyond the Standard Model?

Maybe the new physics sector implies signals we did not expect when the LHC was first designed!

Focus on long-lived particles (LLPs).

A large lifetime implies a macroscopic decay length:

Non-standard signals!



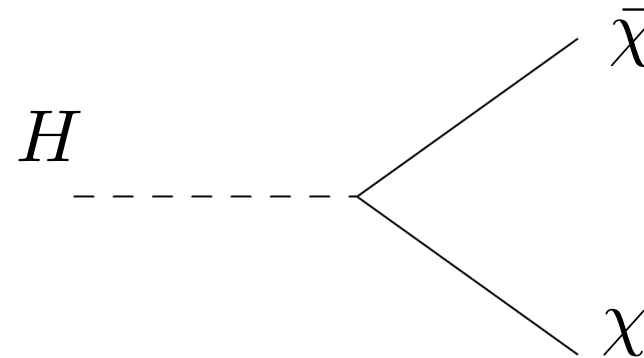
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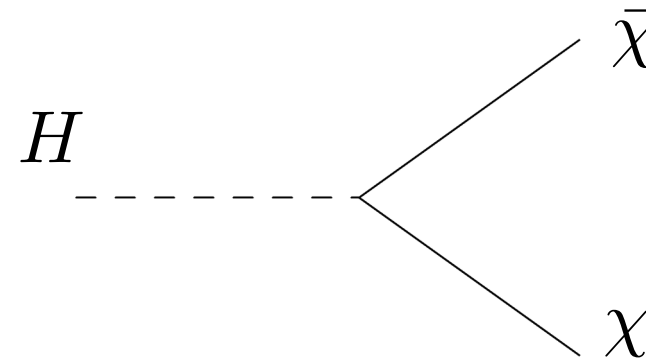
In this work, we are interested in looking for neutral LLPs, assuming production via Higgs decays



## How do we create LLPs at colliders?

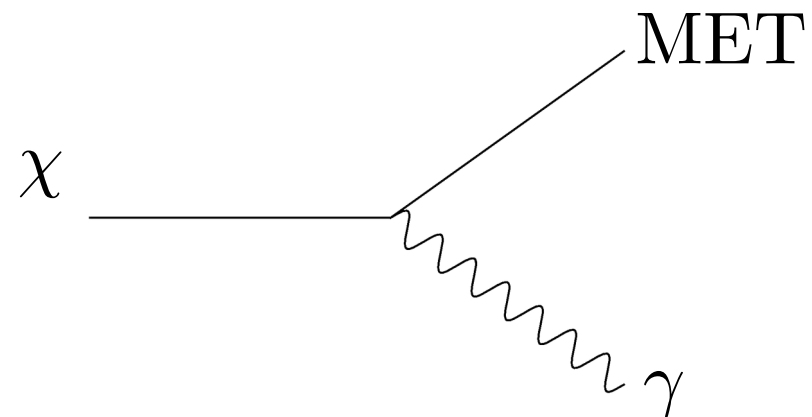
- A GeV mass particle with a long lifetime implies a very small coupling.
- It is better if production and decay processes are not related.

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We are also interested in models where the LLP decays into a photon and an invisible state:

**displaced photon!**



# The search for displaced photons

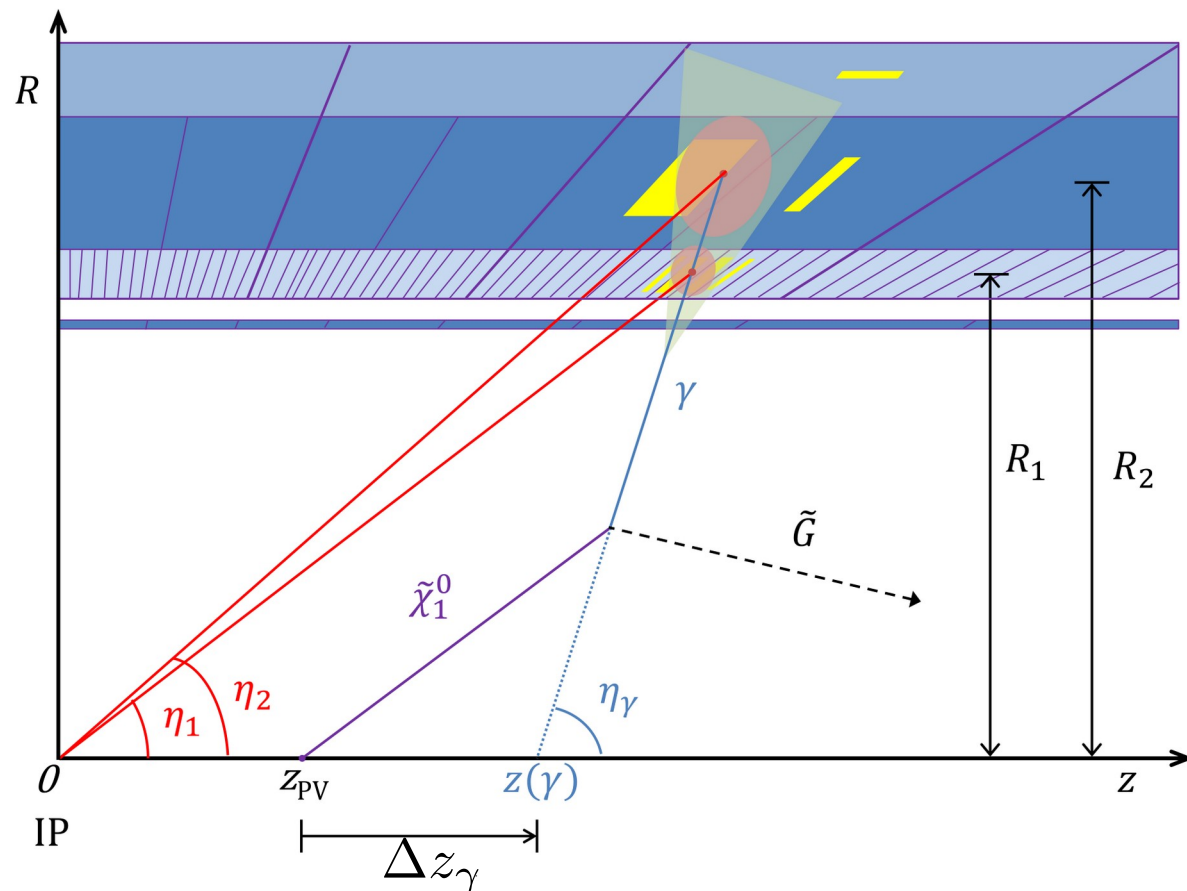


## How do you distinguish displaced photons?

Photons coming from long-lived particles take longer to reach the ECAL, and do not point towards the primary vertex.

Important variables:

$$t_\gamma \quad |\Delta z_\gamma|$$

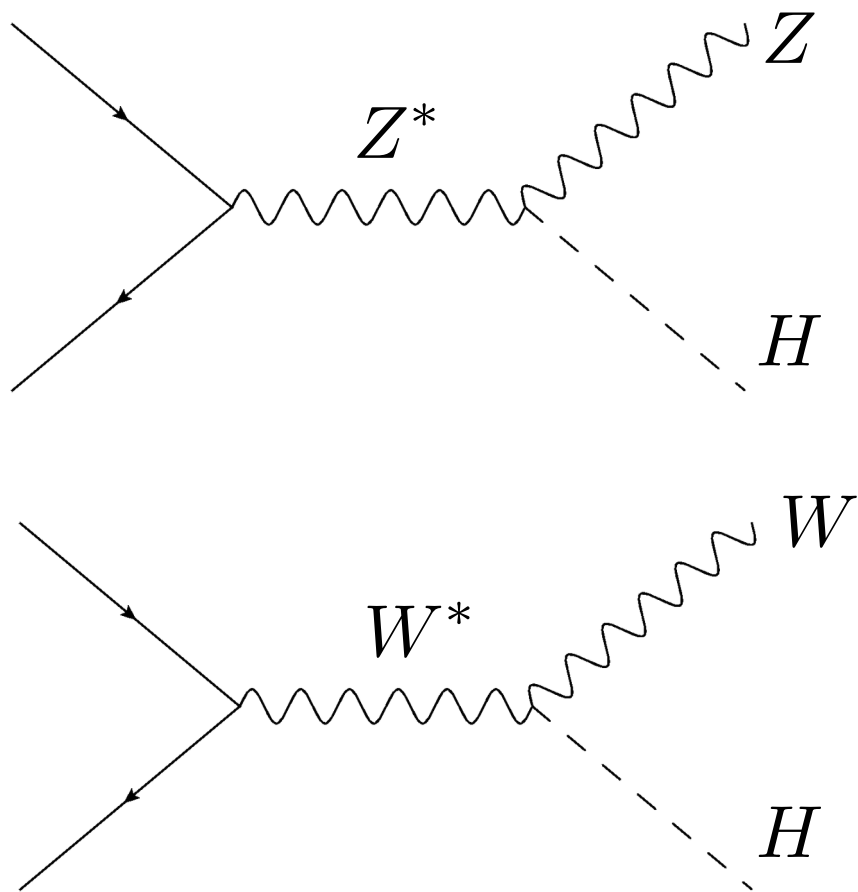


## Displaced photon search

- Trigger: isolated lepton with  $p_T > 27$  GeV.

## Displaced photon search

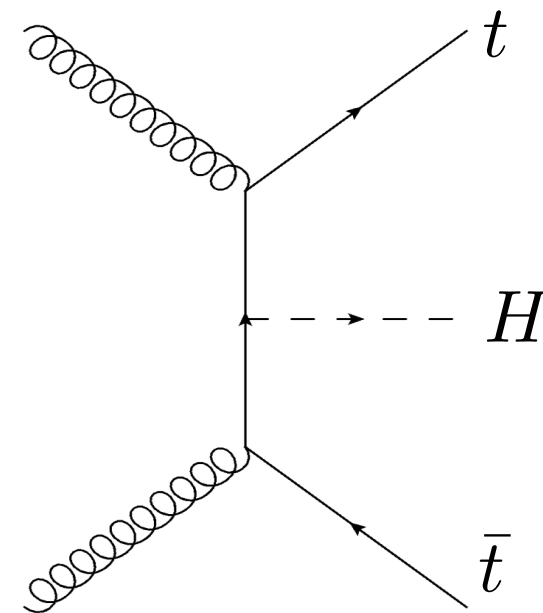
- Trigger: isolated lepton with  $p_T > 27$  GeV.



### Idea:

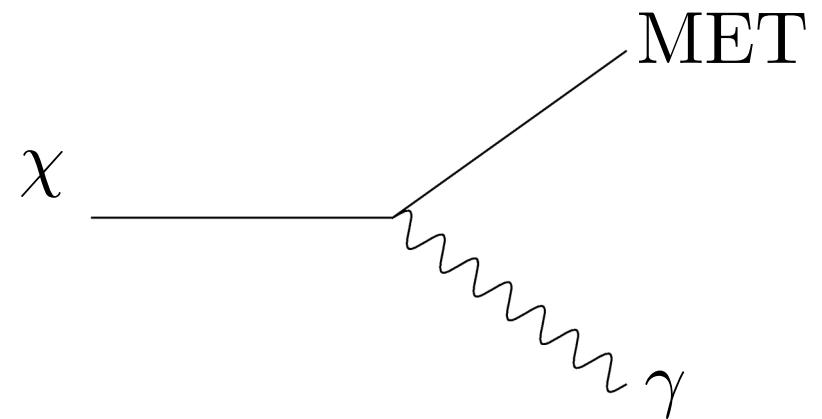
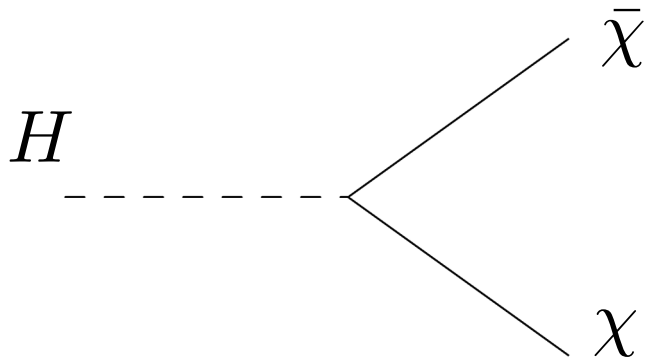
LLPs come from Higgs decay.

Trigger on leptons from Z, W or t.



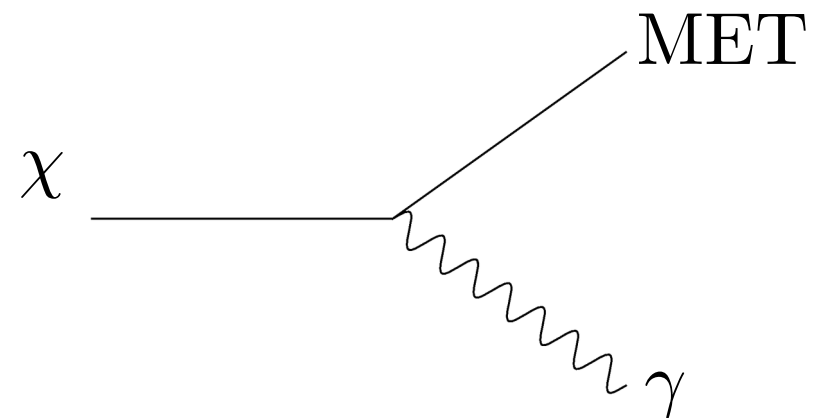
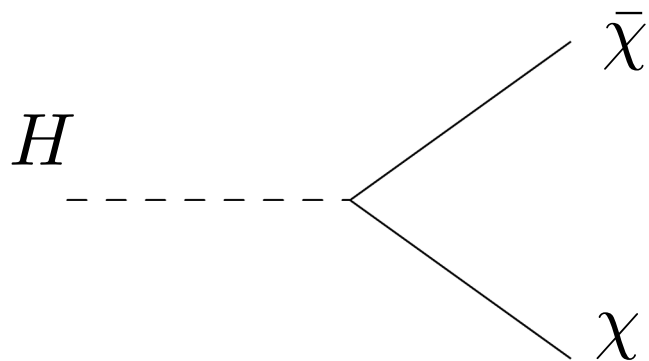
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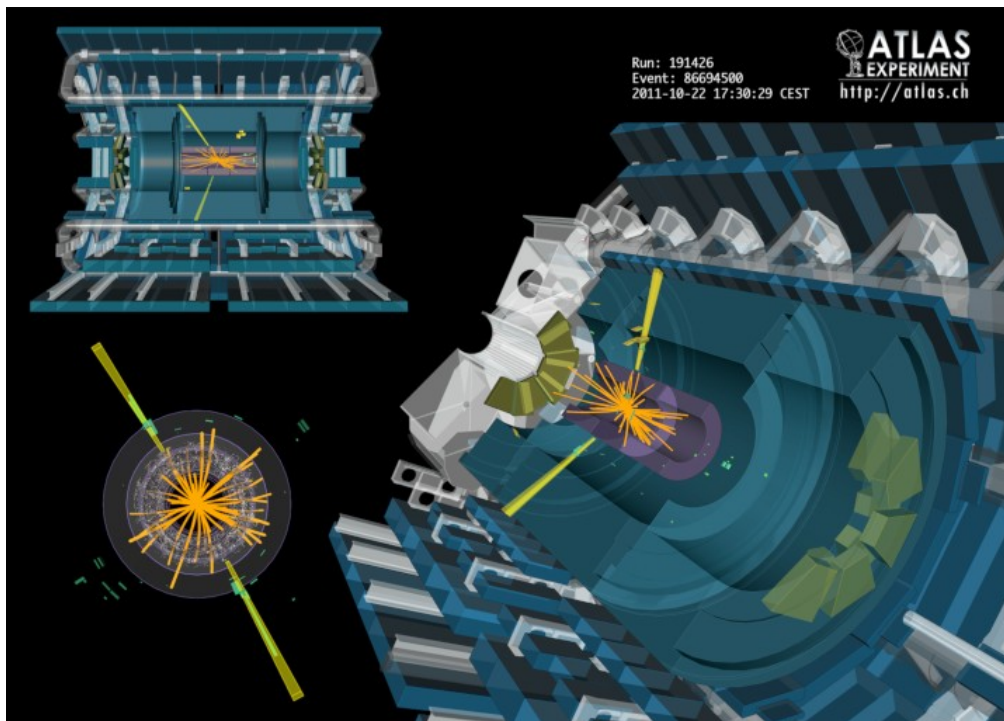
### Issue:

Final state photons are somewhat soft!

Probably not much MET

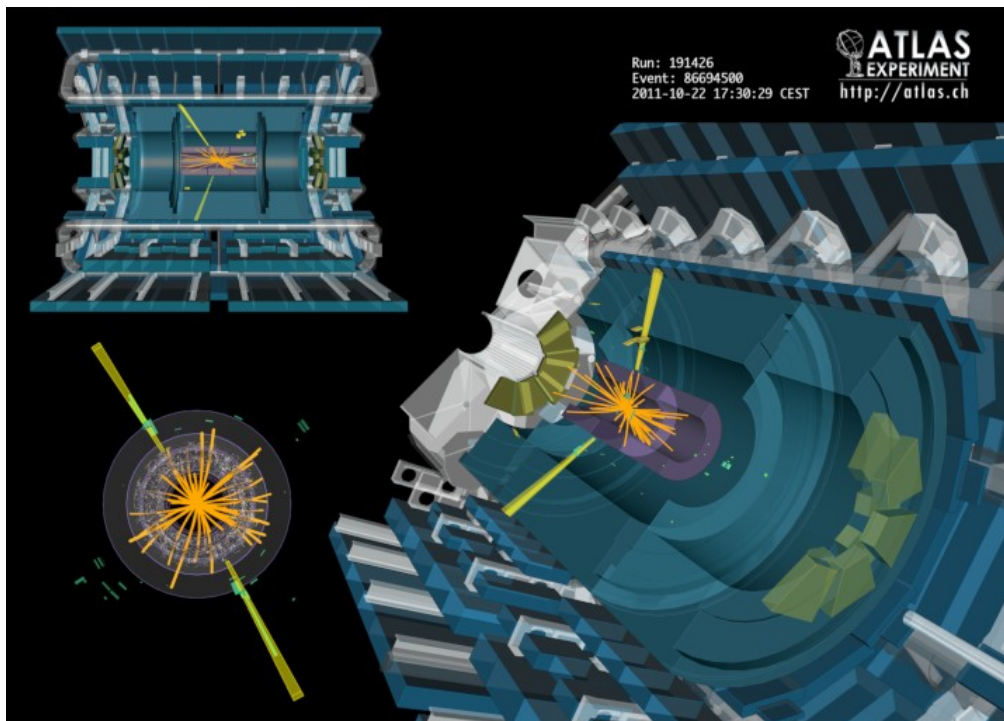
## Displaced photon search

- Trigger: isolated lepton with  $p_T > 27$  GeV.
- At least one “loose” photon with energy larger than 10 GeV.
- Isolation criteria: no deposits larger than 5% - 6.5% of energy within  $\Delta R = 0.2$ .



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### Issue:

Backgrounds are huge! Need to cut using timing, displacement, and MET.

## Displaced photon search

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- One of the photons must be in barrel region (used for analysis).
- If more than one photon in barrel region, use the one with largest energy.
- Require  $E_{cell}$  larger than 10 GeV.

**$E_{cell}$ :**

Maximum energy deposit in a  
middle-layer ECAL cell  
(about 0.2 – 0.5 of total energy)





## Displaced photon search

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- If more than one photon in barrel region, use the one with largest energy.
- Require  $E_{cell}$  larger than 10 GeV.
- Define signal region depending on  $MET > 50$  GeV.
- Place cuts on  $t_\gamma$  and  $|\Delta z_\gamma|$ . Distinguish single and multi-photon samples.

$$(1) \quad 1.5 \text{ ns} < t_\gamma < 12 \text{ ns} \qquad 1 \text{ ns} < t_\gamma < 12 \text{ ns} \quad (2+)$$

$$|\Delta z_\gamma| > 300 \text{ mm}$$

## What did the search find?

|          | 1             | 2+              | 1+            |
|----------|---------------|-----------------|---------------|
| Expected | $3.8 \pm 1.6$ | $0.28 \pm 0.04$ | $4.1 \pm 1.7$ |
| Observed | 4             | 0               | 4             |



Let us put bounds on models... recast the search!

# Dimension Five Seesaw Portal

F. Delgado, L. Duarte, JJP, C. Manrique-Chavil, S. Peña (2205.13550)  
L. Duarte, JJP, C. Manrique-Chavil (2311.17989)

## Dimension-5 Type-I Seesaw Portal

We are interested in an extension of Type-I Seesaw model with  $d=5$  operators, involving the sterile neutrino states and neutral SM bosons.

We again add only two sterile neutrinos.

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + Y_\nu \left( \bar{L} \cdot \tilde{\phi} \nu_R \right) + \frac{1}{2} M_R (\bar{\nu}_R \nu_R^c) + \text{h.c.}$$

$$+ \left( \frac{(\alpha_{N\phi})_{ss'}}{\Lambda} (\phi^\dagger \phi) \bar{\nu}_{Rs} \nu_{Rs'}^c + \frac{(\alpha_{NB})_{ss'}}{\Lambda} \bar{\nu}_{Rs} \sigma^{\mu\nu} \nu_{Rs'}^c B_{\mu\nu} + \text{h.c.} \right)$$

Light neutrinos interact via these operators through “sterile-light” mixing.

## Decay chain

The new couplings allow:

$$pp \rightarrow H \xrightarrow[(\alpha_{N\phi})_{45}]{} N_4 N_5$$

# Decay chain

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[illegible]

$$(\alpha'_{NB})_{\ell h} \equiv U_{s\ell} (\alpha_{NB})_{ss'} U_{s'h}$$

# Decay chain

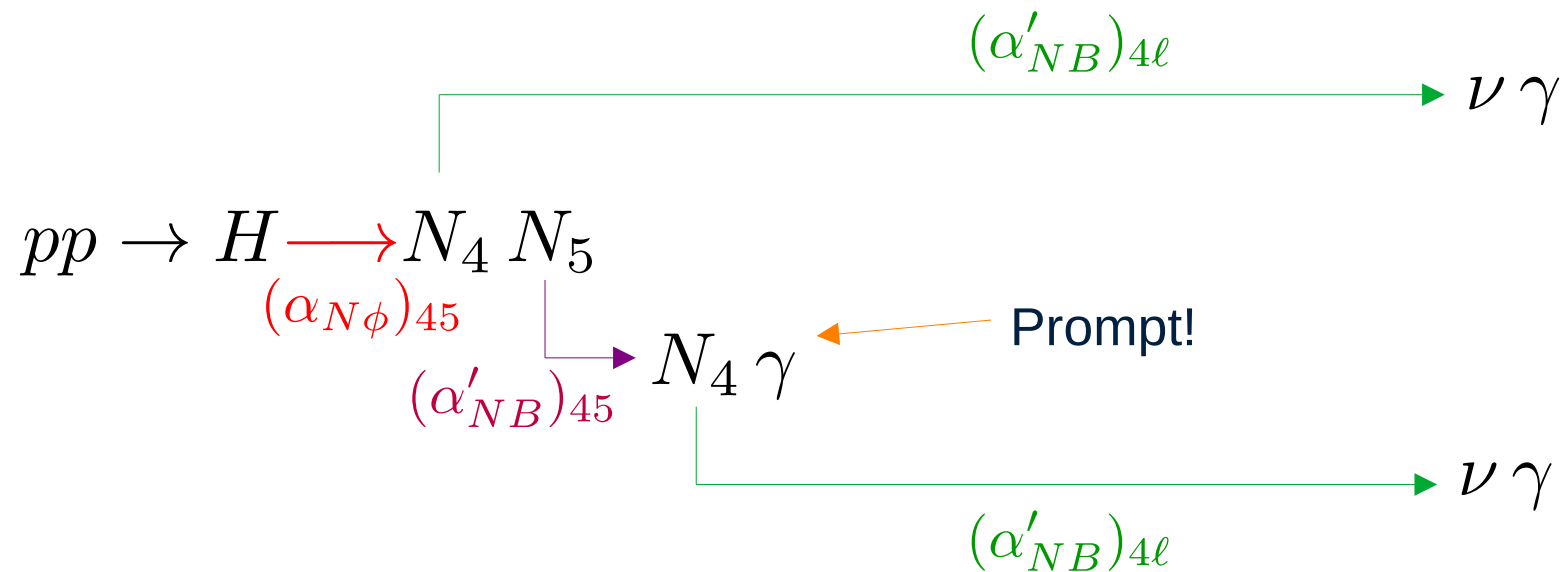
## The new couplings allow:

$$pp \rightarrow H \xrightarrow{(\alpha_{N\phi})_{45}} N_4 N_5 \xrightarrow{(\alpha'_{NB})_{45}} N_4 \gamma \quad \leftarrow \text{Prompt!}$$

$$(\alpha'_{NB})_{\ell h} \equiv U_{s\ell} (\alpha_{NB})_{ss'} U_{s'h}$$

## Decay chain

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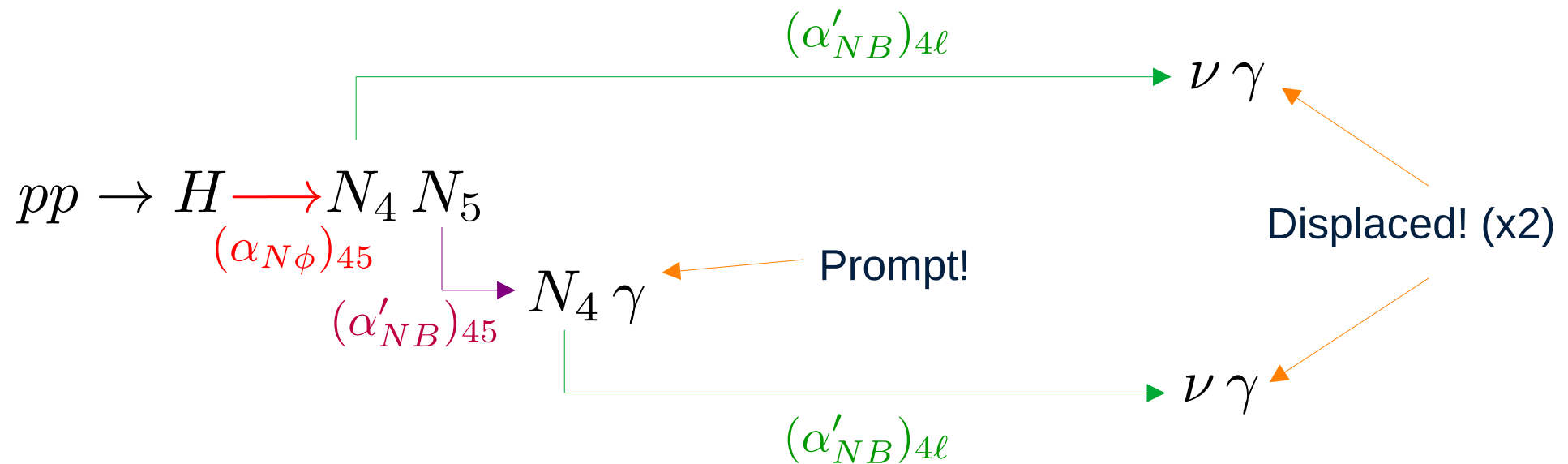


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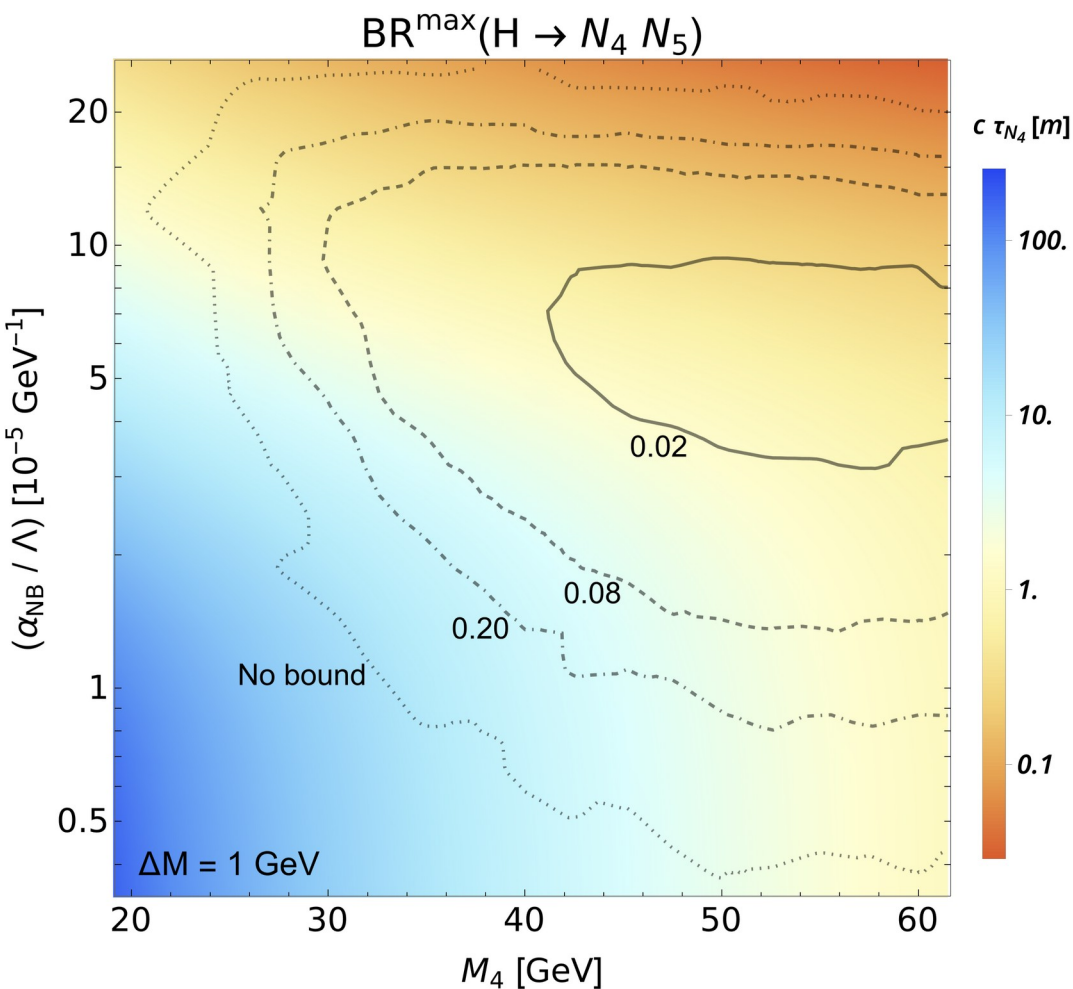
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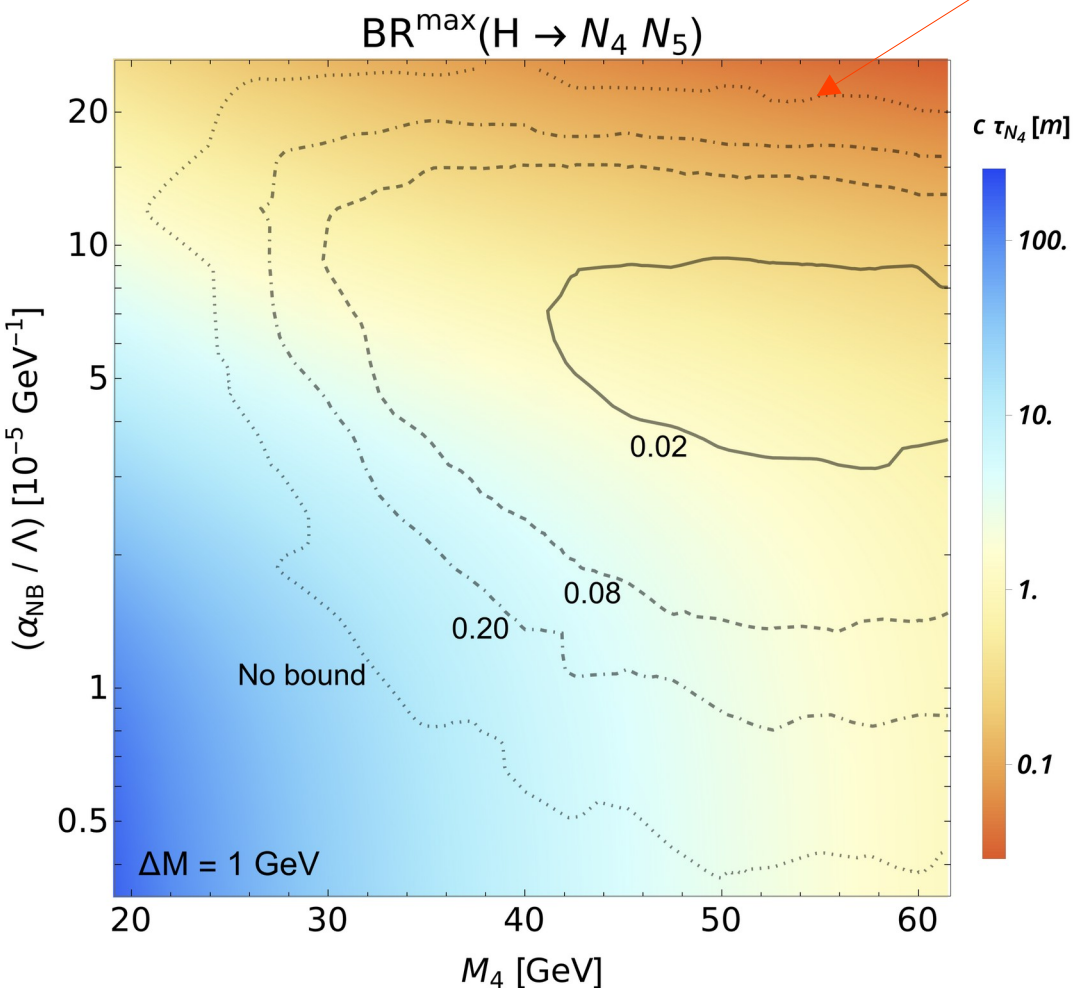
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## Bounds on $\text{BR}(H \rightarrow N_4 N_5)$ !

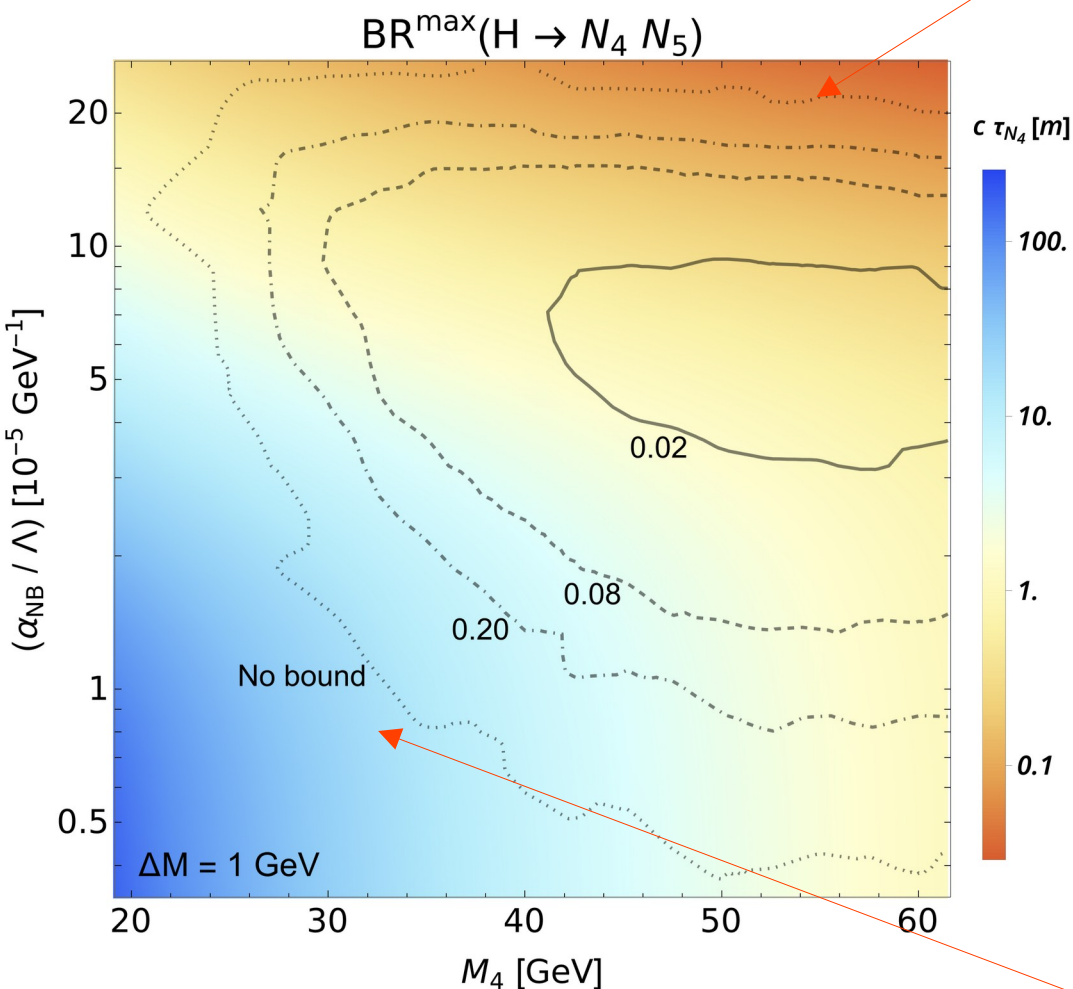


## Bounds on $\text{BR}(H \rightarrow N_4 N_5)$ !

Large coupling implies that  $N_4$  decay is prompt.



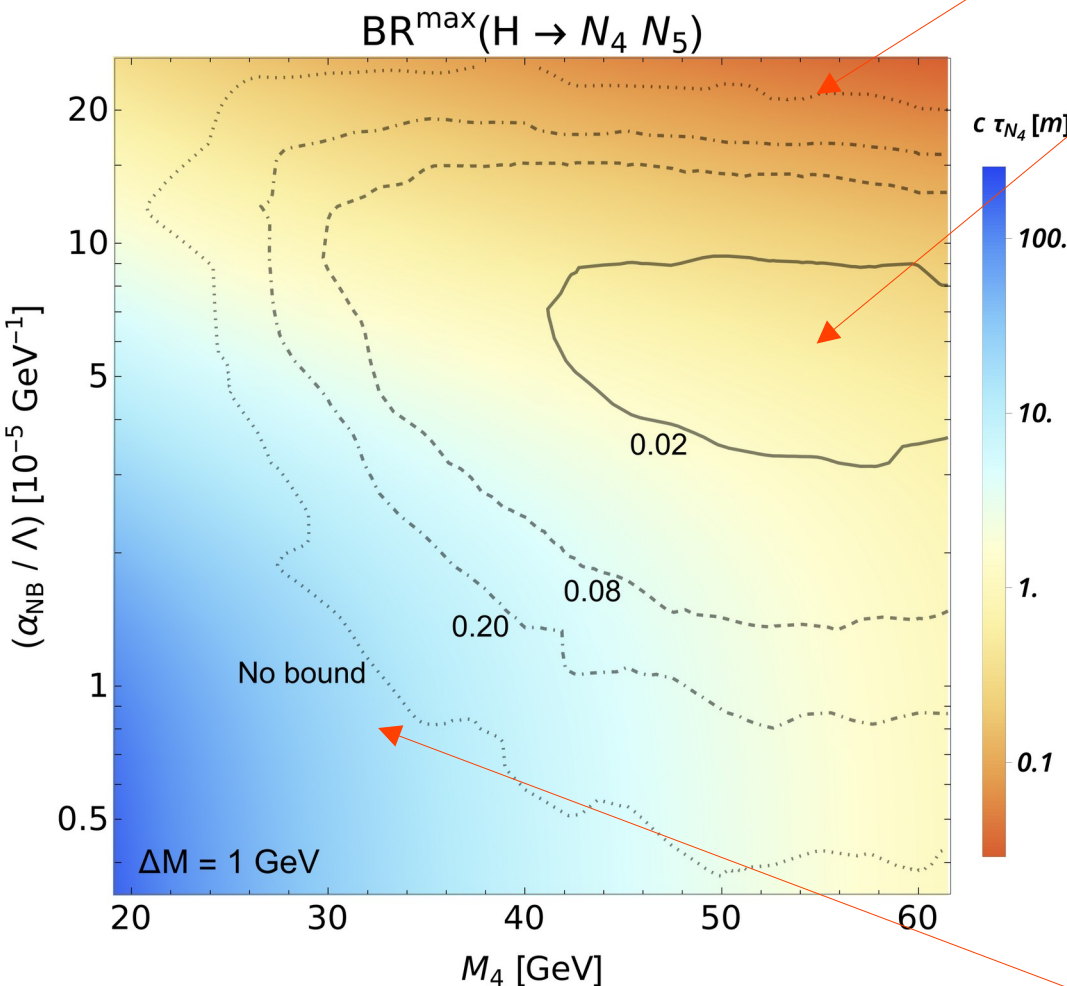
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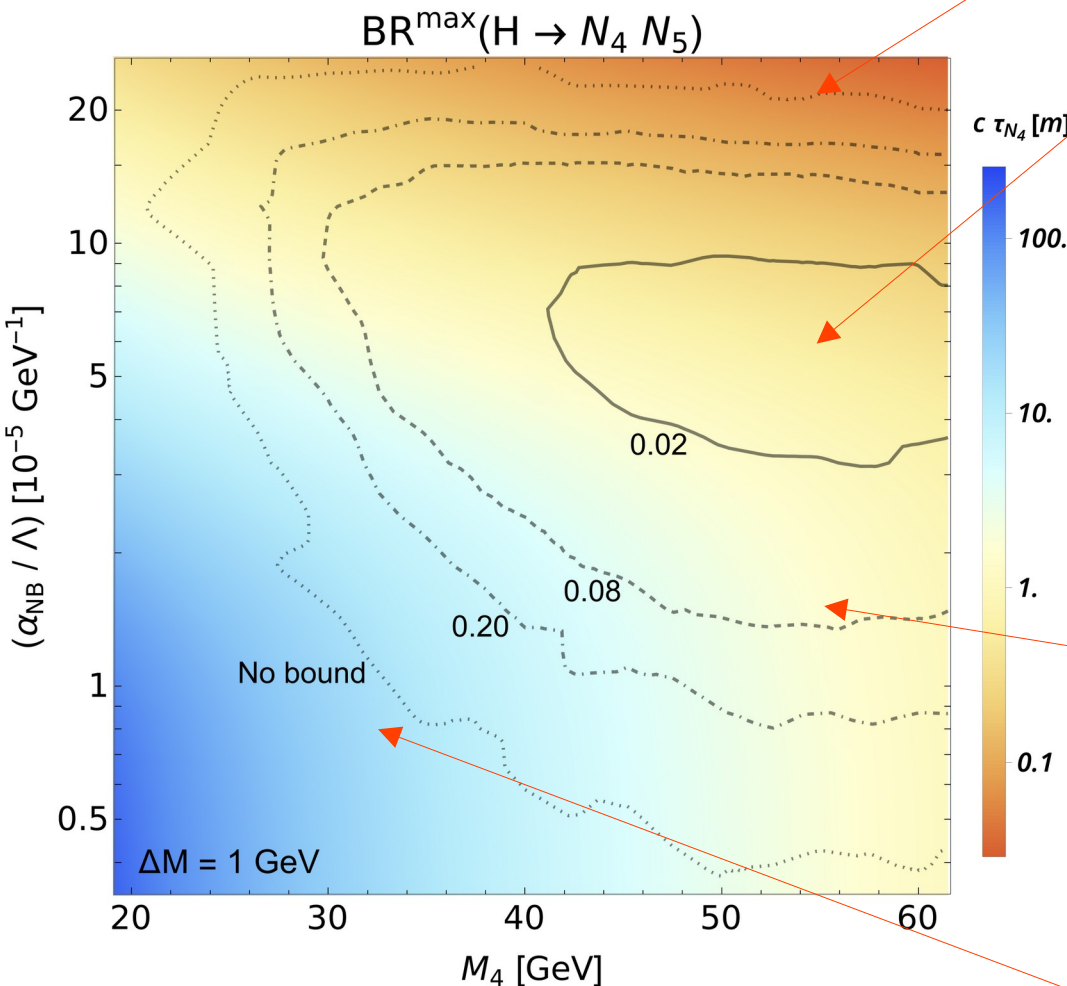
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Region with heavier masses is favoured, as  $N_4$  is less boosted:

- $N_4$  moves more slowly, so  $t_\gamma$  is larger.
- Photon and neutrino can have larger angular separation, so  $|\Delta z_\gamma|$  can be larger (and maybe MET?)

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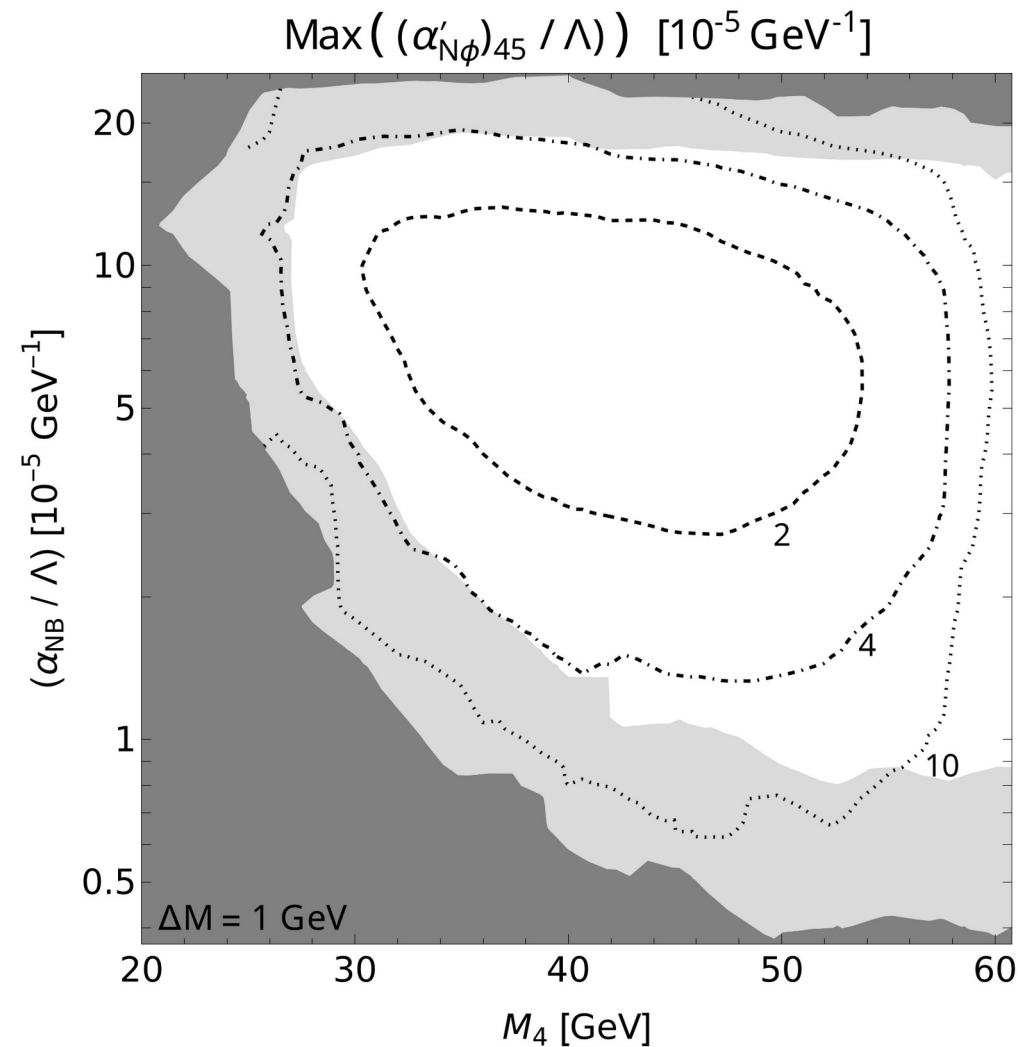
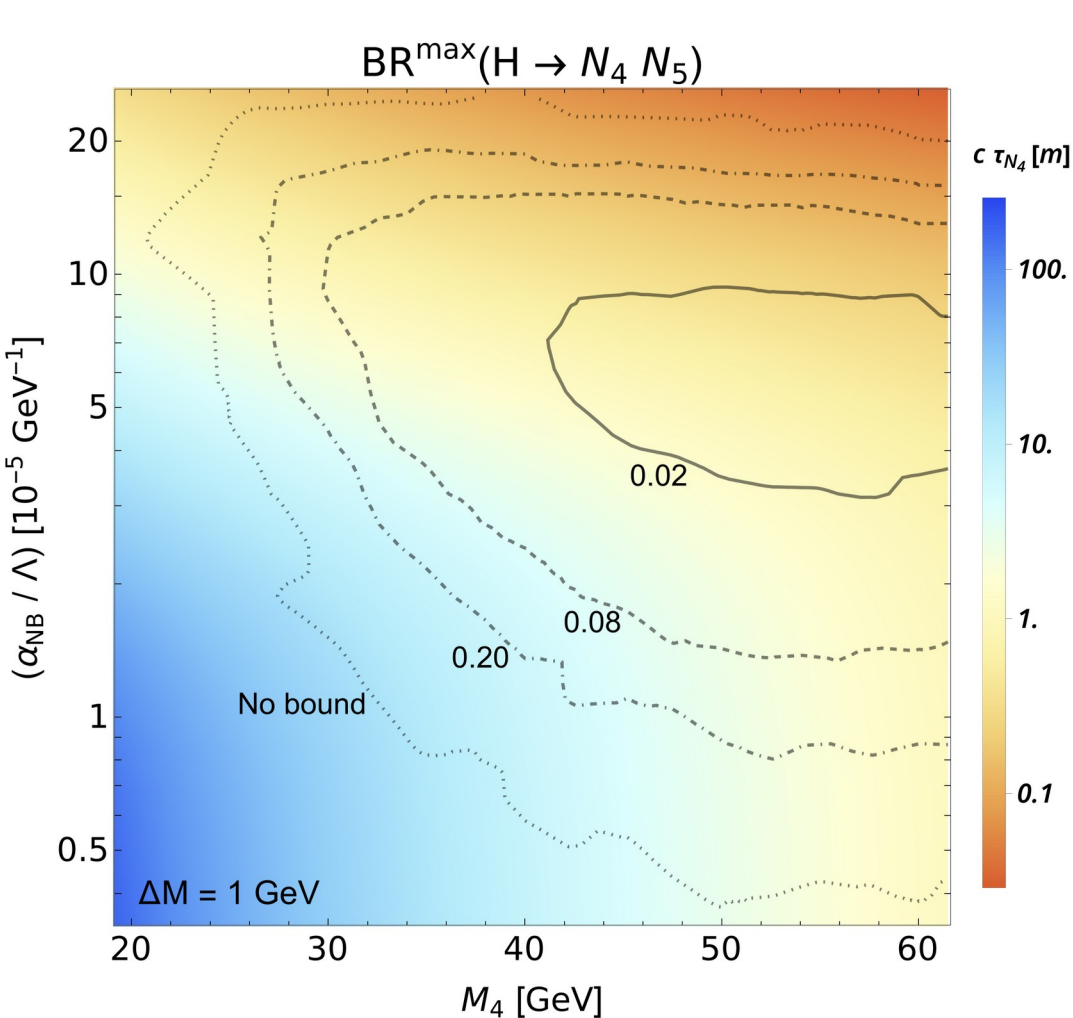
- $N_4$  moves more slowly, so  $t_\gamma$  is larger.
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Standard seesaw decays start to dominate, so lifetime does not depend on  $\alpha_{\text{NB}}$ .

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## Bounds on $\alpha_{N\phi}$ !



# Dark Photon Model

B. Díaz, L. Duarte, JJP, W. Rodriguez, D. Zegarra (25xx.xxxxx)



## Dark photons

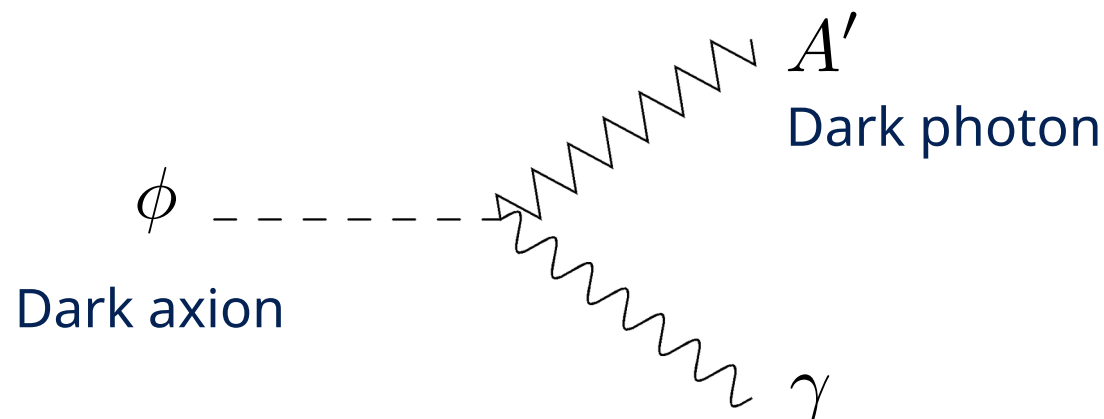
- Can we think of any other model that could leave a displaced photon signal?
- Inspiration: dark axion models!

$$\mathcal{L}_{\text{d.a.}} = \frac{G_1}{4} \phi F_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{G_2}{4} \phi F'_{\mu\nu} \tilde{F}^{\mu\nu} + \frac{G_3}{4} \phi F'_{\mu\nu} \tilde{F}'^{\mu\nu}$$

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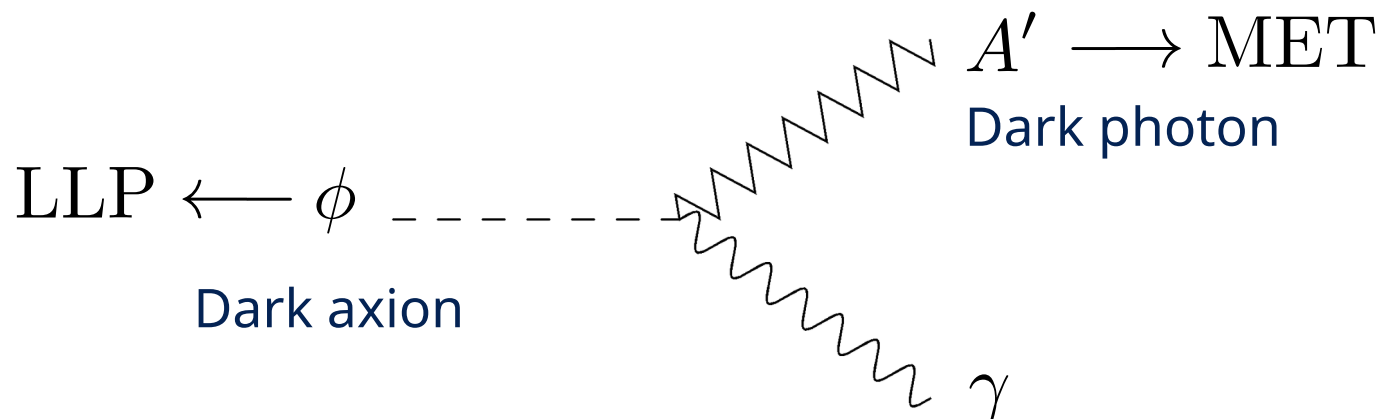
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## Dark photons

- Need to keep  $A'$  stable. Or at least very long lived!
- Add  $Z_2$  symmetry:

$$\phi \rightarrow -\phi \qquad A'_\mu \rightarrow -A'_\mu$$

$$\Rightarrow \mathcal{L}_{\text{d.a.}} = \frac{g_D}{\Lambda} \phi F'_{\mu\nu} \tilde{B}^{\mu\nu}$$

- This is the only interaction term of dark photon, with no kinetic mixing!
- The field  $\phi$  does not transform under  $U(1)$  symmetry of dark photon.

## Dark photons

- Next: add an interaction term for  $\phi$  with the Higgs:

$$V = -\frac{\mu_H^2}{2}|H|^2 + \frac{\lambda_H}{4}(H^\dagger H)^2 + \frac{\mu_\phi}{2}\phi^2 + \frac{\lambda_\phi}{4}\phi^4 + \lambda_{h\phi}|H|^2\phi^2$$

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Should be large-ish...  
Not really an axion?

### REBRANDING:

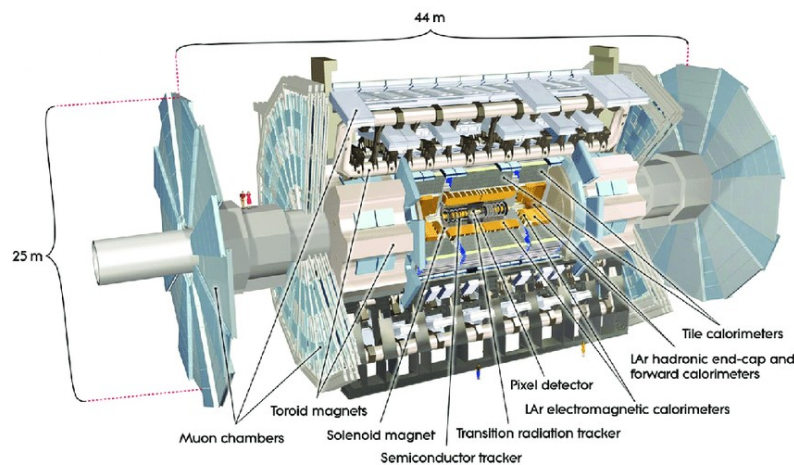
Dark photon with a scalar portal!

### OBJECTIVE:

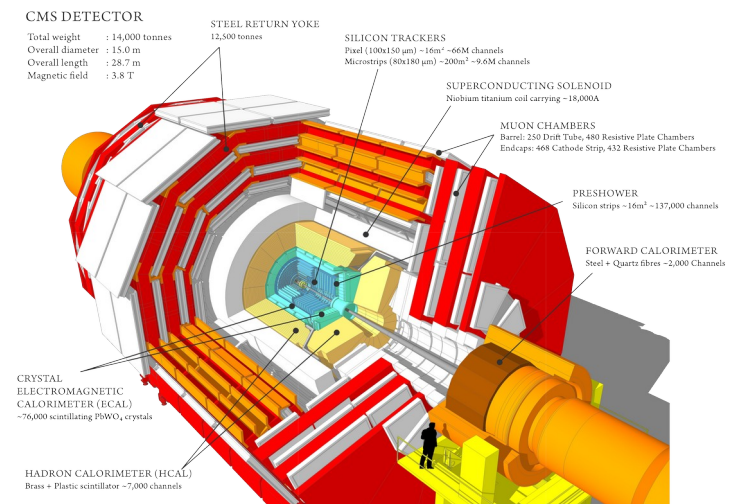
- Bound the model with displaced photons
- Check if good DM candidate

## Additional constraint: Higgs invisible decays

- If the lifetime of the LLP is very long, the particle could escape the detector before it decays.
- In this case, one can constrain models with searches for Higgs invisible decays.

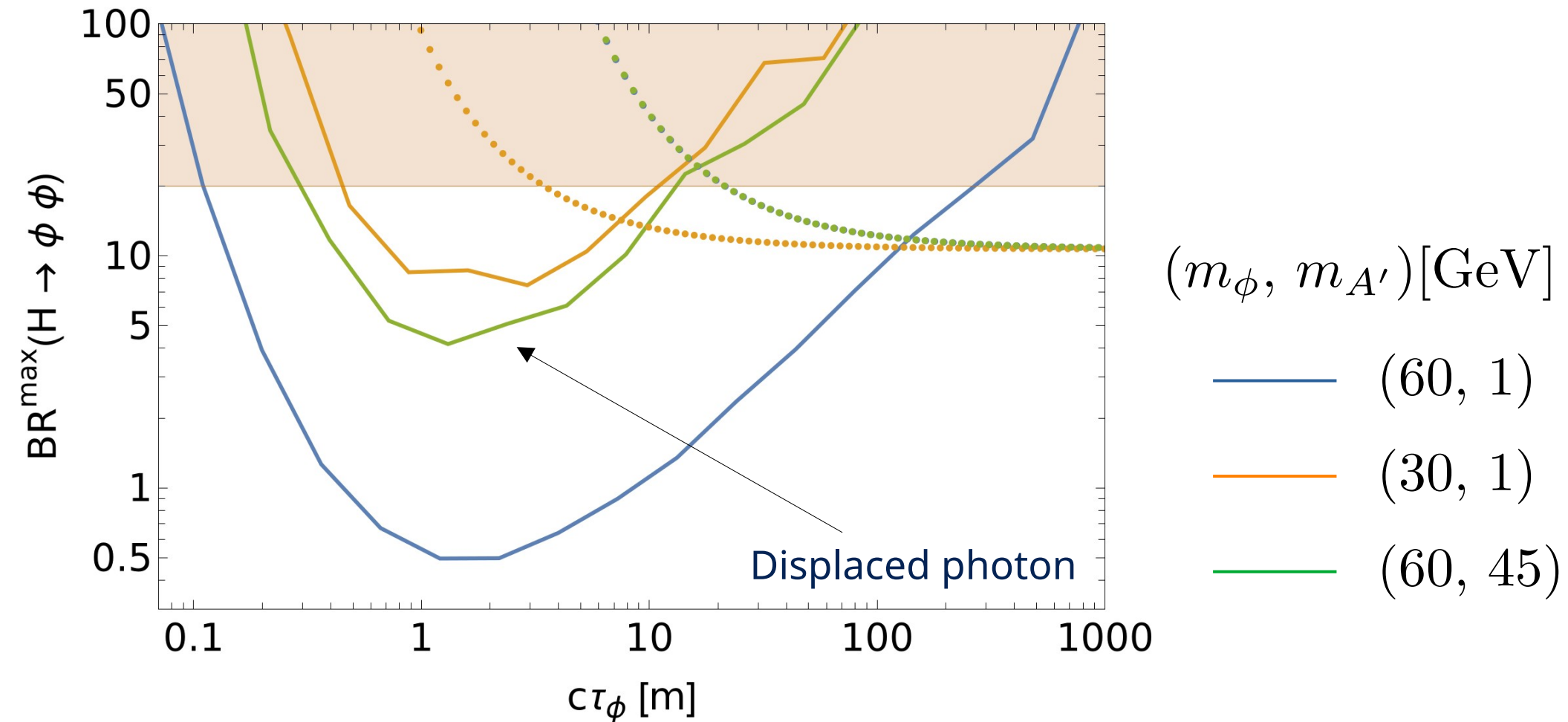


$$\text{BR}(H \rightarrow \text{inv}) < 10.7\%$$



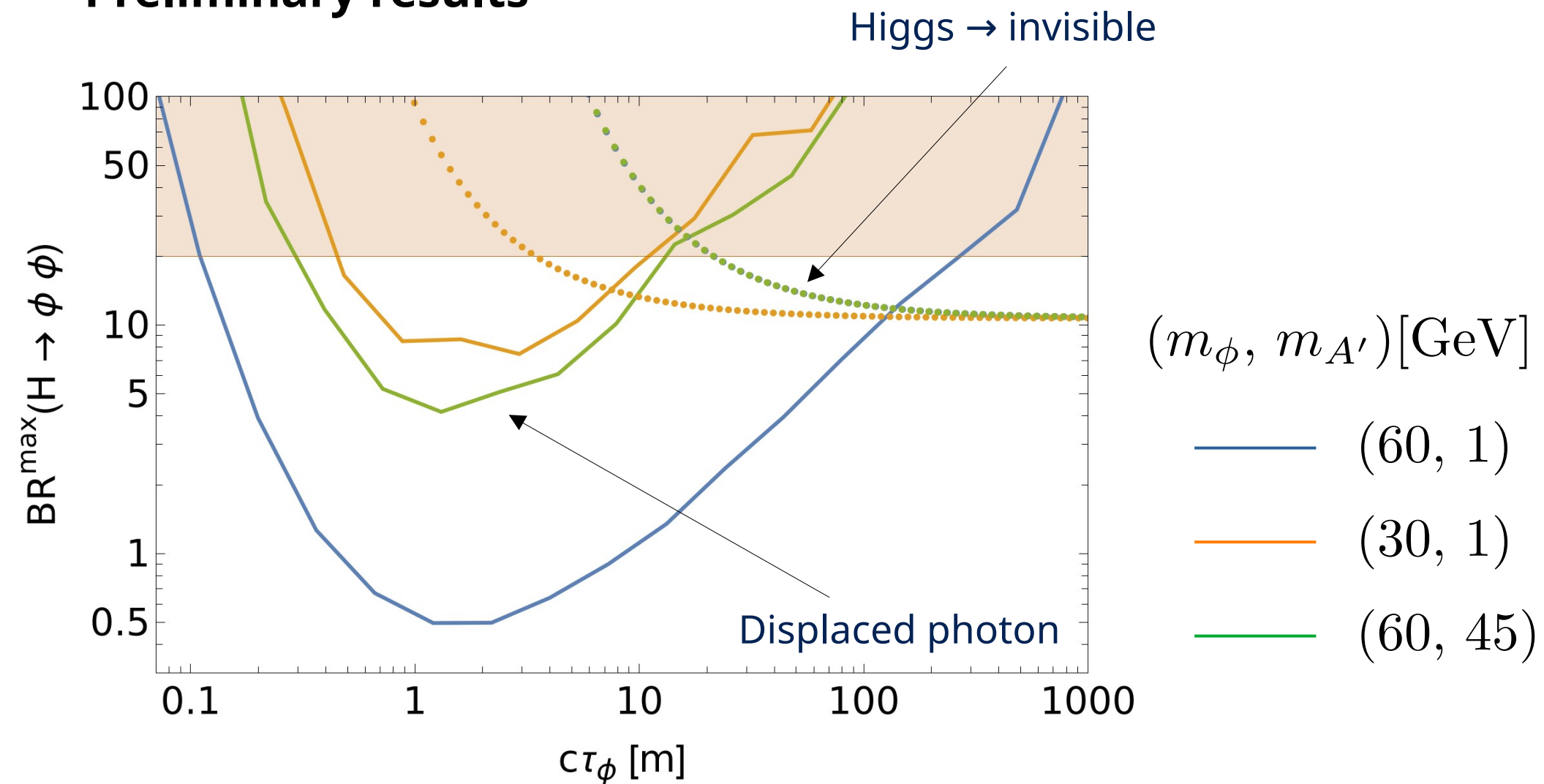
$$\text{BR}(H \rightarrow \text{inv}) < 15\%$$

## Preliminary results

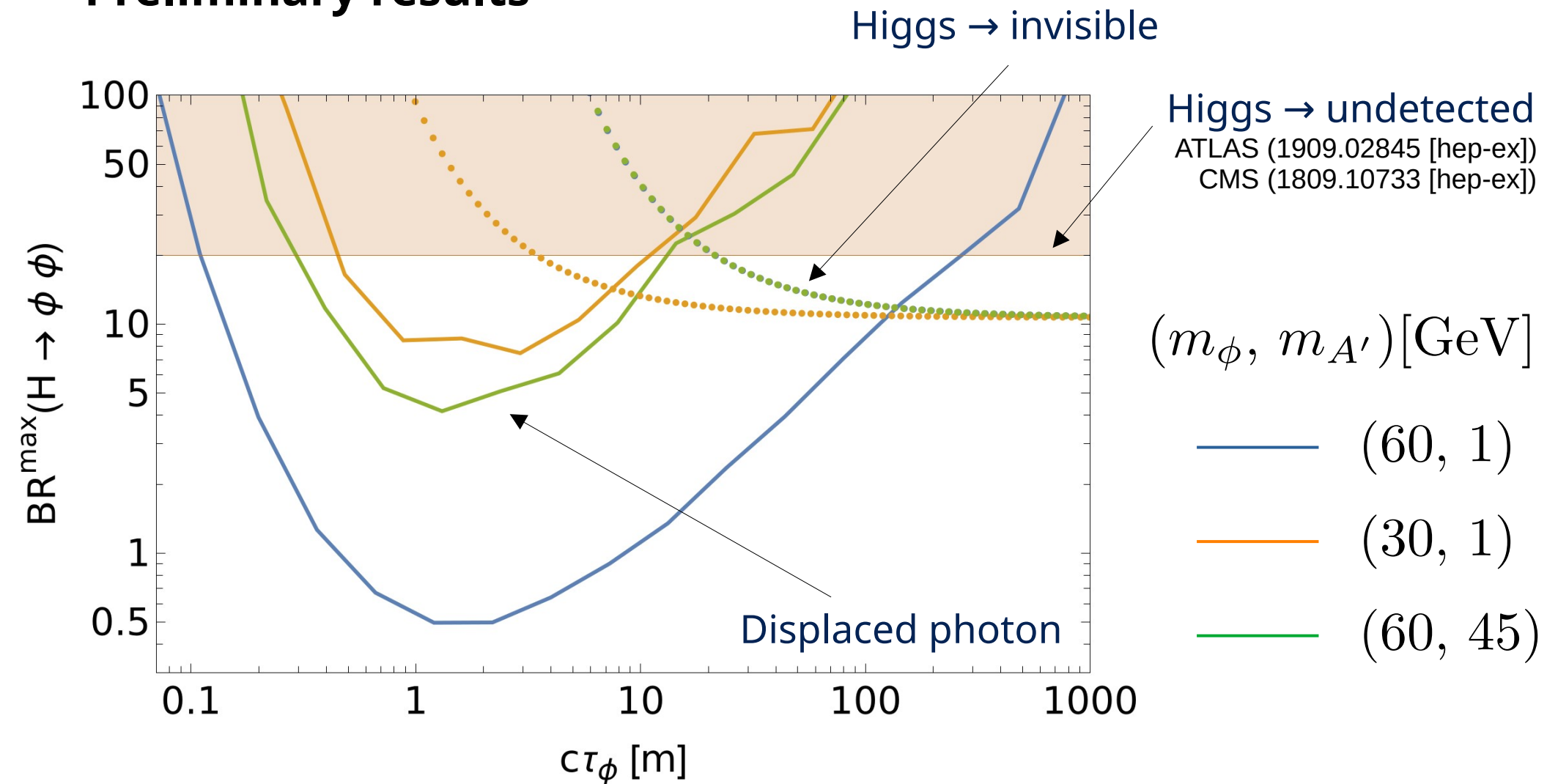




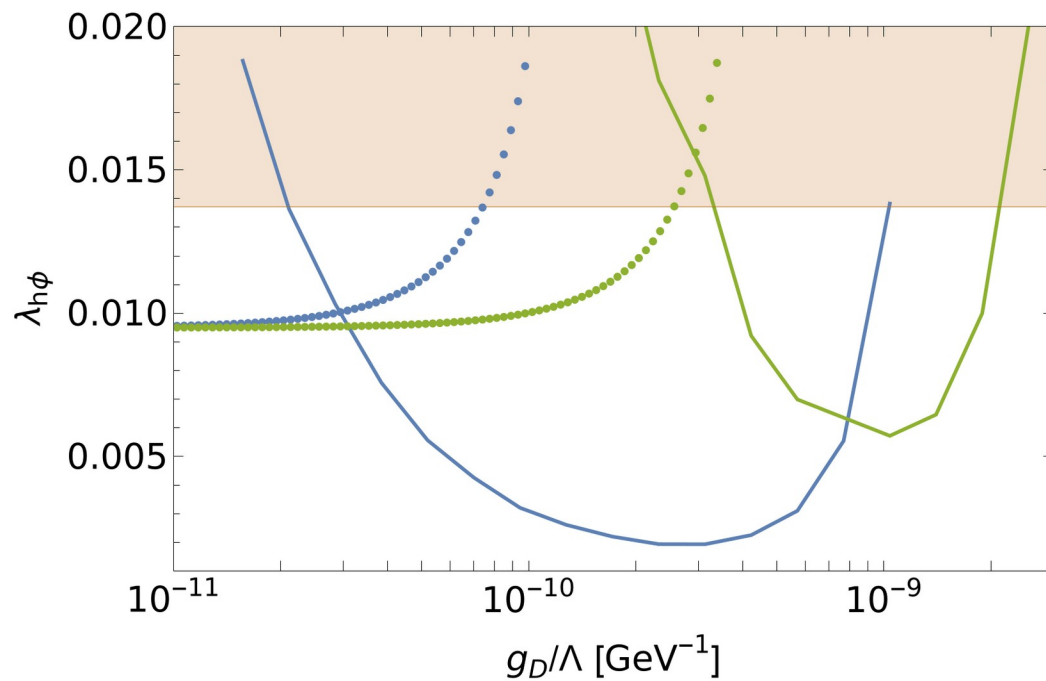
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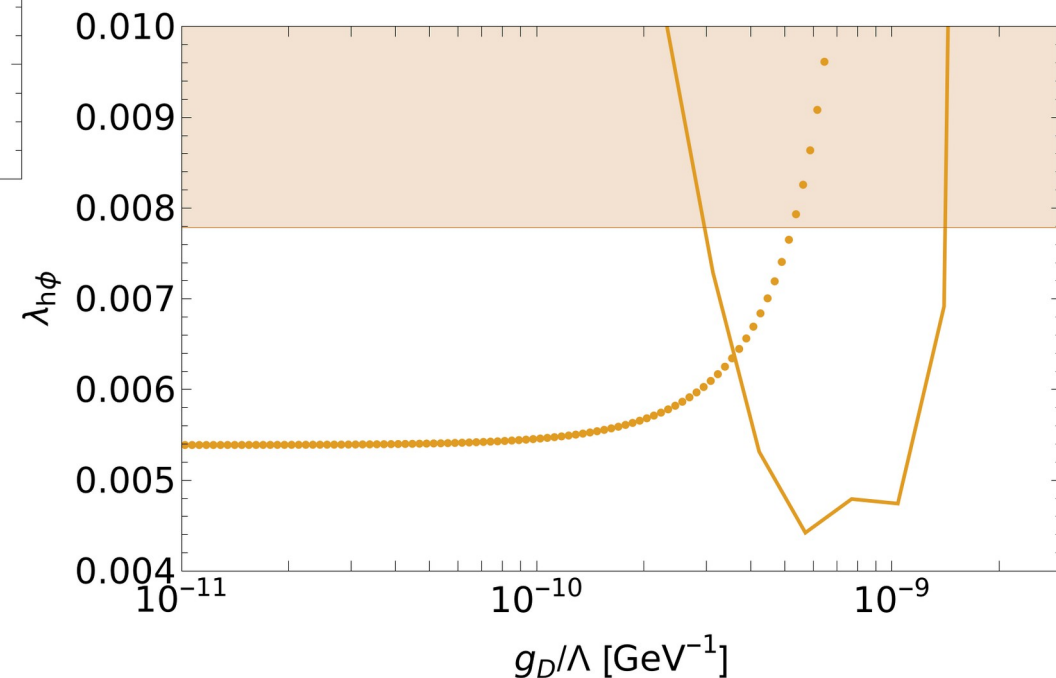
## Preliminary results



— (60, 1)  
— (60, 45)

$(m_\phi, m_{A'})$  [GeV]

— (30, 1)



## Implications for Dark Photon dark matter

To be continued! Some musings on the topic:

- The dark photon is a **FIMP**, as  $g_D$  is very small.
- The scalar should be **thermal**, depends on  $\lambda_{h\phi}$ .
- It seems  $\phi$  **decays** dominate freeze-in production of  $A'$ . Then, as long as  $g_D$  is small, the value of  $\lambda_{h\phi}$  is crucial for prediction. Maybe also **SuperWIMP**?
- Region of parameter space seems to **overproduce** dark matter, assuming **Standard Cosmology**.
- This works seems to point in the direction of displaced photon searches bounding **Non-Standard Cosmology** scenarios.

## Take-home messages

- Searches for displaced photons can place bounds on models with LLPs decaying into photons and MET.
- These searches have largest sensitivity for  $c\tau \sim \mathcal{O}(1 \text{ m})$
- For larger lifetimes, Higgs  $\rightarrow$  invisible searches present better bounds.
- Recast of search in the context of d=5 Seesaw Portal place strong bounds on coefficients of effective operators  $\sim \mathcal{O}(10^{-5} \text{ GeV}^{-1})$
- Recast of search in the context of dark photon with scalar portal places bounds  $\lambda_{h\phi} \sim \mathcal{O}(10^{-3})$
- Possibility of dark photon generating the relic density in this region of parameter space still under consideration. Stay tuned!

## Placing Bounds on New Physics with Displaced Photons



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**UNIVERSIDAD**  
**CATÓLICA**  
DEL PERÚ

# ¡Gracias!

Funded by:



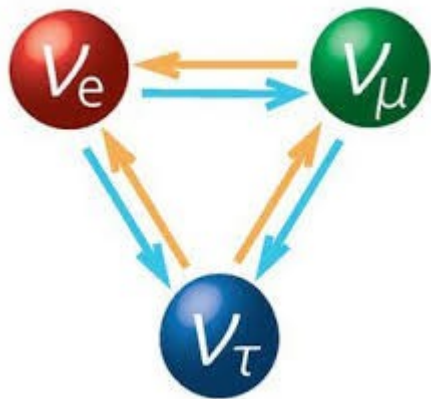
**PUCP**



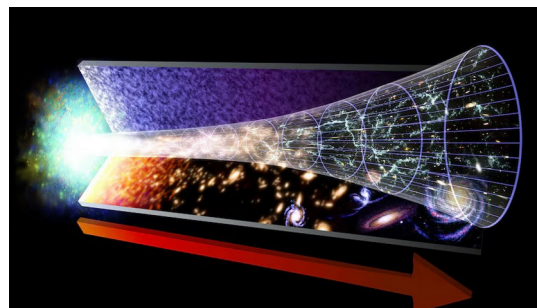
# Backup

## Why expand the Standard Model?

- No mechanism for neutrino masses
- No dark matter candidate
- No explanation for baryon – antibaryon asymmetry
- The Higgs is theoretically uncomfortable (naturalness, vacuum stability)

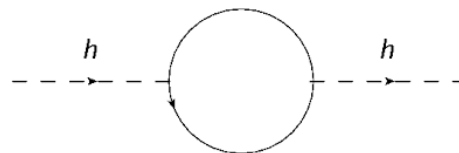


Neutrino oscillations



Where did antimatter go?

Loop corrections  
to Higgs mass



Mass and gas in *Bullet cluster*

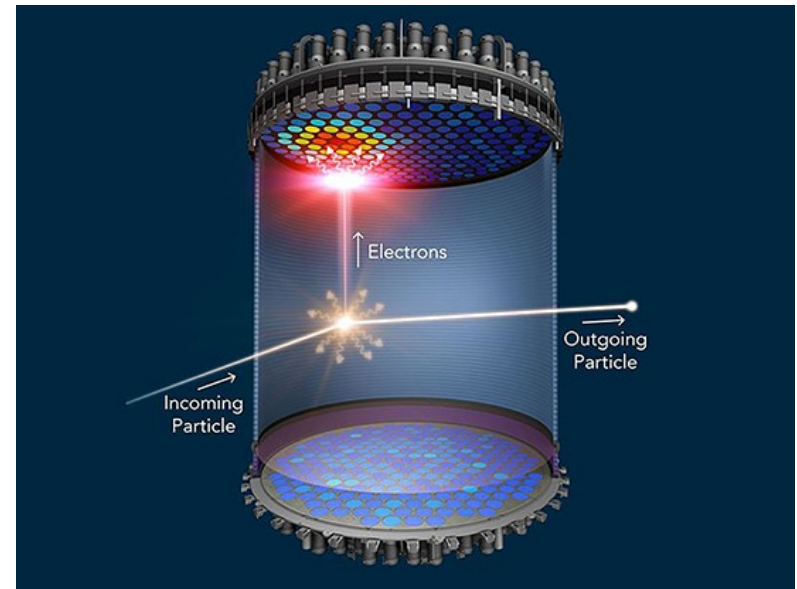


## Where are we looking for New Physics?

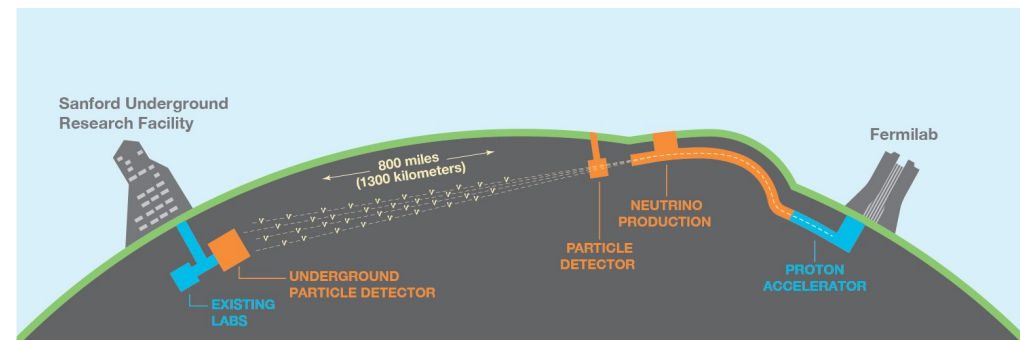
- Colliders
- Neutrino oscillation experiments
- Dark matter experiments
- So many others!



Large Hadron Collider



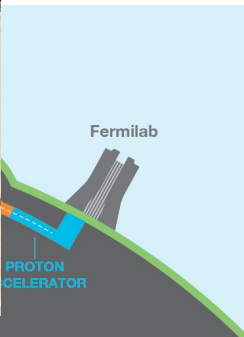
XENON1T Experiment



DUNE experiment

## Where are we looking for New Physics?

- Colliders
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- Dark ma
- So many



DUNE experiment

## Recasting the Displaced Photon search

- Events are generated in MadGraph, with parton shower and hadronization by Pythia.
- Need to calculate  $t_\gamma$  and  $|\Delta z_\gamma|$  from truth-level information. Must include this info in HepMC files, and pass on to Delphes for detector simulation.
- Arrival time  $t_\gamma$ :  
Simulated “prompt LLPs”, and calculated arrival time as a function of pseudorapidity. Subtract this from long-lived case.

## Recasting the Displaced Photon search

- Photon non-pointing variable: separation between primary vertex and projected origin of photon, along the beamline.

$$|\Delta z_\gamma| = r_{\gamma_z^0} - \frac{p_{\gamma_z}}{p_{\gamma_T}^2} (r_{\gamma_x^0} p_{\gamma_x} + r_{\gamma_y^0} p_{\gamma_y})$$

- Not exactly what ATLAS does. Measurement actually relies on information of first two layers of ECAL, disregards information on  $\phi$ . Differences are fortunately small!

## Post-Delphes Cuts

- Apply proper isolation criteria, based on tracks and energy deposits.
- Applied gaussian smear on  $|\Delta z_\gamma|$ , use this for efficiency.
- Apply  $p_T$  and  $\eta$  cuts on photons, separate into 1 or multi-photon channels.
- Apply ID efficiency on electrons, and ID efficiency + isolation on muons.
- Apply  $p_T$  and  $\eta$  cuts on electrons, muons, jets.
- Implement overlap removal: photons > electrons > jets > muons.
- Evaluate if trigger is fired.
- Apply gaussian smear on  $t_y$ .
- Assign to signal region if  $MET > 50$  GeV.
- Separate in different  $|\Delta z_\gamma|$  categories, with a binning based on  $t_y$ .
- Apply statistical analysis on largest  $t_y$  bin, on largest  $|\Delta z_\gamma|$  category.

## Type-I Seesaw

Type I Seesaw is probably most popular mechanism for neutrino masses

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + Y_\nu \left( \bar{L} \cdot \tilde{\phi} \nu_R \right) + \frac{1}{2} M_R (\bar{\nu}_R \nu_R^c) + \text{h.c.}$$



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Mass matrix:

$$M_\nu = \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \quad m_D = \frac{1}{\sqrt{2}} v Y_\nu^*$$





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Mass matrix:

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Light neutrino mass matrix:

$$m_\nu \sim m_D M_R^{-1} m_D^T$$





## Couplings

$$\mathcal{L}_W = \frac{g}{\sqrt{2}} W_\mu^- \bar{\ell}_a \gamma^\mu U_{ai} P_L n_i + h.c.$$

$$\begin{aligned} \mathcal{L}_Z = & \frac{g}{4c_W} Z_\mu \bar{n}_i \gamma^\mu (C_{ij} P_L - C_{ij}^* P_R) n_j \\ & - \frac{s_W}{\Lambda} (\partial_\mu Z_\nu - \partial_\nu Z_\mu) \bar{n}_i \sigma^{\mu\nu} [(\alpha'_{NB})_{ij} P_L - (\alpha'^*_{NB})_{ij} P_R] n_j \end{aligned}$$

$$\mathcal{L}_\gamma = \frac{c_W}{\Lambda} (\partial_\mu A_\nu - \partial_\nu A_\mu) \bar{n}_i \sigma^{\mu\nu} [(\alpha'_{NB})_{ij} P_L - (\alpha'^*_{NB})_{ij} P_R] n_j$$

$$\begin{aligned} \mathcal{L}_h = & -\frac{1}{v} h \bar{n}_i \left[ \frac{1}{2} (C_{ij} m_{n_j} + C_{ij}^* m_{n_i}) - \frac{v^2}{\Lambda} (\alpha'^*_{N\phi})_{ij} \right] P_R n_j \\ & -\frac{1}{v} h \bar{n}_i \left[ \frac{1}{2} (C_{ij} m_{n_i} + C_{ij}^* m_{n_j}) - \frac{v^2}{\Lambda} (\alpha'_{N\phi})_{ij} \right] P_L n_j \end{aligned}$$

$$\mathcal{L}_{hh} = \frac{1}{2\Lambda} h^2 \bar{n}_i [(\alpha'_{N\phi})_{ij} P_L + (\alpha'^*_{N\phi})_{ij} P_R] n_j$$

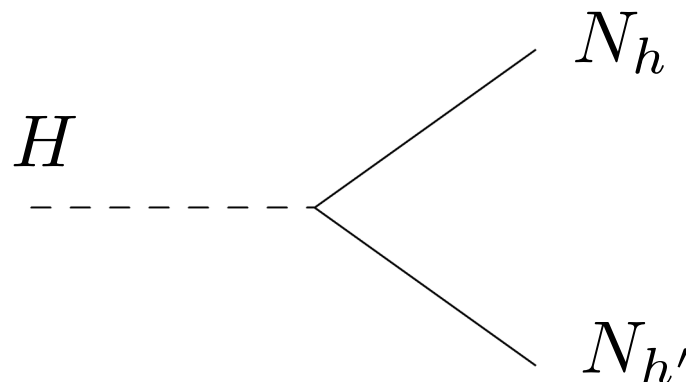
## Dimension-5 Type-I Seesaw Portal

New interactions:

$$\mathcal{L}_{hNN} = \frac{v}{\Lambda} H \bar{N}_h \left[ (\alpha'_{N\phi})_{hh'} P_R + (\alpha'_{N\phi})_{hh'} P_L \right] N_{h'}$$

$$(\alpha'_{N\phi})_{hh'} = U_{sh} (\alpha_{N\phi})_{ss'} U_{s'h'}$$

Allows the decay of the Higgs into two heavy neutrinos (tremendously suppressed in standard Seesaw)



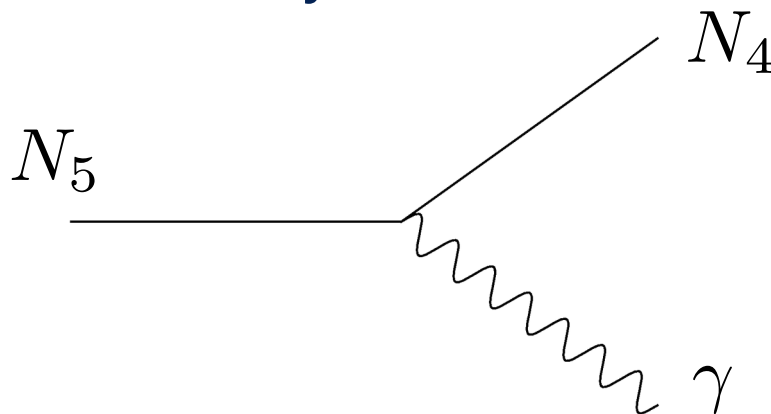
## Dimension-5 Type-I Seesaw Portal

New interactions:

$$\mathcal{L}_{\gamma NN} = \frac{c_W}{\Lambda} F_{\mu\nu} \bar{N}_4 \sigma^{\mu\nu} [(\alpha'_{NB})_{45} P_L - (\alpha'_{NB})_{45}^* P_R] N_5 + h.c.$$

$$(\alpha'_{NB})_{45} = U_{s4} (\alpha_{NB})_{ss'} U_{s'5}$$

Allows the decay of one heavy neutrino into another heavy neutrino and a photon. Usually dominates decay width.



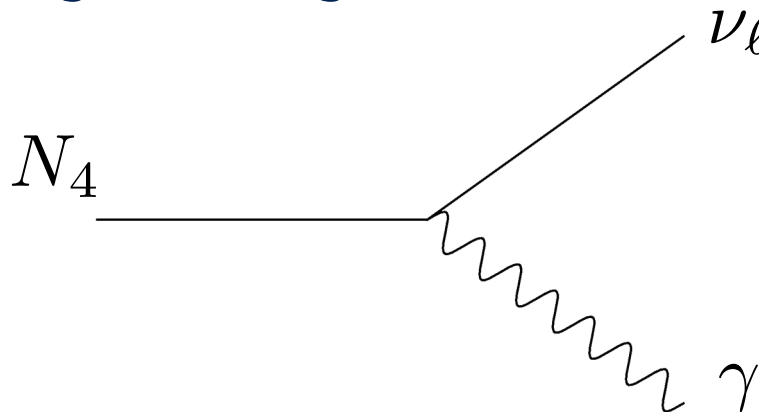
## Dimension-5 Type-I Seesaw Portal

New interactions:

$$\mathcal{L}_{\gamma N \nu} = \frac{c_W}{\Lambda} F_{\mu\nu} \bar{\nu}_\ell \sigma^{\mu\nu} [(\alpha'_{NB})_{\ell 4} P_L - (\alpha'^*_{NB})_{\ell 4} P_R] N_4 + h.c.$$

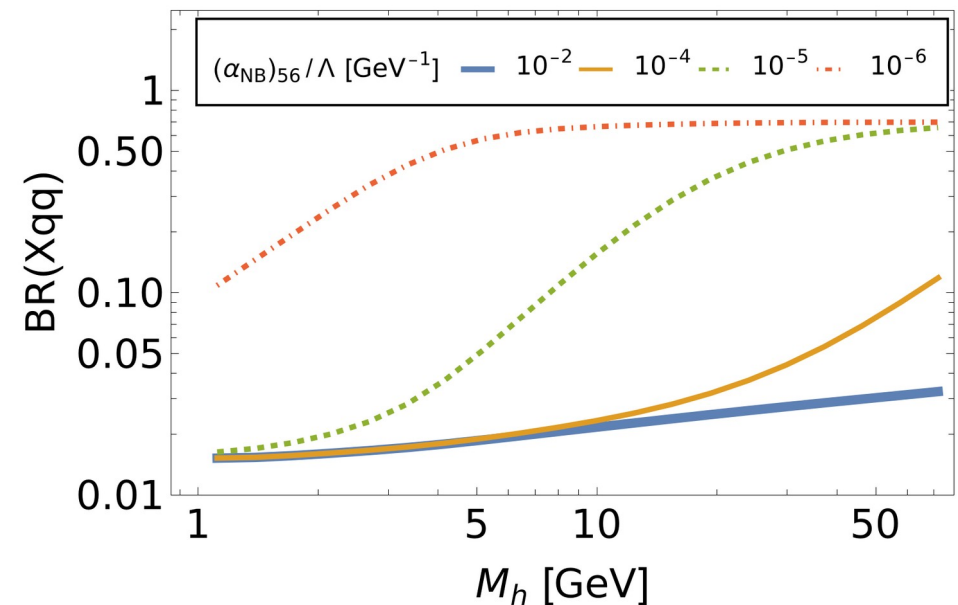
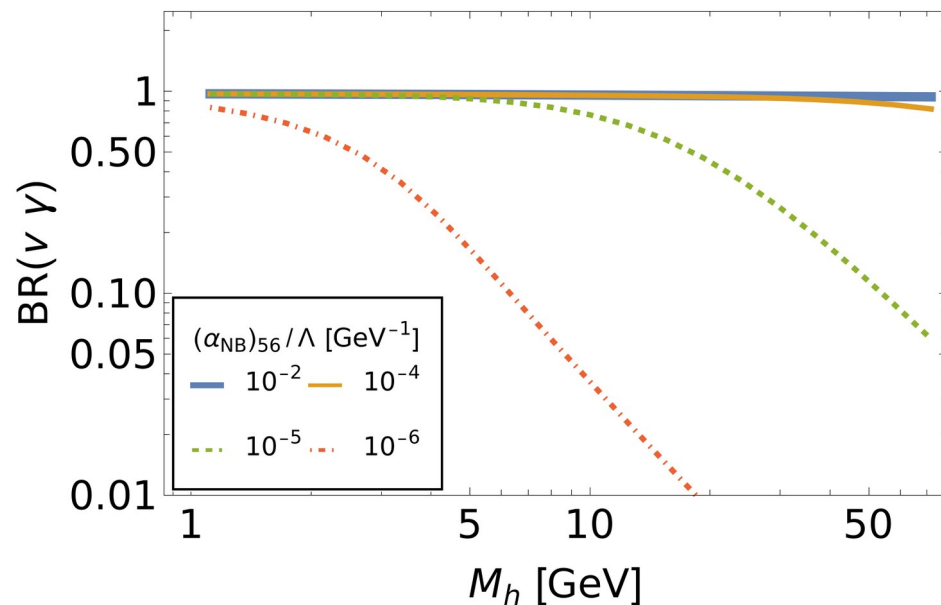
$$(\alpha'_{NB})_{\ell 4} = U_{s\ell} (\alpha_{NB})_{ss'} U_{s'4}$$

Allows the decay of a heavy neutrino into light neutrino and a photon.  
Depends on “sterile-light” mixing.



## Modifications to Heavy Neutrino Width

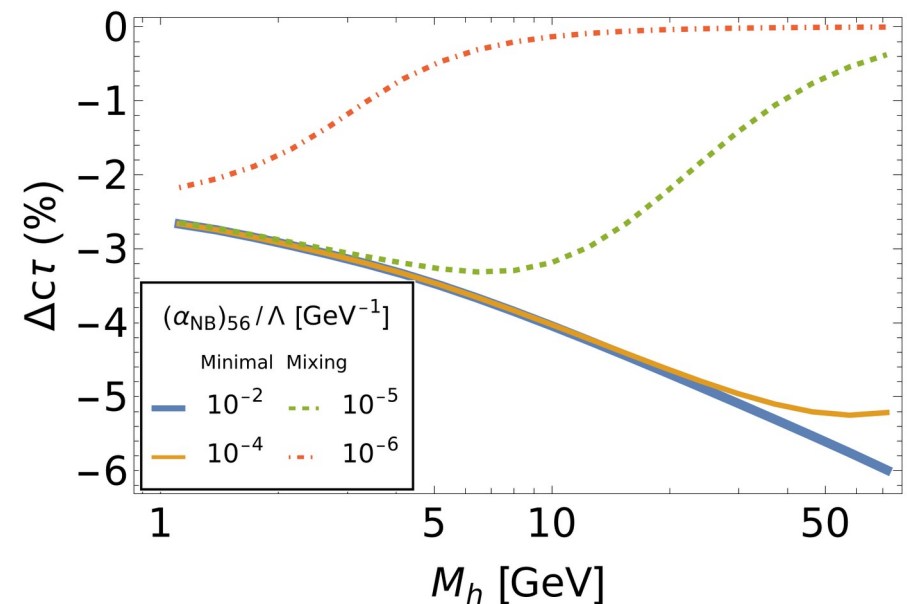
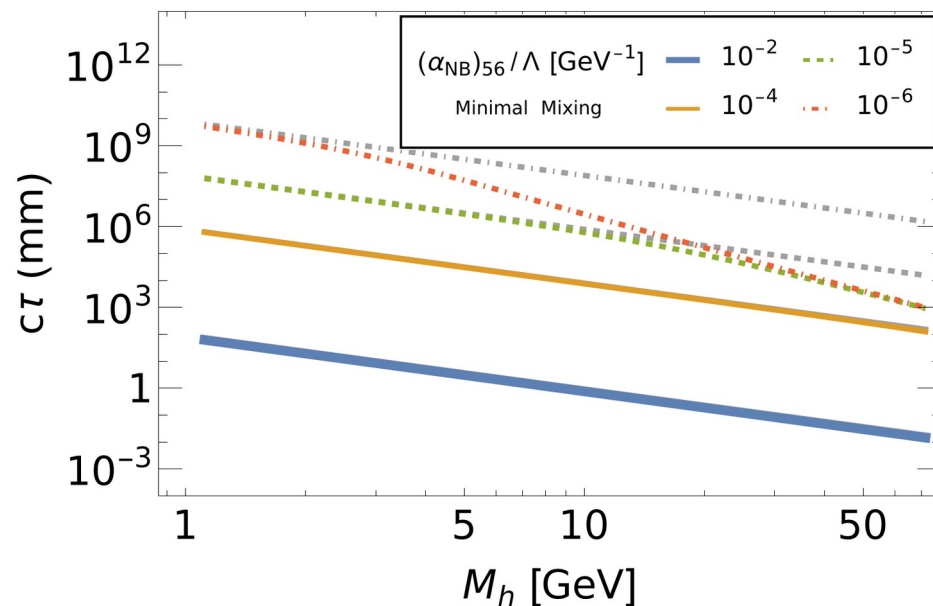
New branching ratios:



Photon +  $\nu$  final state will usually dominate over small masses, but on the GeV regime the other decays are also relevant.

## Modifications to Heavy Neutrino Width

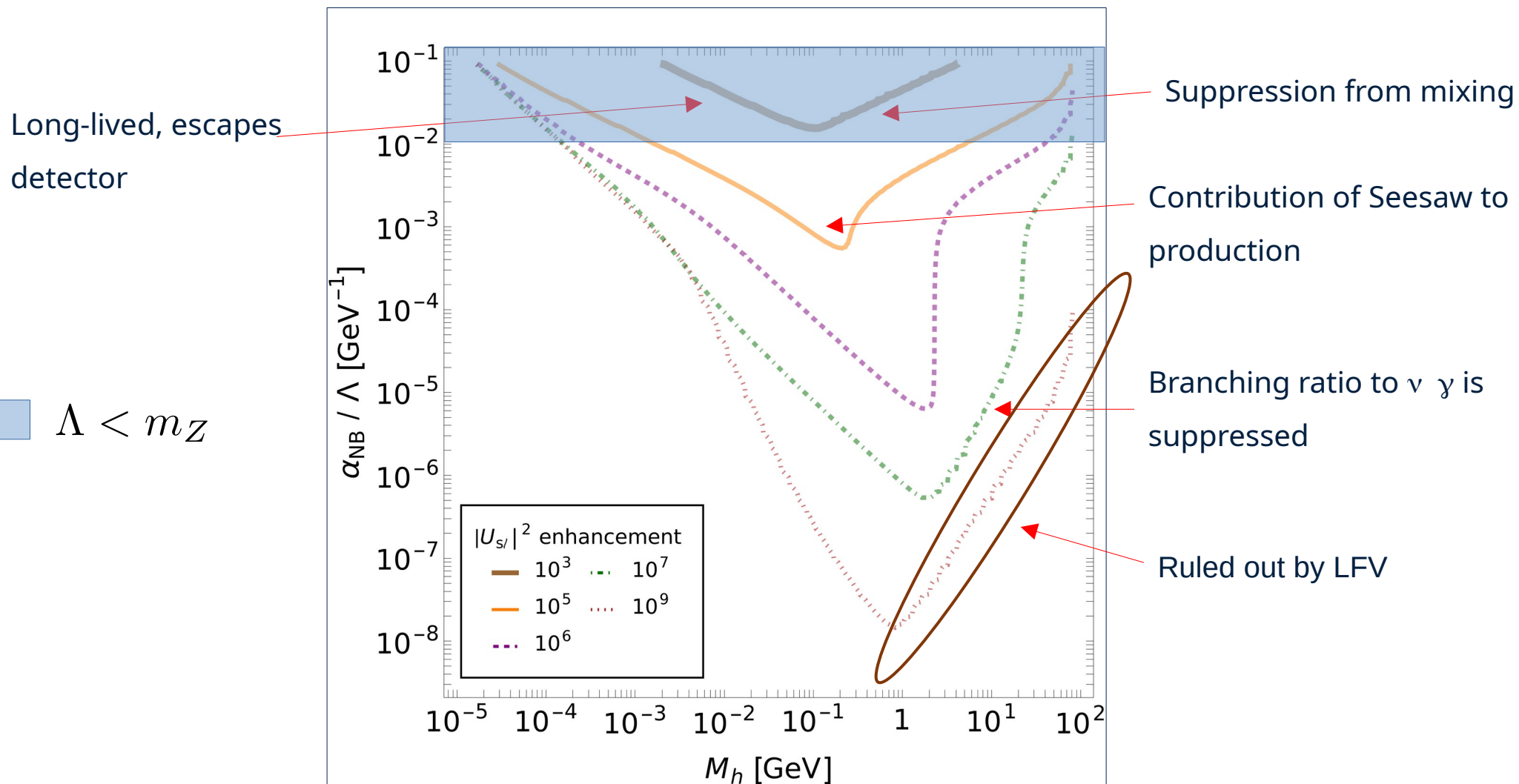
New lifetimes:



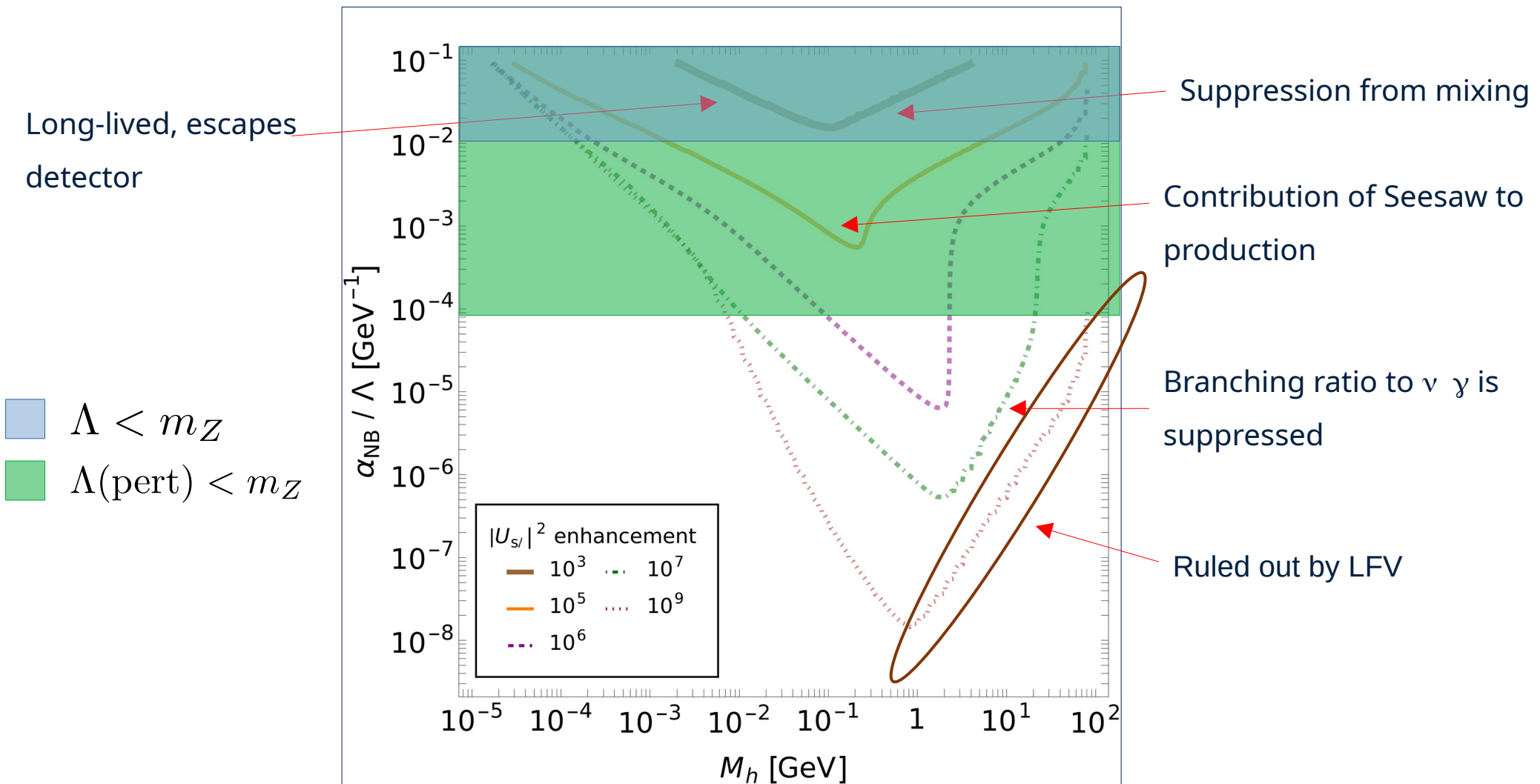
It is important to include at least standard Seesaw three body decays!!

Modifications to three-body widths have small impact, might be relevant after a putative discovery.

## Constraints from LEP

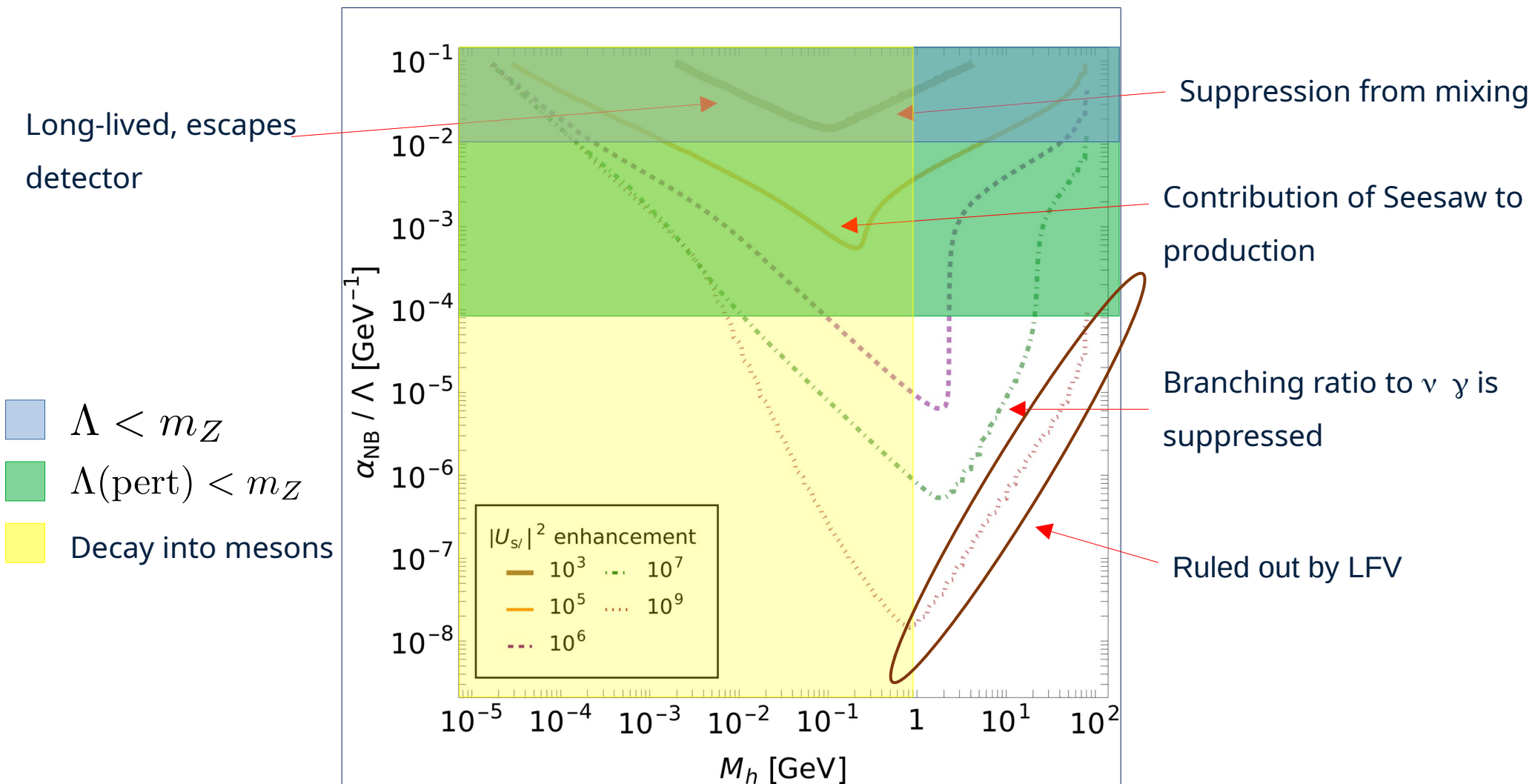


## Constraints from LEP





## Constraints from LEP

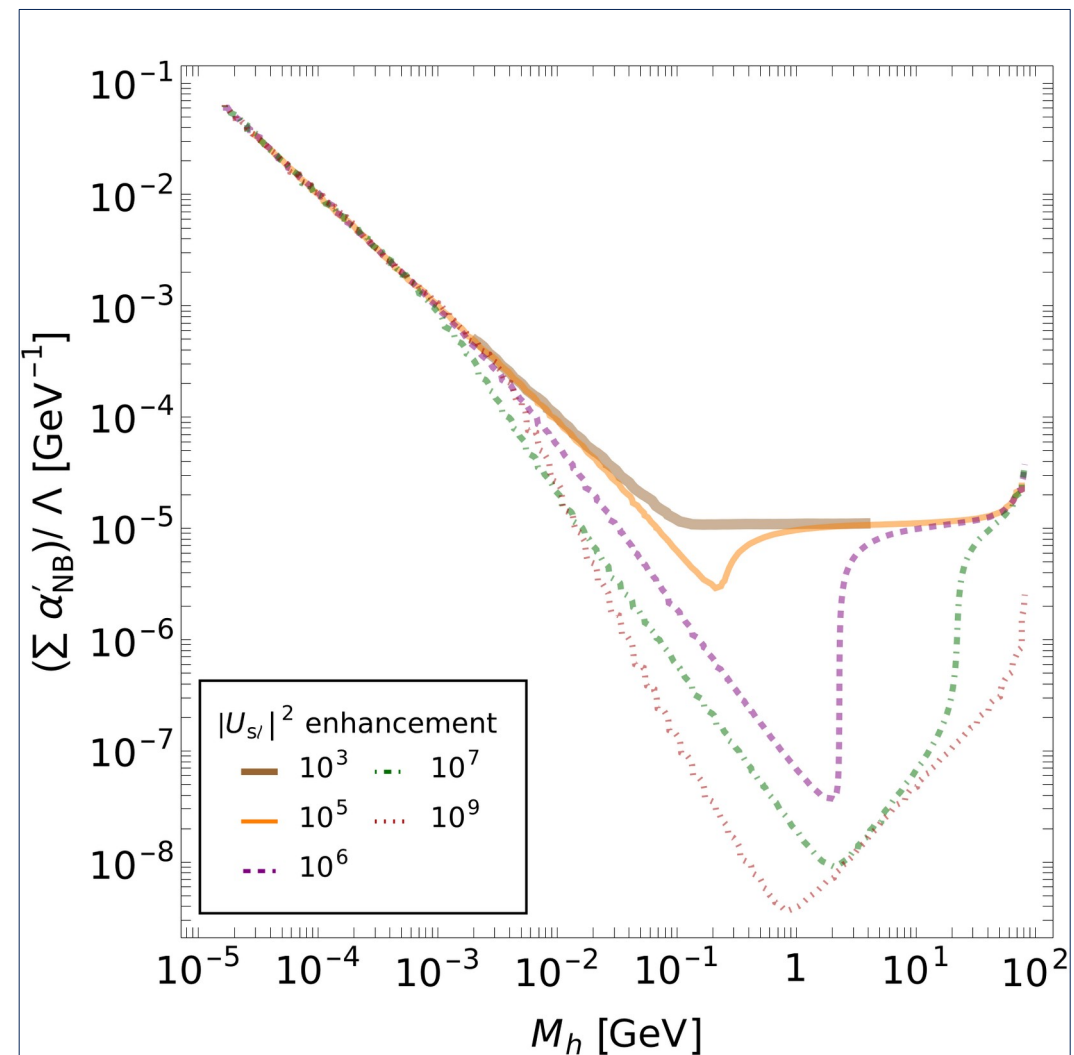


## Constraints from LEP

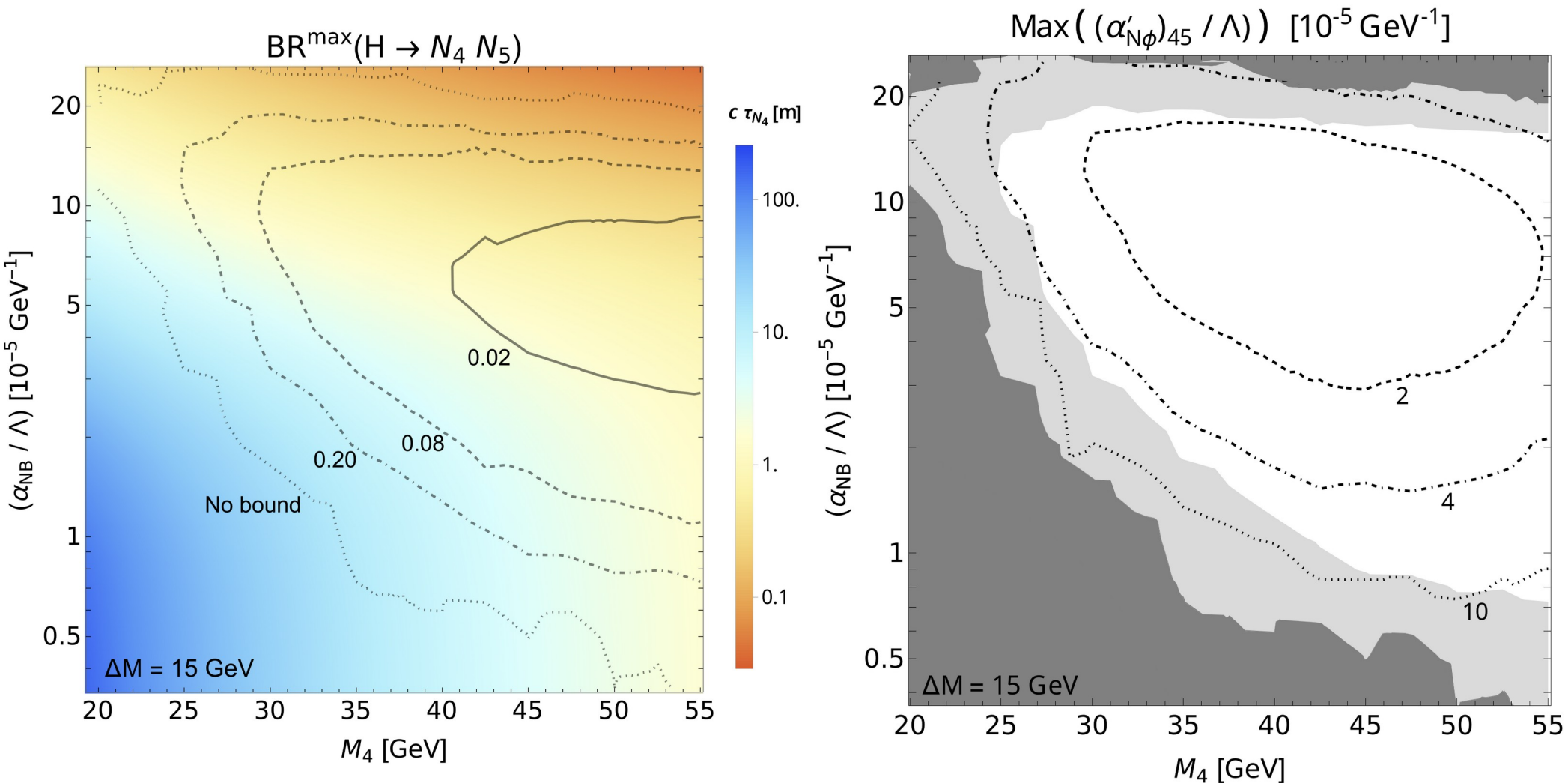
The bound can be written in terms of  $\alpha'_{NB}$ , so can be applied to  $d=6$  operator.

This has been done before, but not in combination with Seesaw contribution.

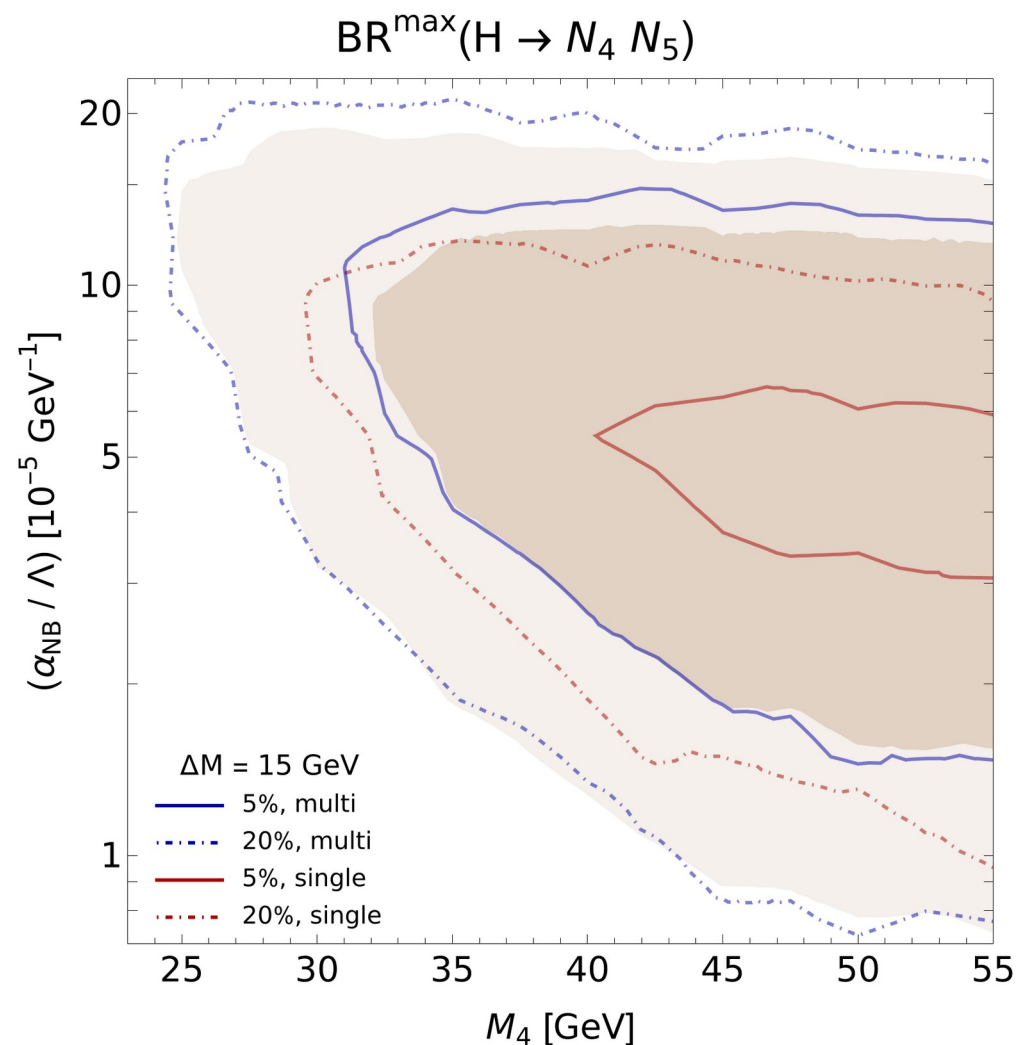
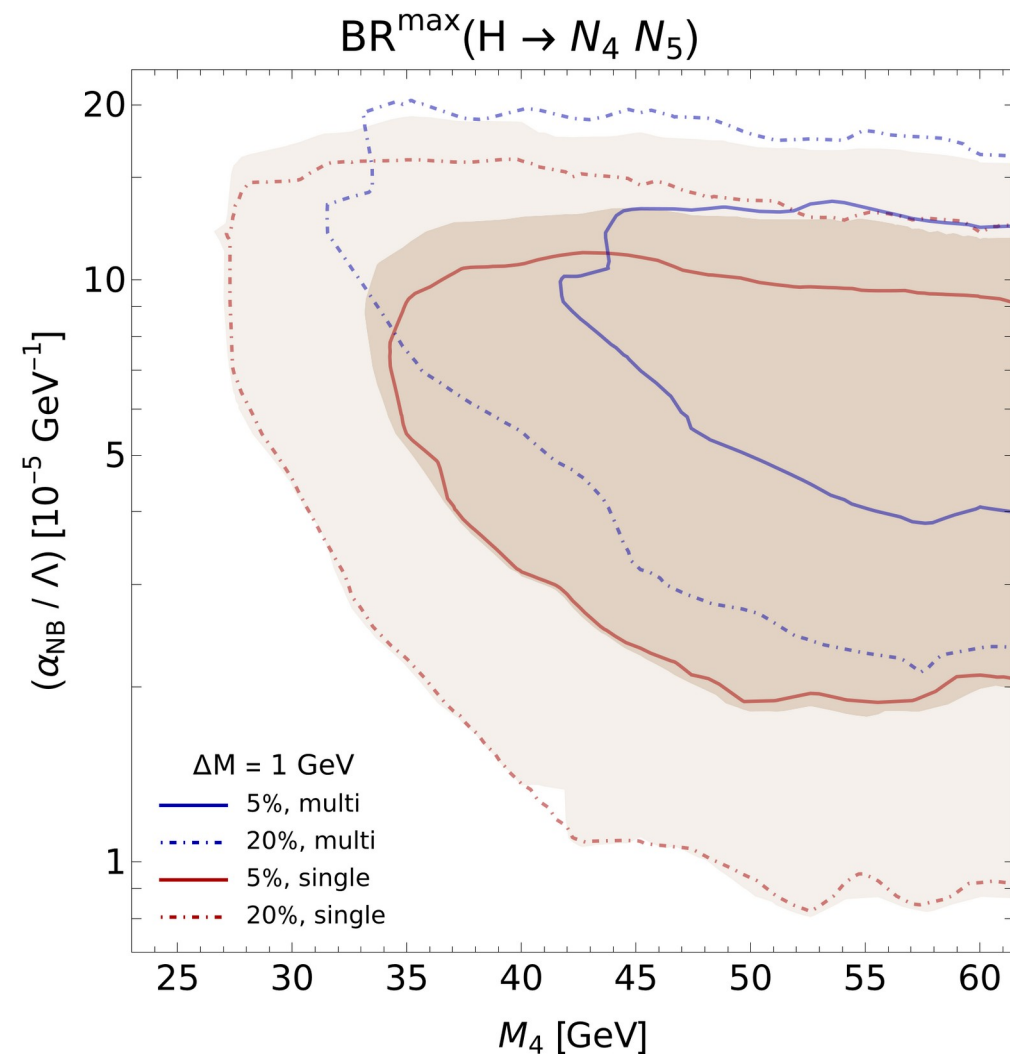
$$(\alpha'_{NB})_{\ell h} \equiv U_{s\ell} (\alpha_{NB})_{ss'} U_{s'h}$$



## Bounds on $\alpha_{N\phi}$ !



## 1+ channel is not always the most sensitive one!



## Application of Higgs $\rightarrow$ invisible limit

- One must take into account the probability of the LLP decaying outside the detector.

$$\text{Prob}(x_1, x_2) = \exp \left[ -\frac{x_1}{\gamma_{\text{rel}} \beta_{\text{rel}} c \tau_\chi} \right] - \exp \left[ -\frac{x_2}{\gamma_{\text{rel}} \beta_{\text{rel}} c \tau_\chi} \right]$$

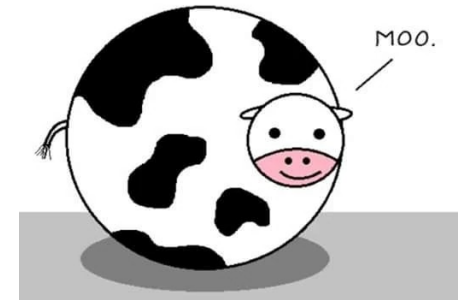
$$\gamma_{\text{rel}} = E_\chi / m_\chi \quad \beta_{\text{rel}} = |\vec{p}_\chi| / E_\chi$$

$$\tau_\chi = \text{Lifetime of } \chi \text{ in rest frame.}$$

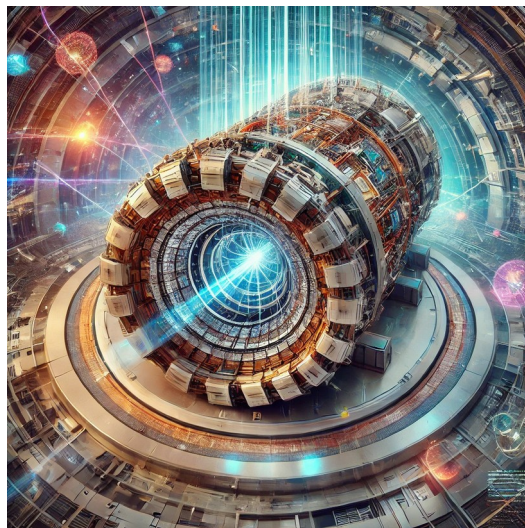


## Simplifications

- Take a spherical detector.

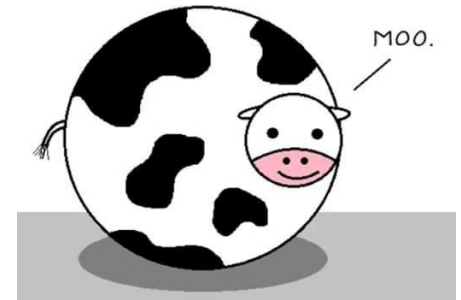


$$\text{BR}(H \rightarrow \chi\bar{\chi}) \times [\text{Prob}(L_{\text{det}}, \infty)]^2 < \text{BR}_{\text{limit}}$$



## Simplifications

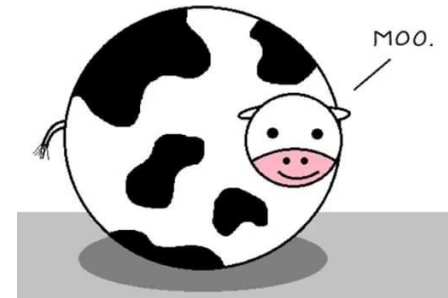
- Take a spherical detector.
- Assume the Higgs decays at rest.



$$E_\chi = m_H/2 \quad \Rightarrow \quad \gamma_{\text{rel}} \beta_{\text{rel}} = \sqrt{\frac{m_H^2}{4m_\chi^2} - 1}$$

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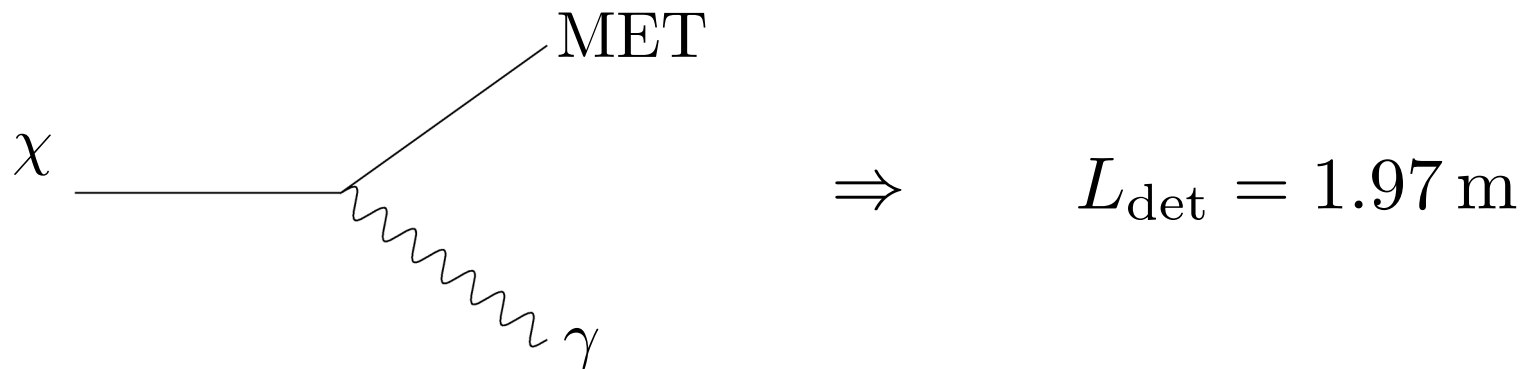
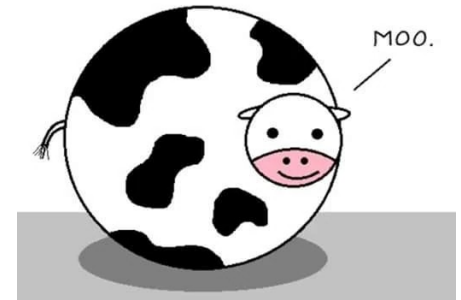
So far:

$$\text{BR}(H \rightarrow \chi\bar{\chi}) \times \exp\left[-\frac{2L_{\text{det}}}{\gamma_{\text{rel}} \beta_{\text{rel}} c \tau_\chi}\right] < \text{BR}_{\text{limit}}$$



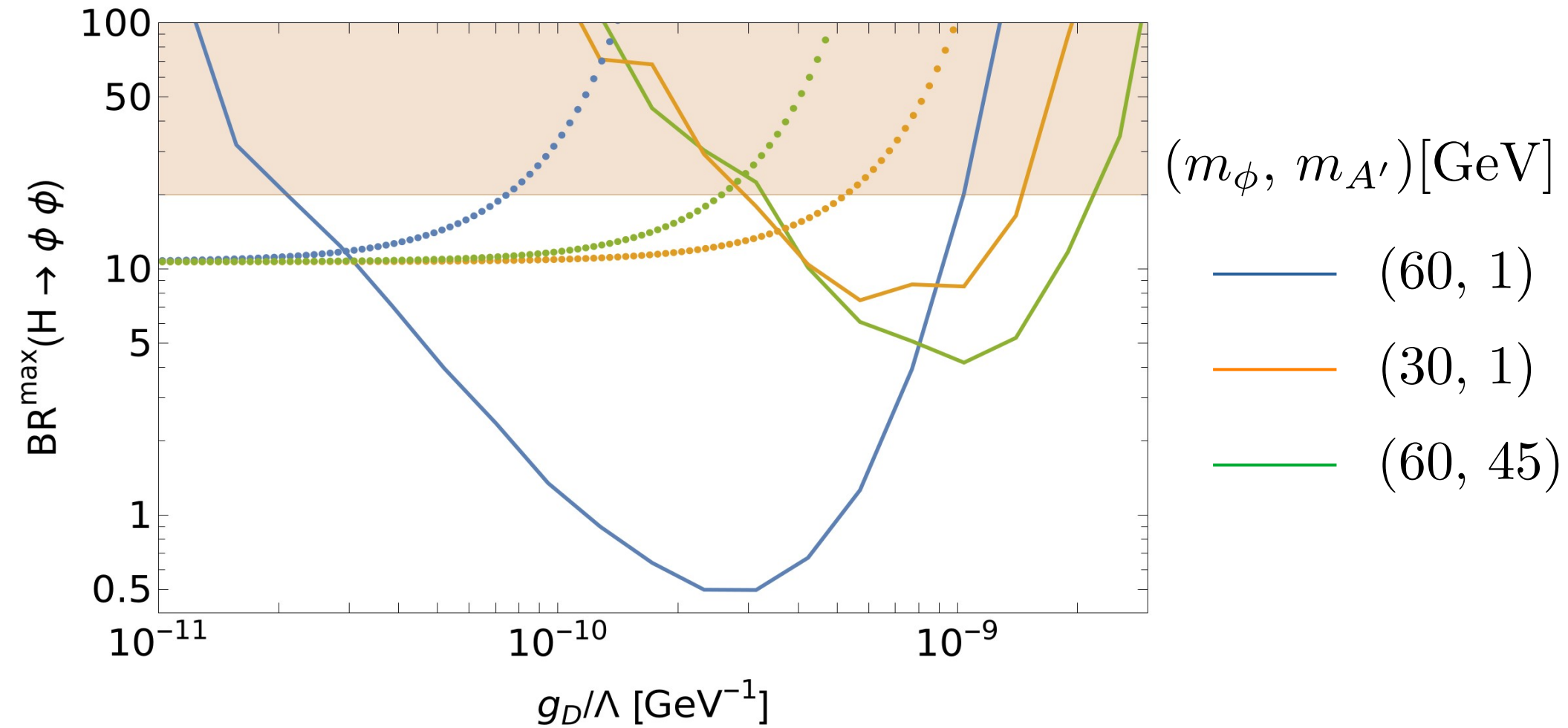
## Simplifications

- Take a spherical detector.
- Assume the Higgs decays at rest.
- Assume the decay is invisible if the LLP decays after the ECAL.



$$\text{BR}(H \rightarrow \chi \bar{\chi}) \times \exp \left[ -\frac{2L_{\text{det}}}{\gamma_{\text{rel}} \beta_{\text{rel}} c \tau_{\chi}} \right] < \text{BR}_{\text{limit}}$$

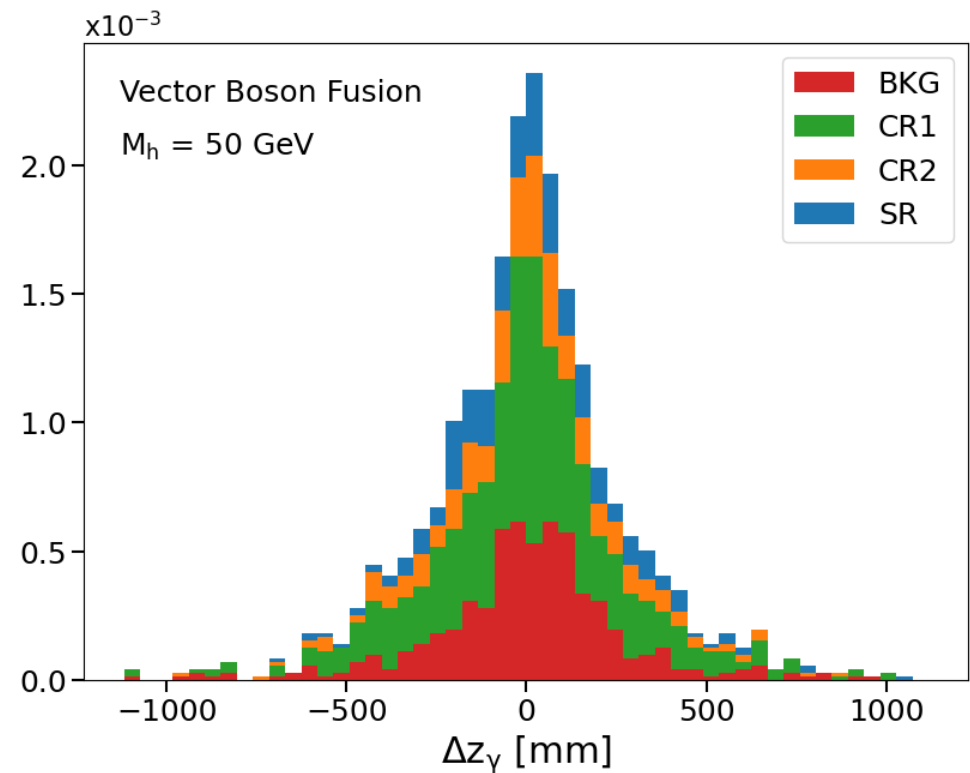
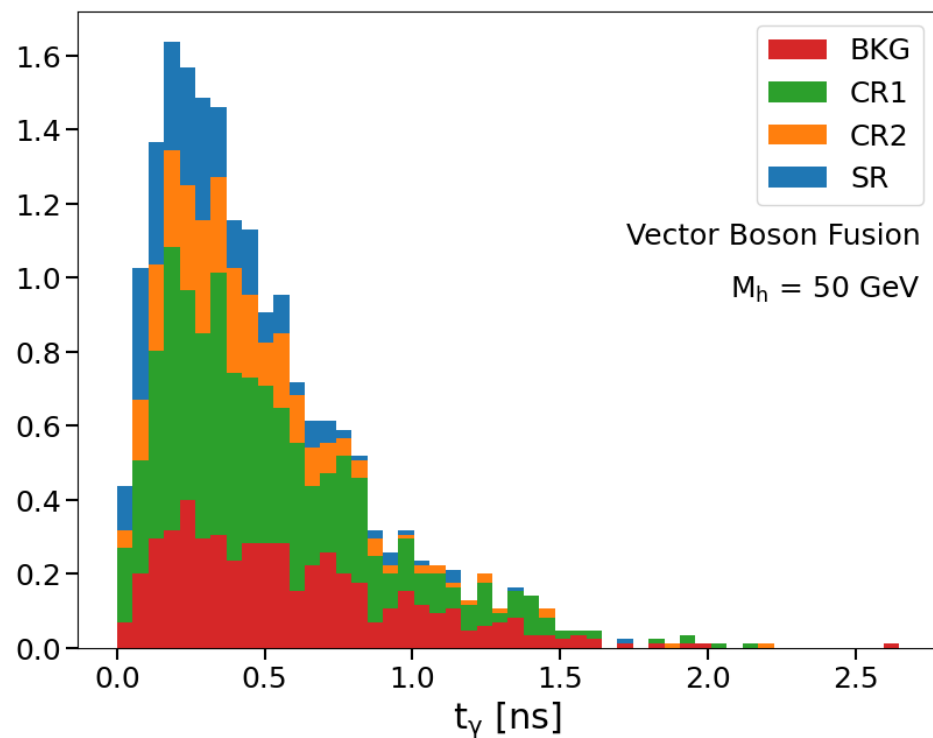
## Preliminary results



## 8 TeV Search, for $20.3 \text{ fb}^{-1}$

- Two “loose” photons with energy larger than 50 GeV.
- One of the photons must be in barrel region.
- Isolation criteria: no deposits larger than 4 GeV within  $\Delta R = 0.4$ .
- If more than one photon in barrel region, use the one with largest  $t_y$ .
- Define background, control and signal region depending on MET.
- Use bin-based analysis considering  $t_y$  and  $|\Delta z_\gamma|$

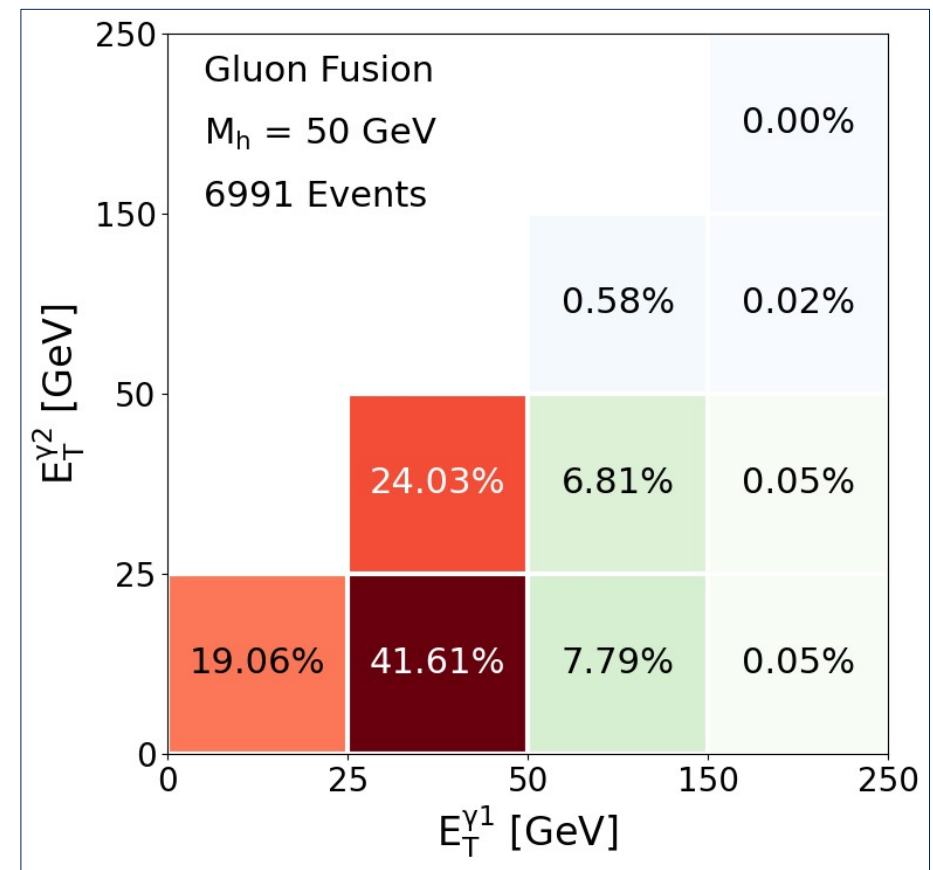
## 8 TeV Search, for $20.3 \text{ fb}^{-1}$



## 8 TeV Search

Old searches for non-pointing photons were designed for heavy particles.

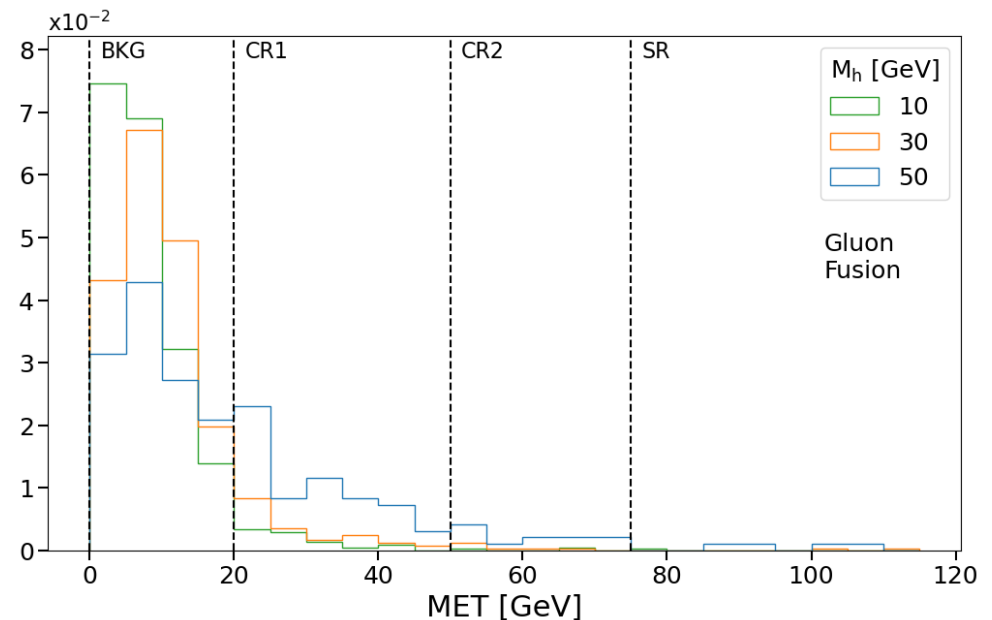
- Triggered using high pT photons.



## 8 TeV Search

Old searches for non-pointing photons were designed for heavy particles.

- Triggered using high  $p_T$  photons.
- Non optimal signal regions.



## 8 TeV Search

Old searches for non-pointing photons were designed for heavy particles.

- Triggered using high  $p_T$  photons.
- Non optimal signal regions.

Even if a photon pair from long-lived  $N_h$  passed the energy cuts, and even if they also had large  $t_y$  and  $|\Delta z_\gamma|$ , they are likely to be assigned to the background or control region. Thus, this strongly suggests the 8 TeV search is not optimal for studying our model.

