

# Mass dependence of the effective action in gauge-theories

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# Outline

General aspects, computation

The method

New results

# The effective action in QED

For a classical  $A_\mu(x)$  field, the effective action is given by :

- ▶ Spinor :

$$\Gamma = -i \ln \det(i\not{D} - m),$$

- ▶ Scalar :

$$\Gamma = \frac{i}{2} \ln \det(-D_\mu^2 - m^2),$$

where  $\not{D} \equiv \gamma^\mu D_\mu = \gamma^\mu(\partial_\mu - ieA_\mu(x))$ .

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These expressions correspond to a perturbative series :



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$$\ln \det(\not{D}^2 - m^2) = \text{Tr} \ln(\not{D}^2 - m^2) \sim - \int_0^\infty \frac{ds}{s} e^{-m^2 s} \text{Tr} e^{-s(-\not{D}^2)}$$

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However, for  $m \rightarrow 0$ , there is no general approach.

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t'Hooft (1976)

$$m = 0 \quad : \quad \Gamma_{\text{spinor}}(0) = \alpha(1/2) \approx 0.145873 + \mathcal{O}(m^2)$$

Kwon, Lee and Min (2000)

$$m \rightarrow 0 \quad : \quad \Gamma_{\text{spinor}}(m) = \alpha(1/2) + \frac{m^2}{2} [\ln m + \gamma - \ln 2] + \mathcal{O}(m^4)$$

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# The Gel'Fand-Yaglom theorem

Lets calculate the determinant

$$\left[ -\frac{d^2}{dx^2} + V(x) \right] \psi(x) = \lambda \psi(x) \quad ; \quad \psi(0) = \psi(L) = 0$$

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$$\Rightarrow \det \left[ -\frac{d^2}{dx^2} + V(x) \right] = \psi(L).$$

## Example: The Helmholtz operator

$$\hat{H} = \left[ -\frac{d^2}{dx^2} + m^2 \right] ; \quad \lambda_n = m^2 + \left( \frac{n\pi}{L} \right)^2 ,$$

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$$\Rightarrow \frac{\det \left[ -\frac{d^2}{dx^2} + m^2 \right]}{\det \left[ -\frac{d^2}{dx^2} \right]} = \frac{\phi(L)}{\phi_0(L)} = \frac{\sinh(mL)}{mL}$$

# Rially Symmetric Backgrounds

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We have an  $O(2) \times O(3)$  symmetry and we can set up a *partial-wave decomposition* :

$$\ln \left[ \frac{\det(\not{D}^2 - m^2)}{\det(\not{\partial}^2 - m^2)} \right] = \sum_{l=0}^{\infty} \Omega(l) \ln \left[ \frac{\det(\mathcal{H}_l + m^2)}{\det(\mathcal{H}_l^0 + m^2)} \right]$$

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GY Theorem (initial value problem):

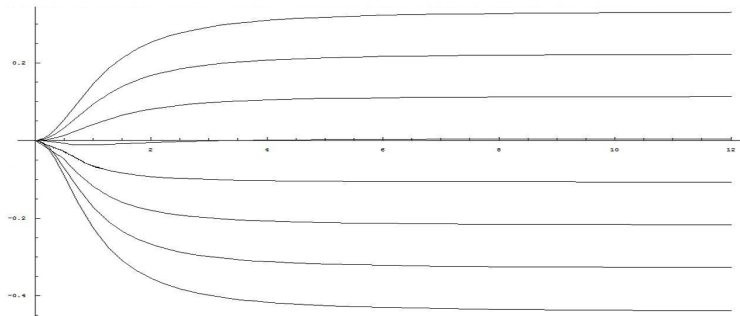
$$\frac{d^2 S_l(r)}{dr^2} + \left( \frac{dS_l(r)}{dr} \right)^2 + \left( \frac{1}{r} + 2m \frac{l'_{2l+1}(mr)}{l_{2l+1}(mr)} \right) \frac{dS_l(r)}{dr} = V(r)$$

$$\{ S_l(0) = 0, S'_l(0) = 0 \}$$

where  $S_l(r) \equiv \ln \frac{\psi(r)}{\psi^0(r)}$  and  $V(r)$  depends on  $g(r)$ .



We can find  $S_l(r)$  numerically



► Example:  $S_l(r)$  ,  $\{l = 4, l_3 = -4, \dots, 4, s_3 = 1/2\}$

# Is that simple?

However,

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$$\sum_{l=0}^L \Omega(l) S_l(\infty) \sim L^2$$

- ▶ Not really a surprise, in more than 1 + 1 dimensions, we need renormalization.

# The strategy

$$\sum_{l=0}^{\infty} \Omega(l) \mathcal{S}_l(\infty) = \sum_{l=0}^L \Omega(l) \mathcal{S}_l(\infty) + \sum_{l=L+1/2}^{\infty} \Omega(l) \mathcal{S}_l(\infty) = \Gamma_{\text{Low}} + \Gamma_{\text{High}}$$

Low-modes :  $\Rightarrow$  GY Theorem (numerical solution)

High-modes:  $\Rightarrow$  WKB series (analytic calculation), perform renormalization.

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$$\Delta_l(r, r; s) = \frac{e^{-s\mathcal{V}_l(r)}}{\sqrt{4\pi s}} \left[ 1 + \left( \frac{s^3}{12} (\mathcal{V}'_l(r))^2 - \frac{s^2}{6} \mathcal{V}''_l(r) \right) + \dots \right]$$

where  $\mathcal{V}_l(r)$  includes a centrifugal term that depends on  $l$ .

## The calculation

- ▶ First we perform the infinite sum over the angular momentum  $l$ . We use the Euler-Maclaurin formula for this:

$$\sum_{n=a}^b f(n) = \int_a^b f(x) dx + \frac{f(a) + f(b)}{2} + \dots$$



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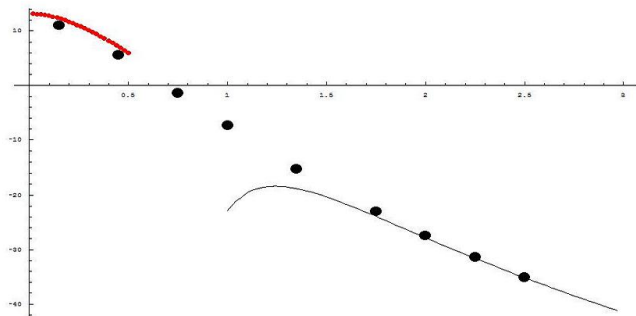
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- ▶ Next we integrate over  $ds$ , we can perform renormalization at this point.
- ▶ We are left with an integral over  $dr$

$$\Gamma_{\text{High}}^{\text{ren}} = \int_0^\infty dr \left( Q_{\log}(r) \ln L + \sum_{n=0}^2 Q_n(r) L^n + \sum_{n=1}^N Q_{-n}(r) \frac{1}{L^n} \right) + \mathcal{O}\left(\frac{1}{L^{N+1}}\right)$$

The result:  $\Gamma^{\text{ren}} = \Gamma_{\text{Low}} + \Gamma_{\text{High}}^{\text{ren}} < \infty$



- ▶ This shows an example of  $\Gamma^{\text{ren}}(m)$  for Scalar QED, with

$$g(r) = B(1 - \text{Tanh}[\beta\sqrt{Br} - \xi]) \quad ; \{B = 1, \beta = 1, \xi = 3\}$$

- ▶ The small-mass expansion (red), the large-mass expansion (line), our method (dots).

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# The spinor case

- ▶ The partial-wave decomposition is more subtle for Spinors.
- ▶ We must account for zero-modes.
- ▶ Our method allows  $A_\mu(r) = \eta_{\mu\nu}^3 x_\nu g(r)$  with arbitrary  $g(r)$ .  
The fall rate of  $g(r)$  determines the existence of zero-modes.
- ▶ We can aim to find properties that are independent of the precise form of  $g(r)$ .

## One application

$$\begin{aligned}\mathcal{L}_{\text{spinor}}^{(1)}(a, b) &= -\frac{1}{8\pi^2} \int_0^\infty \frac{ds}{s^3} e^{-m^2 s} \left\{ (ab)s^2 \coth(as)\coth(bs) \right. \\ &\quad \left. - 1 - \frac{s^2}{3} (a^2 - b^2) \right\}\end{aligned}$$

where  $a^2 - b^2 = -\frac{1}{2}F_{\mu\nu}F^{\mu\nu}$  and  $ab = -\frac{1}{4}F_{\mu\nu}\tilde{F}^{\mu\nu}$ .

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$$\begin{aligned}\mathcal{L}_{\text{spinor}}^{(1)}(a, b) &\sim \frac{1}{48\pi^2} [(a+b)^2 - 5(a-b)^2] \ln m + [\text{finite}] \\ &\quad + [\text{terms that vanish as } m \rightarrow 0]\end{aligned}$$

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- ▶ However :  $[\text{finite}] = f(a, b)$  remains unknown.

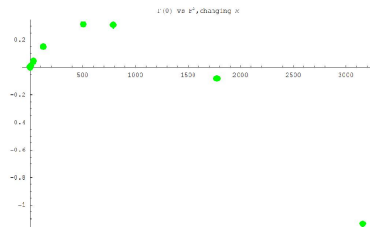
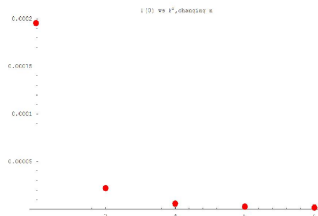


# Some tests

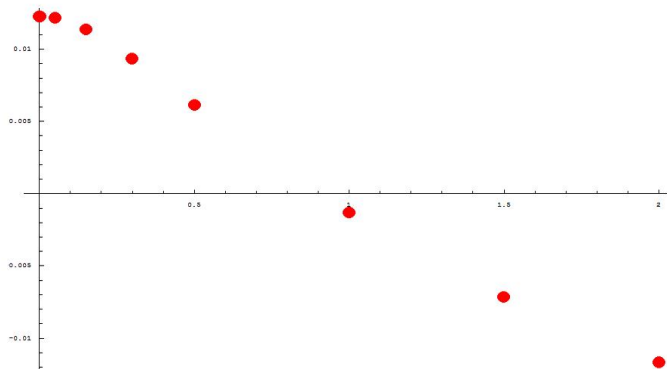
- ▶ Set  $m = 0$ , and consider

$$g(r) = \frac{\kappa}{(1 + r^2)^n}$$

- ▶ Now we may observe  $\Gamma^{\text{ren}}(m = 0)$  as a function of  $n$  (red) or  $\kappa$  (green) :



# Conclusion



- ▶ We can now accurately calculate  $\Gamma(m)$  for spinor abelian theories, in radially symmetric backgrounds.
- ▶ There are interesting questions to investigate, now within our reach.
- ▶ Thanks ■