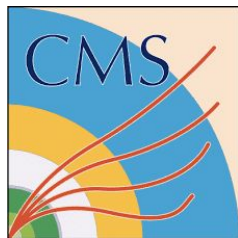


Luminosity Measurement at the CMS Experiment



On Behalf of the CMS Collaboration & BRIL team

Ashish Sehrawat / Universidad de Sonora
RADPyC2023 (12-14 June 2023)
June 13, 2023

OUTLINE

- Introduction
 - Luminosity
 - Precision importance
 - Luminometers
 - Calibration
- Latest results on luminosity
 - 2015/16
 - 2017/18
 - 2022
- Luminosity from Z bosons

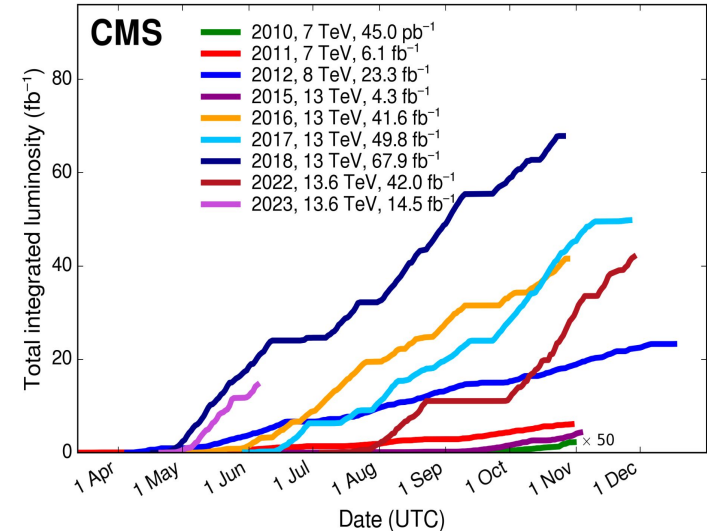
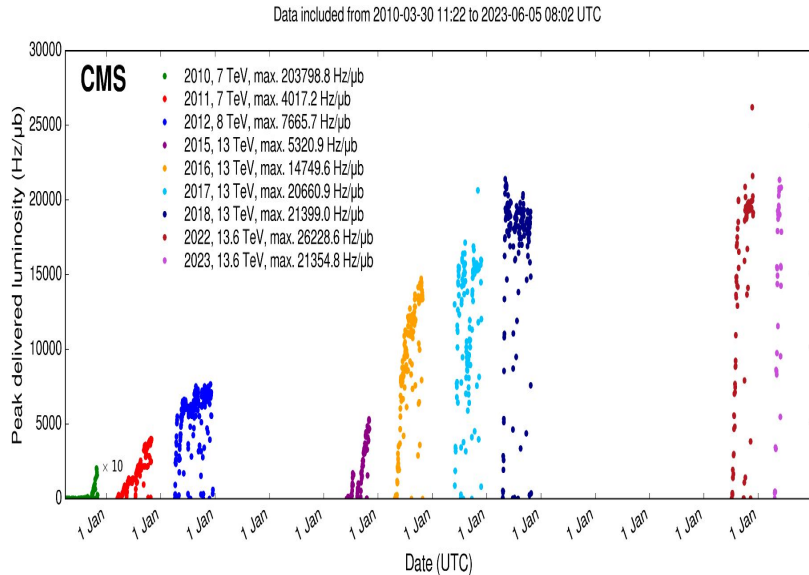
Luminosity

<https://twiki.cern.ch/twiki/bin/view/CMSPublic/LumiPublicResults>

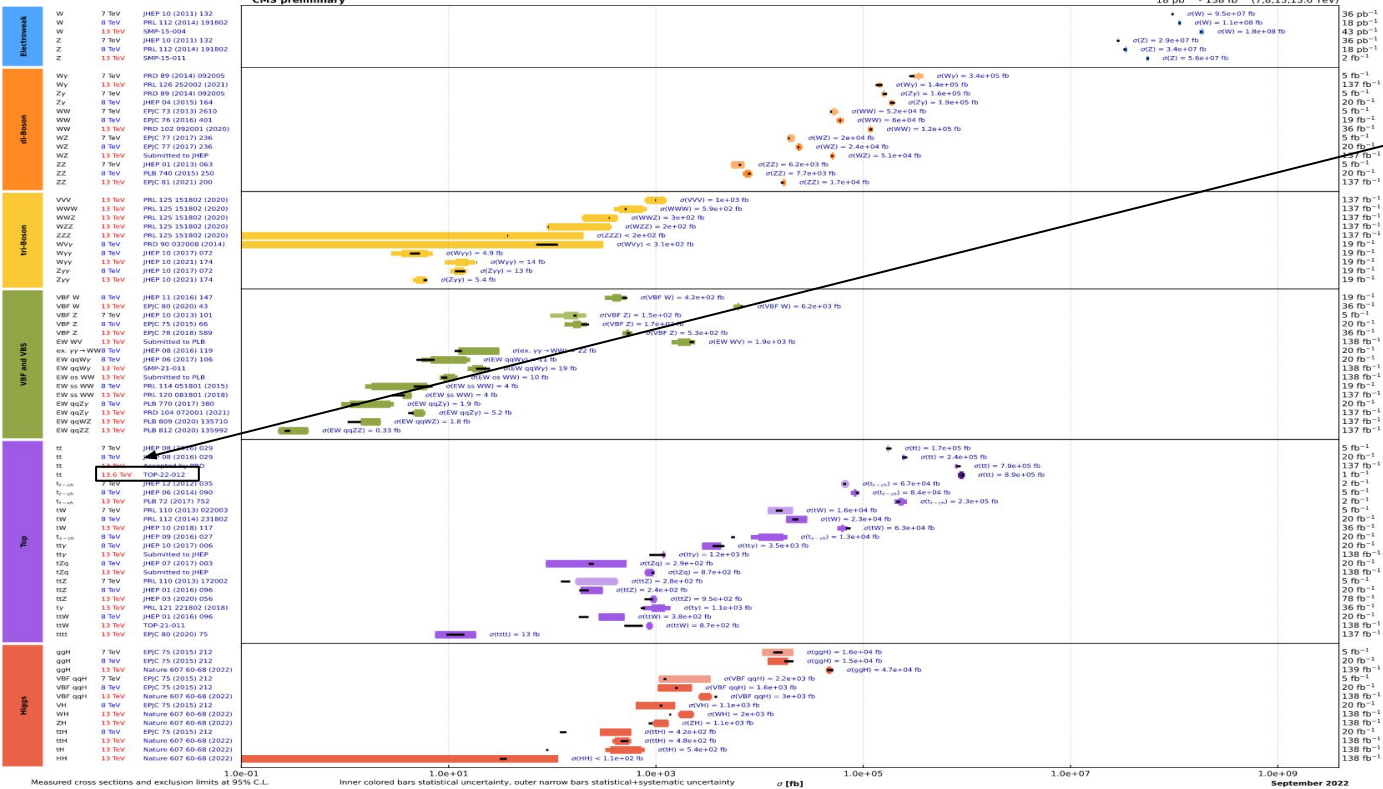
Luminosity measures collisions/unit area/unit time

Higher luminosity increases event rates, aiding rare interaction observations

Luminosity precision facilitates precise measurements for couplings, cross sections



Latest cross section results



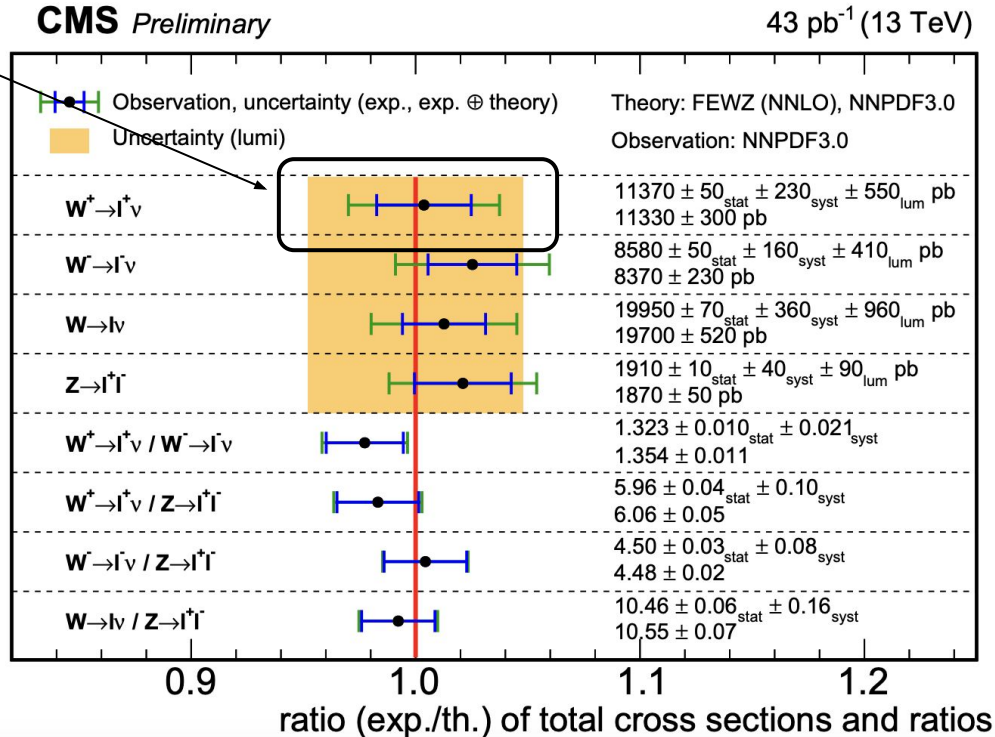
CMS-TOP-22-012
 Combined uncertainty 2.6% (excluding lumi uncertainty)
 luminosity uncertainty 2.3%

Precision on luminosity measurement directly propagates to precision on cross section measurement

Need for luminosity precision measurement

<https://cds.cern.ch/record/2093537>

Uncertainty in luminosity measurement is more than statistical and other systematics uncertainty



Luminometers

<https://cerncourier.com/a/counting-collisions-precisely-at-cms/>

Different luminometers ensure redundancy, backup for device failures

Allow cross-checks, consistency in results

Better control on systematic uncertainties

Employ different methods, complementary information

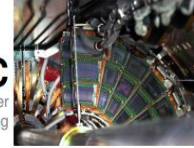
Some with low rates can be calibrated using others

Luminometers

CMS DETECTOR

Total weight : 14,000 tonnes
Overall diameter : 15.0 m
Overall length : 28.7 m
Magnetic field : 3.8 T

PCC
pixel cluster counting



PLT pixel luminosity telescope



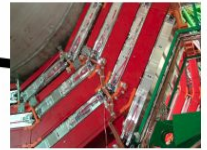
BCM1F fast beam conditions monitor



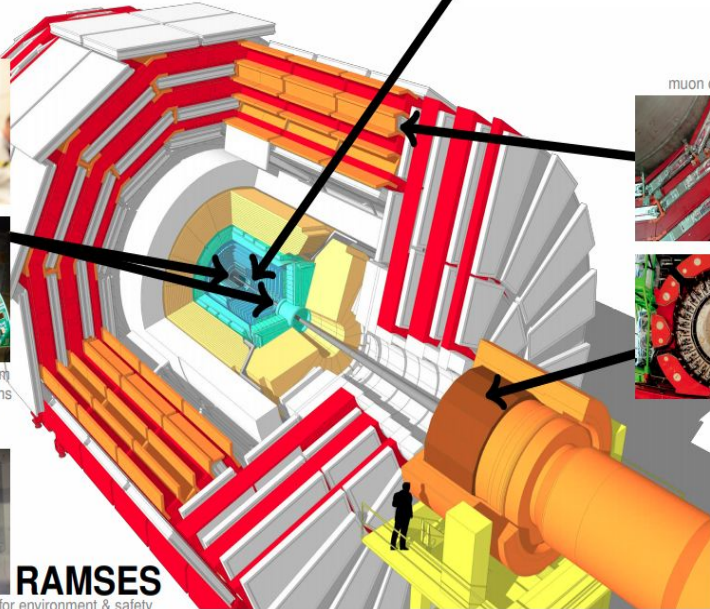
RAMSES

radiation monitoring system for environment & safety

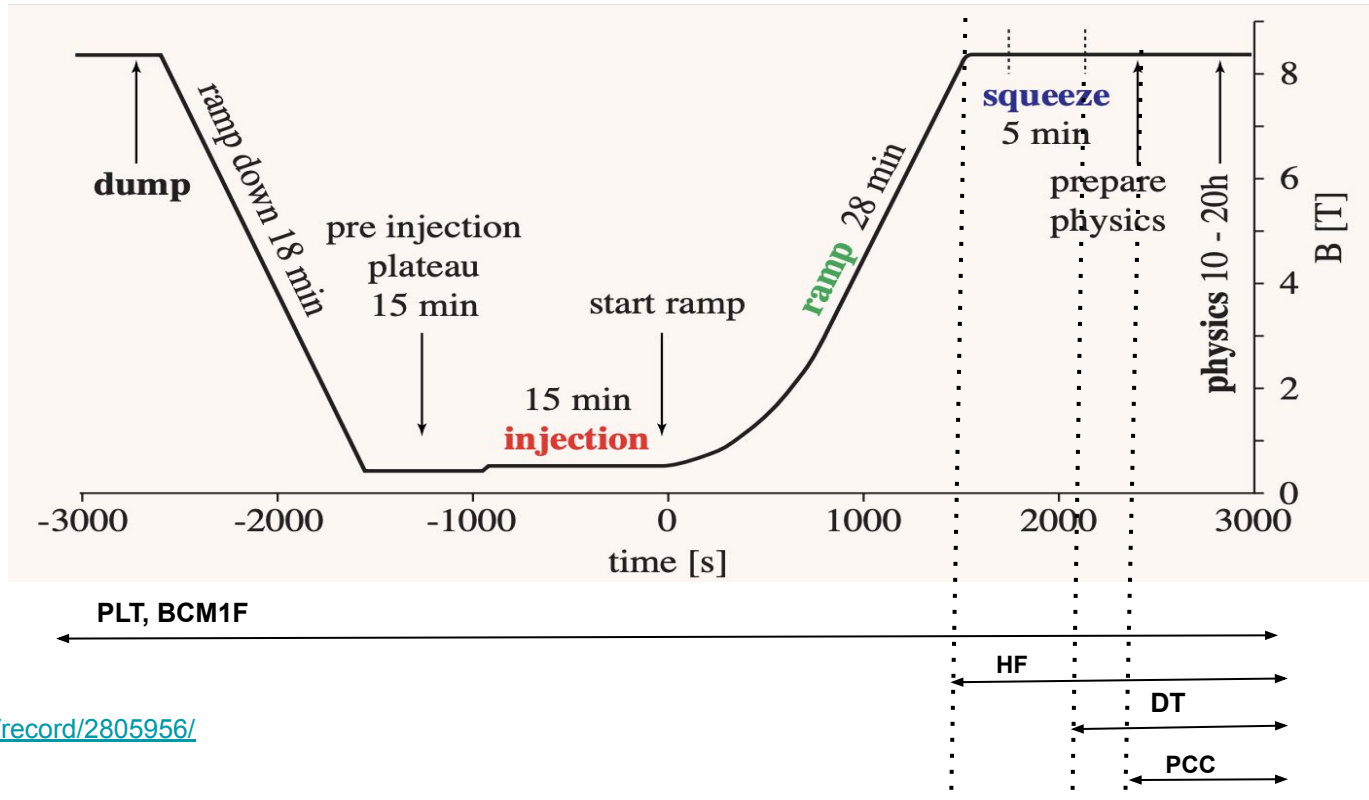
muon drift tubes **DT**



forward hadron calorimeter **HF**



Luminometer status during LHC beam cycle



<https://cds.cern.ch/record/2805956/>

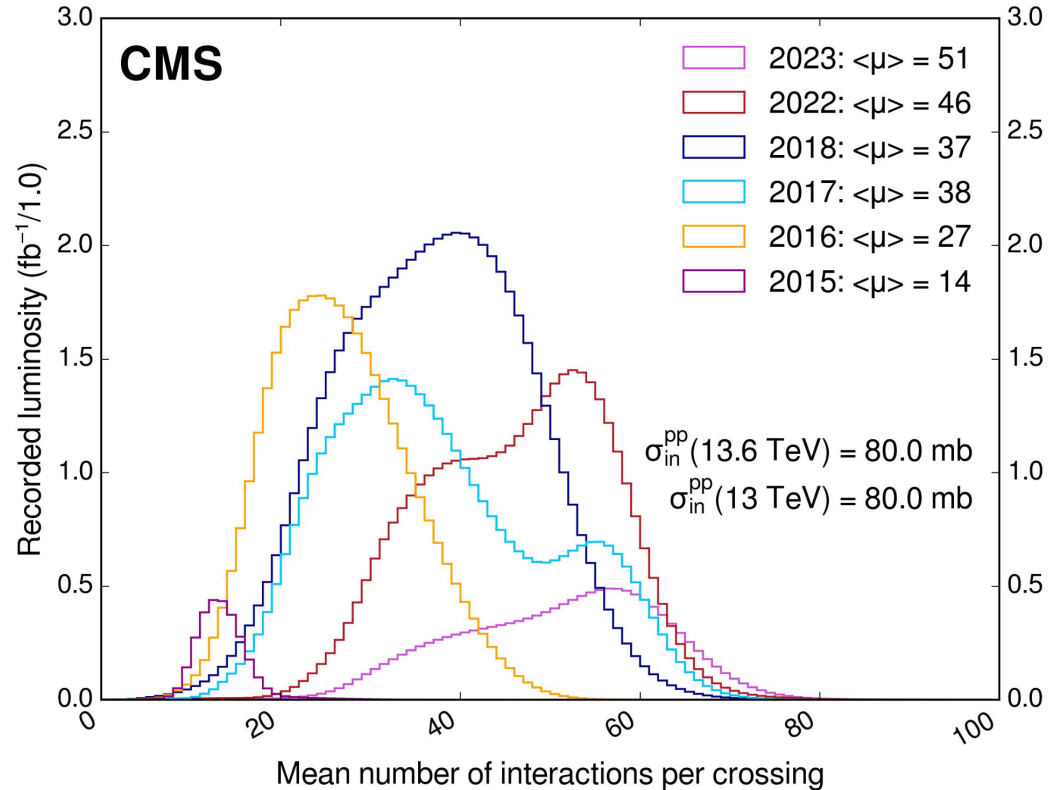
Pileup

Pileup is multiple collisions occurring within a single bunch crossing.

Pileup values/distributions are derived from the measured luminosity by scaling by pp cross section.

$$R = \sigma L$$

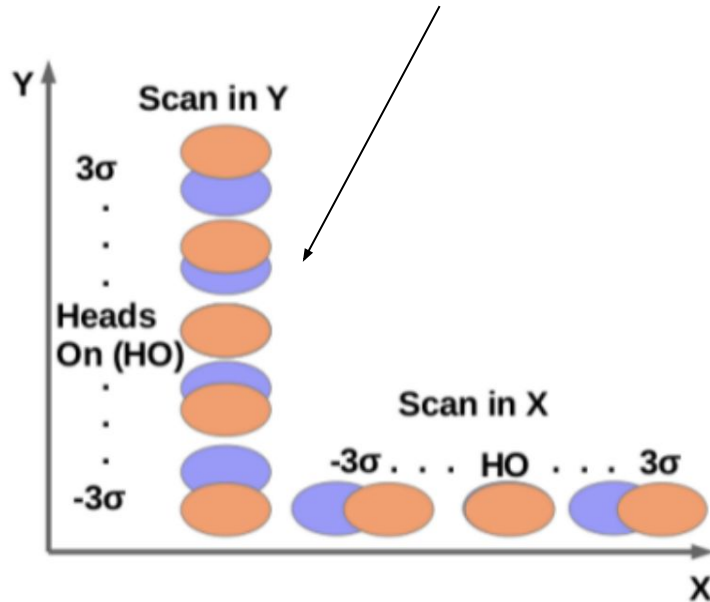
<https://twiki.cern.ch/twiki/bin/view/CMSPublic/LumiPublicResults>



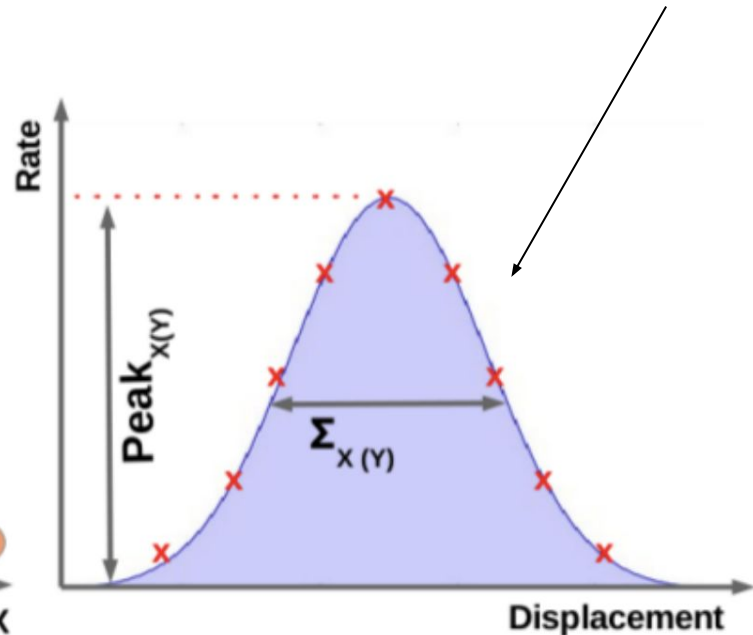
Luminometer Calibration

van der Meer scan

Scanning the beams across each other ($\pm 3\sigma$) in the transverse plane, horizontal (x-scan) and vertical (y-scan) directions

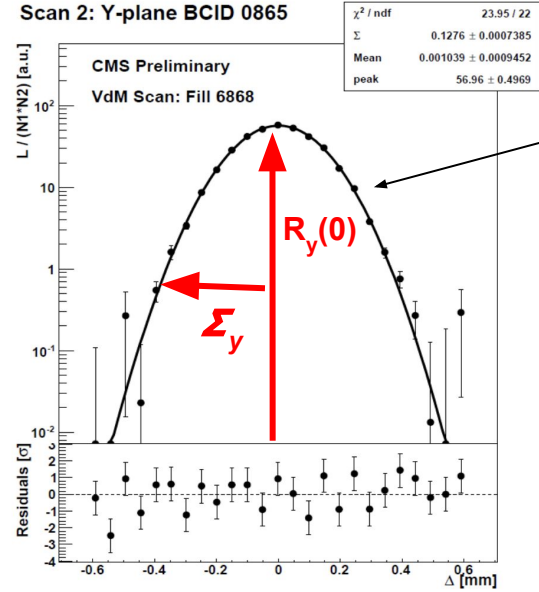
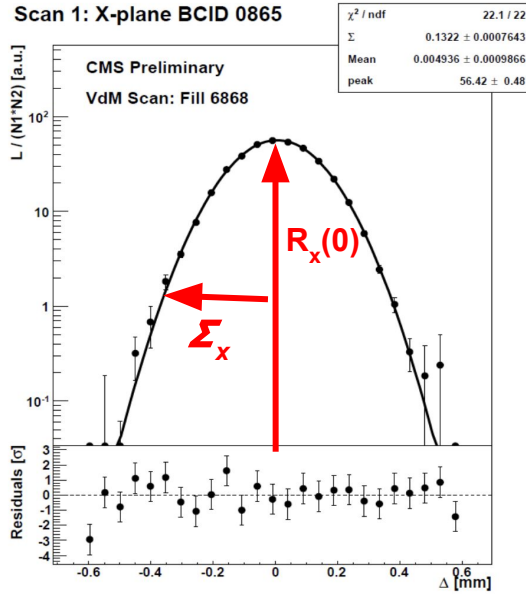


Rate is measured as a function of beam separation



Rate fit

Beam overlap widths and Peak rates are obtained from the fit



Single gaussian fit

Gaussian proton densities
x-y factorization
zero crossing angle

- Bunch intensities (N_1, N_2), N_b number of bunches
- Beam overlap widths (Σ_x, Σ_y)
- Peak rate $R(0)$
- Beam-beam interactions
- LHC orbit frequency f

Visible cross section

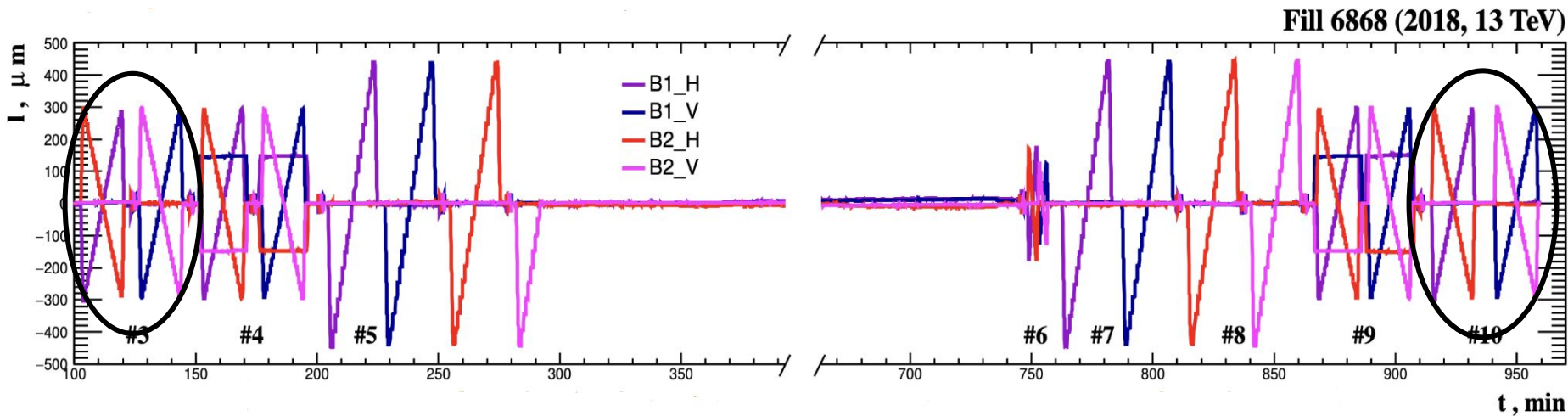
$$\sigma_{vis} = \frac{2\pi \Sigma_x \Sigma_y R(0)}{N_1 N_2 N_b f}$$

Absolute luminosity scale

$$L = \frac{N_1 N_2 f}{2\pi \Sigma_x \Sigma_y} N_b$$

$$R(0) = \frac{R_x(0) + R_y(0)}{2}$$

vdM program



- #3, #10 are vdM scans
- #4, #9 are offset scans
- #5, #7, #8 are beam imaging scans
- #6 is emittance scan

vdM scans are typically conducted at least once a year to ensure the luminometer is correctly calibrated. Results of these scans are used to normalize the data collected by the experiments during physics run.

Luminosity precision

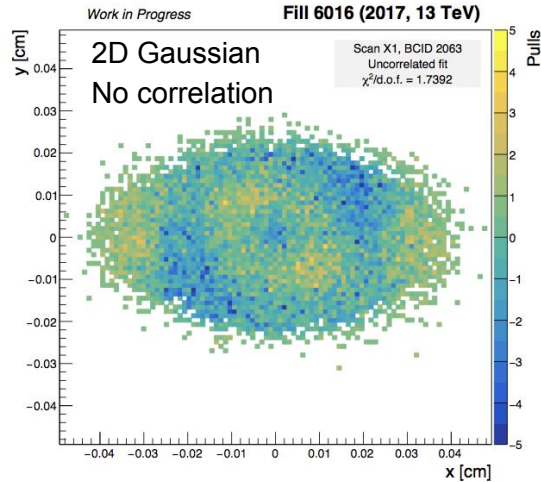
XY non-factorization

<https://cds.cern.ch/record/2621960/>

Beam overlap width transversal to scanning direction is not constant

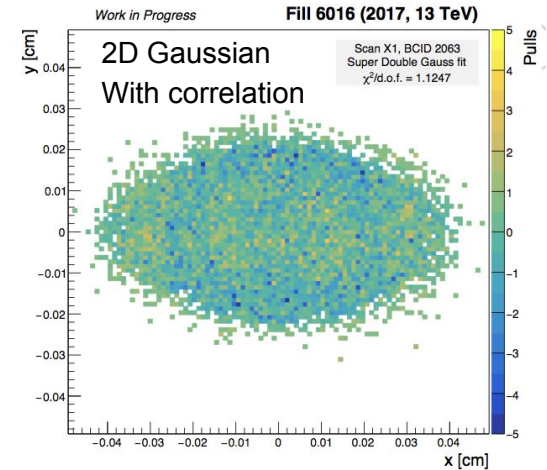
Beam Imaging scans: Beam 2 fixed and Beam 1 moved in steps, beam shape from reconstructed primary vertices

Offset scans: Beams at constant separation in non-scanning direction, luminous region profile from fit to rate vs 2D beam separation



Pull distribution
of global fit to all
scan points

Pull is number of
measured vertices and
vertices from fit,
divided by the
statistical uncertainty

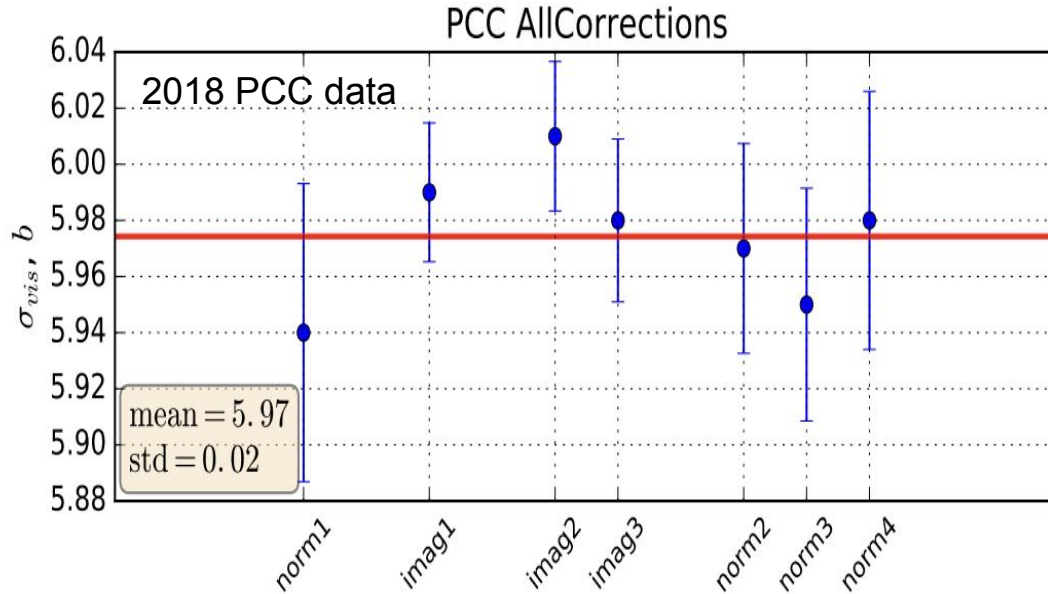


$$g(x) \cdot g(y) = \exp \left[-\frac{1}{2} \left(\frac{x^2}{\sigma_x^2} + \frac{y^2}{\sigma_y^2} \right) \right]$$

$$g(x, y) = \exp \left[-\frac{1}{2(1-\rho^2)} \left(\frac{x^2}{\sigma_x^2} + \frac{y^2}{\sigma_y^2} - \frac{2\rho xy}{\sigma_x \sigma_y} \right) \right] \quad 14$$

Scan to Scan variation

[CMS LUM-18-002-pas](#)



Scan to scan variation of visible cross section averaged over all BCIDs

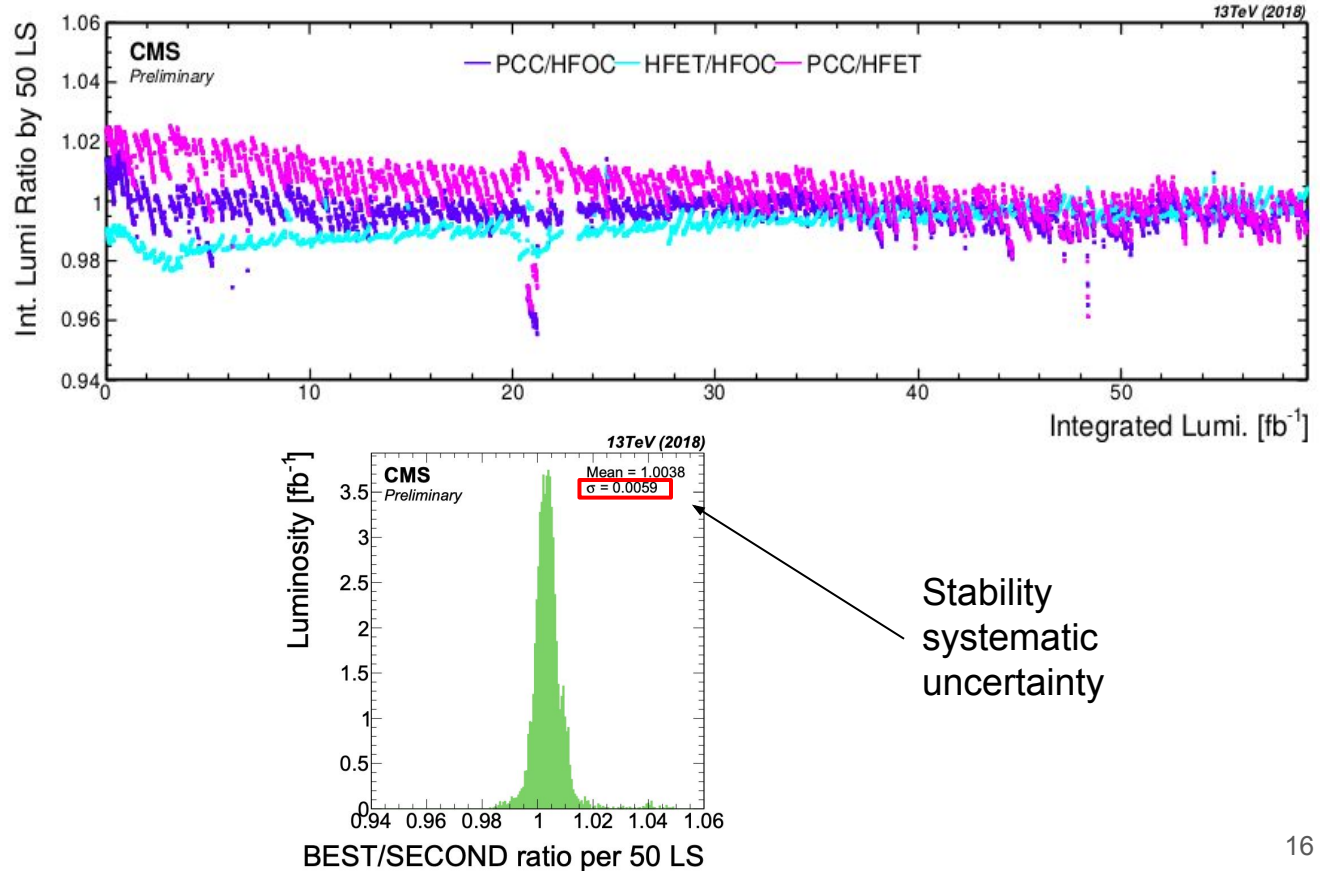
Statistical uncertainty is shown for each scan

Stability

Time-dependent changes in detector performance or environmental conditions.

Detector aging, radiation damage, fluctuation in noise.

The standard deviation is used as stability systematic uncertainty.



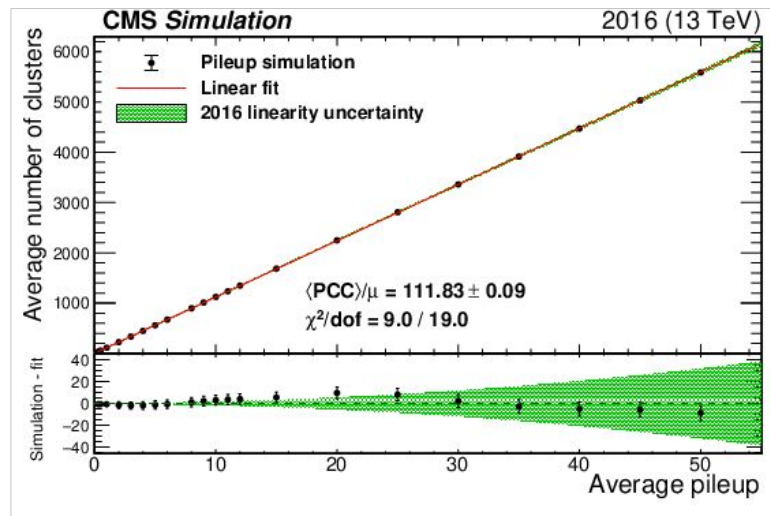
Linearity

Detector response variation at different luminosity levels.

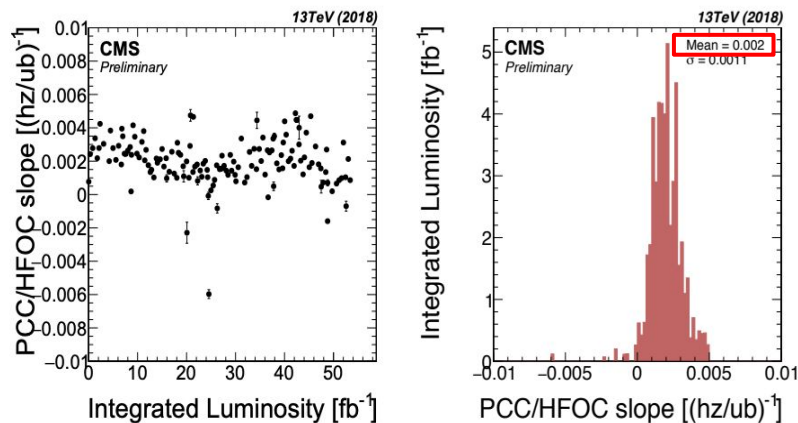
Pileup conditions affect the linearity of luminosity measurement.

High pileup can cause over- or underestimation of luminosity.

The product of slope and luminosity range is measure of linearity systematic uncertainty.



CMS-PAS-LUM-18-002



Afterglow

Type 1 Afterglow Noise: Electronic spillover

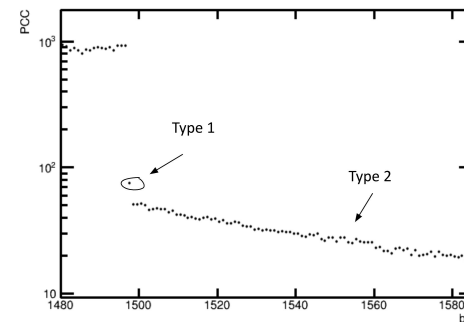
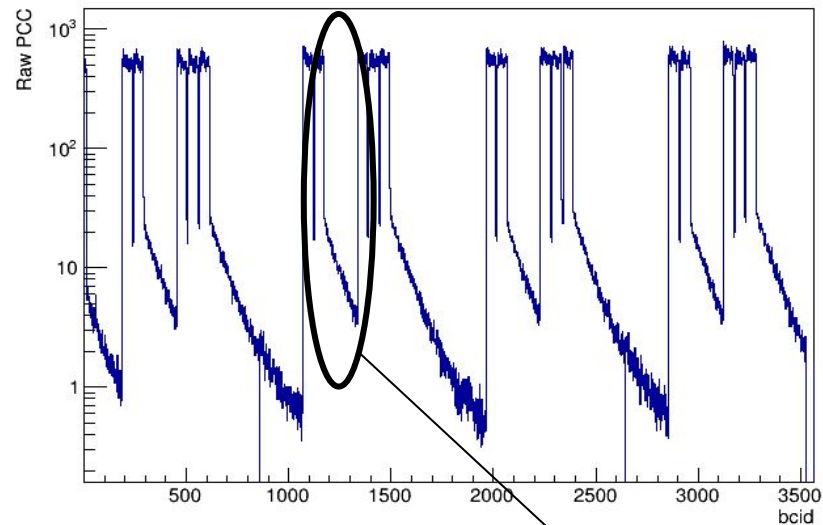
Uncertainty is mean type 1 residual

$$\textit{Type1 residual} = \frac{\textit{first empty bunch lumi}}{\textit{colliding bunch lumi}}$$

Type 2 Afterglow Noise: Exponential decay of detector activation

Uncertainty is mean type 2 residual

$$\textit{Type2 residual} = \frac{\textit{second, third,.. empty bunches lumi}}{\textit{colliding bunch lumi}}$$



Mathematical model:

$$F(x) = \sum_{k=0}^{N_{bcid}} N_k [x - k == 0] + A[x - k == 1] + B \exp(-C(x - k - 1)) [x - k >= 1]$$

2017

Uncertainty in Run 2 Luminosity

<https://doi.org/10.1140/epjc/s10052-021-09538-2>

2015-2016

Source	2015 [%]	2016 [%]	Corr
Normalization uncertainty			
<i>Bunch population</i>			
Ghost and satellite charge	0.1	0.1	Yes
Beam current normalization	0.2	0.2	Yes
<i>Beam position monitoring</i>			
Orbit drift	0.2	0.1	No
Residual differences	0.8	0.5	Yes
<i>Beam overlap description</i>			
Beam-beam effects	0.5	0.5	Yes
Length scale calibration	0.2	0.3	Yes
Transverse factorizability	0.5	0.5	Yes
<i>Result consistency</i>			
Other variations in σ_{vis}	0.6	0.3	No
Integration uncertainty			
<i>Out-of-time pileup corrections</i>			
Type 1 corrections	0.3	0.3	Yes
Type 2 corrections	0.1	0.3	Yes
<i>Detector performance</i>			
Cross-detector stability	0.6	0.5	No
Linearity	0.5	0.3	Yes
<i>Data acquisition</i>			
CMS deadtime	0.5	<0.1	No
Total normalization uncertainty	1.3	1.0	—
Total integration uncertainty	1.0	0.7	—
Total uncertainty	1.6	1.2	—

	Systematic	Correction (%)	Uncertainty (%)
Normalization	Length scale	-0.9	0.3
	Orbit drift	—	0.2
	x - y correlations	+0.8	0.8
	Beam-beam deflection	+1.6	0.4
	Dynamic- β^*	—	0.5
	Beam current calibration	—	0.3
	Ghosts and satellites	—	0.1
	Scan to scan variation	—	0.9
	Bunch to bunch variation	—	0.1
	Cross-detector consistency	0.4-0.6	0.6
Integration	Afterglow (HF)	—	0.2 \oplus 0.3
	Cross-detector stability	—	0.5
	Linearity	—	1.5
	CMS deadtime	—	0.5
Total			2.3

2018

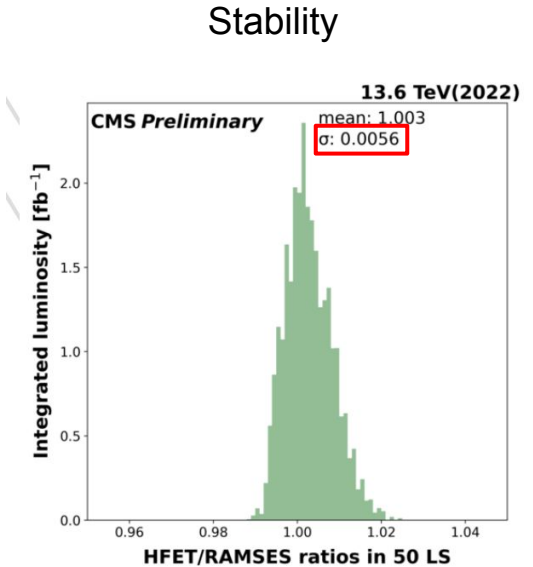
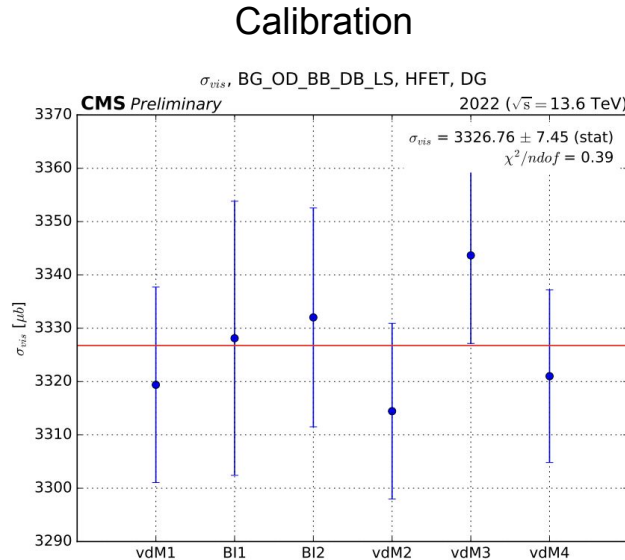
	Systematic	Correction (%)	Uncertainty (%)
Normalization	Length scale	-0.8	0.2
	Orbit drift	0.2	0.1
	x - y nonfactorization	0.0	2.0
	Beam-beam deflection	1.5	0.2
	Dynamic- β^*	-0.5	—
	Beam current calibration	2.3	0.2
	Ghosts and satellites	0.4	0.1
	Scan to scan variation	—	0.3
	Bunch to bunch variation	—	0.1
	Cross-detector consistency	—	0.5
Integration	Background subtraction	0 to 0.8	0.1
	Afterglow (HFOC)	0 to 4	0.1 \oplus 0.4
	Cross-detector stability	—	0.6
	Linearity	—	1.1
	CMS deadtime	—	<0.1
Total			2.5

Systematic uncertainty expected to improve due to ongoing work

[LUM-18-002-pas](#)

Run 3 preliminary results

Source	Uncertainty (%)
Calibration	
Bunch current measurement	0.3
Orbit drift	0.3
Residual beam positions	1.0
Beam-beam effects	0.5
Length scale	0.1
Factorization bias	1.4
Scan-to-scan variation	0.6
Bunch-to-bunch variation	0.6
Cross-detector consistency	0.5
Integration	
HFET OOT pileup corrections	0.2
Cross-detector stability	0.6
Cross-detector linearity	0.6
Total	
Calibration	2.1
Integration	0.7
Total	2.2



Luminosity measurement from Z boson

Z counting method

Muons penetration ensures accurate identification, low background.

Large sample Z events allow in situ efficiency calibration.

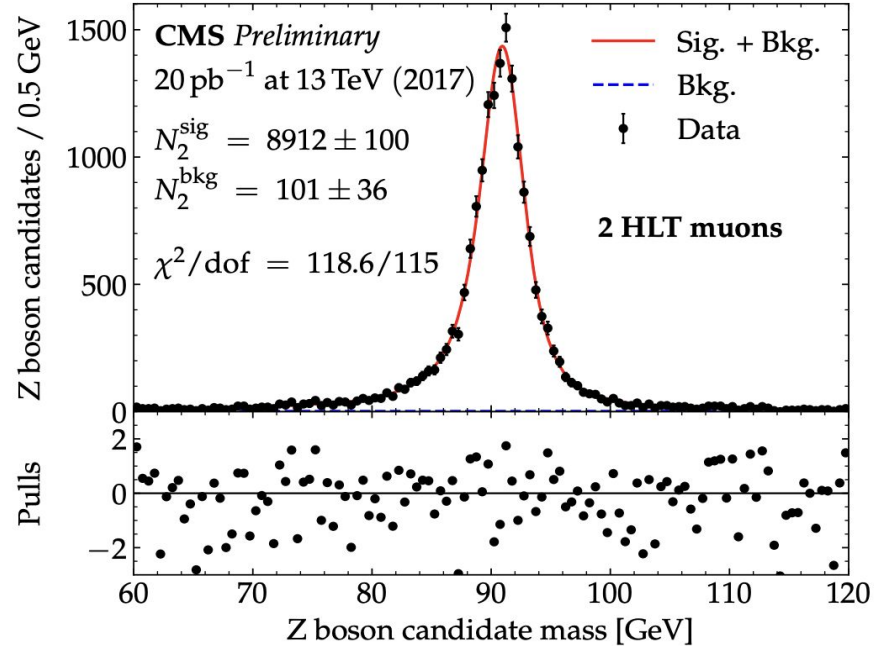
Comparison of Z counts under varied pileup conditions cancels systematic uncertainties, enhancing precision.

Two datasets considered: low and high pileup samples.

High pileup luminosity is calculated using low pileup sample.

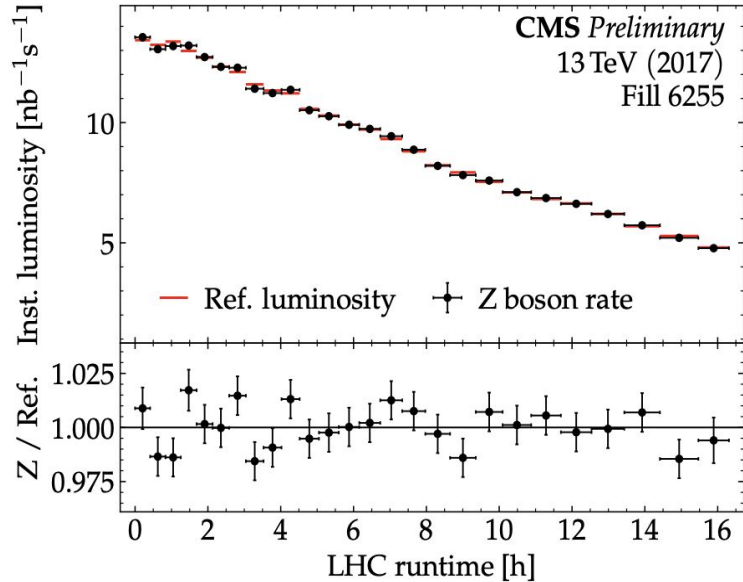
$$L_{highPU} = \frac{N_{highPU}^Z}{N_{lowPU}^Z} L_{lowPU}$$

[LUM-21-001-pas](#)



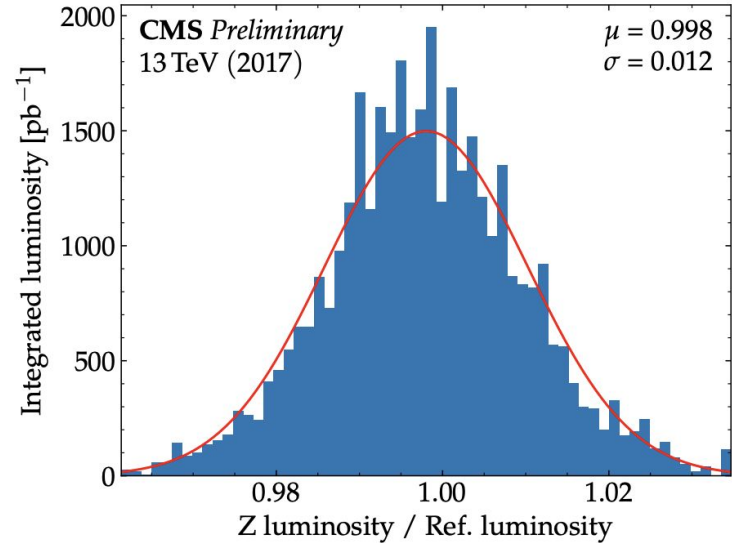
Fit used to get number of Z boson events (N^Z) and muon trigger efficiency.
Muon trigger correlation coefficient from Z boson and trigger efficiency.

Z luminosity



Efficiency corrected Z boson rate compared with HF

[LUM-21-001-pas](#)



Consistency between two methods within 0.2%

Summary

- Precision on luminosity measurement directly propagates to precision in coupling, cross section measurement of various physics processes.
- Several luminometers are used to have backup for luminometer failure, better estimation of systematic uncertainty, carry out luminosity measurement using different techniques.
- Dominant uncertainty comes from xy non-factorization during 2018, scan to scan variation during 2017 and residual differences between the measured beam positions and the ones provided by the operational settings of the LHC magnets in 2015.
- Best precision on luminosity measurement during Run 2 is 1.2 % for 2016 data taking.
- Due to several improvements in reanalysis, the expected precision for 2017-2018 is 1%.
- Phase II HL-LHC will increase luminosity by 10 times its current value. CMS detector will have extended tracker, upgraded data acquisition system, improved muon system, upgraded trigger system, robustness to radiation and new readout electronics.