Description of the Hyperon global polarization in Heavy-Ion Collisions at HADES, NICA and RHIC energies from the Core-Corona Model

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XIX Mexican School of Particles and Fields









2 Core Corona Model

 ${f 3}$ Excitation Function for the Global Λ and $ar\Lambda$ Polarization







Motivation: Why we are interested in measure Hyperon Global Polarization?





The fluid at midrapidity has a whirling substructure oriented (on average in the direction of the total angular momentum, J. [Nature 548,62-65(2017)]

- The Λ and $\bar{\Lambda}$ polarization are linked to the properties of the medium produced in relativistic heavy-ion collisions
- For semicentral collisions, Angular momentum can be quantified in terms of the thermal vorticity
- The global polarization can be measured using the self-analising $\Lambda/\bar{\Lambda}$ decays.



Global Polarization as a function of energy





Energy range $\sqrt{s_{NN}} = \{2, 11\}$ GeV can be covered by ongoing/future experiments

STAR BES-II + FXT: 3-19 GeV HADES:2-3 GeV NICA:4-11 GeV \rightarrow MPD



Energy dependence of kinematic vorticity predicted by a transport model (UrQMD) ^a

^oX.-G. Deng et al., PRC101.064908(2020)



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9.08.2021 – XIX MS of P&F 5/2





Core Corona Model: Two-component source

In heavy-ion collisions, Λ and $\bar{\Lambda}$ come from different density regions

• Core: Via QGP processes like

 $q\bar{q}\rightarrow s\bar{s} \text{ and } gg\rightarrow s\bar{s}$

• **Corona**: Via n + n reactions by recombination-like processes

The number of Λs can be written

$$N_{\Lambda} = N_{\Lambda_{QGP}} + N_{\Lambda_{REC}}$$

The polarization

$$\mathcal{P} = \frac{N^{\uparrow} - N^{\downarrow}}{N^{\uparrow} + N^{\downarrow}}$$

can be rewritten in terms of the number of Λ s (or $\overline{\Lambda}$ s) produced in the different density regions



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Rewriting Polarization



$$\mathcal{P}^{\Lambda} = \frac{(N^{\uparrow}_{\Lambda_{QGP}} + N^{\uparrow}_{\Lambda_{REC}}) - (N^{\downarrow}_{\Lambda_{QGP}} + N^{\downarrow}_{\Lambda_{REC}})}{(N^{\uparrow}_{\Lambda_{QGP}} + N^{\uparrow}_{\Lambda_{REC}}) + (N^{\downarrow}_{\Lambda_{QGP}} + N^{\downarrow}_{\Lambda_{REC}})}$$

$$\mathcal{P}^{\bar{\Lambda}} = \frac{(N^{\uparrow}_{\bar{\Lambda}_{QGP}} + N^{\uparrow}_{\bar{\Lambda}_{REC}}) - (N^{\downarrow}_{\bar{\Lambda}_{QGP}} + N^{\downarrow}_{\bar{\Lambda}_{REC}})}{(N^{\uparrow}_{\bar{\Lambda}_{QGP}} + N^{\uparrow}_{\bar{\Lambda}_{REC}}) + (N^{\downarrow}_{\bar{\Lambda}_{QGP}} + N^{\downarrow}_{\bar{\Lambda}_{REC}})}$$

After some algebra, we get:

$$\mathcal{P}^{\Lambda} = rac{\left(\mathcal{P}^{\Lambda}_{REC} + rac{N^{\uparrow}_{\Lambda_{QGP}} - N^{\downarrow}_{\Lambda_{QGP}}}{N_{\Lambda_{REC}}}
ight)}{\left(1 + rac{N_{\Lambda_{QGP}}}{N_{\Lambda_{REC}}}
ight)}$$

$$\mathcal{P}^{\bar{\Lambda}} = \frac{\left(\mathcal{P}_{REC}^{\bar{\Lambda}} + \frac{N_{\bar{\Lambda}_{QGP}}^{\uparrow} - N_{\bar{\Lambda}_{QGP}}^{\downarrow}}{N_{\bar{\Lambda}_{REC}}}\right)}{\left(1 + \frac{N_{\bar{\Lambda}_{QGP}}}{N_{\bar{\Lambda}_{REC}}}\right)}$$

Where the polarization along the angular momentum produced in the corona is:

$$\mathcal{P}^{\Lambda}_{REC} = rac{N^{\uparrow}_{\Lambda_{REC}} - N^{\downarrow}_{\Lambda_{REC}}}{N^{\uparrow}_{\Lambda_{REC}} + N^{\downarrow}_{\Lambda_{REC}}}$$

$$\mathcal{P}_{REC}^{\bar{\Lambda}} = \frac{N_{\bar{\Lambda}_{REC}}^{\uparrow} - N_{\bar{\Lambda}_{REC}}^{\downarrow}}{N_{\bar{\Lambda}_{REC}}^{\uparrow} + N_{\bar{\Lambda}_{REC}}^{\downarrow}}$$



208.2021 - XIX MS of P&F 8/2



Assumptions: Polarization of Λ ($ar\Lambda$) from the Corona





- Nucleon-Nucleon scattering not enough to align the spin in the direction of the angular momentum
- Polarization of Λ and $ar{\Lambda}$ averages to zero.

$$\mathcal{P}^{\Lambda}_{REC}=\mathcal{P}^{\bar{\Lambda}}_{REC}=0$$





$$\mathcal{P}^{\Lambda} = \frac{\left(\mathcal{P}^{\Lambda}_{REC} + \frac{N^{\dagger}_{\Lambda_{QGP}} - N^{\downarrow}_{\Lambda_{QGP}}}{N_{\Lambda_{REC}}}\right)}{\left(1 + \frac{N_{\Lambda_{QGP}}}{N_{\Lambda_{REC}}}\right)}$$
$$\mathcal{P}^{\bar{\Lambda}} = \frac{\left(\mathcal{P}^{\bar{\Lambda}}_{REC} + \frac{N^{\dagger}_{\Lambda_{QGP}} - N^{\downarrow}_{\Lambda_{QGP}}}{N_{\bar{\Lambda}_{REC}}}\right)}{\left(1 + \frac{N_{\bar{\Lambda}_{QGP}}}{N_{\bar{\Lambda}_{REC}}}\right)}$$

We define ${\bf z}$ and $\bar{{\bf z}}$ which represent the Λ and $\bar{\Lambda}$ intrinsic polarization respectively

$$egin{aligned} &N^{\uparrow}_{\Lambda_{QGP}}-N^{\downarrow}_{\Lambda_{QGP}}= extbf{z}N_{\Lambda_{QGP}}\ &N^{\uparrow}_{ar{\Lambda}_{QGP}}-N^{\downarrow}_{ar{\Lambda}_{QGP}}=ar{ extbf{z}}N_{ar{\Lambda}_{QGP}} \end{aligned}$$



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9.08.2021 – XIX MS of P&F 10/27





The number of $\overline{\Lambda}$ s are proportional to an energy-dependent coefficient $\mathbf{w}(\mathbf{w}')$ times the number of Λ s in the corona(core)

$$egin{array}{rcl} N_{ar{\Lambda}_{REC}} &=& oldsymbol{w} N_{\Lambda_{REC}} \ N_{ar{\Lambda}_{QGP}} &=& oldsymbol{w}' N_{\Lambda_{QGP}} \end{array}$$



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Λ and $\bar{\Lambda}$ global polarization



With this assumptions

$$\mathcal{P}^{\Lambda}_{REC}=\mathcal{P}^{\bar{\Lambda}}_{REC}=0$$

$$egin{array}{lll} N^{\uparrow}_{\Lambda_{QGP}} &- N^{\downarrow}_{\Lambda_{QGP}} &= & z N_{\Lambda_{QGP}} \ N^{\uparrow}_{ar{\Lambda}_{QGP}} &- N^{\downarrow}_{ar{\Lambda}_{QGP}} &= & ar{z} N_{ar{\Lambda}_{QGP}} \end{array}$$

$$egin{array}{rcl} N_{ar{\Lambda}_{REC}} &=& w N_{\Lambda_{REC}} \ N_{ar{\Lambda}_{QGP}} &=& w' N_{\Lambda_{QGP}} \end{array}$$

Global polarization depends on the coefficients w, w', z, \bar{z} and the ratio $\frac{N_{\Lambda_{QGP}}}{N_{\Lambda_{REC}}}$ that can be estimated from data or calculated



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09.08.2021 - XIX MS of P&F 12/2



The ratio $w=N_{ar{\Lambda}_{REC}}/N_{\Lambda_{REC}}$

Model as p + p collisions

- Experimental data obtained from p + p collisions at different energies¹
- w is defined only for $\sqrt{s} > 4.1 {\rm GeV}$. The threshold energy for $p+p \rightarrow p+p+\Lambda+\bar{\Lambda}$
- w is smaller than 1 except for energies $\sqrt{s} > 1$ TeV



¹M. Gazdzicki and D. Rohrich, Z. Phys. C 71 (1996) 55; V. Blobel et al. Nucl. Phys. B 69(1974), 454–492; J. W. Chapman et al., Phys. Lett. 47B (1973) 465; D. Brick et al., Nucl. Phys. B 164 (1980) 1; C. Höhne, CERN-THESIS-2003-034; J. Baechler et al. [NA35 Collaboration], Nucl. Phys. A 525 (1991) 221C; G. Charlton et al., Phys. Rev. Lett. 30 (1973) 574; F. Lopinto et al., Phys. Rev. D 22 (1980) 573; H. Kichimi et al., Phys. Rev. D 20 (1979) 37; F. W. Busser et al., AID (1976) 309; S. Erhan, et al., Phys. Lett. 35B(1979) 447; B. L. Abelev et al., [STAR Collaboration], Phys. Rev. C 75 (2007) 06490); E. Abbas et al. [ALICE Collaboration], Eur. Phys. J. C 73 (2013) 2496



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The ratio $w' = N_{ar{\Lambda}_{QGP}}/\overline{N_{\Lambda_{QGP}}}$

• The coefficient w' is computed as the ratio of the equilibrium distributions of \overline{s} to s-quark for a given temperature and chemical potential $\mu = \mu_B/3$

$$w' = \frac{e^{(m_s - \mu)/T} + 1}{e^{(m_s + \mu)/T} + 1}$$

 m_s the mass of the s-quark and μ_B and T along the maximum chemical potential at freeze out.





Production of Λ in the core and the corona



$$N_{\Lambda_{QGP}} = c N_{p_{QGP}}^2$$

in which

$$N_{p_{QGP}} = \int n_p(\mathbf{s},\mathbf{b}) heta[n_p(\mathbf{s},\mathbf{b}) - n_c] d^2s$$

with $n_c = 3.3$ fm⁻², the critical density required to form the QGP

Number of Λ s in the corona

$$N_{\Lambda_{REC}} = \sigma_{NN}^{\Lambda} \int T_B(\mathbf{b}\!-\!\mathbf{s}) T_A(\mathbf{s}) \theta[n_c\!-\!n_p(\mathbf{s},\!\mathbf{b})] d^2s$$

where σ_{NN}^{Λ} is obtained from experimental





$\Lambda {f s}$ in the Core and Corona



At low energies $N_{\Lambda_{QGP}}$ depends on σ_{NN} , different parametrizations impact on the strenght of polarization

 σ_{NN} affects the ratio $N_{\Lambda QGP}/N_{\Lambda_{REC}}$ and the value of b at which the ratio is smaller than 1



Hyperon global polarization





Intrinsic Polarization

The intrinsic polarizations are given by:

$$z = 1 - e^{-\Delta \tau_{QGP}/\tau}$$

$$\bar{z} = 1 - e^{-\Delta au_{QGP}/ar{ au}}$$

in terms of the relaxation times τ and $\bar{\tau}$ an the QGP life-time $\Delta \tau_{QGP}$ The relaxation time can be computed as the inverse of the interaction rate

$$\tau \equiv 1/\Gamma$$

given by

$$\Gamma = V \int \frac{d^3 p}{2\pi^3} \Gamma(p_0)$$
, with $V = \pi R^2 \Delta \tau_{QGP}$

where V is the volume of the core region, related with the QGP life-time $\Delta\tau_{QGP}$ in the scenario of a Bjorken expansion

$$\Delta \tau_{QGP} = \tau_f - \tau_0 = \tau_0 \left[\left(\frac{T_0}{T_f} \right)^3 - 1 \right]$$





Volume and QGP life-time





 T_0 is estimated from p_T of ϕ mesons, $\tau_0=0.35-0.60$ fm to incorporate the effect of collision centrality, and T_f is taken as the value along the maximum chemical potential curve at freeze-out



μ_B and T at freeze out





Eur.Phys.J. 52 (2016) 218-219

Maximum freeze-out baryon density in nuclear collisions

$$T(\mu_B) = 166 - 139\mu_B^2 - 53\mu_B^4$$
$$\mu_B(\sqrt{s_{NN}}) = \frac{1308}{1000 + 0.273\sqrt{s_{NN}}}$$

Phys.Rev.C 74 (2006) 047901

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Relaxation time and intrinsic polarization as a function of energy



For energies below the Λ production threshold energy, the τ and $\bar{\tau}$ increase dramatically, as expected, since the interaction rate should vanish below these energies.



Section 3

Excitation Function for the Global Λ and $\bar{\Lambda}$ Polarization





10 - 40 %







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9.08.2021 – XIX MS of P&F 22/27

Λ and $\bar{\Lambda}$ polarization in Au+Au at BES centrality



20 - 50 %

- Similar trend to the case of the analysis with smaller centrality
- Magnitude of global polarization increases for larger centrality as a consequence of the angular velocity increase

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- The two component source model describes the main characteristics features of the Λ and $\bar{\Lambda}$ polarization excitation functions in semi central relativistic heavy-ion collisions
- The change of the relative abundance of Λ s coming from the core versus those coming from the corona as a function of the collision energy has as a consequence that both polarization peak at collision energies $\sqrt{s_{NN}} \lesssim 10 \text{ GeV}$
- The relaxation time can be obtained can be obtained from a field theoretical approach that links the alignment of the strange quark spin with the thermal vorticity, modeling the QGP volume and life-time using a simple scenario
- The model predicts a maximum for the Λ and $\bar{\Lambda}$ polarizations which should be possible to be measured in the NICA and HADES energy range.







- Extend the analysis for Ag+Ag collisions to compare with HADES preliminary SQM2021
- Trend looks similar for 0-40% of centrality







Thank You



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