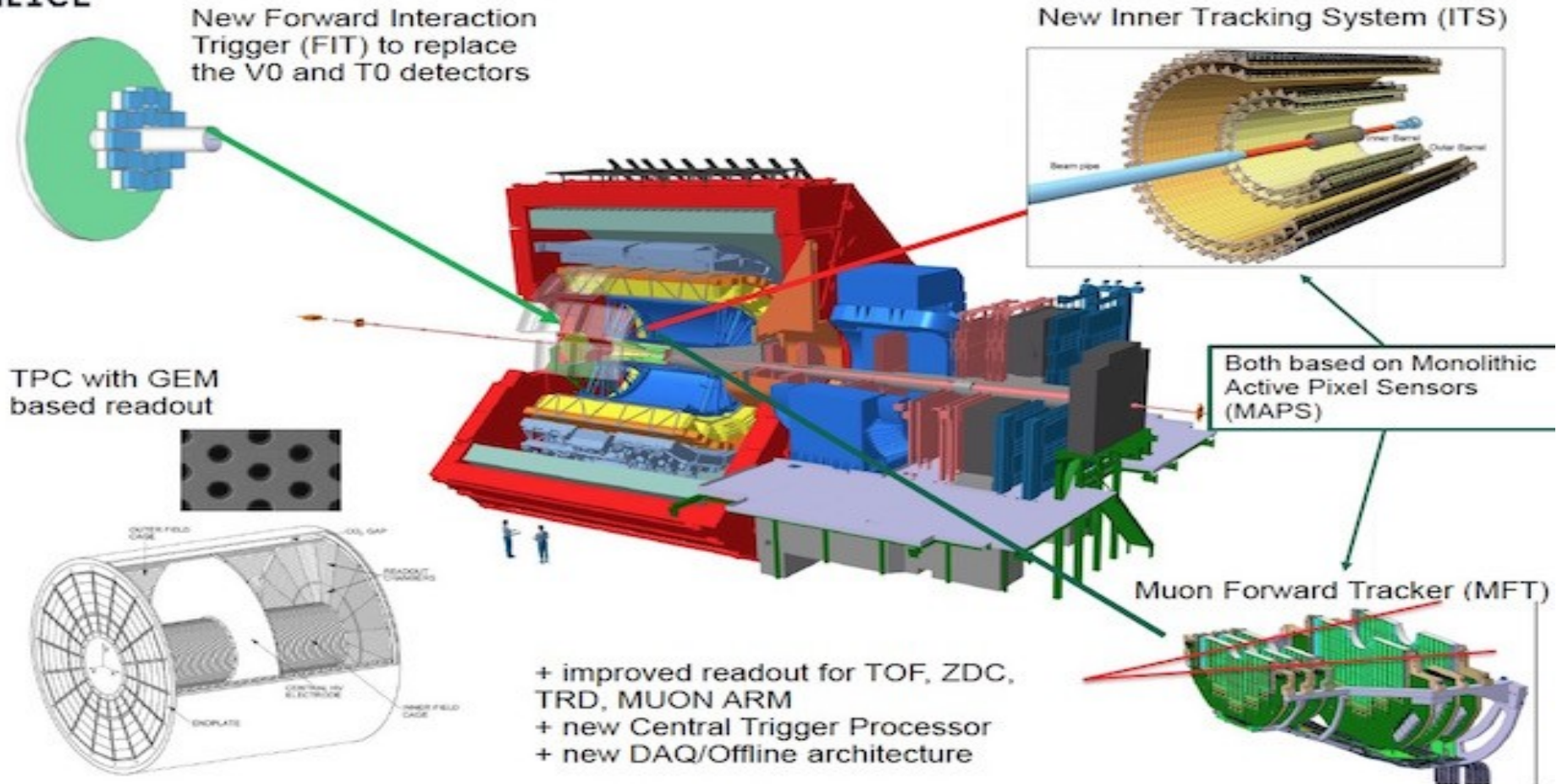


FV0 detector for ALICE FIT

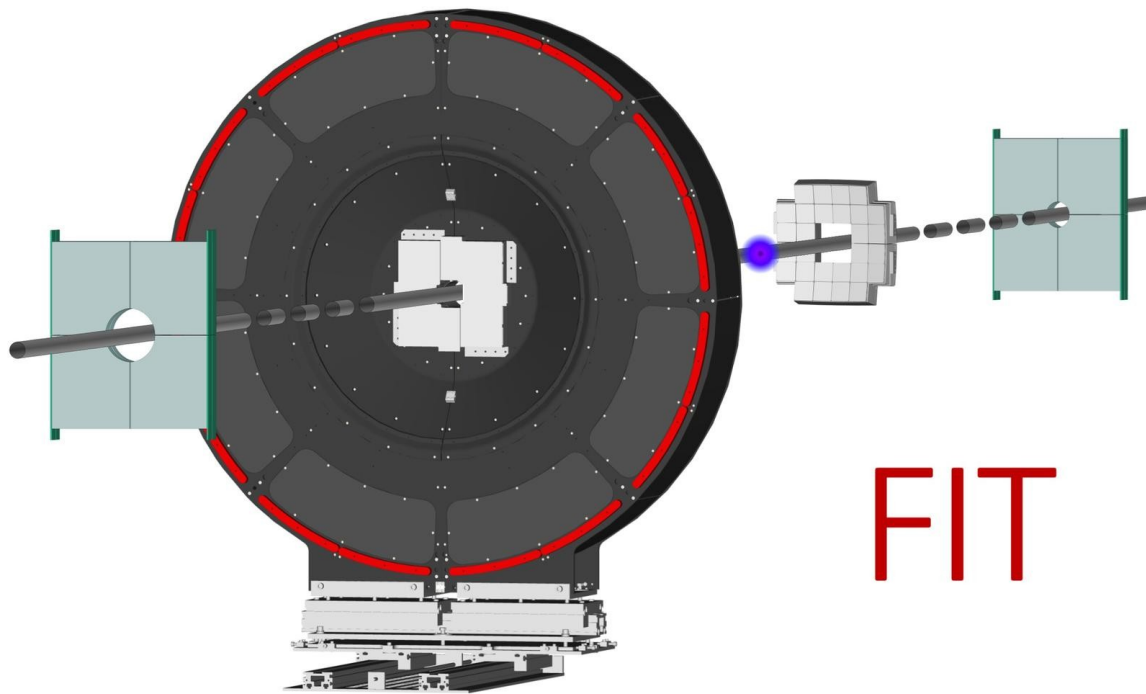
V. Grabski
IFUNAM



ALICE detector upgrade



Collaboration ALICE FIT



FIT

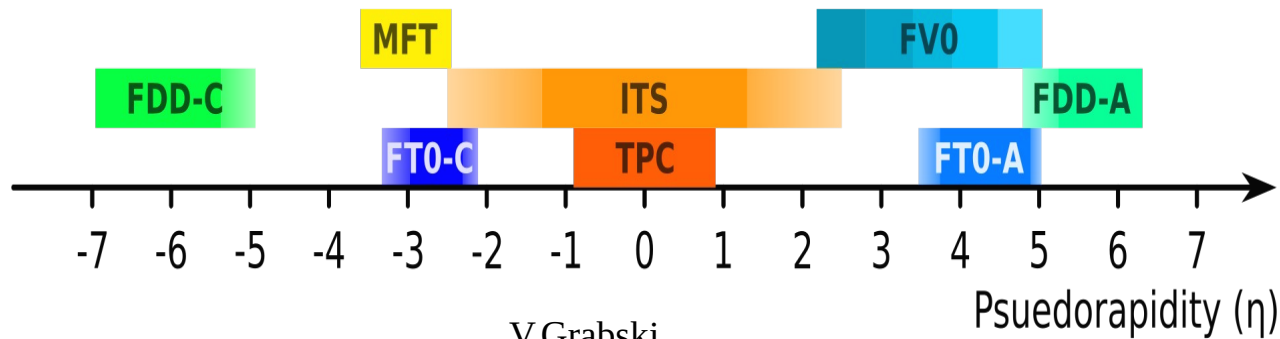
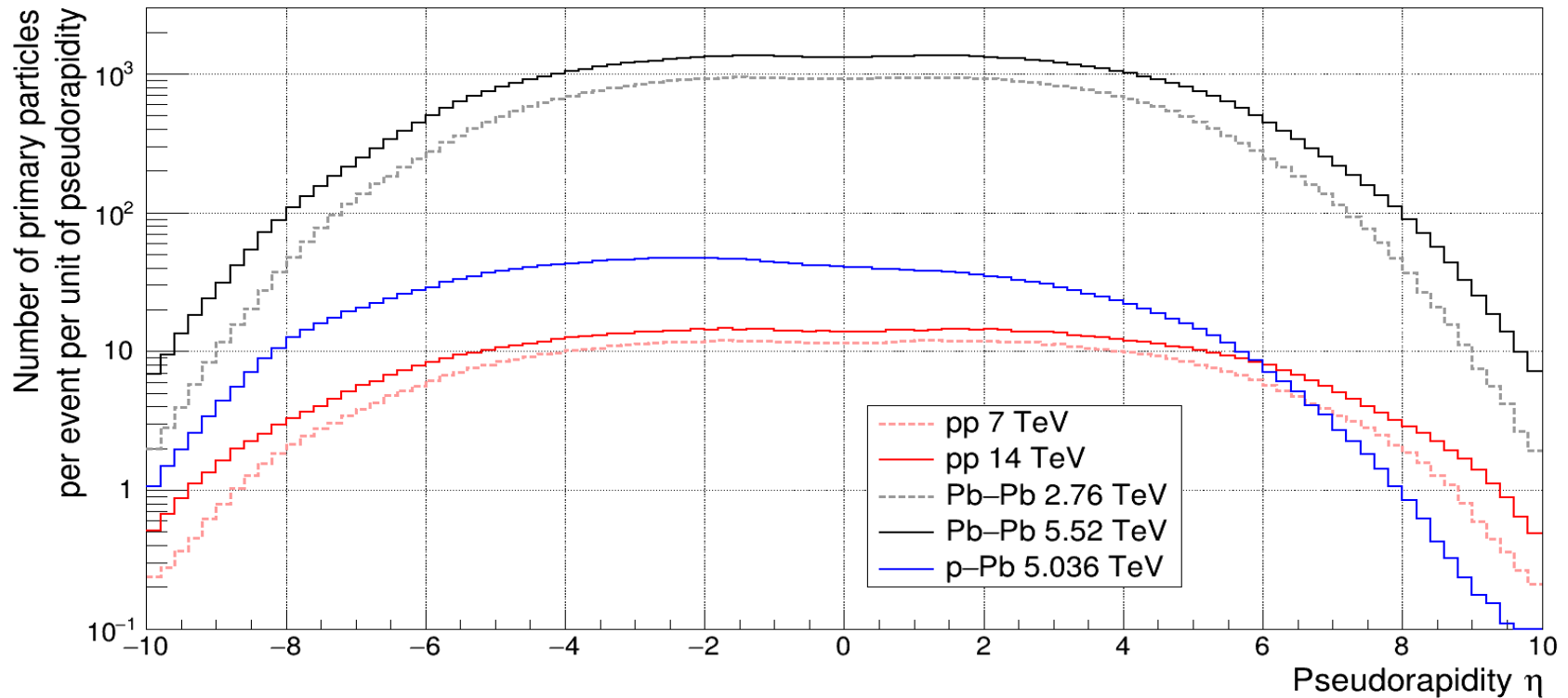
Austria,Vienna,Stefan Meyer Institute
 Czech,Praha, Technical University
 Denmark,Copenhagen,Niels Bohr Institute, University of Copenhagen
 Finland,Helsinki Institute of Physics (HIP) and Univ. of Jyväskylä
 Finland Helsinki
 Helsinki Institute of Physics (HIP) and Univ. of Helsinki
 Mexico Mexico City Instituto de Física, UNAM
 Mexico Mexico City and Merida
 CINVESTAV, Mexico Mexico City
 Instituto de Ciencias Nucleares, UNAM
 Mexico,Puebla,Benemérita Universidad Autónoma de Puebla
 Mexico,Culiacan, Sinaloa,Universidad Autónoma de Sinaloa
 Peru Lima,Pontificia Universidad Católica del Perú
 Poland Krakow ,AGH,Poland,Krakow IFJ-PAN
 Poland Warsaw ,NCBJ
 Russia Moscow,Institute for Nuclear Research
 Russia Moscow,Moscow Engineering Physics Institute
 Russia Moscow,Russian Research Centre Kurchatov Institute
 USA Chicago ,Chicago State University
 USA San Luis Obispo,California Polytechnic State University

https://alice-collaboration.web.cern.ch/menu_proj_items/FIT

The main purpose of the FIT

- Time reference for the other detectors;
- Vertex and collision rate estimation;
- Luminosity Monitor
- Residual gas background estimation;
- Collision centrality estimation;
- Estimation of the reaction plane;
- Different triggers for the data analysis;

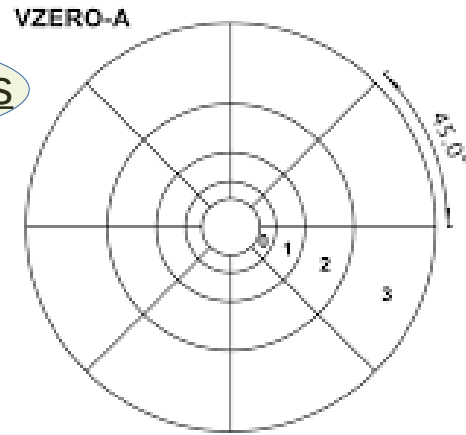
Detectors acceptances in Pseudorapidity



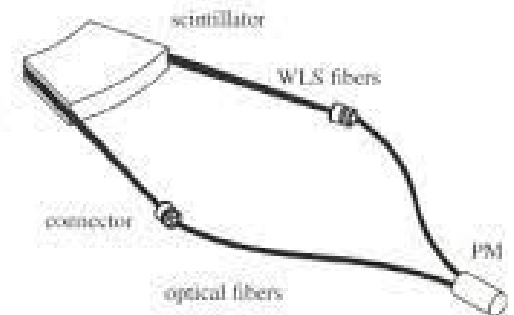
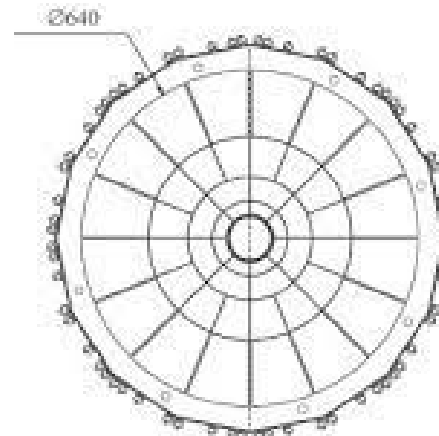
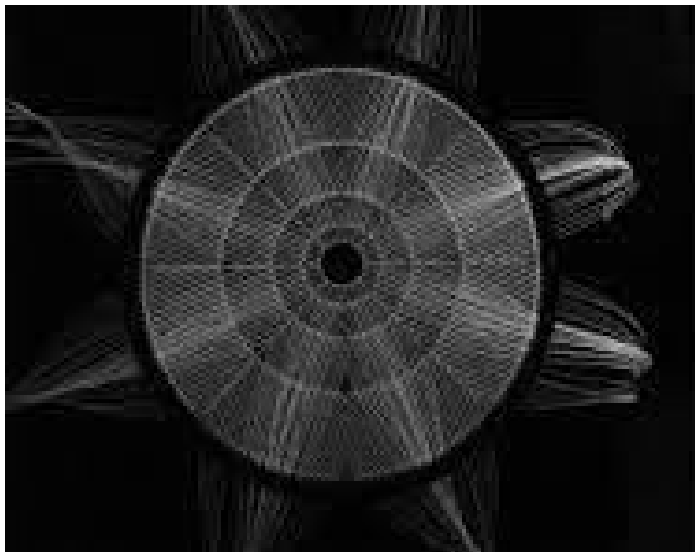
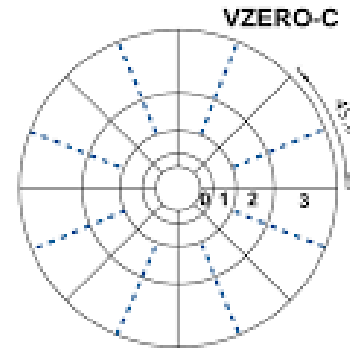
V0A and V0C detectors

Time resolution 200ps

Mexico



Francia



A new design for the FV0 detector

New fiber read-out design for the large area scintillator detectors: providing good amplitude and time resolutions

V. Grabski¹

Instituto de Fisica Universidad Nacional Autonoma de Mexico, Mexico

^aAv Universidad 3000, Instituto de Fisica, Coyoacan, Mexico

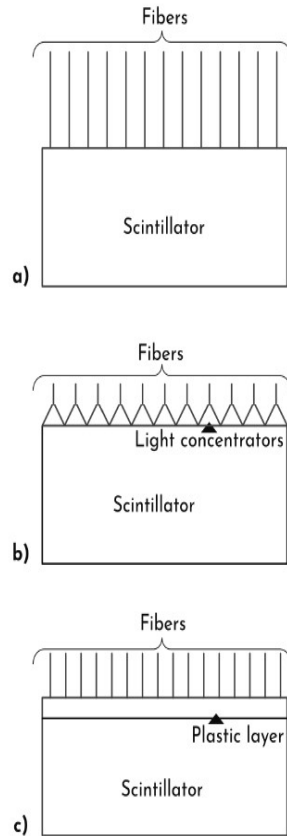


Figure D.1: Three different light collection schemes.

arXiv:1909.01184v1 [physics.ins-det] 3 Sep 2019

Abstract

Most of the time-of-flight systems as well as the fast interaction trigger detectors have large surfaces and the principal requirements for the above mentioned detectors is a good time resolution of the order $100 - 200ps$. The easiest solution here is to split the large surface into the small tiles with a read-out directly attached to the photo-sensor. This solution is expensive because the number of the channels grows proportional to the surface. Although if the coverage of the whole surface by the sensors is not sufficient, one can obtain non uniform response, which is not acceptable if uniformity is required. Our suggestion is based on the usage of clear fibres mounted perpendicular on the surface of the scintillator as a matrix providing uniform surface reduction and keeping most of the benefices of the small tiles with the photo-sensors. In this work the design and some analytic estimations of the light collection efficiency for direct and reflected photons will be presented. Also some simple estimations are presented for the time-spread and for the light pulse wave form dependent on the lateral sizes and the thickness of the scintillator.

Keywords: Scintillation detector, Time-of-flight detector, Particle identification, Fast Interaction Trigger, Optical Fibre

^{*}Corresponding author
Email address: varlen.grabski@cern.ch (V. Grabski)

A new design for the FV0 detector

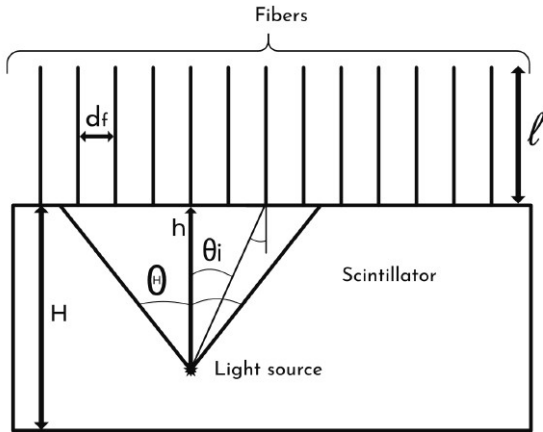


Figure D.2: Light collection scheme from the point like source for direct photons. Description of all notations are in the text.

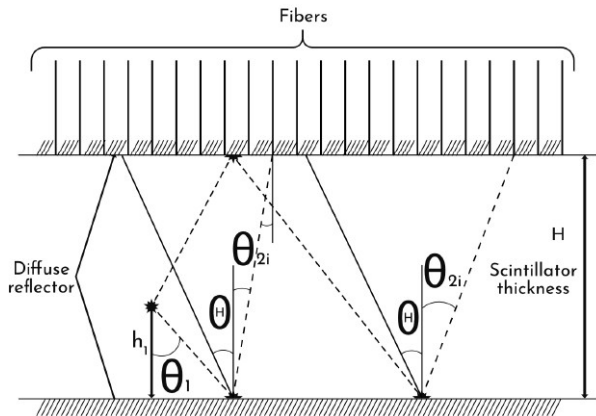


Figure D.3: Light collection scheme from the point like source and for the reflected photons. Description of all notations are in the text.

$$N_{DP} = \int_0^H N_p(h) dh = \frac{N_0 \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta}{4 \times H} \times \int_0^H \left(\frac{1}{N_f} \times \sum_{i=0}^{N_f} \exp\left(-\frac{a(h,l)}{\cos \theta_i}\right) \times \cos^3 \theta_i \right) dh \quad (5)$$

The dependence of N_{DP} from the distance h that is included in $a(h,l)$ is as h/λ_{sc} . If $h \ll \lambda_{sc}$ then N_{DP} is approximately independent from the distance of the matrix plane. Of course the condition of $h \geq d_f$ will be fulfilled having at least few fibres involved in the light collection. Most of the fast scintillators have attenuation lengths larger than 100cm, so this approximation could be good enough for the scintillator thickness up to 10cm. For the designs shown in Fig 1(b and c) the condition of $h \geq d_f$ is already fulfilled so it can be achieved better volume uniformity for entire scintillator volume. The calculation of the integral in (5) to estimate of N_{DP} is performed in Appendix A. Obtained analytic expression is a large one, that's why we present here as a function $\Phi_{dp}(\cos \Theta, a_0, a_1)$:

$$N_{DP} = \frac{N_0 \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta}{4 \times H} \times \Phi_{dp}(\cos \Theta, a_0, a_1)$$

where $a_0 = a(0, l)$ is the fibre longitude in attenuation longitude units and $a_1 = a(H, l)$ is the sum of the scintillator thickness and the fibre longitude in attenuation longitude units. The light collection efficiency for DP can be estimated as $\Upsilon_{DP}(\cos \Theta, a_0, a_1) = N_{DP}/N_0$.

$$\Upsilon_{DP} = \frac{N_0 \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta}{4 \times H} \times \Phi_{dp}(\cos \Theta, a_0, a_1) \quad (6)$$

, which converts a relation independent from H if we ignore attenuation in the scintillator:

$$\Upsilon_{DP} = \frac{N_0 \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta}{4} \times \Phi_{dp}(\cos \Theta, a_0, a_0) \quad (7)$$

Ignoring also light attenuation in the fibres we obtain a simple expression.

$$\Upsilon_{DP} = \frac{F_{tr} \times S_f \times D_f \times \tan^2 \Theta (1 - \cos^4 \Theta)}{16(1 - \cos \Theta)} \quad (8)$$

A new design for the FV0 detector

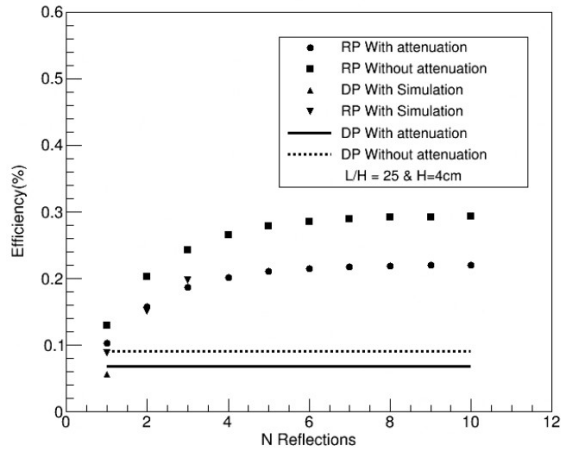


Figure D.4: Light collection efficiency vs number of reflections. Lines show efficiency for DP with and without attenuation in the scintillator and in the fibres

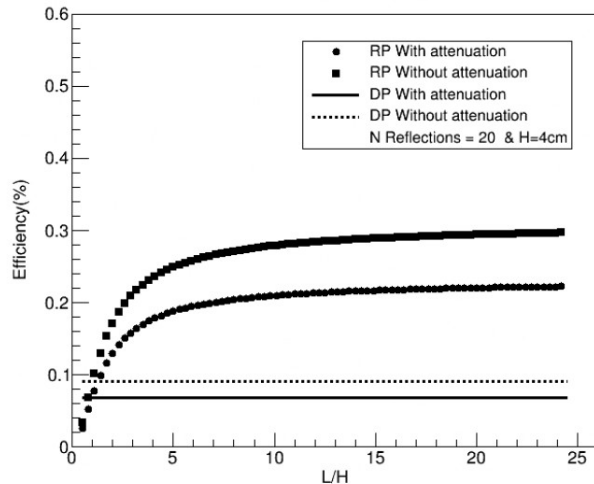


Figure D.5: Light collection efficiency vs detector lateral sizes in the scintillator thickness units.

The light collection efficiency with diffuse reflection for n reflections and for infinite lateral sizes can be estimated as:

$$\Upsilon_{DFR} = \frac{\epsilon_{df_r} \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta}{4H} \times I(\Psi, H) \times (S_{df_r}(a_1, \Theta) + \epsilon_{df_r}^2 \times \epsilon_{sfs}^3 \times \epsilon_{fm} \times S_{df_r}(a_2, \Theta) + \dots + \epsilon_{df_r}^{2n-2} \times \epsilon_{sfs}^{2n-1} \times \epsilon_{fm}^{n-1} \times S_{df_r}(a_n, \Theta)) \quad (13)$$

where $a_n = a(nH, l)$. In case of short fibre lengths the attenuation can be ignored and for the light collection efficiency can be obtained a simple expression for the first n reflections that is shown below (see Appendix B):

$$\Upsilon_{DFR} = \epsilon_{df_r} \times F_{tr} \times S_f \times D_f \times \tan^2 \Theta \times (1 - \cos \Psi) \times \frac{(1 - \cos^5 \Theta)}{10 \times (1 - \cos^2 \Theta)} (1 + \epsilon_{df_r}^2 \times \epsilon_{sfs}^3 \times \epsilon_{fm} + \dots + \epsilon_{df_r}^{2n-2} \times \epsilon_{sfs}^{2n-1} \times \epsilon_{fm}^{n-1}) \quad (14)$$

For the average reflection efficiency about $\epsilon_{df_r} = 0.8$ up to 7-8 reflection will give significant contributions. So using diffuse reflective coating the amount of the light can be increased up to 4 times compared with the direct photons. So for the tasks when the amplitude resolution is important, the usage of diffuse reflection is preferable. If the detector lateral sizes are limited, which is usual case in the real life, it is also possible to perform analytic estimations considering reflections from the lateral planes. This can be performed in a similar way and it should be done for the already defined geometry. The total light collection efficiency Υ_{TOT} can be estimated as a sum of the direct and the reflected light collection efficiencies:

$$\Upsilon_{TOT} = \Upsilon_{DF} + \Upsilon_{DFR}$$

All analytic expressions except when the attenuation is ignored include the Exponential Integral, which can be evaluated only numerically. For this reason we make two plots to show the dependencies of the light collection efficiency from the number of reflections and from the lateral size. Here we ignore the reflections

A new design for the FV0 detector

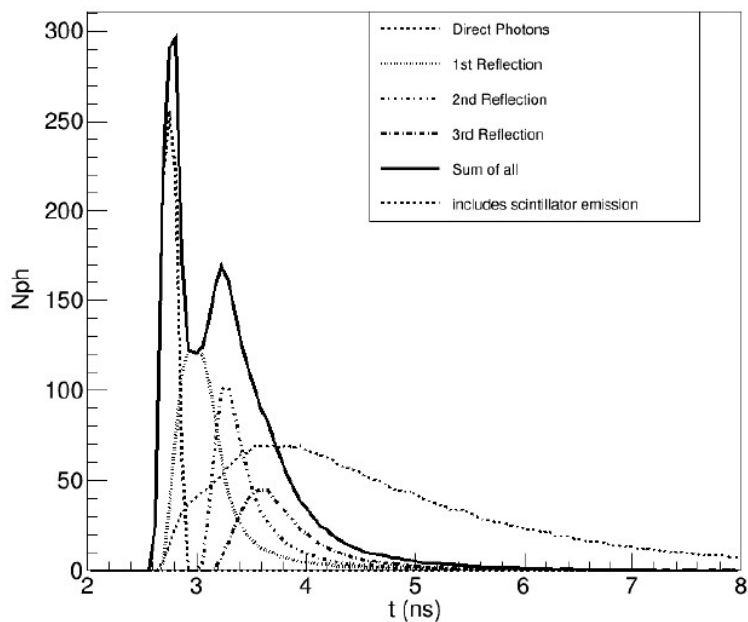


Figure D.6: Pulse time wave forms for direct and reflected photons.

to average value) assuming that the light speed is the same in the scintillator and in the fibre as it is shown below:

$$\frac{\sigma_t}{\bar{t}} = \sqrt{\frac{(1 + \frac{H}{l}) \times (\cos \Theta - 1)^2}{((1 + \frac{H}{2l})^2 \times \cos \Theta \times \ln^2 \cos \Theta) - 1}} \quad (15)$$

where Θ is the value of the fibre aperture angle, \bar{t} is the average value of the time transition to the photo-sensor, l is the fibre length and H is the scintillator thickness. As it can be seen from the formula() the main time spread comes from the fibre length if the thickness of the scintillator is much smaller than the fibre longitude. In this case the time spread is proportional to average transition time which in its turn is proportional to the fibre longitude. For the single clad fibre with NA(numerical aperture) 0.5 the coefficient is about 0.02. For the multi-clad fibres this coefficient will increase up to 0.03(using formula C6). So this means that for the time performance is better to use higher density of the fibres and a single clad than a multi-clad fibres. Of course this estimation should be considered as a limit that can be achieved for the given H and l .

For the specular reflection from the large surface of the scintillator the thickness of the scintillator will be doubled. In case of the diffuse reflection the time spread estimation is not as easy as in the previous case. To study the time spread for the diffuse reflection a simple simulation geometry have been used (shown in Fig 6). All important simulation parameters are mentioned in the previous chapter. The simulation results for the pulse time wave forms of direct and reflected photons with the large statistics just to have smooth curve is shown in Fig. 6. As it can be seen from the figure the number of the reflected photons is significantly larger(factor of 3) than DP and the rise time of fast photons is less than 100ps. The delay of the twice reflected photons is about 400ps, which corresponds to the thickness of scintillator. The standard deviation of the time spread of DP is about 112 ps, which is similar to the approximate value obtained from formula (C6). In the same figure is also shown the modification of the time wave forms when the scintillator stochastic time emission is included. As it can be seen from the figure the rise time of the light

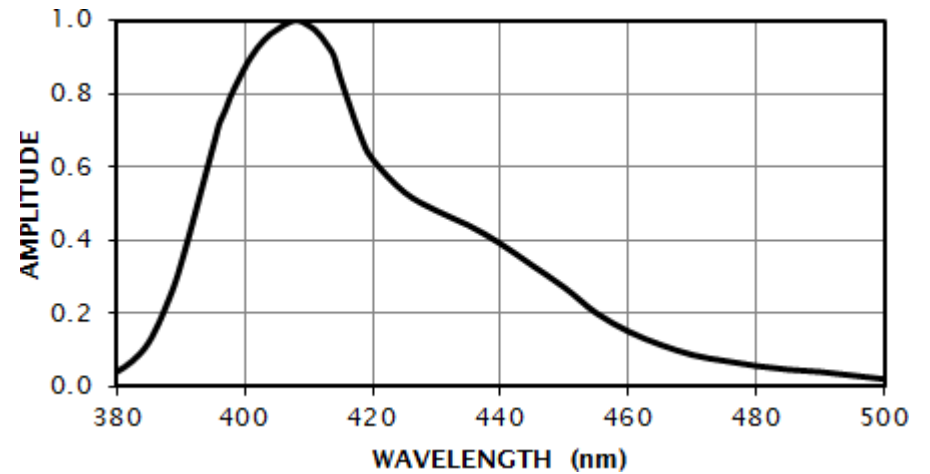
Detector Construction

Choice of construction materials

- Scintillator
- Optical Fiber
- Photo-sensor

Scintillator

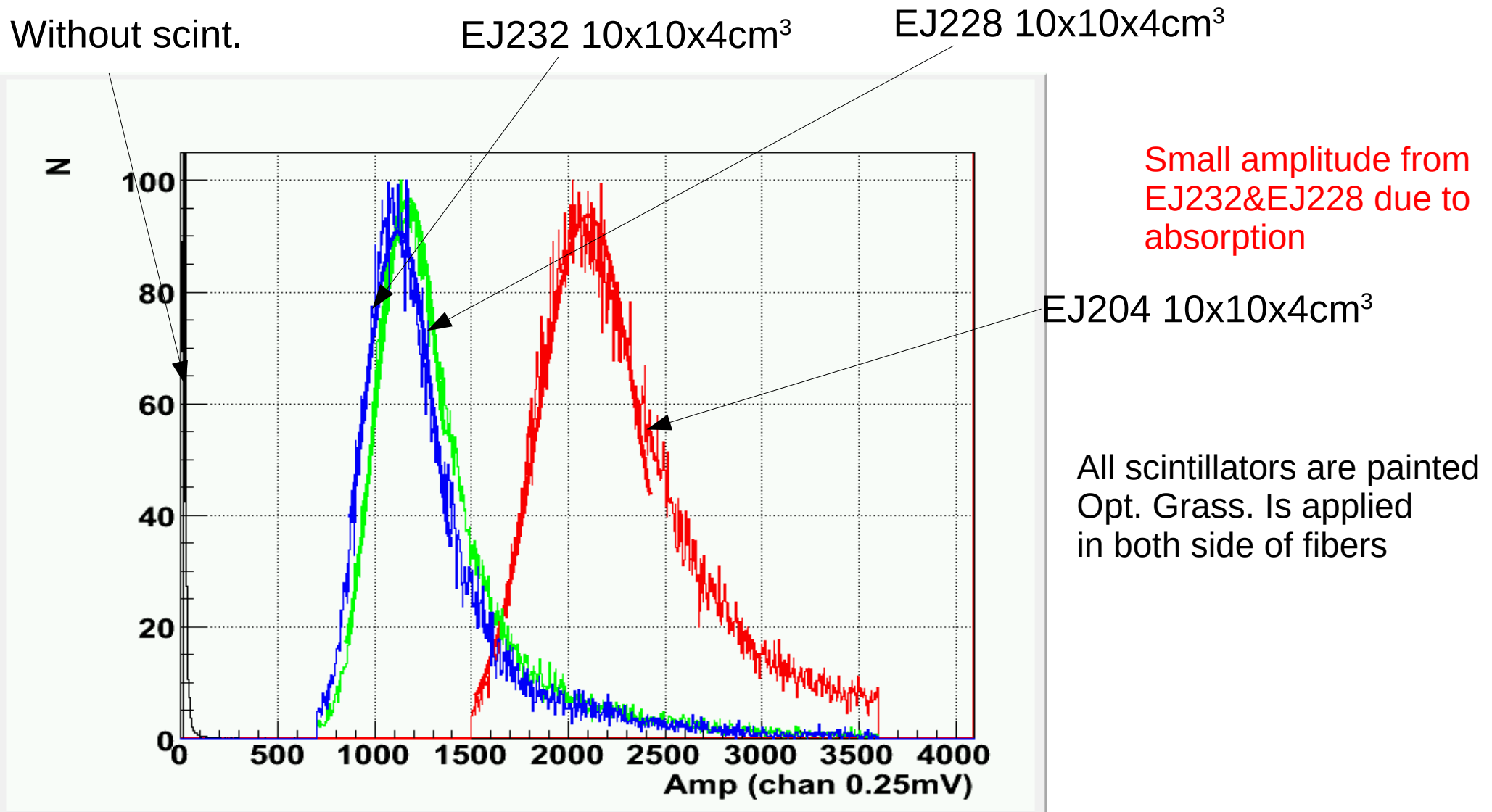
| PROPERTIES | EJ-200 | EJ-204 | EJ-208 | EJ-212 |
|---|--|--------|--------|--------|
| Light Output (% Anthracene) | 64 | 68 | 60 | 65 |
| Scintillation Efficiency (photons/1 MeV e ⁻) | 10,000 | 10,400 | 9,200 | 10,000 |
| Wavelength of Maximum Emission (nm) | 425 | 408 | 435 | 423 |
| Light Attenuation Length (cm) | 380 | 160 | 400 | 250 |
| Rise Time (ns) | 0.9 | 0.7 | 1.0 | 0.9 |
| Decay Time (ns) | 2.1 | 1.8 | 3.3 | 2.4 |
| Pulse Width, FWHM (ns) | 2.5 | 2.2 | 4.2 | 2.7 |
| No. of H Atoms per cm ³ (x10 ²²) | 5.17 | 5.15 | 5.17 | 5.17 |
| No. of C Atoms per cm ³ (x10 ²²) | 4.69 | 4.68 | 4.69 | 4.69 |
| No. of Electrons per cm ³ (x10 ²³) | 3.33 | 3.33 | 3.33 | 3.33 |
| Density (g/cm ³) | 1.023 | 1.023 | 1.023 | 1.023 |
| Polymer Base | Polyvinyltoluene | | | |
| Refractive Index | 1.58 | | | |
| Softening Point | 75°C | | | |
| Vapor Pressure | Vacuum-compatible | | | |
| Coefficient of Linear Expansion | 7.8 x 10 ⁻⁵ below 67°C | | | |
| Light Output vs. Temperature | At 60°C, L.O. = 95% of that at 20°C No change from 20°C to -60° | | | |
| Temperature Range | -20°C to 60°C | | | |



| PROPERTIES | EJ-228 | EJ-230 |
|---|--------|--------|
| Light Output (% Anthracene) | 67 | 64 |
| Scintillation Efficiency (photons/1 MeV e⁻) | 10,200 | 9,700 |
| Wavelength of Maximum Emission (nm) | 391 | 391 |
| Light Attenuation Length (cm) | - | 120 |
| Rise Time (ns) | 0.5 | 0.5 |
| Decay Time (ns) | 1.4 | 1.5 |
| Pulse Width, FWHM (ns) | 1.2 | 1.3 |
| No. of H Atoms per cm³ (x10²²) | 5.15 | 5.15 |
| No. of C Atoms per cm³ (x10²²) | 4.69 | 4.69 |
| No. of Electrons per cm³ (x10²³) | 3.33 | 3.33 |
| Density (g/cm³) | 1.023 | 1.023 |

Prototype 2 with cosmic muons

New cheap EJ204 tests



Optical Fibers

Kuraray and Bycron for the science

Kuraray and Bycron better quality, but more expensive

Commercials
Asahi, Mitsubishi
Etc.

The commercial ones from 5 to 10 times cheaper

Asahi

Configuration: Simplex

Core Diameter: 1.0mm

Core Material: PMMA (Polymethyl Methacrylate)

Numerical Aperture: 0.6

Attenuation: 0.19 dB/m at 650 nm

Operating Temperature Range: -55°C to 85°C

Mitsubishi

Configuration: Simplex

Core Diameter: 1.0mm

Core Material: PMMA (Polymethyl Methacrylate)

Numerical Aperture: 0.5

Attenuation: 0.20 dB/m at 650 nm

Operating Temperature Range: -55°C to 70°C

Kuraray

Configuration: Simplex

Core Diameter: 1.0mm

Core Material: ?

Numerical Aperture: 0.55

Attenuation: ?

Operating Temperature Range: ?

Photo-sensors

FINE MESH PMT SERIES for HIGH MAGNETIC FIELD ENVIRONMENTS

| Tube Diameter | Type Number | Spectral Response Range(nm) & Curve Code | ① Outline No. | ② Socket | No. of Stages | Cathode Sensitivity | | | Anode Sensitivity | | | | | | | |
|---------------|-------------|--|---------------|----------|---------------|------------------------------------|---------------------------------|-----------------------|----------------------|-------------------|-------------------|-------------------|--|---------|---------------------------------------|-----------|
| | | | | | | Luminous Typ. ($\mu\text{A/lm}$) | Blue Sens. Index (CS 5-58) Typ. | Q.E. at Peak Typ. (%) | Luminous Typ. (A/Lm) | Nominal Gain ④ | | | Supply Voltage ④ for Nominal Gain at 0 T | | Dark Current at ④ Nominal Gain at 0 T | |
| | | | | | | | | | | (at 0 T) | (at 0.5 T) | (at 1 T) | Typ.(V) | Max.(V) | Typ. (nA) | Max. (nA) |
| 25 mm (1") | R5505 | 300 to 650 400K | ① | E678-17A | 15 | 80 | 9.5 (7.0) | 23 | 40 | 5.0×10^5 | 2.3×10^5 | 1.8×10^4 | 1850 | 2300 | 5 | 30 |
| 38 mm (1.5") | R5946 | 300 to 650 400K | ② | E678-19D | 16 | 80 | 9.5 (7.0) | 23 | 80 | 1.0×10^6 | 4.3×10^5 | 2.9×10^4 | 1800 | 2300 | 5 | 30 |
| 38 mm (1.5") | R7761 | 300 to 650 400K | ③ | - | 19 | 80 | 9.5 (7.0) | 23 | 800 | 1.0×10^7 | 3.0×10^5 | 1.5×10^5 | 1800 | 2300 | 15 | 100 |
| 51 mm (2") | R5924 | 300 to 650 400K | ④ | - | 19 | 70 | 9.0 (7.0) | 22 | 700 | 1.0×10^7 | 4.1×10^5 | 2.0×10^5 | 1750 | 2300 | 30 | 200 |
| 64 mm (2.5") | R6504 | 300 to 650 400K | ⑤ | - | 19 | 70 | 9.0 (7.0) | 22 | 700 | 1.0×10^7 | 4.1×10^5 | 2.0×10^5 | 1750 | 2300 | 50 | 300 |

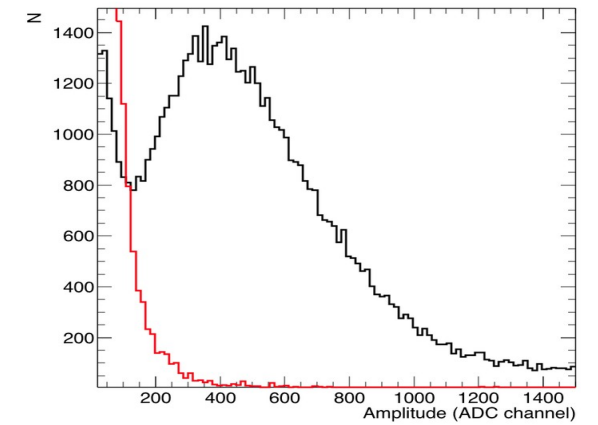
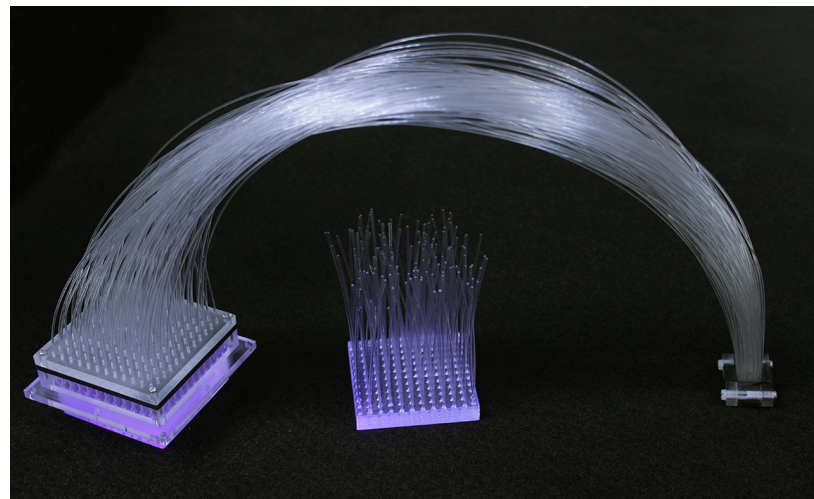
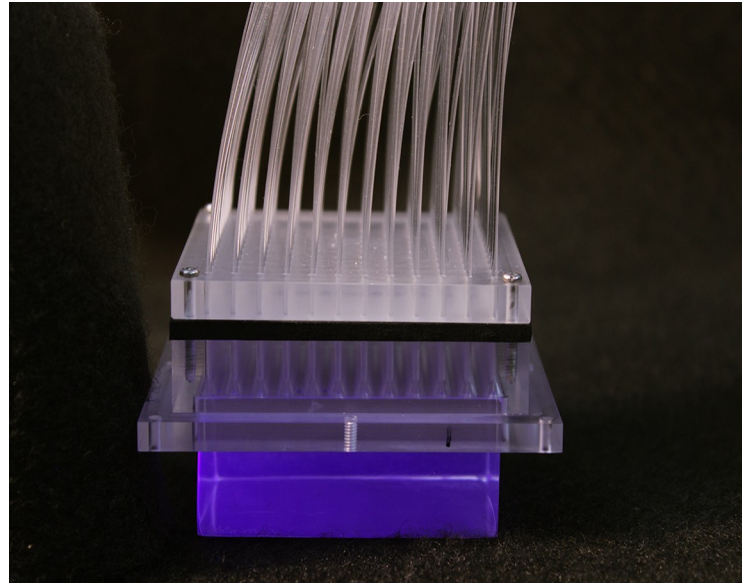
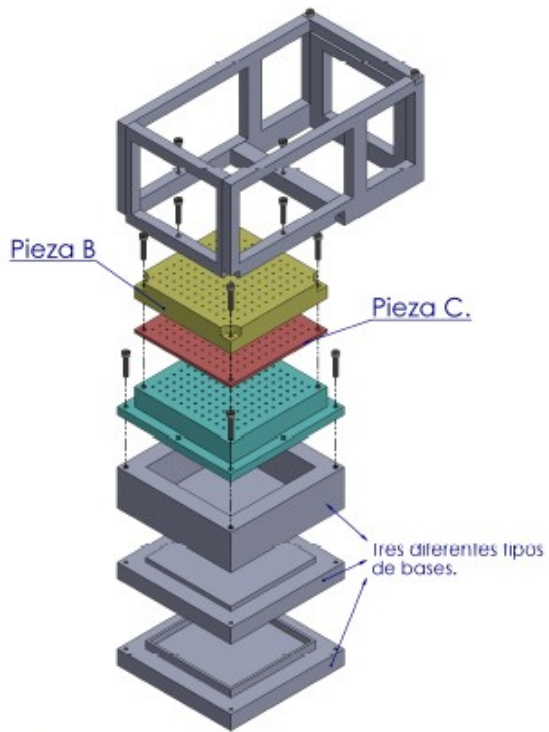


SiPM
Radiation
Cheap
FEE

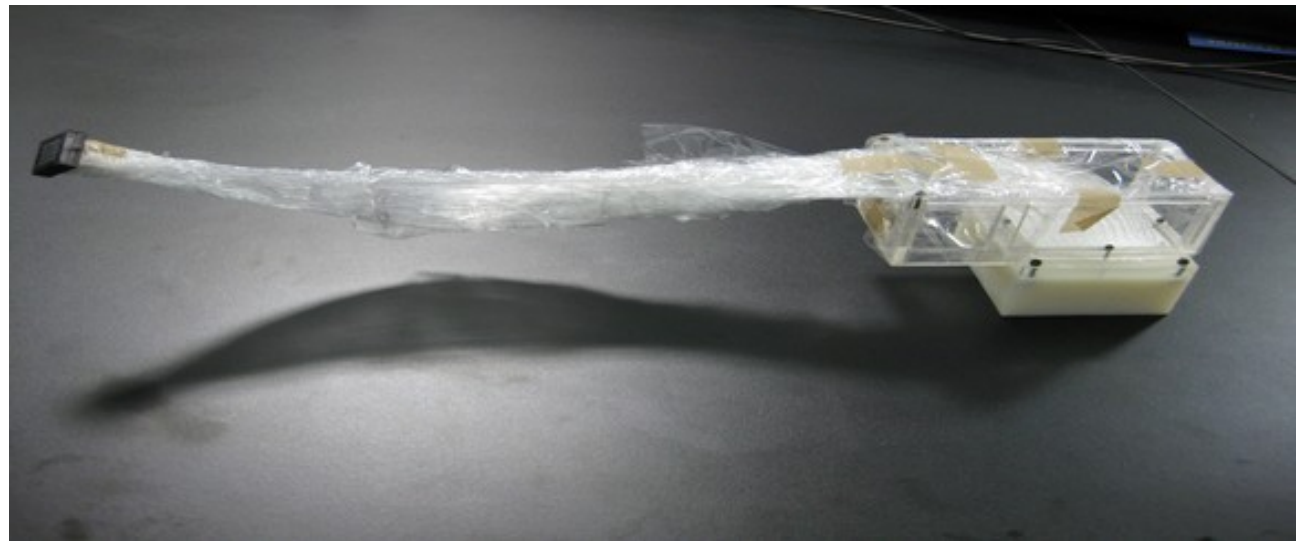
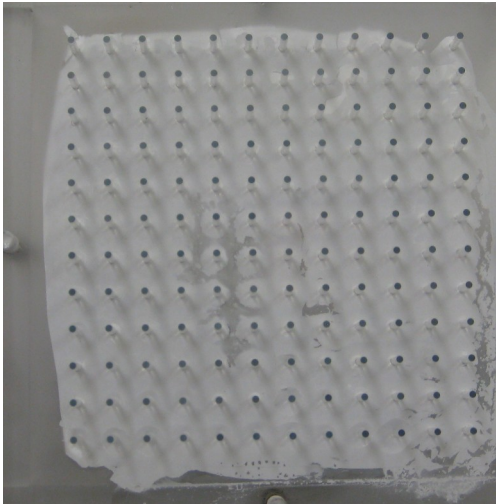
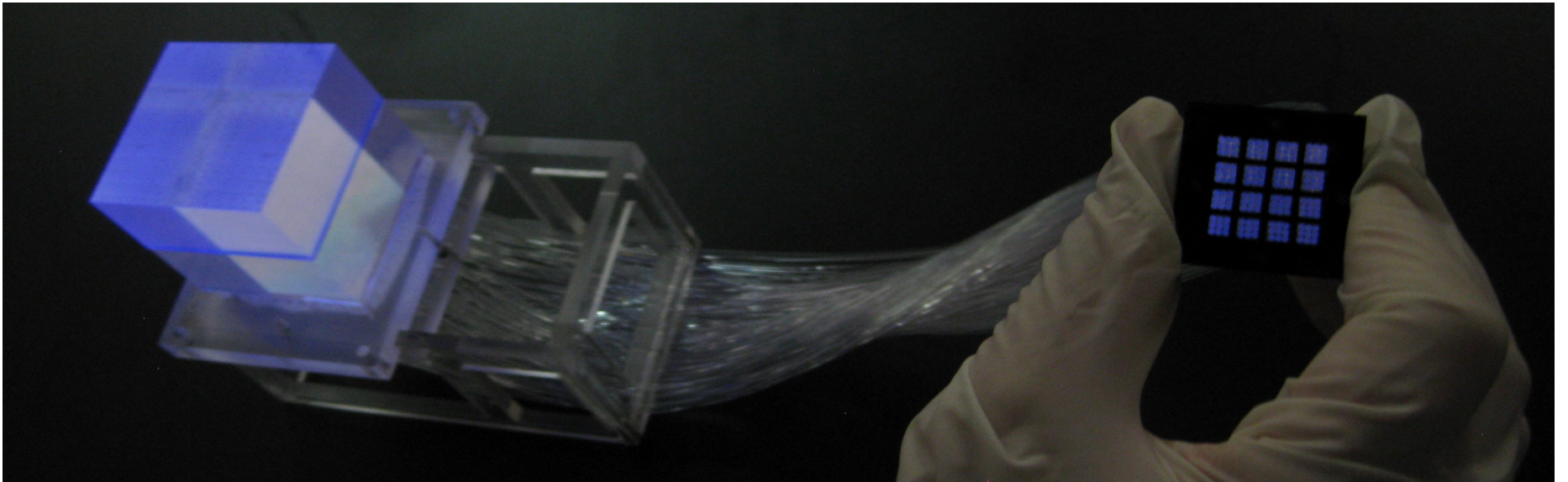
MCP
More expensive
Limited accumulated
charge
Limited anode current
Fast and short pulses



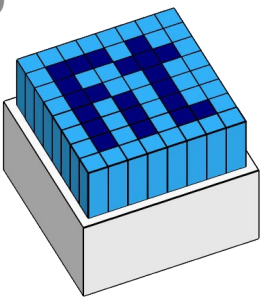
First Prototype



Second Prototype



0
1
2
3
4
5
6
7
8
9
A
B
C
D
E
F
G
H
I
J
K
L
M
N
O
P
Q
R
S
T
U
V
W
X
Y
Z

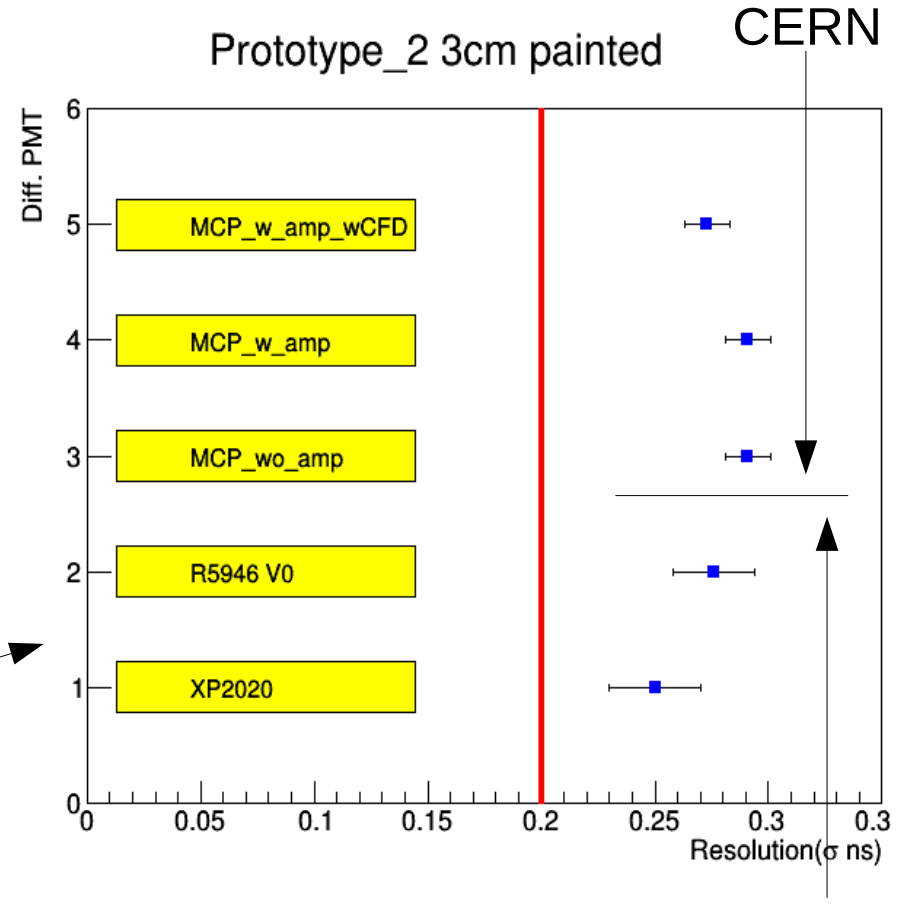
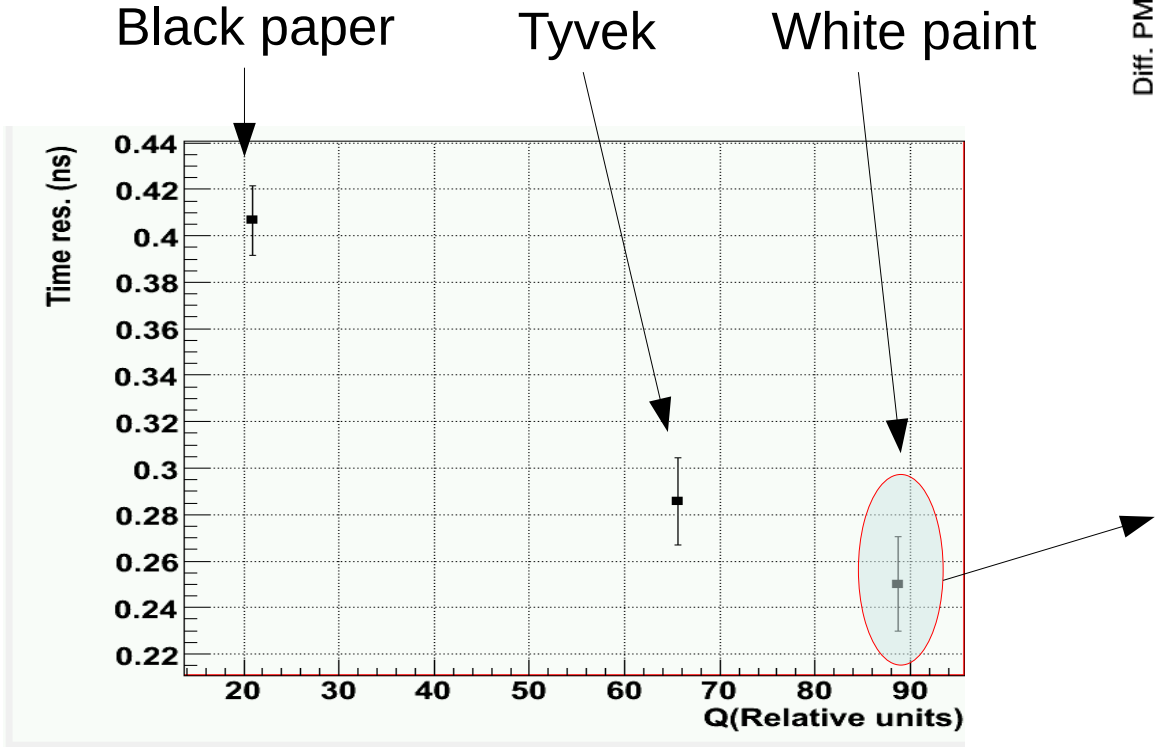


Scintillator coating study



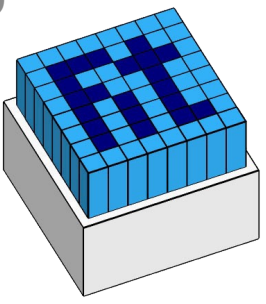
- To reach 200ps limit should be increase the light
- The thickness of scintillator
 - The density of the fibers

Variation between 30 ps

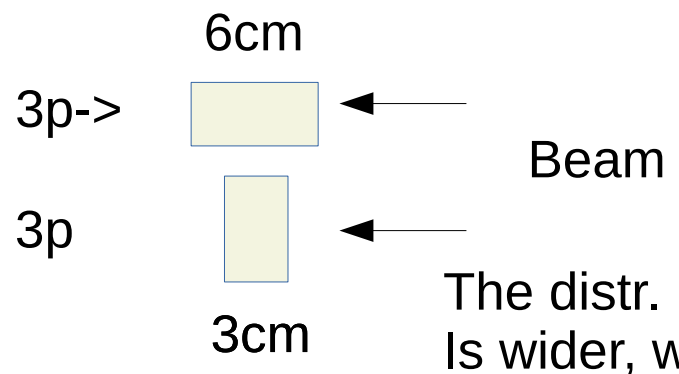


cosmic

0
1
2
3
4
5
6
7
8
9
A
B
C
D
E
F
G
H
I
J
K
L
M
N
O
P
Q
R
S
T
U
V
W
X
Y
Z

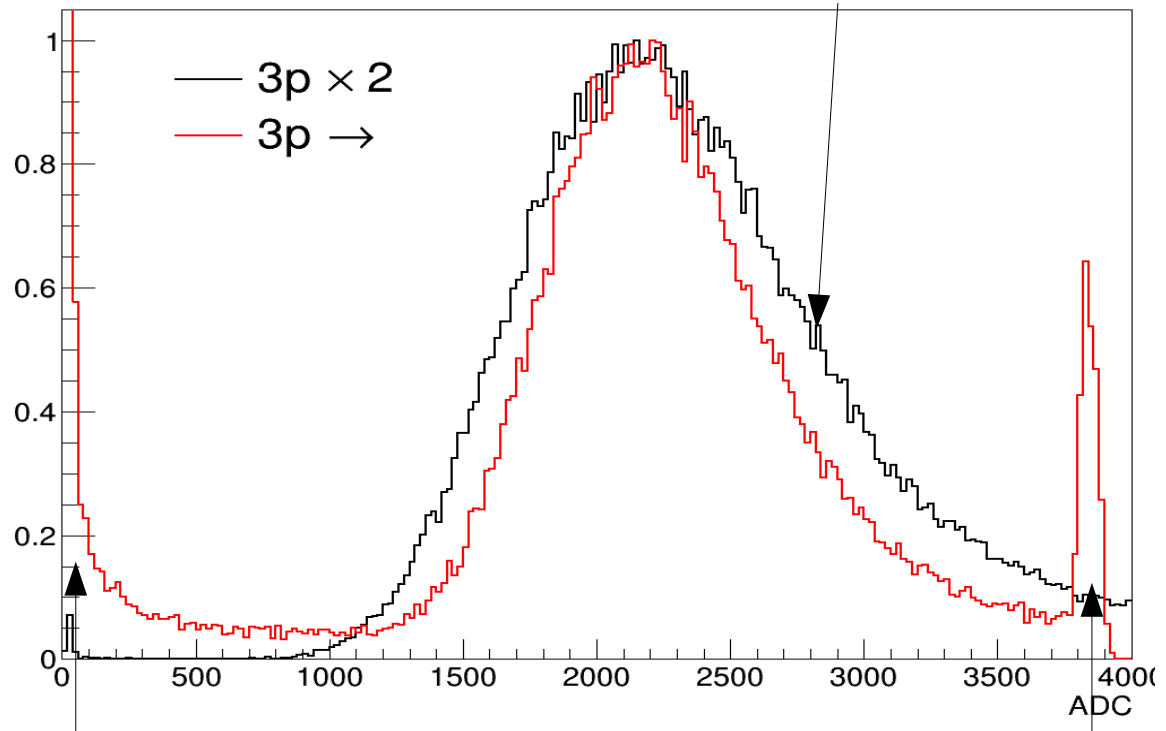


Scintillator thickness study



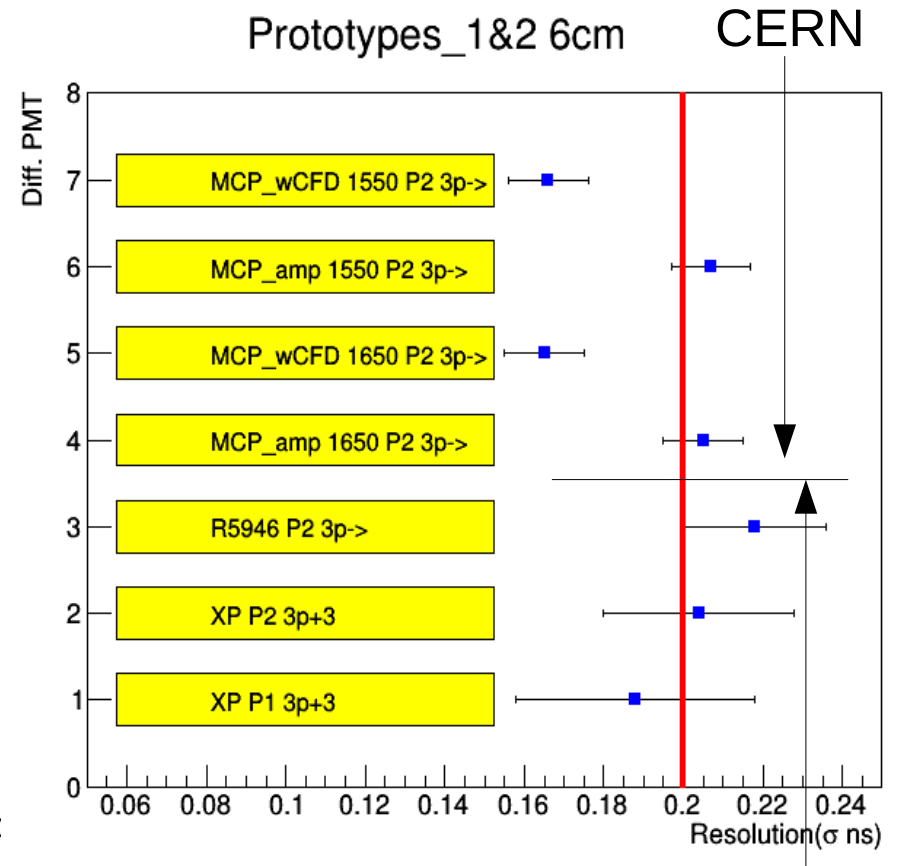
Significant improvement ~40 ps
Using CFD ?
Should be checked our time estimation

The distr. In case of 3p
Is wider, which is expected



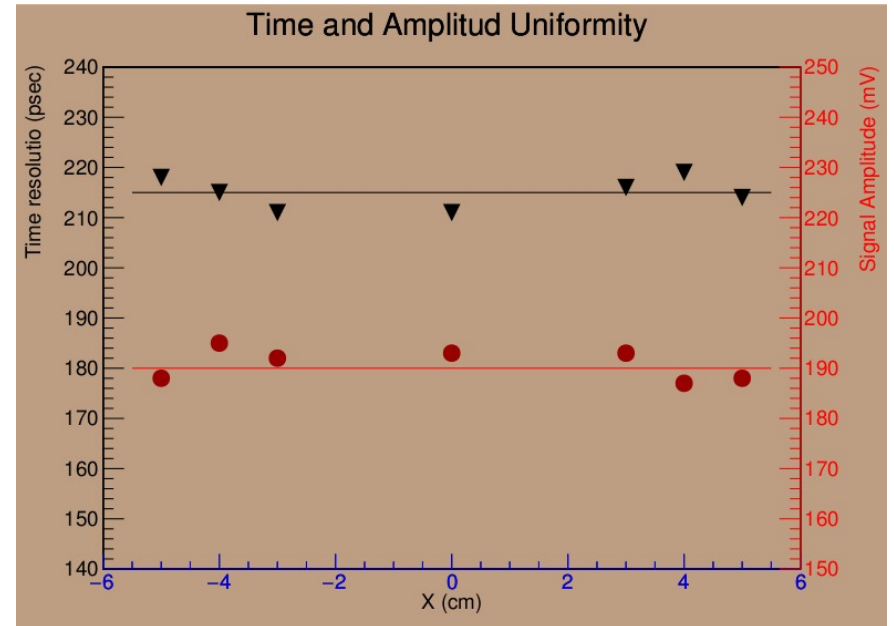
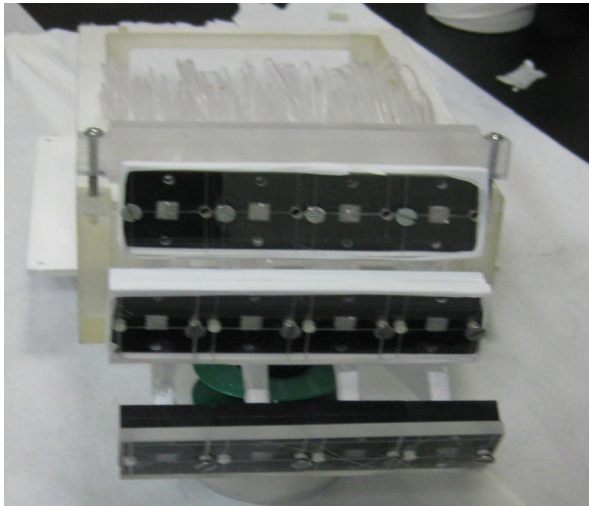
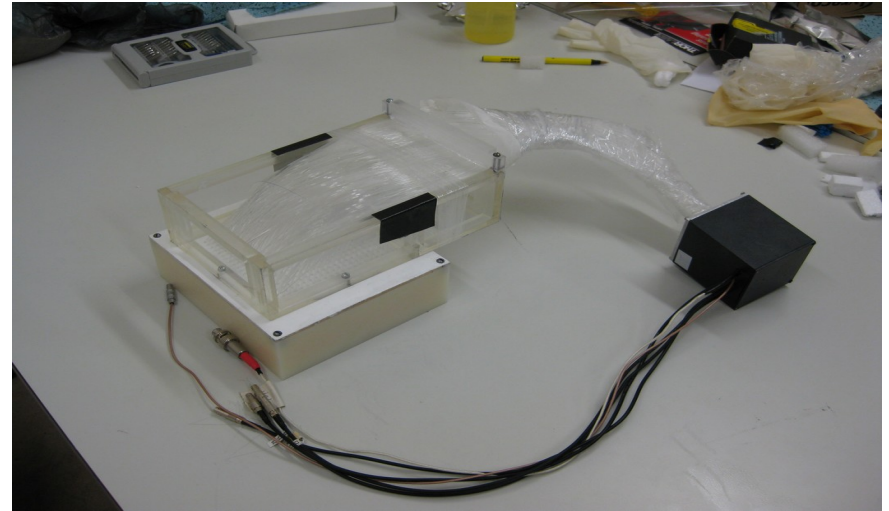
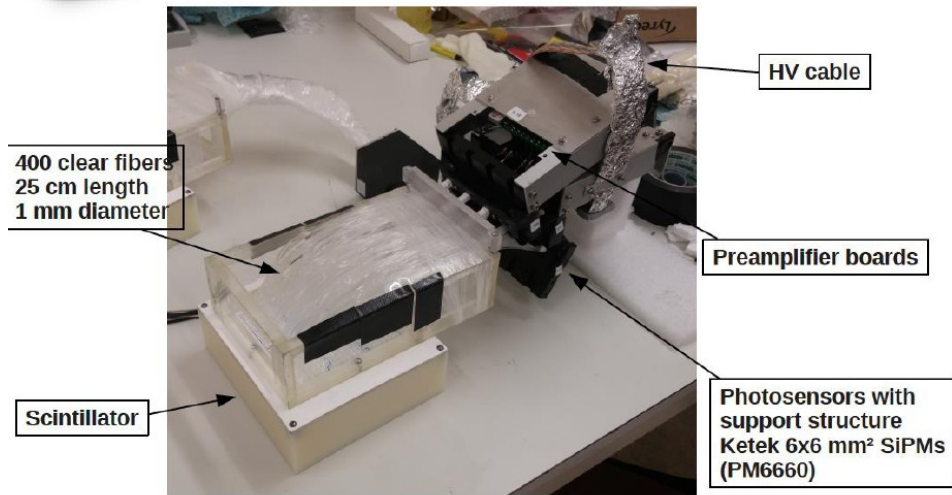
Part of beam is out

ADC saturation

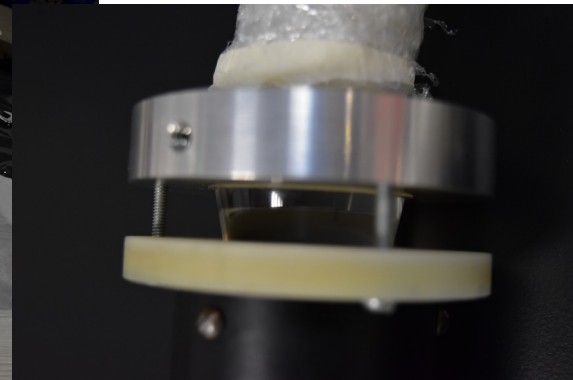
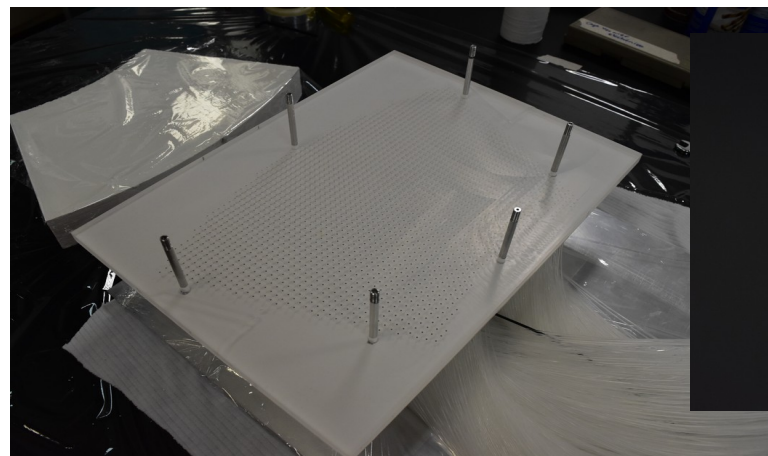
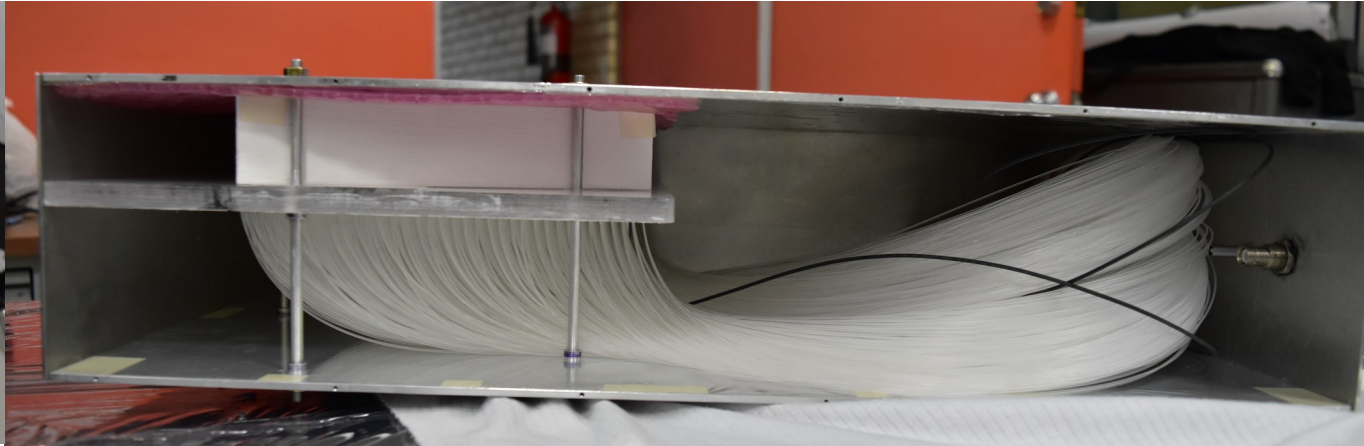
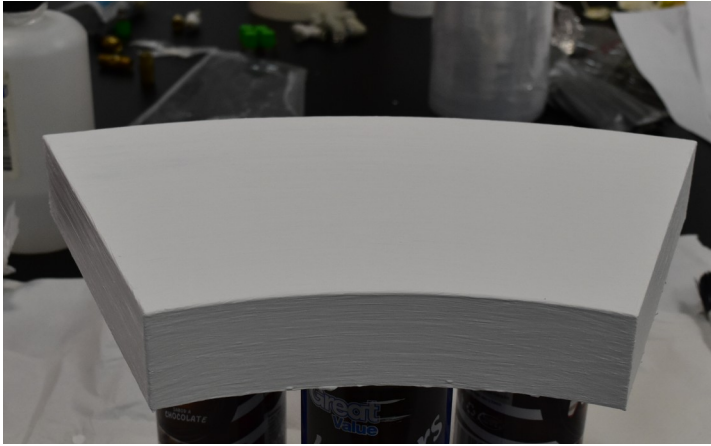


Cosmic muons

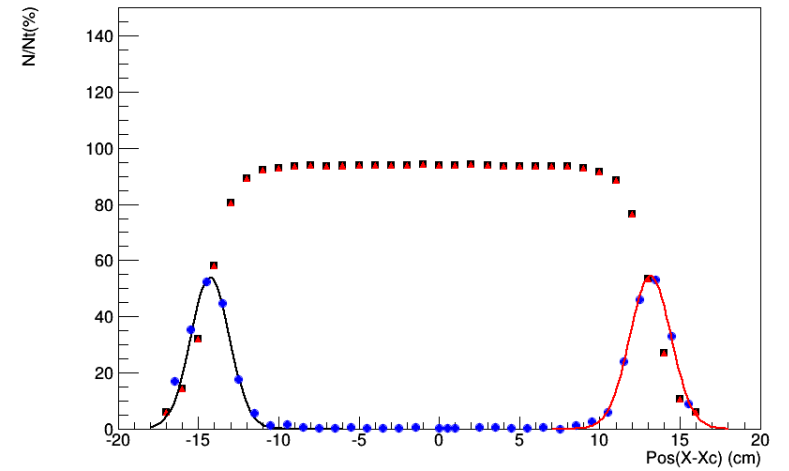
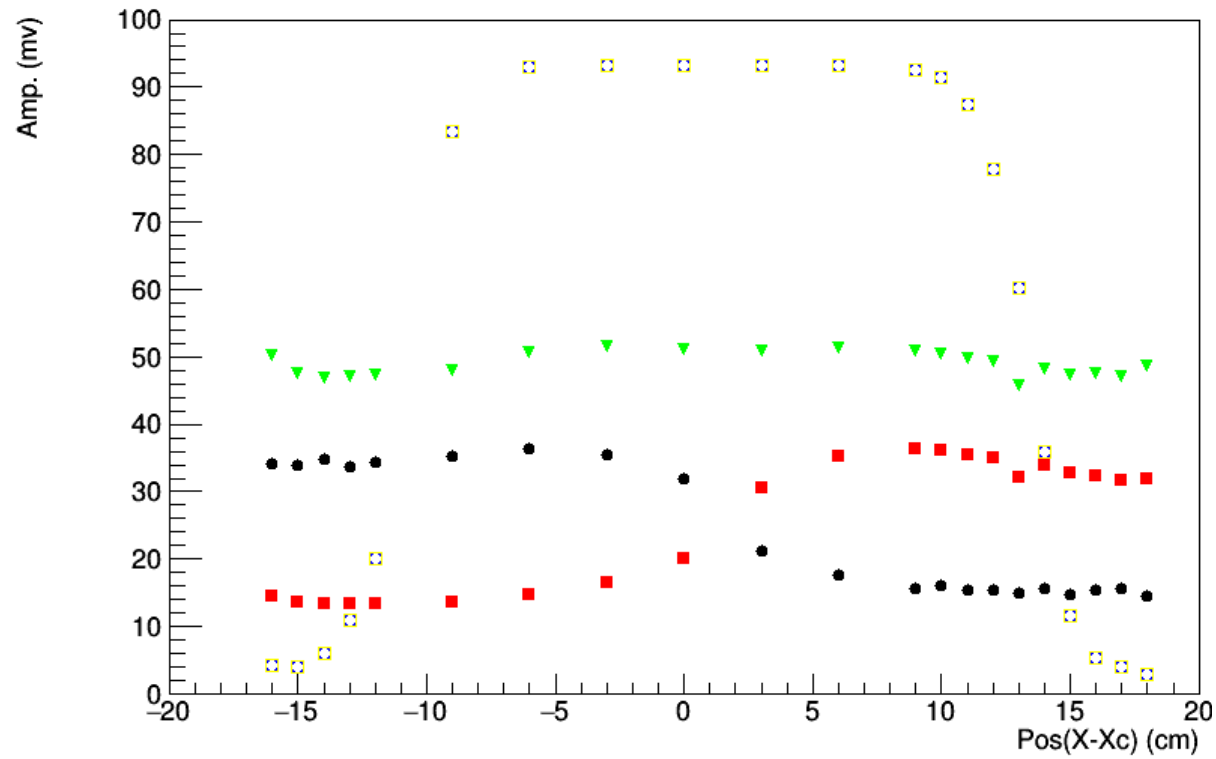
Third Prototype



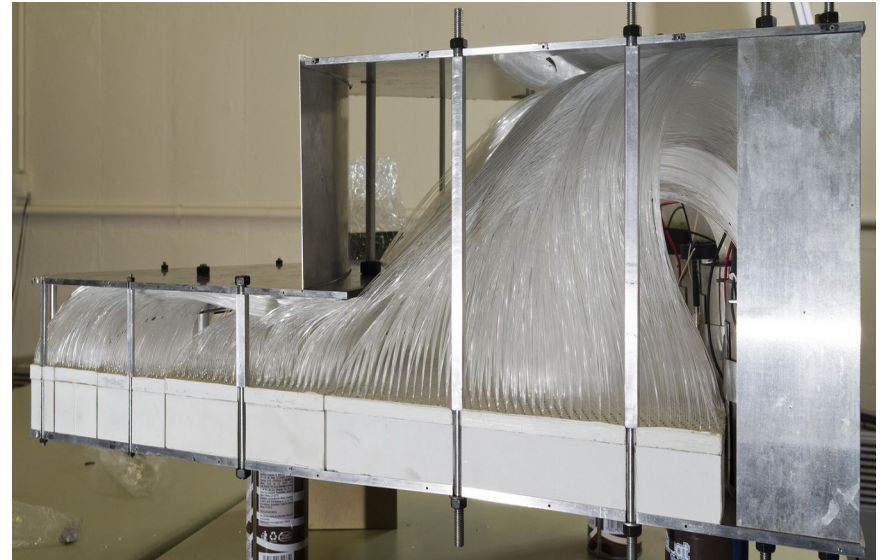
Forth Prototype



Test beam results for the forth Prototype

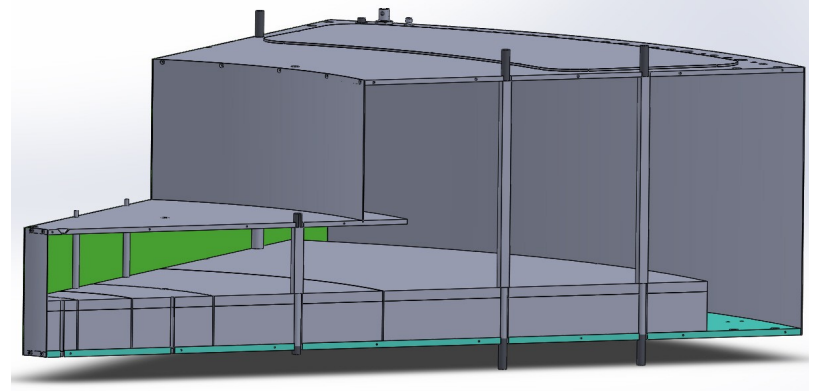
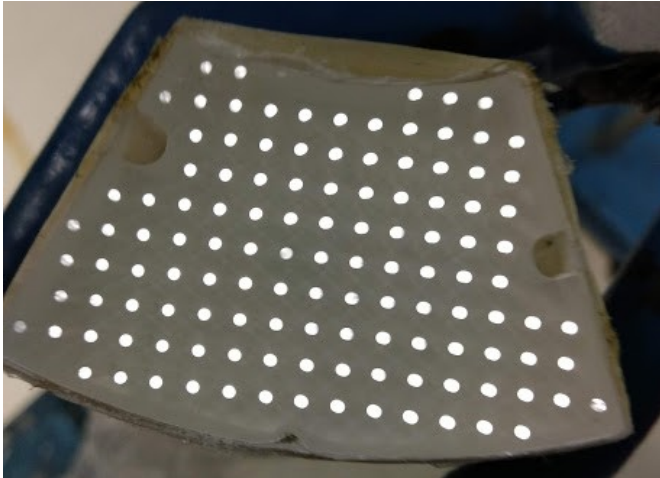


Sector Prototype

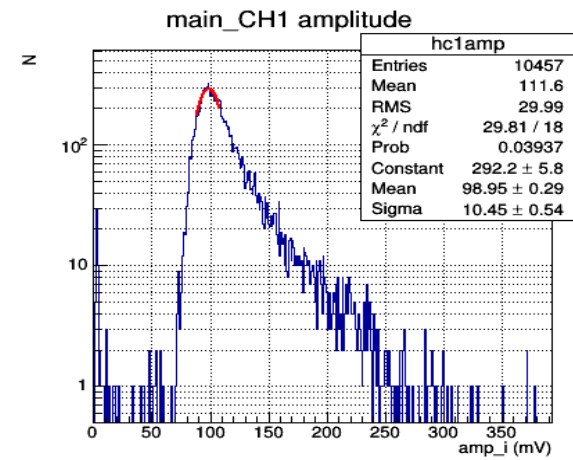
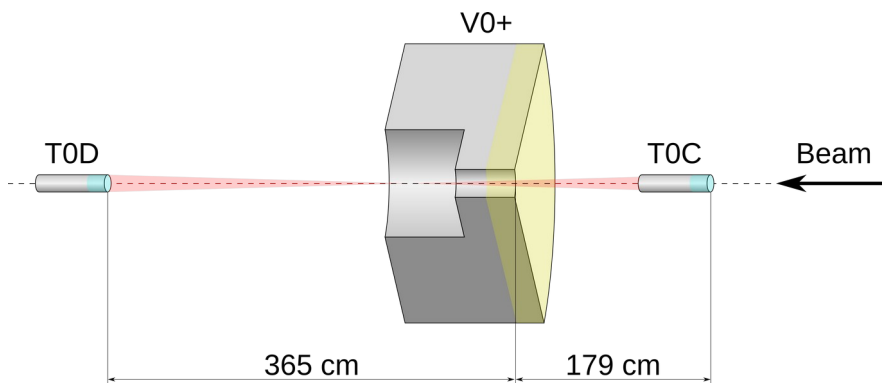
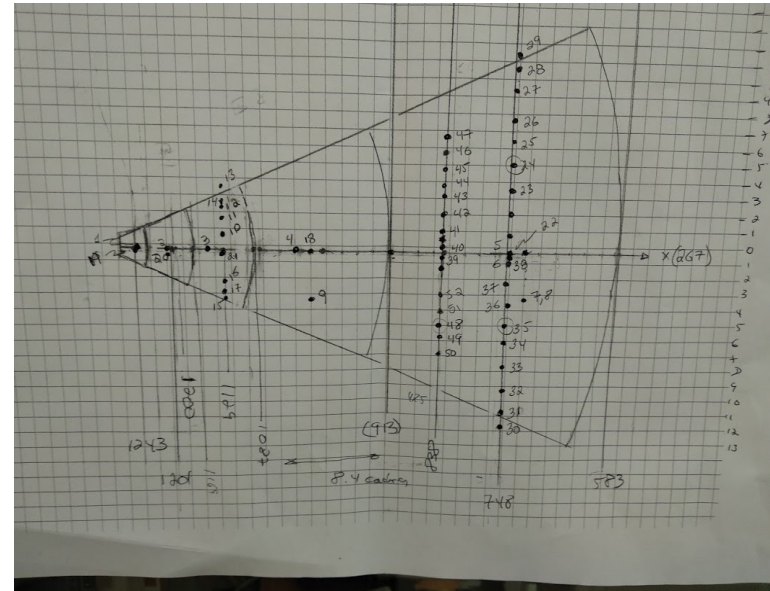


V.Grabski

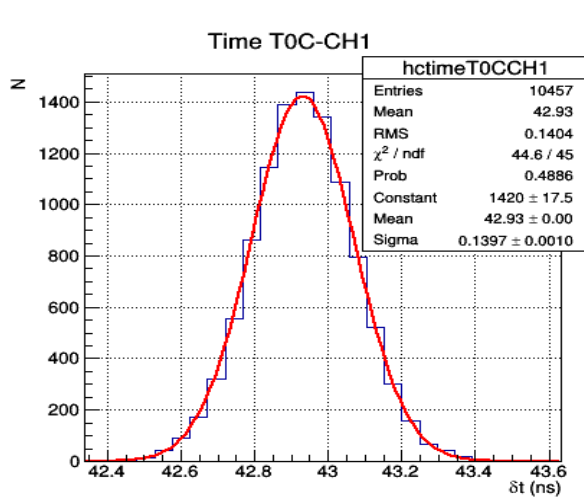
Sector Prototype



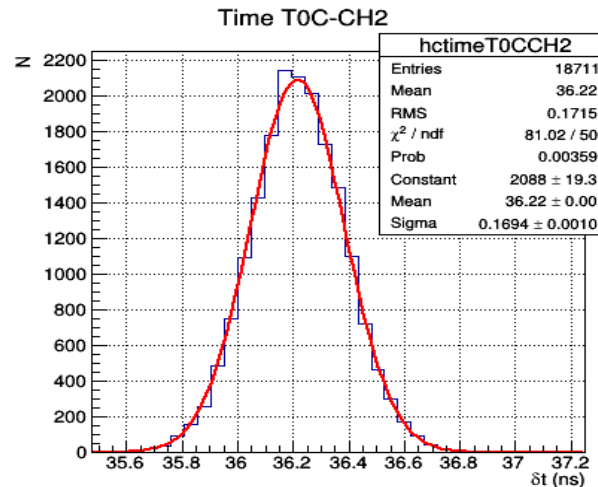
Test beam setup



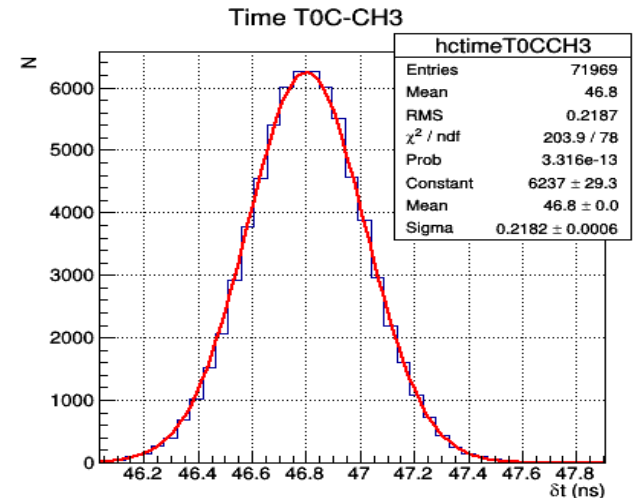
Test beam results for the Sector Prototype



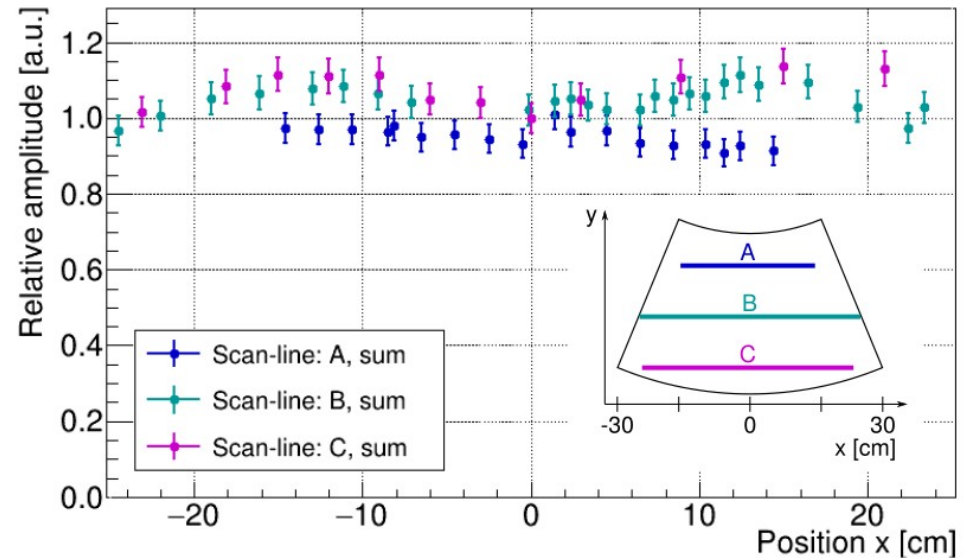
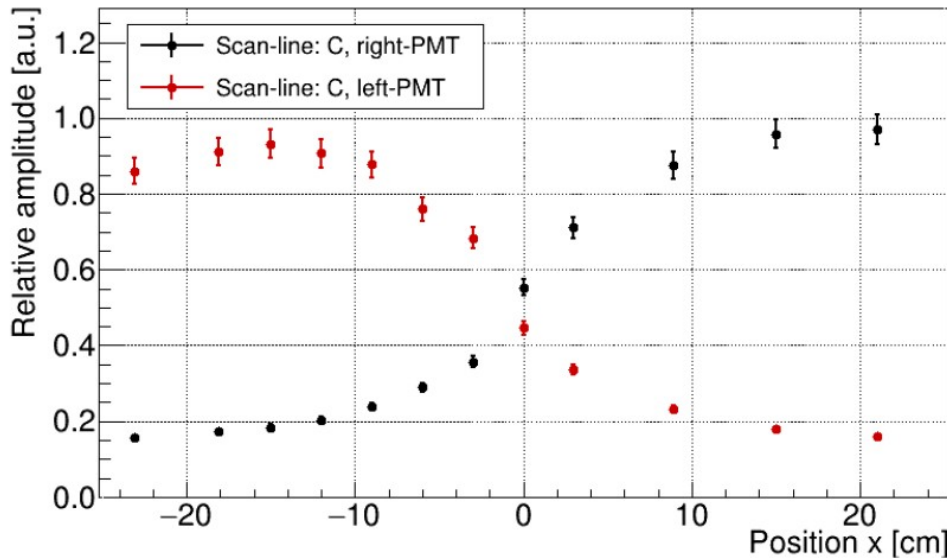
Fiber distance 3mm



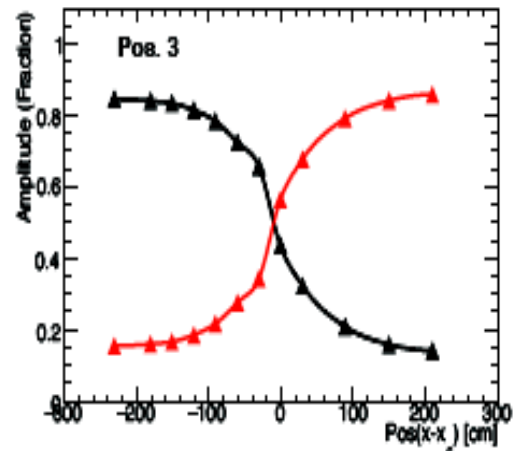
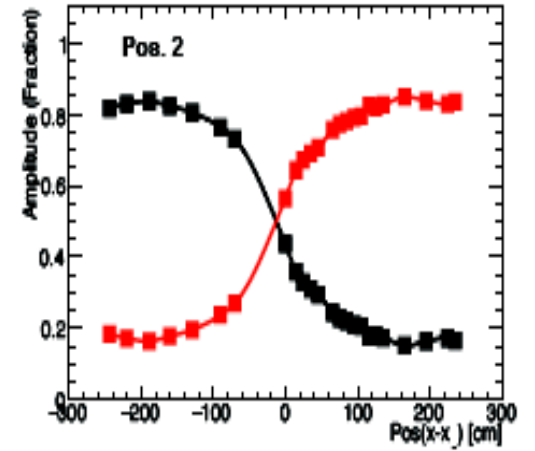
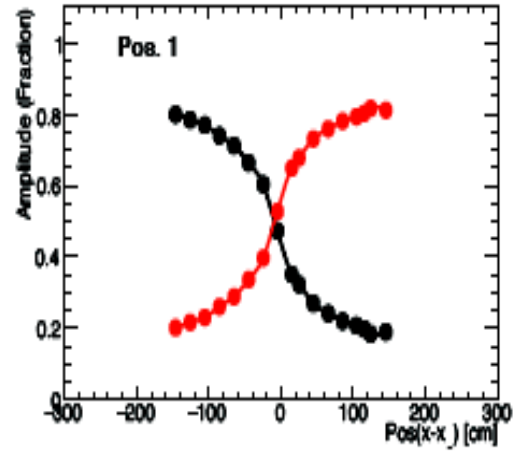
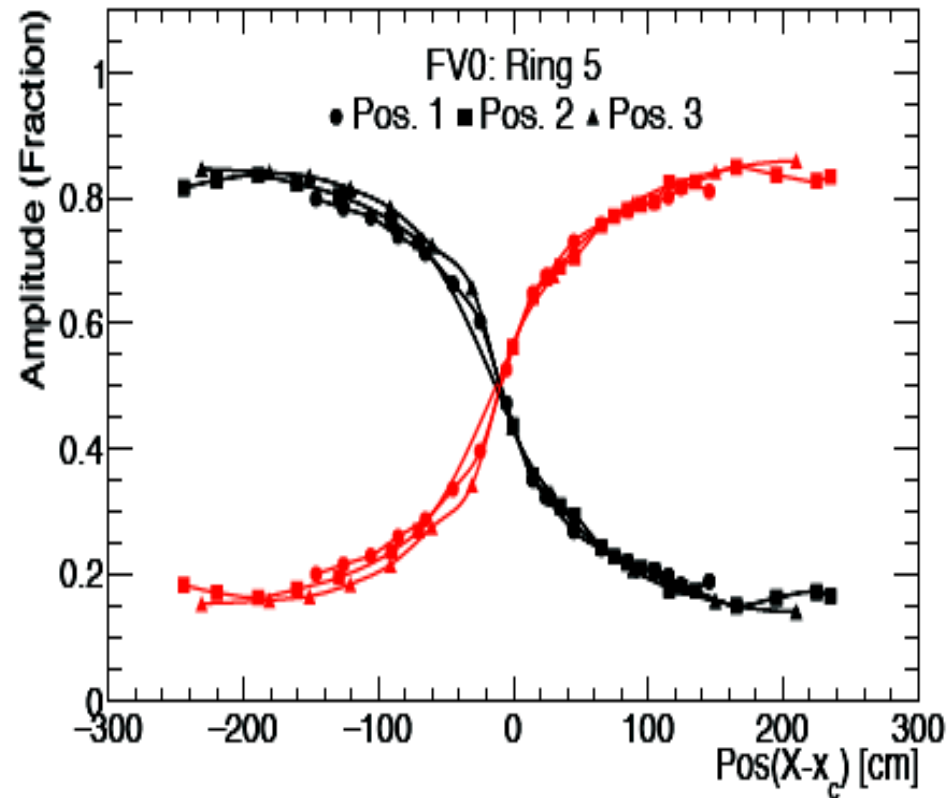
Fiber distance 4mm



Fiber distance 5mm



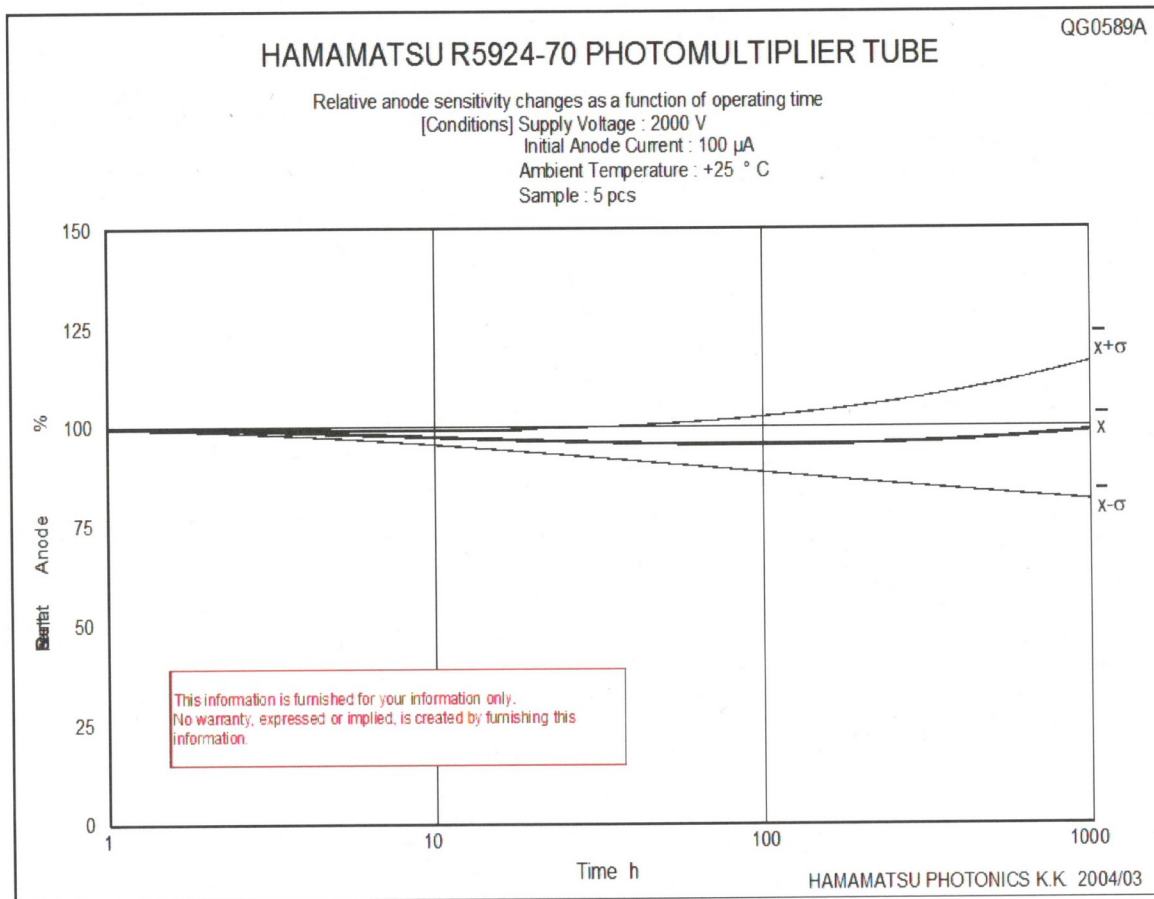
Normalized amplitudes in different positions



FV0: Ring 5
● Pos. 1 ■ Pos. 2 ▲ Pos. 3

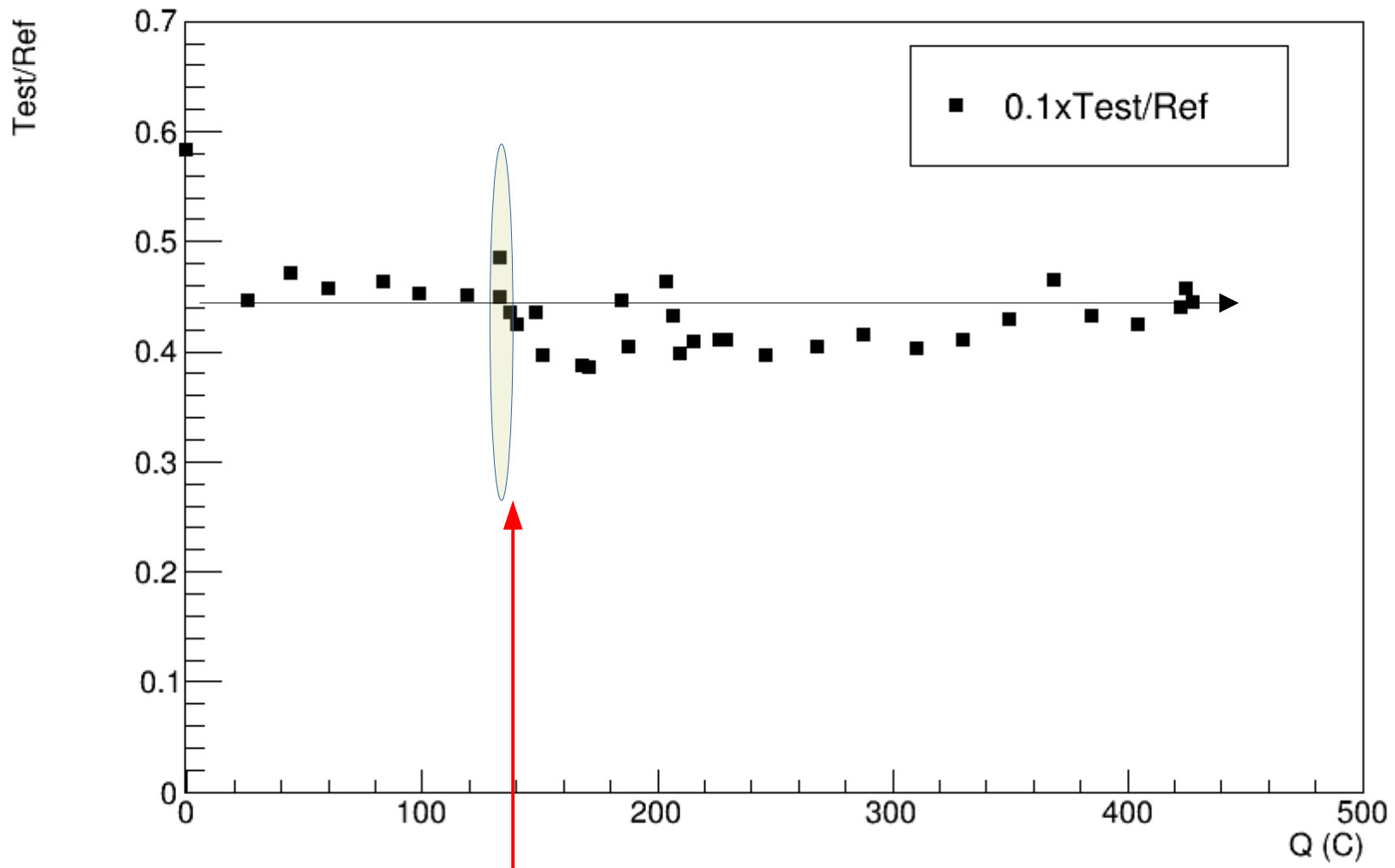
Life time tests for 2" fine mesh

Similar to typical one below 1000 hours



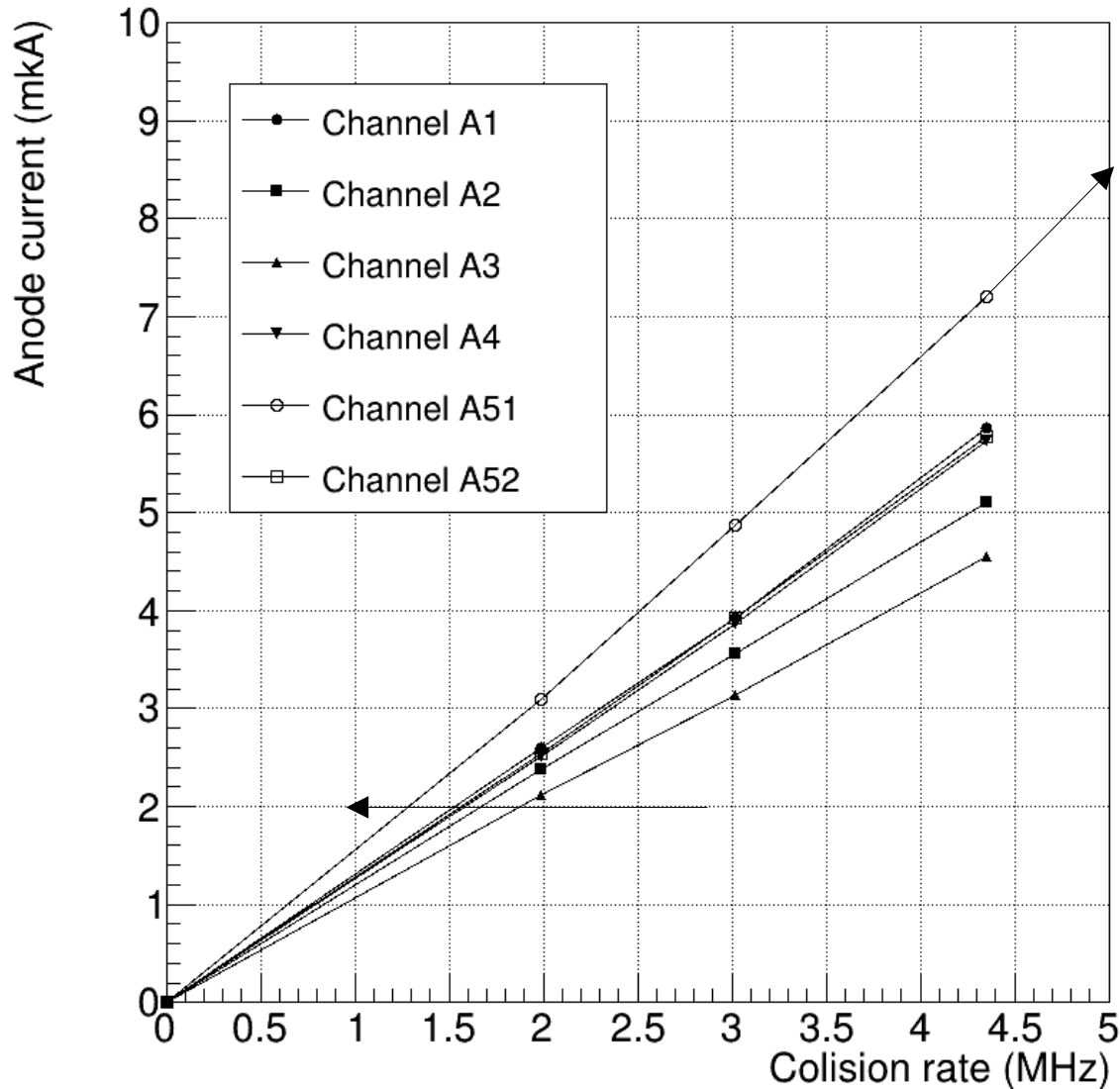
In average non stability is less than 5% after 1000hours or ~ 360C accumulated charge
Another +/- 15% is the spread.
Comes from the sample of 5pcs

Test/Ref



After 10 days stop

Anode current from rate scan



←

If Pb-Pb equivalent 5MHz pp
Then for 1month Pb-Pb Accumulated
charge is less than 15C(50%
efficiency). 10Years < 150C.
1MHz pp 10 months per year with 50%
Max. curr. 2MkA. Accum. Q < 26C, 10
years <260C. Sum is less than 410C

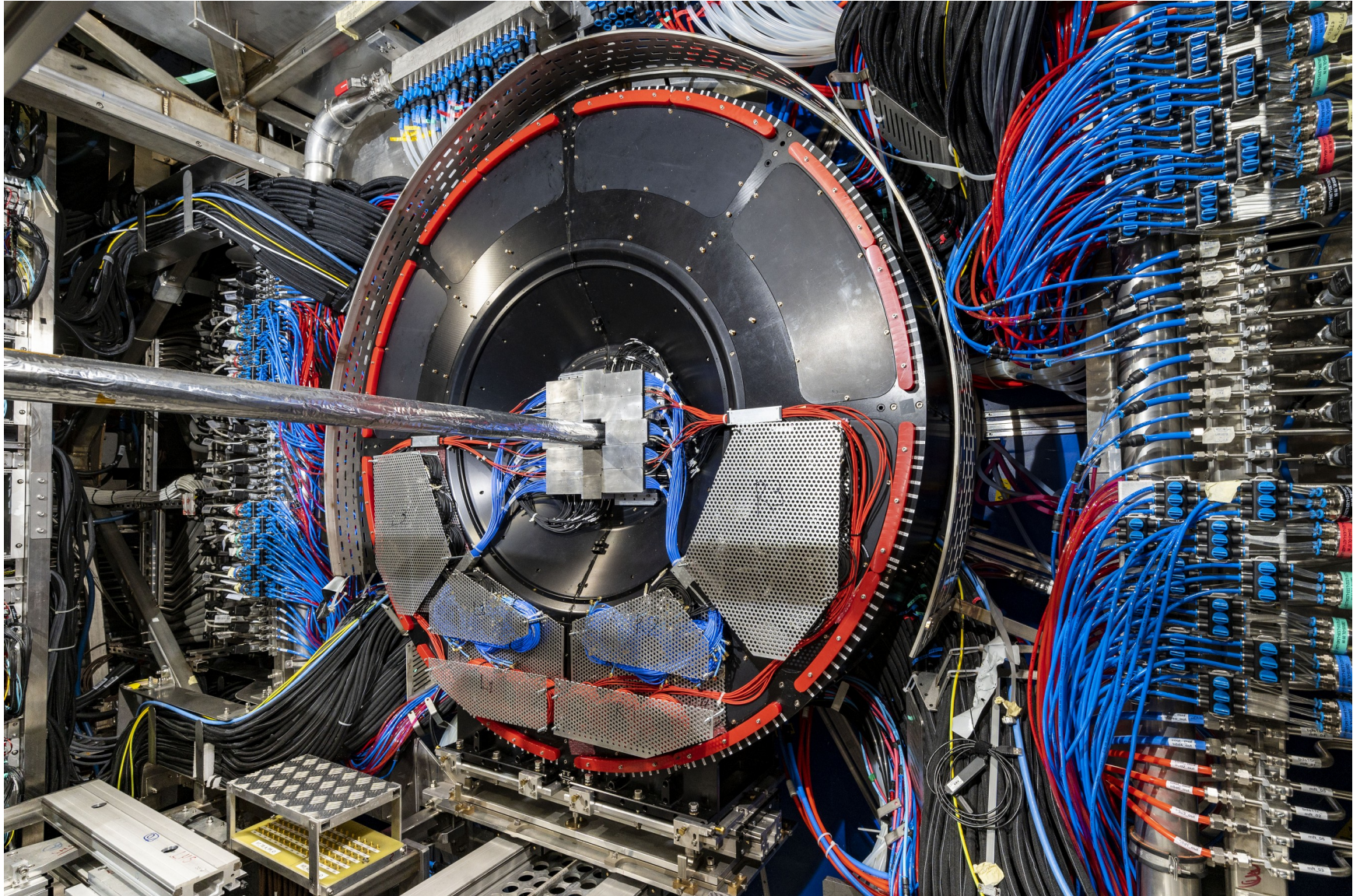
Construction participants



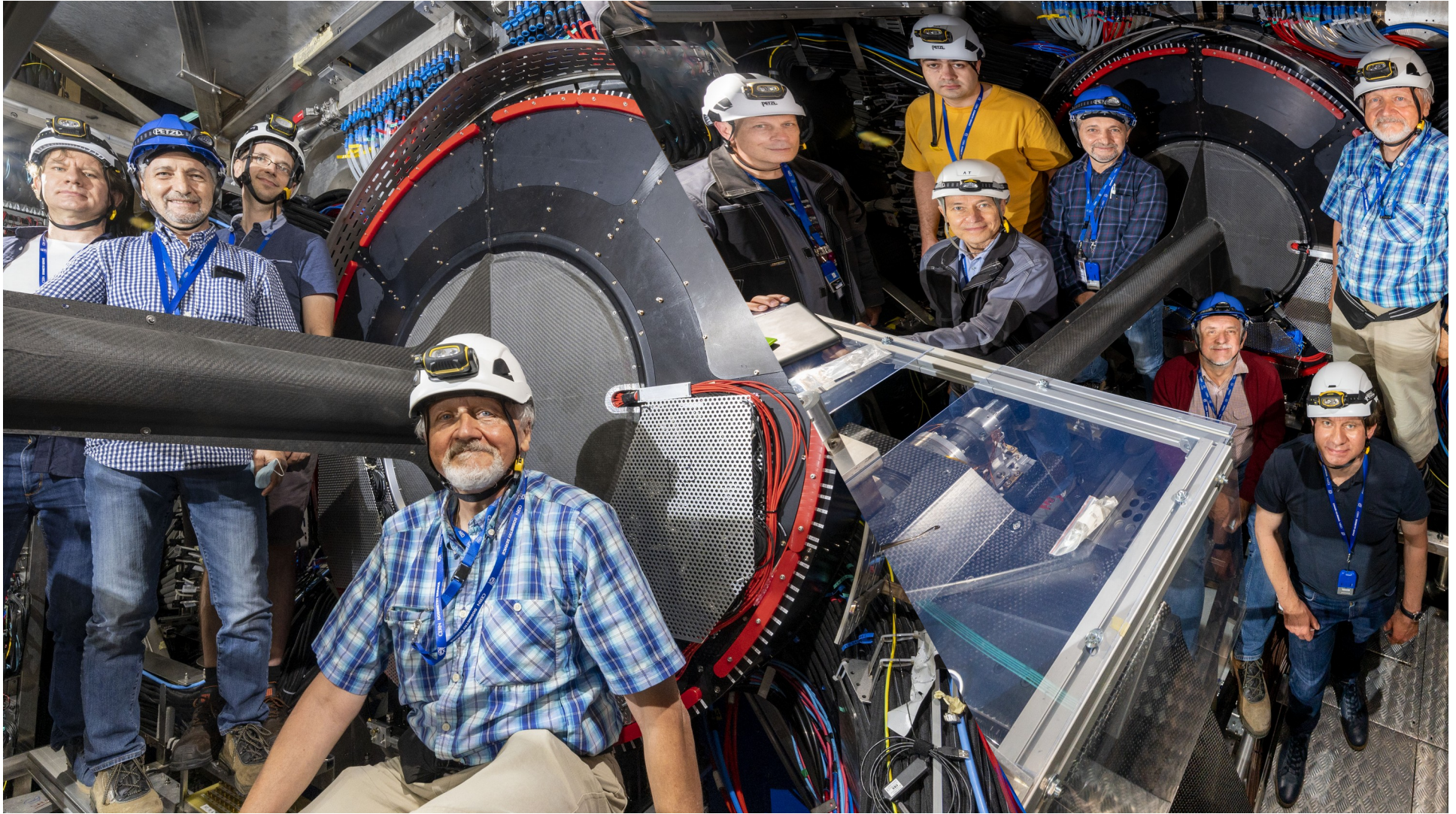
Detector ensemble



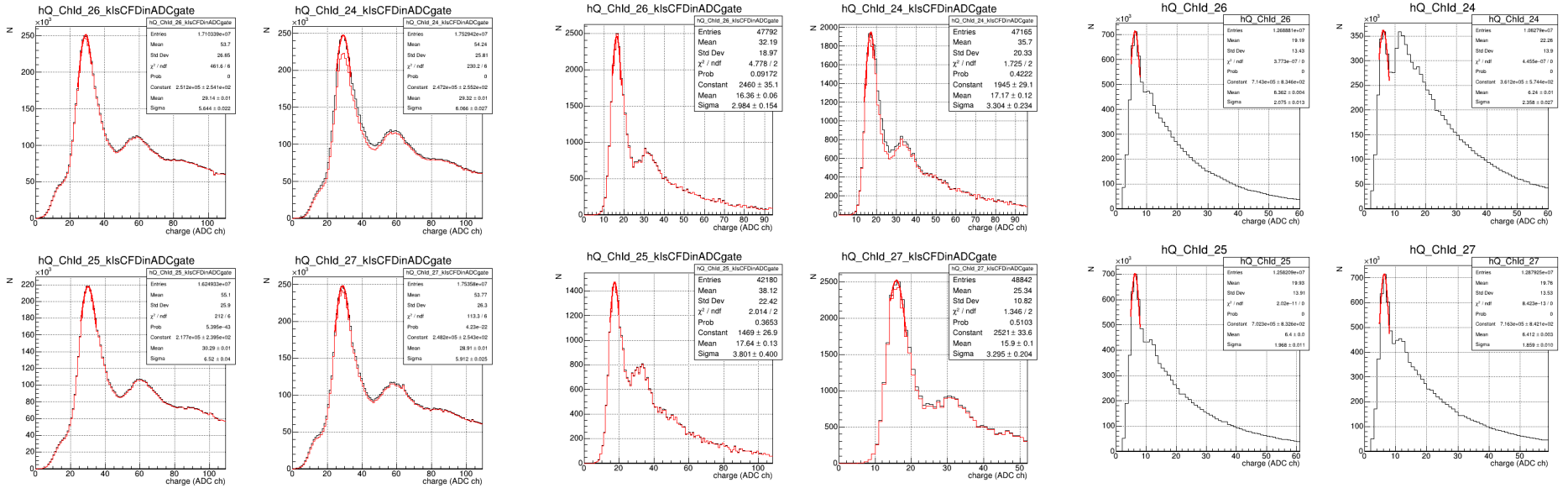
After the installation



After the installation



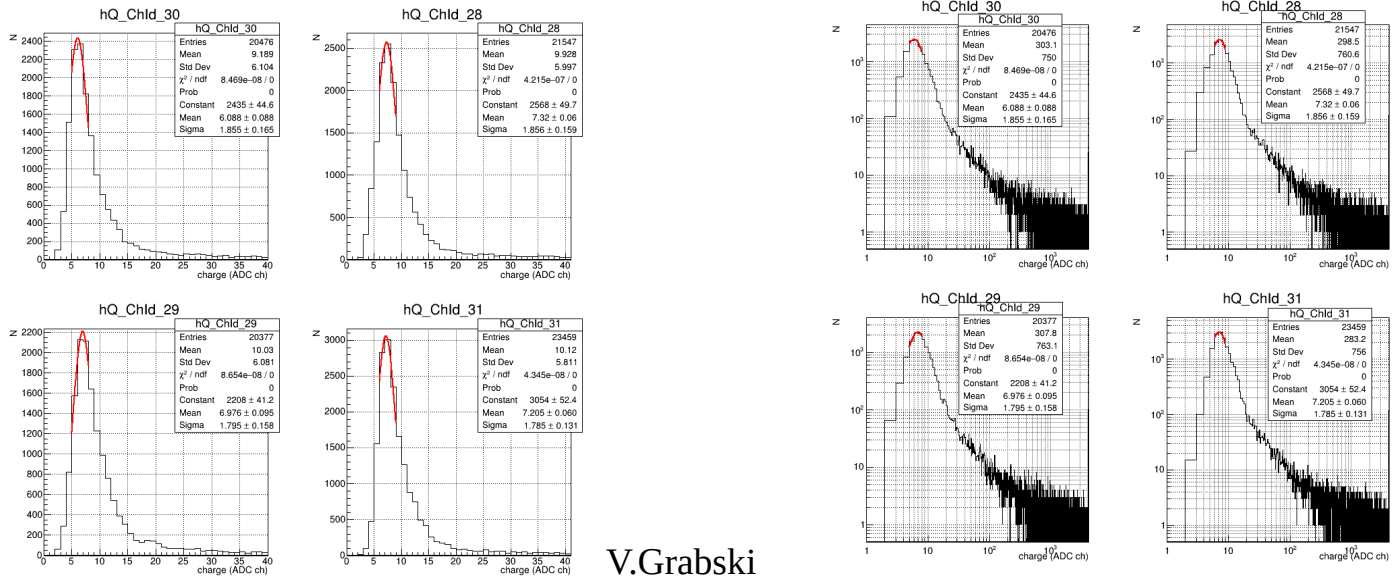
Some Beam Results



MIP = 32ch

MIP = 17ch

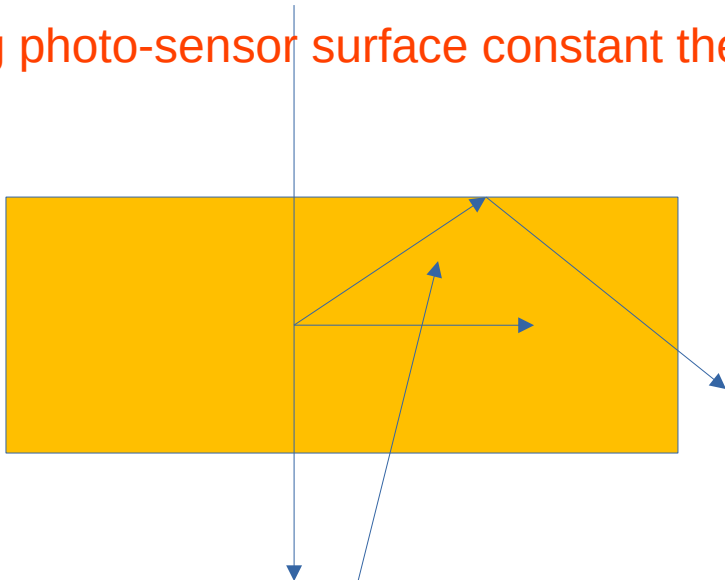
MIP = 6ch



MIP = 6ch

Some ideas on muonID

Keeping photo-sensor surface constant the collected light is independent from scintillator thickness

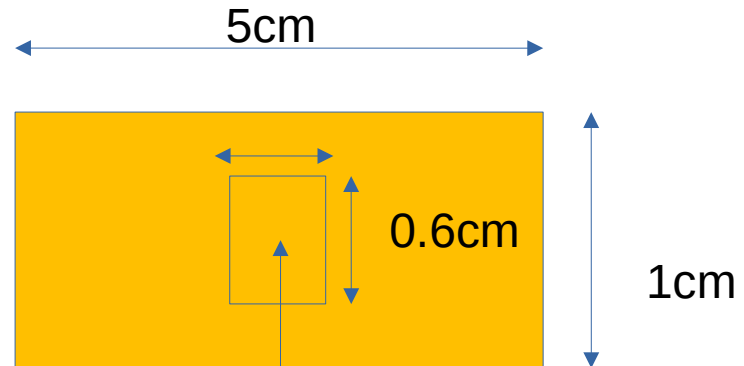


$\theta=50.7^\circ$ internal reflection angle

$(1-\cos\theta)/2 \sim 0.18$, in one side 18% of light is inside internal reflection.

1cm scint. $\sim 10000\text{ph}$

We should expect more light using reflective covers



SiPM-6x6mm²

Surface ratio = $0.36/5 = 0.072$

Total min eff = $0.18 \times 0.072 = 0.013$

130ph \rightarrow 32-50PE

Some ideas on muonID



Two SiPM or find a masking to make more uniform

Some ideas on muonID



Two scintillators with sizes
5x100x1cm & 5x100x1.7cm
was cut and should be tested
in the next year