## Dark Sector of a Higgs Portal with Q4 Symmetric Matter\*

#### Catalina Espinoza

Cátedra Conacyt – IFUNAM

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Outline

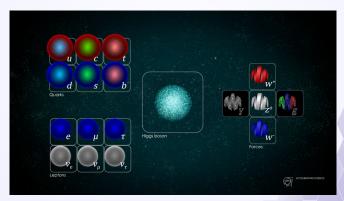
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- Motivation
- 2 Higgs portals
- Q4 Model

# **MOTIVATION**

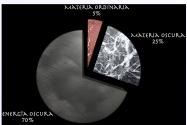
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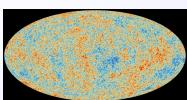
The Standard Model of Particle Physics can be regarded as one of the most important achievements of fundamental science.



Despite the Standard Model's outstanding successes, by now it is clear that a more fundamental theory is required to tackle several shortcomings and unexplained evidences such as (e.g. J. Ellis, hep-th/9812235):

Dark matter and dark energy





Motivation

- Why aren't the particle masses much closer to the Planck mass (10<sup>19</sup> GeV)?
- Why some particle masses are so small  $(m_{\nu})$  and others relatively much larger  $(m_t)$ ?
- Can all the particle interactions be unified in a simple gauge grou
- What is the origin of the 6 flavours each of quarks and lepton
- Baryon asymmetry, mixing patterns, quantization of gravity

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In this talk I will describe one model built with the aim of:

- Provide one (particle) dark matter (DM) candidate.
- Generate the mass hierarchy of quarks and leptons through a discrete symmetry.

In particular, this Beyond the SM (BSM) theory has the following characteristics:

- An extended scalar sector
- DM sector coupled through a Higgs portal

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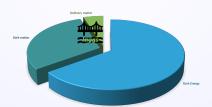
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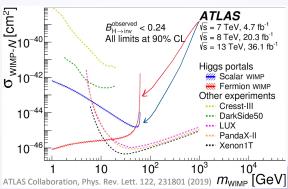
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### **HIGGS PORTALS**

- In a Higgs portal model the DM fields couple only to the scalar fields of the visible sector.
- An immediate consequence of these couplings is that an effective coupling between dark matter particles and quarks is induced, leading to the possibility of dark matter - nucleon scattering.



 LHC constraints from SM invisible Higgs decays nicely complement Direct Detection limits.



- For DM masses above 100 GeV up to several TeV, collider constraints from heavy scalar searches "propagate" to the DM sector.
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 The couplings and mass spectrum of such scalars are constrained thanks to decades of collider searches (Tevatron, LEP and LHC).



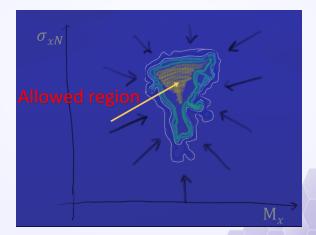
This has immediate repercussions in the DM sector since the scalars
 mediate e.g. the scattering DM - nucleon cross section.

$$L_{\text{eff}} = \sum_{k} \overline{\Psi_{R}^{C}} c_{\Psi}^{k} \Psi_{R} \ h_{k} + \sum_{k,q} \overline{q} c_{q}^{k} q \ h_{k}$$

 The scattering amplitude will depend directly on the scalar masses which are directly constrained from collider searches.

$$\mathcal{M}_k = \frac{4M_{\Psi}m_N}{q^2 + m_{h_k}^2} c_{\Psi}^k c_N^k \ \delta_{ss'} \delta_{rr'}$$

This effectively helps to limit the parameter space of the DM sector.



# Q4 MODEL

- The discrete  $D_N$  symmetry is that of a regular polygon of N sides, and occurs in nature e.g. in poly-atomic molecules.
- The discrete non-abelian group  $Q_4$  (or binary dihedral group with N=4) can be seen as the group cover of  $D_4$ , and has pseudo-real representations which is advantageous for chiral theories.
- In this model the Q4 symmetry is imprinted in the fermionic sector in order to predict the mass and mixing patterns of quarks and neutrinos.
- We propose a scalar sector with 2 Higgs doublets  $H_1$ ,  $H_2$  and 1 real scalar singlet  $\phi$  that mixes with the CP-even scalars.
- The scalar singlet is further coupled to a right handed heavy neutrino DM candidate Ψ

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• The scalar potential is given by the expression  $V = V_1 + V_2$  with:

$$V_{1} = m_{11}^{2} H_{1}^{\dagger} H_{1} + m_{22}^{2} H_{2}^{\dagger} H_{2} - m_{12}^{2} \left( H_{1}^{\dagger} H_{2} + H_{2}^{\dagger} H_{1} \right) + \frac{\lambda_{1}}{2} \left( H_{1}^{\dagger} H_{1} \right)^{2} + \frac{\lambda_{2}}{2} \left( H_{2}^{\dagger} H_{2} \right)^{2} + \lambda_{3} H_{1}^{\dagger} H_{1} H_{2}^{\dagger} H_{2} + \lambda_{4} H_{1}^{\dagger} H_{2} H_{2}^{\dagger} H_{1} + \frac{\lambda_{5}}{2} \left[ \left( H_{1}^{\dagger} H_{2} \right)^{2} + \left( H_{2}^{\dagger} H_{1} \right)^{2} \right],$$

$$(1)$$

$$V_{2} = \mu_{\varphi}^{2} \varphi^{2} + \frac{\lambda_{\varphi}}{2} \varphi^{4} + \lambda_{7} \varphi^{2} H_{1}^{\dagger} H_{1} + \lambda_{9} \varphi^{2} H_{2}^{\dagger} H_{2} + h.c.,$$
 (2)

with all parameters real.

From the scalar potential we obtain the mass matrices for the different scalar particles. The physical particles A and  $H^{\pm}$  have masses given by:

$$M_A^2 = m_{12}^2 \csc \beta \sec \beta - v^2 \lambda_5 \tag{3}$$

$$M_{H^{\pm}}^2 = m_{12}^2 \csc \beta \sec \beta - \frac{1}{2} v^2 (\lambda_4 + \lambda_5)$$
 (4)

where  $\tan \beta = v_{H_2}/v_{H_1}$ . For the CP-even neutral scalars we can write the mass matrix as:

$$M_{\text{scalar}}^2 = \begin{pmatrix} a & d & f \\ d & b & e \\ f & e & c \end{pmatrix}, \tag{5}$$

with:

$$a = m_{12}^2 \tan \beta + \lambda_1 v^2 \cos^2 \beta$$

$$b = m_{12}^2 \cot \beta + \lambda_2 v^2 \sin^2 \beta$$

$$c = \lambda_\phi v_\phi^2$$

$$d = -m_{12}^2 + \lambda_{345} v^2 \cos \beta \sin \beta$$

$$e = \lambda_9 v v_\phi \sin \beta$$

$$f = \lambda_7 v v_\phi \cos \beta$$

where  $\lambda_{345}$  is short for  $(\lambda_3 + \lambda_4 + \lambda_5)$ . The neutral scalar mass matrix is diagonalized by the mixing matrix  $Z^H$  such that

$$Diag(m_h^2, m_H^2, m_{H3}^2) = Z^H M_{scalar}^2 Z^{HT}$$
 (6)

In this model we have 3 physical CP-even scalars one of which corresponds to a SM Higgs-like h, the other two are denoted H<sub>3</sub> and H. We find for the masses:

$$m_h^2 = \frac{1}{3} (a+b+c-2\sqrt{x_1}\cos[\Xi_s/3])$$

$$m_H^2 = \frac{1}{3} (a+b+c+2\sqrt{x_1}\cos[(\Xi_s-\pi)/3])$$

$$m_{H_3}^2 = \frac{1}{3} (a+b+c+2\sqrt{x_1}\cos[(\Xi_s+\pi)/3])$$
(7)

where

$$x_1 = a^2 + b^2 + c^2 - ab - ac - bc + 3(d^2 + f^2 + e^2)$$
 (8)

and:

$$\Xi_{s} = \begin{cases} \arctan\left(\frac{\sqrt{4x_{1}^{3} - x_{2}^{2}}}{x_{2}}\right) &, \quad x_{2} > 0\\ \pi/2 &, \quad x_{2} = 0\\ \arctan\left(\frac{\sqrt{4x_{1}^{3} - x_{2}^{2}}}{x_{2}}\right) + \pi &, \quad x_{2} < 0 \end{cases}$$
(9)

with

$$x_2 = -(2a - b - c)(2b - a - c)(2c - a - b) +9[(2c - a - b)d^2 + (2b - a - c)f^2 + (2a - b - c)e^2] - 54def$$
 (10)

Note that  $\Xi_s \in [-\pi/2, 3\pi/2]$  so  $m_H^2$  is always grater than  $m_h^2$  but  $m_{H3}^2$  can be smaller than  $m_h$ .

### Q4: Likelihood analysis

- We present several numerical results based on a scan of the parameter space of the model where we construct likelihood profiles involving observables of interest by comparing predictions with experimental measurements.
- We limit ourselves to include the information from the measured values of the relic density  $\Omega h_{\mathrm{Planck}}^2$  and Higgs mass  $m_h$  as basic Gaussian likelihoods  $\mathcal{L}_{\Omega}$ ,  $\mathcal{L}_{m_h}$  respectively.

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### Q4: Likelihood analysis

 We also include a likelihood function \( \mathcal{L}\_{DD} \) based on recent results from the XENON1T Direct Detection Experiment, we then maximize over the model's parameter space the composite log-likelihood:

$$\log \mathcal{L} = \log \mathcal{L}_{DD} + \log \mathcal{L}_{\Omega} + \log \mathcal{L}_{m_h}$$
 (11)

Note that in the high statistic limit, twice the negative of the composite log-likelihood approaches a  $\chi$ -square function so this procedure is equivalent to minimizing such function.

#### Numerical analysis

- Using public tools\* we impose hard cuts discarding points not complying with positivity and stability of the scalar potential, and exclusion limits from scalar searches at Tevatron, LEP and LHC. \* EVADE, JHEP 03 (2019), 109, HiggsBounds, Eur. Phys. J. C 80 (2020) no.12, 1211.
- DM-nucleon scattering cross sections. (Micromegas, Comput. Phys. Commun 231 (2018), 173-186)
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#### Experimental information

We compare the predictions of the model with **publicly available data from different experiments** and implemented in diverse public tools:

- XENON 1T for DM direct detection limits (XENON 1T, Phys. Rev. Lett. 121 (2018) no.11, 111302, DDcalc, Eur. Phys. J. C 77 (2017) no.12, 831, and references therein)
- PLANCK for DM relic density ( PLANCK, Astron. Astrophys. 641 (2020), A6)
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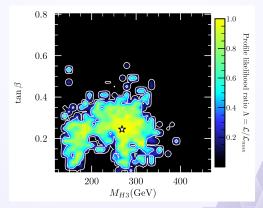
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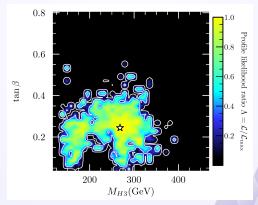
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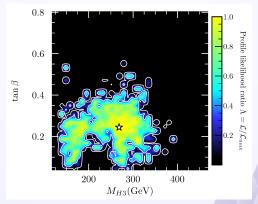
This is the likelihood profile for **the**  $H_3$  **scalar** showing the dependence of the quotient of the Higgs vevs  $(\tan \beta)$  in the mass of  $H_3$ .



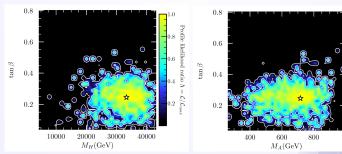
Solid lines are contours of 68% and 95% of C.L. and the star points to the best fit point (BFP) or maximum of the composite likelihood function.



For this model the mass of the  $H_3$  scalar that maximizes the composite likelihood is found to be around **260 GeV** and contained in an interval **in-between 150 - 350 GeV** at 68% of C.L.



• These are the likelihood profiles for the scalar H and the pseudo-scalar A. The mass of A turns out to be around 720 GeV at the BFP while the scalar H is very heavy around 33 TeV at the BFP.

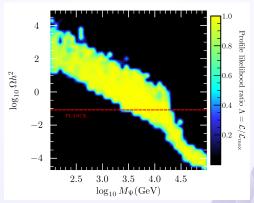


 The dark matter candidate is a spin 1/2 fermion which in general communicates with the visible sector through all the scalar mass eigenstates due to the coupling:

$$L_{
m yuk}\supset y_{\Psi}\overline{\Psi_R^C}\Psi_R\varphi$$

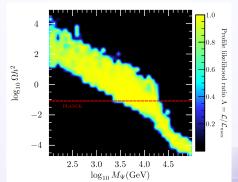
# Q4 dark sector: Relic density

This is the likelihood profile of the **relic density** as a function of the mass of the DM fermion (not including the likelihood function for the relic density).



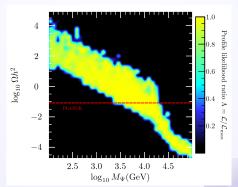
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As this plot evidences, DM fermion masses below around 2.5 TeV would be overproduced during the freeze-out epoch, so they are excluded despite being consistent with e.g. direct detection constraints (as some regions below 2 TeV have large likelihood function value with respect to the direct detection data).



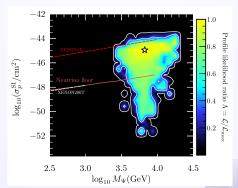
# Q4 dark sector: Relic density

From this analysis we find that this DM candidate can accommodate the observed DM abundance if its mass is in-between 2.5 and 20 TeV, we further find the BFP (not shown) corresponds to a DM mass of 6 TeV.



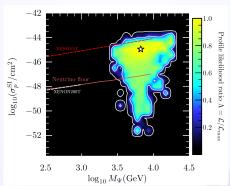
# Q4 dark sector: Elastic scattering cross section

This is the likelihood profile corresponding to the elastic spin-independent DM-nucleon scattering cross section. We compare with exclusion limits of the XENON 1T experiment and its projection to 200 tons.



# Q4 dark sector: Elastic scattering cross section

 The analysis shows that more than half the currently allowed region will be probed with the high sensitivity of the 200 ton future experiment.



# **CONCLUSIONS**

- We have built a predictive and viable extended 2HDM, where the scalar and fermion sectors are enlarged by the inclusion of a gauge singlet scalar and a right handed Majorana neutrino, respectively.
- The model contains a Higgs portal to the DM sector and the constraints from the scalar sector influence also the phenomenology of the DM candidate.

#### Conclusions

- The consistency of our model with the constraints arising from collider searches for heavy scalars, stability of the scalar potentials, the dark matter relic density and current and future direct detection experiments sets the mass of the scalar dark matter candidate to be around 6 TeV.
- The results of these analysis are complementary and enrich the respective matter sector analysis (not discussed here) such as those coming from the generation of the fermion masses and mixing patterns.

#### Acknowledgments

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- I would like to thank the support of Conacyt Cátedra No. 341 and IF-UNAM.

# Thank you for your attention!