SYMMETRIES AND THEIR MASSIVE CONSEQUENCES

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SYMMETRIES

Modern physics is built on the observation that there are symmetries in Nature (exact or broken)

Symmetry is a transformation that leaves the system invariant



Fig: Wikipedia

SYMMETRIES

- ► Quantum field theory combines quantum mechanics and special relativity
- Space-time symmetries: rotations, translations, Lorenz and Poincaré transformations
- ➤ Internal symmetries: transformation of the fields in the theory → gauge symmetries
- ► Global → spacetime momentum, angular momentum, spin
- ➤ Local → gauge symmetries
- ➤ Continuous symmetries→ conserved quantities
 - rotational symmetry angular momentum conservation
 - translational symmetry momentum and energy conservation
- ➤ Discrete → charge and parity conjugation CP
- Label and classify particles
- ► Determine interactions among particles → they must respect the symmetries
- Exact, broken, a little bit broken (softly), hidden

DO YOU NOT UNDERSTAND?

 $\bar{G}^{a}\partial^{2}G^{a} + g_{s}f^{abc}\partial_{\mu}G^{a}G^{b}g^{c}_{\mu} - \partial_{\nu}W^{+}_{\mu}\partial_{\nu}W^{-}_{\mu} - M^{2}W^{+}_{\mu}W^{-}_{\mu} - \frac{1}{2}\partial_{\nu}Z^{0}_{\mu}\partial_{\nu}Z^{0}_{\mu} - \frac{1}{2c^{2}}M^{2}Z^{0}_{\mu}$ $\frac{1}{2}\partial_{\mu}A_{\nu}\partial_{\mu}A_{\nu} - \frac{1}{2}\partial_{\mu}H\partial_{\mu}H - \frac{1}{2}m_{h}^{2}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{-} - M^{2}\phi^{+}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{-} - M^{2}\phi^{+} - \frac{1}{2}\partial_{\mu}\phi^{-} - \frac{1}{2$ $\frac{1}{2c_{w}^{2}}M\phi^{0}\phi^{0}-\beta_{h}\left[\frac{2M^{2}}{g^{2}}+\frac{2M}{g}H+\frac{1}{2}(H^{2}+\phi^{0}\phi^{0}+2\phi^{+}\phi^{-})\right]+\frac{2M}{g^{2}}\alpha_{h}-igc_{w}[\partial_{v}Z_{\mu}^{0}(W_{\mu}^{+}W_{\nu}^{-}-W_{\mu}^{-})]$ $W_{\nu}^{+}W_{\mu}^{-}) - Z_{\nu}^{0}(W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+}) + Z_{\mu}^{0}(W_{\nu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}A_{\mu}(W_{\mu}^{-}W_{\nu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - igs_{\omega} \partial_{\nu}A_{\mu}(W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\nu}A_{\mu}(W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{-})] - igs_{\omega} \partial_{\nu}A_{\mu}(W_{\mu}^{-} - W_{\mu}^{-})] - igs_{\omega} \partial_{\mu}A_{\mu}(W_{\mu}^{-} - W_{\mu}^{-})] - igs_{\omega} \partial_{\mu}A_{\mu}(W_{\mu}^{-}$ $W_{\nu}^{+}W_{\mu}^{-}) - A_{\nu}(W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\mu}^{-}\partial_{\nu}W_{\mu}^{+}) + A_{\mu}(W_{\nu}^{+}\partial_{\nu}W_{\mu}^{-} - W_{\nu}^{-}\partial_{\nu}W_{\mu}^{+})] - \frac{1}{2}g^{2}W_{\mu}^{+}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-} + \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}^{-}W_{\nu}^{+}W_{\nu}^{-} + \frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}^{-}W_{\nu}^{+}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{+}W_{\nu}^{-}W$ $W_{\nu}^{-}+g^{2}c_{\omega}^{2}(Z_{\mu}^{0}W_{\mu}^{+}Z_{\mu}^{0}W_{\nu}^{-}-Z_{\mu}^{0}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-})+g^{2}s_{\omega}^{2}(A_{\mu}W_{\mu}^{+}A_{\nu}W_{\nu}^{-} A_{\mu}A_{\mu}W_{\nu}^{+}W_{\nu}^{-}) + g^{2}s_{\omega}c_{\omega} A_{\mu}Z_{\nu}^{0}(W_{\mu}^{+}W_{\nu}^{-} - W_{\nu}^{-}W_{\mu}^{-}) - 2A_{\mu}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}) + g^{2}s_{\omega}c_{\omega}^{-}A_{\mu}Z_{\nu}^{0}(W_{\mu}^{+}W_{\nu}^{-} - W_{\nu}^{-}W_{\mu}^{-}) - 2A_{\mu}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}W_{\nu}^{-}] - g\alpha[H^{3} + M_{\nu}^{-}W_{$ $H\phi^{0}\phi^{0} + 2H\phi^{+}\phi^{-}] - \frac{1}{2}g^{2}\alpha_{h} H^{4} + (\phi^{0})^{4} + 4(\phi^{+}\phi^{-})^{2} + 4(\phi^{0})^{2}\phi^{+}\phi^{-} + 4H^{2}\phi^{+}\phi^{-}$ $2(\phi^{0})^{2}H^{2}] - 9MW_{\mu}^{+}W_{\mu}^{-}H - \frac{1}{2}g\frac{4}{2}Z_{\mu}^{0}Z_{\mu}^{0}H - \frac{1}{2}ig[W_{\mu}^{+}(\phi^{0}\partial_{\mu}\phi^{-} - \phi^{-}\partial_{\mu}\phi^{0}) - W_{\mu}^{-}(\phi^{0}\partial_{\mu}\phi^{-}) + \frac{1}{2}ig[W_{\mu}^{+}(\phi^{0}\partial_{\mu}\phi^{-}) + \frac{1}{2}ig$ $\phi^{+}\partial_{\mu}\phi^{0})]_{2} + \frac{1}{2}g[W_{\mu}^{+}(H\partial_{\mu}\phi^{-} - \phi^{-}\partial_{\mu}H) - W_{\mu}^{-}(H\partial_{\mu}\phi^{+} - \phi^{+}\partial_{\mu}H)] + \frac{1}{2}g\frac{1}{c}(Z_{\mu}^{0}(H\partial_{\mu}\phi^{0} - \phi^{-}\partial_{\mu}H) - W_{\mu}^{-}(H\partial_{\mu}\phi^{0} - \phi^{-}\partial_{\mu}H)] + \frac{1}{2}g\frac{1}{c}(Z_{\mu}^{0}(H\partial_{\mu}\phi^{0} - \phi^{-}\partial_{\mu}H)) + \frac{1}{c}(U_{\mu}^{0}(H\partial_{\mu}\phi^{0} - \phi^{-}\partial_{\mu}H)) + \frac{1}{c}(U_$ $\phi^{0}\partial_{\mu}H) \quad ig \frac{s_{\omega}}{s_{\omega}}MZ^{0}_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) + igs_{\omega}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) + igs_{\omega}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) + igs_{\omega}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) + igs_{\omega}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) + igs_{\omega}MA_{\mu}(W^{+}_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-W^{-}_{\mu}\phi^{+}) - ig\frac{1-2c_{\omega}}{2c_{\omega}}Z^{0}_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}) - ig\frac{1-2c_{\omega}$ $(+igs_{W}A_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-\phi^{-}\partial_{\mu}\phi^{+})-\frac{1}{2}g^{2}W_{\mu}^{+}W_{\mu}H^{2}+(\phi^{0})^{2}+2\phi^{+}\phi^{-})$ $[+2(2s_{\omega}^{2}-1)^{2}\phi^{+}\phi^{-}] - \frac{1}{2}g^{2}\frac{s_{\omega}}{s_{\omega}}Z_{\mu}^{0}\phi^{0}(W_{\mu}^{+}\phi^{-})$ $_{\mu}\phi^{0}(W_{\mu}^{+}\phi^{-}+W_{\mu}^{-}\phi^{+})+\frac{1}{2}ig^{2}s_{w}A_{\mu}H(W_{\mu}^{+}\phi^{-})$ $-\bar{u}_{j}^{\lambda}(\gamma\partial + m_{\mu}^{\lambda})u_{j}^{\lambda} - d_{j}^{\lambda}(\gamma\partial + m_{d}^{\lambda}d_{j}^{\lambda} + igs_{\varpi}A_{\mu}[-(\bar{e}^{\lambda}\gamma^{\mu}e^{\lambda}) + \frac{2}{3}(\bar{u}_{j}^{\lambda}\gamma^{\mu}e^{\lambda})]$ $-Z^0_{\mu}[(\bar{\nu}^{\lambda}\gamma^{\mu}(1+\gamma^5)\nu^{\lambda}) + (\bar{e}^{\lambda}\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) - (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) - (\bar{u}^{\lambda}_i\gamma^{\mu}(1+\gamma^5)\nu^{\lambda}) + (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) - (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) + (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) - (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) + (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) - (\bar{u}^{\lambda}_i\gamma^{\mu}(4s^2_{\omega}-1-\gamma^5)e^{\lambda}) + (\bar{u}^{$ $(1 - \frac{8}{3}s_{\omega}^2 - \gamma^5)d_j^{\lambda}) = \pm \frac{19}{2\sqrt{2}}W_{\mu}^+(\nu^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda}) - (u_j^{\lambda}\gamma^{\mu}(1 + \gamma^5)e^{\lambda})$ $\gamma^{5})C_{\lambda\kappa}d_{j}^{s})] + \frac{ig}{2\sqrt{2}}W_{\mu}^{-}[(\bar{e}^{\lambda}\gamma^{\mu}(1+\gamma^{5})\nu^{\lambda}) + (d_{j}^{s}C_{\lambda\kappa}\gamma^{\mu}(1+\gamma^{5})u_{j}^{\lambda})] + \frac{ig}{2\sqrt{2}}\frac{m}{M}[-\phi^{+}$ $\gamma^{5}(e^{\lambda}) + \phi^{-}(\bar{e}^{\lambda}(1+\gamma^{5})\nu^{\lambda})] - \frac{g}{2} \frac{m_{e}^{2}}{M} \left[H(\bar{e}^{\lambda}e^{\lambda}) + i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})) + i\phi^{0}(\bar{e}^{\lambda}\gamma^{5}e^{\lambda})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1-\bar{u}_{j}^{\kappa}) + i\phi^{0}(\bar{u}_{j}^{\kappa})\right] + \frac{ig}{2M\sqrt{2}}\phi^{+}\left[-m_{d}^{\kappa}(\bar{u}_{j}^{$ $\gamma^{5})d_{j}^{\kappa}) + m_{u}^{\lambda}(\bar{u}_{j}^{\lambda}C_{\lambda\kappa}(1+\gamma^{5})d_{j}^{\kappa}] + \frac{iq}{2M\sqrt{2}}\phi^{-}[m_{d}^{\lambda}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1+\gamma^{5})u_{j}^{\kappa}) \quad m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\dagger}(1-\gamma^{5})u_{j}^{\kappa}) + m_{u}^{\kappa}(\bar{d}_{j}^{\lambda}C_{\lambda\kappa}^{\prime}(1-\gamma^{5})u_{j}^{\kappa}) + m_{u}^{\kappa}(\bar{$ $\gamma^5)u_j^{\kappa}] - \frac{2}{2} \frac{m_s^{\kappa}}{M} H(u_j^{\lambda} u_j^{\lambda}) - \frac{2}{2} \frac{m_s^{\lambda}}{M} H(d_j^{\lambda} d_j^{\lambda}) + \frac{42}{2} \frac{m_b^{\kappa}}{M} \phi^0(u_j^{\lambda} \gamma^5 u_j^{\lambda})$ $\frac{10}{2} \frac{m_{10}}{M} \phi^0(a_i^{\lambda} \gamma^5 a_i^{\lambda}) +$ $X^{+}(\partial^{2}-M^{2})X^{+}+X^{-}(\partial^{2}-M^{2})X^{-}+X^{0}(\partial^{2}-\frac{M^{2}}{c^{2}})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}-M^{2})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}Y + igc_{w}W^{+}(\partial_{\mu}X^{0}X^{-}+W^{-})X^{0}+Y\partial^{2}$ $\partial_{\mu}X^{+}X^{0}) + igs_{\mu}W^{+}_{\mu}(\partial_{\mu}\bar{Y}X^{-} - \partial_{\mu}X^{+}Y) + ig_{e_{\mu}}W^{-}_{\mu}(\partial_{\mu}X X^{0} - \partial_{\mu}\bar{X}^{0}X^{+}) +$ $igs_{\omega}W_{\mu}(\partial_{\mu}X^{-}Y - \partial_{\mu}YX^{+}) + igc_{\omega}Z_{\mu}^{0}(\partial_{\mu}X^{+}X^{+} - \partial_{\mu}X^{-}X^{-}) + igs_{\omega}A_{\mu}(\partial_{\mu}X^{+}X^{+} - \partial_{\mu}X^{-}X^{-}) + igs_{\omega}A_{\mu}(\partial_{\mu}X^{+}X^{+}) + igc_{\omega}Z_{\mu}^{0}(\partial_{\mu}X^{+}X^{+} - \partial_{\mu}X^{-}X^{-}) + igs_{\omega}A_{\mu}(\partial_{\mu}X^{+}X^{+}) + igc_{\omega}Z_{\mu}^{0}(\partial_{\mu}X^{+}X^{+}) + igc_{\omega}Z_{\mu}^{0}$ $\partial_{u}\bar{X}^{-}X^{-}) - \frac{1}{2}gM[\bar{X}^{+}X^{+}H + \bar{X}^{-}X^{-}H + \frac{1}{c_{u}^{2}}\bar{X}^{0}X^{0}H] + \frac{1-2c_{u}^{2}}{2c_{u}}igM[\bar{X}^{+}X^{0}\phi^{+} - \frac{1}{2c_{u}}\bar{X}^{0}A^{0}H] + \frac{1-2c_{u}^{2}}{2c_{u}}igM[\bar{X}^{+}X^{0}\phi^{+} - \frac{1}{2c_{u}}igM[\bar{X}^{+}X^{0}\phi^{+} - \frac{1}{2c_{u}}$ $\begin{array}{l} X^{-}X^{0}\phi^{-}] + \frac{1}{2cw} igM[X^{0}X^{-}\phi^{+} - X^{0}X^{+}\phi^{-}] + igMs_{w}[X^{0}X^{-}\phi^{+} - X^{0}X^{+}\phi^{-}] + \\ \frac{1}{2}igM\bar{X}^{+}X^{+}\phi^{0} - X^{-}\bar{X}^{-}\phi^{0}] \end{array}$

WHAT PART OF

 $-\tfrac{1}{2}\partial_{\nu}g^a_{\mu}\partial_{\nu}g^a_{\mu} - g_s f^{abc}\partial_{\mu}g^a_{\nu}g^b_{\nu}g^c_{\nu} - \tfrac{1}{4}g^2_s f^{abc}f^{ade}g^b_{\mu}g^c_{\nu}g^d_{\mu}g^e_{\nu} + \tfrac{1}{2}ig^2_s(\bar{q}^{\sigma}_i\gamma^{\mu}q^{\sigma})g_{\mu}$

STANDARD MODE

+ Fi Jij Fog +h.c.

 $+\left|\mathcal{D}_{\mathcal{A}}\varphi\right|^{2}-\bigvee(\varphi)$

LAGRANGIAN



WHY GO BEYOND?

- ► The hierarchy problem
- Neutrino masses
- ► All masses
- Origin of gauge interactions
- ► Dark matter
- Matter over anti-matter abundance
- Cosmological constant
- ► Inflation

Higgs sector not natural Fermion masses vastly different Origin of electroweak symmetry breaking unknown Dirac or Majorana neutrinos Strong CP problem

Not enough CP in SM for Baryogengesis Value of cosmological constant Inflation inconsistent with non-zero baryon number Is DM a particle, then which, is it only one

HOW TO GO BEYOND THE STANDARD MODEL (BSM)?

- Addition of symmetries
- Addition of particles
- Addition of interactions
- Addition of space-time dimensions
- ► All of the above...

I will concentrate on masses and mixings.

And perhaps dark matter and leptogenesis, and...







Picture from: Neutrino Oscillations in the Atmospheric Parameter Region: From the Early Experiments to the Present BGiacomelli, M. Giorgini, L. Patrizii, M. Sioli. Adv.High Energy Phys. 2013 (2013) 464926.

MASSES ARE VERY DIFFERENT...

- Quark and lepton masses hierarchical and very different
- Neutrino masses enhance the problem, tiny!
- Higgs boson crucial for mass generation





HOW DO WE MOVE UP (OR DOWN) IN ENERGY?

- We know how a QFT behaves at different scales through the renormalization group RG
- The theory has the same structure at different energy scales, but the parameters — couplings and masses — change with energy
- Related to scale invariance and conformal invariance

$$\beta(g) = \mu \frac{\partial g}{\partial \mu}$$

 $\gamma(\phi) = \mu \frac{\partial \ln Z}{\partial \mu}$

Set of differential equations that describe the bahaviour of the theory at different scales — Renormalization Group Equations RGE

SCALE INVARIANCE AND CONFORMAL INVARIANCE



Conformally invariant



Scale invariant

Scale invariance appears in many physical systems

SUSY GUTS: MORE SYMMETRY

- GUTs: at high energies only one fundamental interaction (excluding gravity)
- SUSY: relates bosons (interactions) and fermions(matter)
- At low energies, most studied: MSSM Extended Higgs sector:
 → Hu & Hd, 5 physical Higgs bosons, tanβ = vu/vd
- ► Broken SUSY...

MORE SYMMETRY NOT ALWAYS HELPFUL...

- ➤ Larger symmetry groups → more (unobserved) particles
- ➤ Different ways to break the symmetry → more parameters
- ► Examples:
 - GUTs leptoquarks, heavy bosons
 - SUSY spartners
 - Superstrings all of the above plus extra dimensions that have to be compactified
- Super interesting, but need extra guidance

RENORMALIZATION GROUP INVARIANTS RGI

➤ Search for more fundamental theory → less parameters Renormalization Group Invariants (RGI)

$$\Phi(g_1, \dots, g_N) = 0$$

$$\mu d\Phi/d\mu = \sum_{i=1}^N \beta_i \partial \Phi/\partial g_i = 0$$

Equivalent to solve reduction equations

$$\beta_g \left(dg_i / dg \right) = \beta_i$$

$$i = 1, \ldots, N$$

- Reduced theory has only one coupling and its beta function
- Reduction → power series solution
- Uniqueness of solution can be studied at one-loop Zimmermann (1985); Zimmermann, Oehme, Sibold (1984–1985)

REDUCTION OF COUPLINGS

- May indicate an underlying symmetry
- ➤ There are solutions to the RE without apparent symmetry
- ► SUSY solutions appear often as solutions to the RE
- In SM, reducing the Higgs and top couplings in terms of the strong coupling constant leads to

$$\alpha_t / \alpha_s = \frac{2}{9}$$
; $\alpha_\lambda / \alpha_s = \frac{\sqrt{689} - 25}{18} \simeq 0.0694$
 $M_t = 98.6 \pm 9.2 \ GeV$, $M_h = 64.5 \pm 1.5 \ GeV$

Outside experimental range, but still remarkable... Kubo, Sibold, Zimmermann (1984–1987)

REDUCTION OF COUPLINGS WITHOUT SUSY

- It's also possible to do reduction of couplings without SUSY
- ► Analysis done in the past for the 2HDM Denner
- New systematic analysis for 2HDM done now, facilitates extension to different models M.A. May Pech M.Sc. Thesis (2022)
- ► Miguel Angel May talk on Friday for details...

FINITENESS = SCALE INVARIANCE

- Finiteness ⇒ β = 0
 to all orders in perturbation
 theory
- Scale or conformal invariance
- Couplings do not depend on energy scale
- Based on RGI and reduction of couplings
- Reduces greatly the number of free parameters
 - → new symmetries

$\textbf{FINITESS} \Longrightarrow \textbf{ GAUGE YUKAWA UNIFICATION}$

Grand Unified SUSY N=1, no gauge anomalies:

$$W = \frac{1}{2} m^{ij} \Phi_i \Phi_j + \frac{1}{6} C^{ijk} \Phi_i \Phi_j \Phi_k$$
$$\beta_g^{(1)} = 0 = \gamma_i^{j(1)}$$
$$\sum_i T(R_i) = 3C_2(G), \qquad \frac{1}{2} C_{ipq} C^{jpq} = 2\delta_i^j g^2 C_2(R_i)$$

T Dynkin index of irrep, C₂ Casimir invariant of group

C_{ijk} Yukawa couplings, g gauge coupling

- Restricts the gauge group
- Relates gauge and Yukawa couplings
- Can be made finite to all orders
 - Conformal invariance

- Just analyze one-loop solution
- Isolated and non-degenerate solution Lucchesi, Piguet, Sibold
- Implies extra symmetries, in this case discrete

 $\beta = 0$ non-renormalization of coupling constants, not complete UV finiteness where field renormalization is absent

MANY ASPECTS OF FINITENESS STUDIED

- ► SU(5) models extensively studied
- One coincides with a non-standard Calabi-Yau
- Finite string theories and criteria for branes
- Models with three generations
- > SU(N)^k models finite \Leftrightarrow 3 generations $SU(3)^3$ finite
- Relations non-commutative theories and finiteness
- Proof of conformal invariance (dimensionless part) Kazakov, Bork; MM & Reyes
- Relation between finiteness and QFT in curved space-time & inflation Elizalde et al
- **Recent reviews**

Heinemeyer, M.M., Tracas, Zoupanos, Phys.Rept. 814 (2019); Fortsch.Phys. 68 (2020)

Rabi et al: Kazakov et al: Quirós et al: MM, Zoupanos et al

Babu, Enkhbat, Gogoladze; MM & Jiménez

MM, Zoupanos

lbáñez

MM, Ma, Zoupanos

Jack. Jones

BOUNDARY CONDITIONS AT GUT SCALE

Model A

$$g_t^2 = \frac{8}{5} g^2$$

$$g_{b,\tau}^2 = \frac{6}{5} g^2$$

$$m_{H_u}^2 + 2m_{10}^2 = M^2$$

$$m_{H_d}^2 + m_{\overline{5}}^2 + m_{10}^2 = M^2$$

Model B

$$\blacktriangleright g_t^2 = \frac{4}{5} g^2$$

•
$$g^2_{b, au} = rac{3}{5} g^2$$

•
$$m_{H_u}^2 + 2m_{10}^2 = M^2$$

$$m_{H_d}^2 - 2m_{10}^2 = -\frac{M^2}{3}$$

$$m_{\overline{5}}^2 + 3m_{10}^2 = \frac{4M^2}{3}$$

2 free parameters:
M, m²/₅

Each model implies different discrete symmetries → isolated and non-degenerate solution of the RE

STANDARD MODEL Yukawa couplings $\begin{aligned} \mathcal{J} &= -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\ &+ i F \mathcal{B} \gamma + h.c. \end{aligned}$ free parameters + $\chi_{ij} \chi_{j} \neq h.c.$ + $|D_{\mu} \varphi|^2 - V(\varphi)$ $M_H^2 = \lambda v^2$ $m_f = g_f v / \sqrt{2}$ LAGRANGIAN

PREDICTIONS - NEW 2018 AND 2022 ANALYSES

- Results consistent with B physics constraints (not trivial)
- Predictions for top and bottom quark masses in experimental range
 M_{top} 171-174 GeV exp 173.3 ± 0.76 GeV
 M_{bot} 2.6-2.9 (M_Z) GeV exp 2.8 ± 0.1 GeV
- > Large tan β 48-52
- ► Higgs mass 122-129 GeV exp 125.1 ± 0.3
- ► Heavy SUSY spectrum in TeV region Collider phenomenology, challenging even for FCC
- ► Not all models survive, non minimal FUT SU(5) and SU(3) ^ 3 FUT do

RECENT DEVELOPMENTS

- Updated phenomenological analysis still consistent with 3rd generation masses, large tanβ and very heavy SUSY spectrum S. Heinemeyer, J. Malinowski, W. Kotlarski, M. Mondragón, N. Tracas, G. Zoupanos 2018 and 2022
- Finiteness implies conformal invariance and phase transition
 L.E. Reyes Rodríguez, Lic. Thesis (2018)
- Three generation analysis SU(5): Diagonal quark mass matrix compatible with data and proton decay Luis Odín Estrada, M.Sc. Thesis (2018)
- Three generation solution for SU(5) with Z symmetries compatible with good textures at high energies Luis Odín Estrada,Ph.D. Thesis
- Finiteness in Soft breaking terms lead to anomaly mediated type breaking L.E. Reyes Rodríguez, M.Sc.Thesis (2021)
- ► SU(3) ^ 3 finite split susy model in progress L.E. Reyes Rodríguez, Ph. D. Thesis

OUTLOOK AND TO PACK...

- ► Inclusion of neutrino masses through *K*
- Will not change much the collider phenomenology presented
- > Will impact the DM candidate \Rightarrow gravitino + ?
- Detailed analysis of three gen solutions with discrete symmetries
- Detailed analysis of phase transition and SUSY breaking
- ► Lately: FUTs in curved space time ⇒ successful inflation E. Elizalde, S. Odintsov, E. Pozdeeva, S. Vernov (2015)
- New type of finite theories imply duality between UV and IR fixed points. Connection to FUTs?
 Y. Kawamura (2015)

AND AT LOWER ENERGIES ...?

HOW DO WE CHOOSE A FLAVOUR SYMMETRY?

- ► Several ways:
- Look for inspiration in a high energy extension of SM, i.e. strings or GUTs
- Look at low energy phenomenology
- At some point they should intersect...
- In here:
 - Find the smallest flavour symmetry suggested by data
 - Explore how generally it can be applied (universally)
 - ► Follow it to the end
 - Compare it with the data

SOME ASPECTS OF THE FLAVOUR PROBLEM

 Quark and charged lepton masses very different, very hierarchical

 $m_u: m_c: m_t \sim 10^{-6}: 10^{-3}: 1$

```
m_d: m_s: m_b \sim 10^{-4}: 10^{-2}: 1
```

 $m_e: m_\mu: m_\tau \sim 10^{-5}: 10^{-2}: 1$

- Neutrino masses unknown, only difference of squared masses.
- Type of hierarchy (normal or inverted) also unknown
- Higgs sector under study

Quark mixing angles

 $\theta_{12} \approx 13.0^{o}$ $\theta_{23} \approx 2.4^{o}$ $\theta_{13} \approx 0.2^{o}$

Neutrino mixing angles

 $\Theta_{12} \approx 33.8^{\circ}$ $\Theta_{23} \approx 48.6^{\circ}$ $\Theta_{13} \approx 8.6^{\circ}$

- Small mixing in quarks, large mixing in neutrinos.
 Very different
- Is there an underlying symmetry?

Plot of mass ratios

Logarithmic plot of quark masses

$$\begin{bmatrix} |V_{\rm ud}| & |V_{\rm us}| & |V_{\rm ub}| \\ |V_{\rm cd}| & |V_{\rm cs}| & |V_{\rm cb}| \\ |V_{\rm td}| & |V_{\rm ts}| & |V_{\rm tb}| \end{bmatrix} \approx \begin{bmatrix} 0.974 & 0.225 & 0.003 \\ 0.225 & 0.973 & 0.041 \\ 0.009 & 0.040 & 0.999 \end{bmatrix},$$

Suggests a 2⊕1 structure

S3-3HDM

- Smallest non-Abelian discrete group
- > Has irreducible representations, 2, 1_s and 1_A
- We add three right-handed neutrinos to implement the seesaw mechanism
- We apply the symmetry "universally" to quarks, leptons and Higgs-es
 - First two families in the doublet
 - Third family in symmetric singlet
- ► Three sectors related, we treat them simultaneously

PREDICTIONS, ADVANTAGES?

- Possible to reparametrize mixing matrices in terms of mass ratios, successfully
- CKM has NNI and Fritzsch textures
- ➤ PMNS → fix one mixing angle, predictions for the other two within experimental range
- ▶ Reactor mixing angle
 → θ₁₃ ≠ 0
- Some FCNCs suppressed by symmetry

- Higgs potential has 8 couplings
- Underlying symmetry in quark, leptons and Higgs
 residual symmetry of a more fundamental one?
- Lots of Higgses:
 3 neutral, 4 charged,
 2 pseudoscalars
- Further predictions will come from Higgs sector: decays, branching ratios

A. Mondragón, M. M., F. González, E. Peinado, U. Saldaña, O. Félix, E. Rodríguez, A. Pérez, H. Reyes, C. Espinoza, E. Garcés,...; Das, Dey et al; Teshima et al; E. Barradas, O. Félix, E. Rodríguez; M. Rebelo, P. Osland et al; many many more

Quarks

	3HDM: $G_{SM} \otimes S_3$							
	Possible mass textures		Mass matrix		ψ^f_R	ψ^f_L		
	$ \begin{pmatrix} & & 0 \\ t^2 \end{pmatrix} & \mu_7^f/c \\ & \mu_3^f - \mu_1^f - \mu_2^f c^2 (1 - 3t^2) \end{pmatrix} $	$ \begin{array}{ccc} 0 & \mu_{2}^{f}sc\left(3-t^{2}\right) \\ sc\left(3-t^{2}\right) & -2\mu_{2}^{f}c^{2}\left(1-3t^{2}\right) \\ 0 & \mu_{7}^{f*}/c \end{array} $	$ \begin{array}{c} \mu_6^f \\ \mu_7^f \\ \mu_3^f \end{array} \right) \\$	$\mu_1^f \overset{\mu_4^f}{\underset{\mu_9^f}{\overset{-}}} \mu_2^f$	$\begin{pmatrix} \mu_1^f + \mu_2^f \\ \mu_4^f \\ \mu_8^f \end{pmatrix}$	$2, \mathbf{1_S}$	2 , 1 _{S}	A
NN	$ \begin{array}{ccc} \frac{1}{3}\mu_{2}^{f} & 0\\ 0 & \frac{2}{\sqrt{3}}\mu_{7}^{f}\\ \frac{1}{3}\mu_{9}^{f} & \mu_{3}^{f} - \mu_{1}^{f} \end{array} \right) $	$\begin{pmatrix} 0 & \frac{2}{\sqrt{3}}\mu_2^f \\ \frac{2}{\sqrt{3}}\mu_2^f & 0 \\ 0 & \frac{2}{\sqrt{3}}\mu_9^f \end{pmatrix}$						A^{\prime}
	$ \begin{array}{ccc} 3t^2) & 0 \\ t^2) & -\mu_6^f/c \\ \mu_3^f - \mu_1^f + \mu_4^f sc(3-t^2) \end{array} \right) $	$\begin{array}{ccc} 0 & -\mu_4^f c^2 \left(1 - 3t^2\right) \\ {}^f_4 c^2 \left(1 - 3t^2\right) & 2\mu_4^f sc \left(3 - t^2\right) \\ 0 & -\mu_6^{f*}/c \end{array}$	$ \begin{pmatrix} \mu_7^f \\ -\mu_6^f \\ \mu_3^f \end{pmatrix} $	$\substack{\mu_4^f \\ \mu_1^f - \mu_2^f \\ \mu_8^f}$	$\begin{pmatrix} \mu_1^f + \mu_2^f \\ \mu_4^f \\ - \mu_9^f \end{pmatrix}$	2 , 1 _{A}	2 , 1 _{A}	В
NN	$2\mu_4^f = 0 \ 0 = -2\mu_6^f \ \mu_8^f = \mu_3^f - \mu_1^f \end{pmatrix}$	$egin{pmatrix} 0 & -2\mu_4^f \ -2\mu_4^f & 0 \ 0 & 2\mu_8^f \end{pmatrix}$						$B^{'}$

Table 2: Mass matrices in S_3 family models with three Higgs $SU(2)_L$ doublets: H_1 and H_2 , which occupy the S_3 irreducible representation **2**, and H_S , which transforms as $1_{\mathbf{S}}$ for the cases when both the left- and right-handed fermion fields are in the same assignment. The mass matrices shown here follow a normal ordering of their mass eigenvalues (m_1^f, m_2^f, m_3^f) . We have denoted $s = \sin \theta$, $c = \cos \theta$ and $t = \tan \theta$. The third column of this table corresponds to the general case, while the fourth column to a case where we have rotated the matrix to a basis where the elements (1, 1), (1, 3) and (3, 1) vanish. The primed cases, A' or B', are particular cases of the unprimed ones, A or B, with $\theta = \pi/6$ or $\theta = \pi/3$, respectively.

Mass matrices reproduce the NNI or the Fritzsch forms

F. González et al, Phys.Rev. D88 (2013) 096004

NEW RESULTS S3-3H

- Adriana's talk: full scalar potential analysis
- 2 scenarios:
 - ► A: SM Higgs lightest one
 - B: Neutral scalar lighter than SM~100 GeV possible neutral scalar signal?
- Both compatible with SM limit for trilinear and quartic couplings
- Small deviations from SM in trilinear and quartic couplings compatible with recent phenomenological analyses in the modifier or κ framework

$$g_{H_2H_2H_2} \equiv \lambda_{SM}\kappa_{\lambda} = \frac{m_{H_2}^2}{2v} \left[(1+2\delta^2)\sqrt{1-\delta^2} + \delta^3(\tan\theta - \cot\theta) - \frac{m_{h_0}^2}{m_{H_2}^2} \frac{\delta^3}{9s_\theta c_\theta^3} \right]$$

S3-3H FUTURE ANALYSIS

- Mass spectrum in reach of future runs of the LHC
 - Inclusion of one-loop corrections necessary
 - Calculation of decays and branching rations of scalars
- New neutral scalar not coupled to gauge bosons, DM?
- ➤ Residual symmetry ⇒ problematic in fermionic sector Solutions:
 - ► break S3
 - ► modular S3
 - high energy sector terms
 - ► 4 Higgs doublets
 - Higgs singlets

Possible to return to good previous results for mixing matrices

S3-4H

- > "Saturate" the irreps: add an extra inert Higgs doublet in the 1_A
- Natural DM candidates from inert part the non-inert part same as S3-3H in SM alignment limit
- Full analysis for DM, with Higgs bounds, relic density, indirect detection done
 E. Garcés, C. Espinoza, M.M, H. Reyes, PLB (2018), M.Sc. Thesis H. Reyes
- Multi-component DM in progress With scalar and neutrino M. Valenzuela, Ph.D. Thesis in progress, with E. Espinoza, A. Ramírez With 2 scalars J. Pacheco, Ph.D. Thesis in progress, with E. Espinoza, E. Barradas, T. Valencia
- ► Neutrino portal in progress with E. Espinoza, E. Barradas, T. Valencia
- ► LFV analysis in progress J. Pacheco, Ph.D. Thesis in progress
- S3 with 3H and U(1)_{B-L} multi component DM in progress L. E. Gutiérrez-Luna, Ph.D. Thesis in progress with J.C. Gómez-Izquierdo, C. Espinoza

Figure 1: Mass of the DM candidate as a function of $\tan \theta$ (left panel), and value of the DM relic density as a function of the DM mass. The dark blue points (set A) are the ones that comply with stability and unitarity constraints, the light blue points (set B) are also compatible with the experimental bounds for extra scalar searches (see text), the red points also satisfy the decoupling limit and the green points in the right panel lie within the experimental Planck bound.

C. Espinoza, E. Garcés, M.M., H Reyes-González *Phys.Lett.B* 788 (2019) 185-19 1

Figure 3: Annihilation cross section as a function of the DM mass for small DM masses, the points are colored according to their (normalized) likelihood (with respect to the relic density) value. Also shown is the FermiLAT dwarf spheroidal combined DM exclusion curve.

WHAT CAN GUIDE US IN BETWEEN?

- ➤ SU(5) x Q6 (non-FUT) ⇒ good CKM, similar predictions to S3-3H for neutrino sector J.C. Gómez-Izquierdo, F. González-Canales, MM (2015)
- Breaking of mu-tau symmetry through Q6 SUSY and S3 non-SUSY

J.C. Gómez-Izquierdo, F. González-Canales, M.M. (2017);(2018)

- g-2 solution through LFV in extended MSSM, with discrete symmetry inspired terms M. Gómez-Bock, F. Flores-Báez, MM (2016)
- ► And of course experimental data and observations...

MORE MODELS?

- ► 2+1 successful in quark sector
- Neutrino sector also, but more flexibility
- Q4 2HDM and singlets see Catalina Espinoza's talk on Thursday: quarks, leptons, DM, leptogenesis, g-2
 Gatto-Sartori-Tonin relation between quark masses and mixing angles A. Cárcamo, E. Espinoza, J.C. Gómez-Izquierdo, M.M. Eur.Phys.J.Plus 137 (2022) 11, 1224
- S3-3H with S3 as modular symmetry, nice results without residual Z2 in quark sector M.C. Cerón, M.Sc. Thesis (2021)
- ► SUSY SU(5) non-minimal with modular S3 A.C. Samaniego, M.Sc. Thesis
- S4 3HDM and 4HDM, plus singlets, with very predictive neutrino sector A. Cárcamo, C. Espinoza, J.C. Gomez, J. Marchant, M.M.

Explore other models

OUTLOOK AND CONCLUSIONS

Among the different ways to go BSM finiteness proves to be a good guiding principle.
 Reduces greatly the number of free parameters, RG flow of the third family in the right direction

- ► Needs extended Higgs sector and discrete flavour symmetries
- At low energies S3, S4, Q4, Q6 theories with extended Higgs sector explain well CKM and have predictions for neutrino sector.

Provide baryogenesis through leptogenesis and good DM candidates

Maybe is possible to connect both approaches

RGIs: Reduction of couplings, finiteness Symmetries: SUSY, GUTs, discrete symmetries Allow us to put some order... They all require extended Higgs sectors for success They all have consequences for the particle masses For our journeys Beyond the Standard Model there is a lot to pack...

GRACIAS!