

Observation of radio galaxies with HAWC

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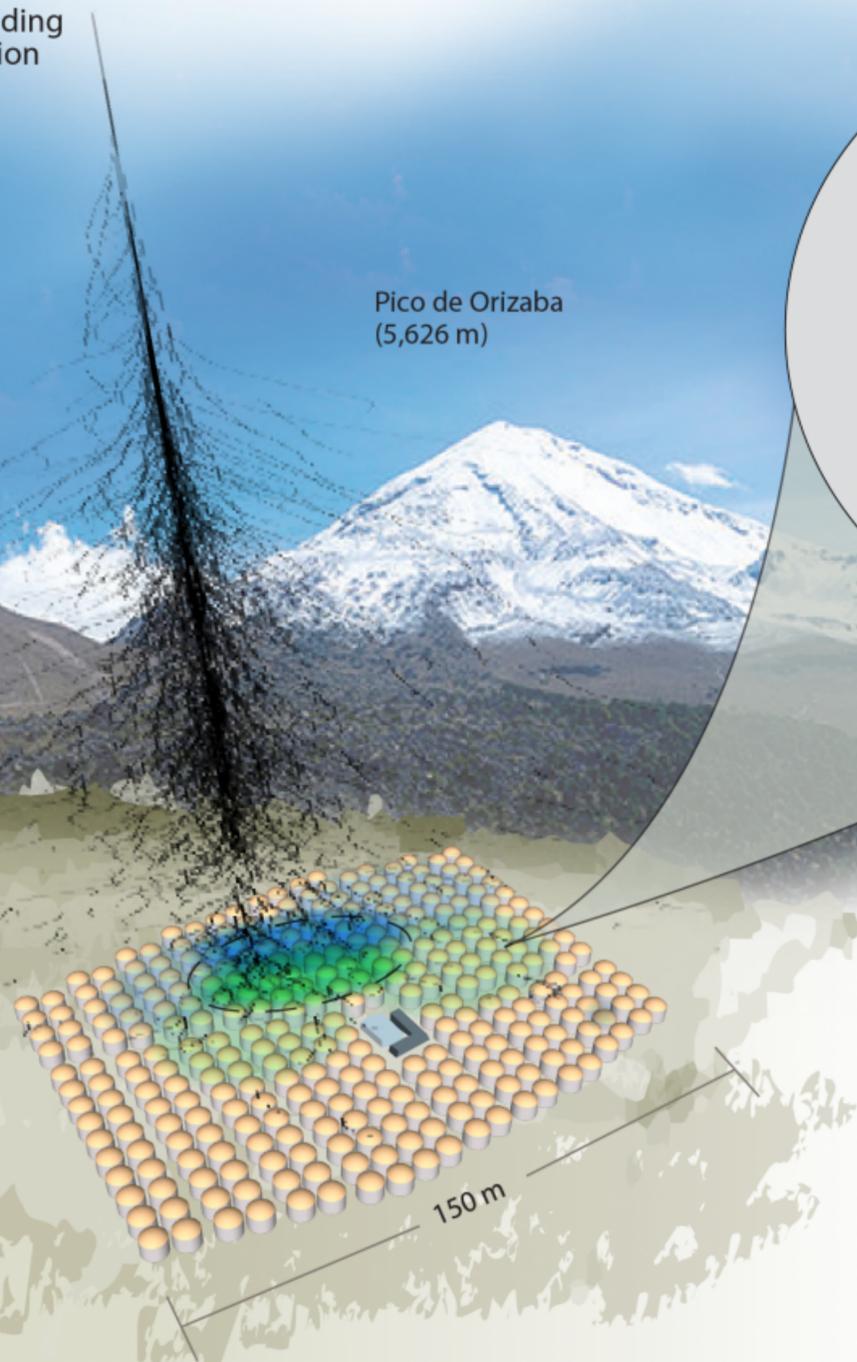
HAWC Observatory

HAWC operates day and night, providing a large field of view for the observation of the highest energy gamma rays.



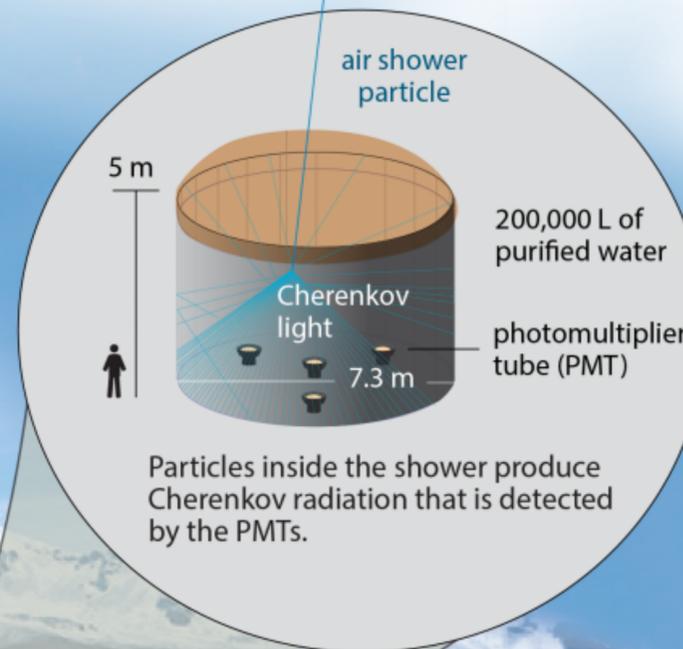
Pico de Orizaba
(5,626 m)

HAWC is located at 4,100 m above sea level, covering an area of 20,000 m².



Water Cherenkov tank

HAWC comprises an array of 300 tanks that record the particles created in gamma-ray and cosmic-ray showers.

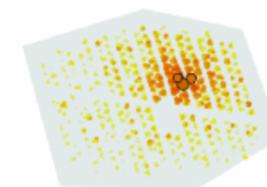


Particles inside the shower produce Cherenkov radiation that is detected by the PMTs.

Gamma rays vs cosmic rays

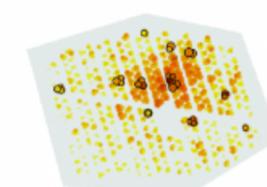
HAWC selects gamma rays from among a much more abundant background of cosmic rays.

gamma-ray shower



"hot" spots concentrate around the core

cosmic-ray shower



"hot" spots are more dispersed

Some characteristics:

- DC > 95%
- Wide F.O.V
- 300 GeV - 100 TeV

=> HAWC is a good experiment to source monitoring.

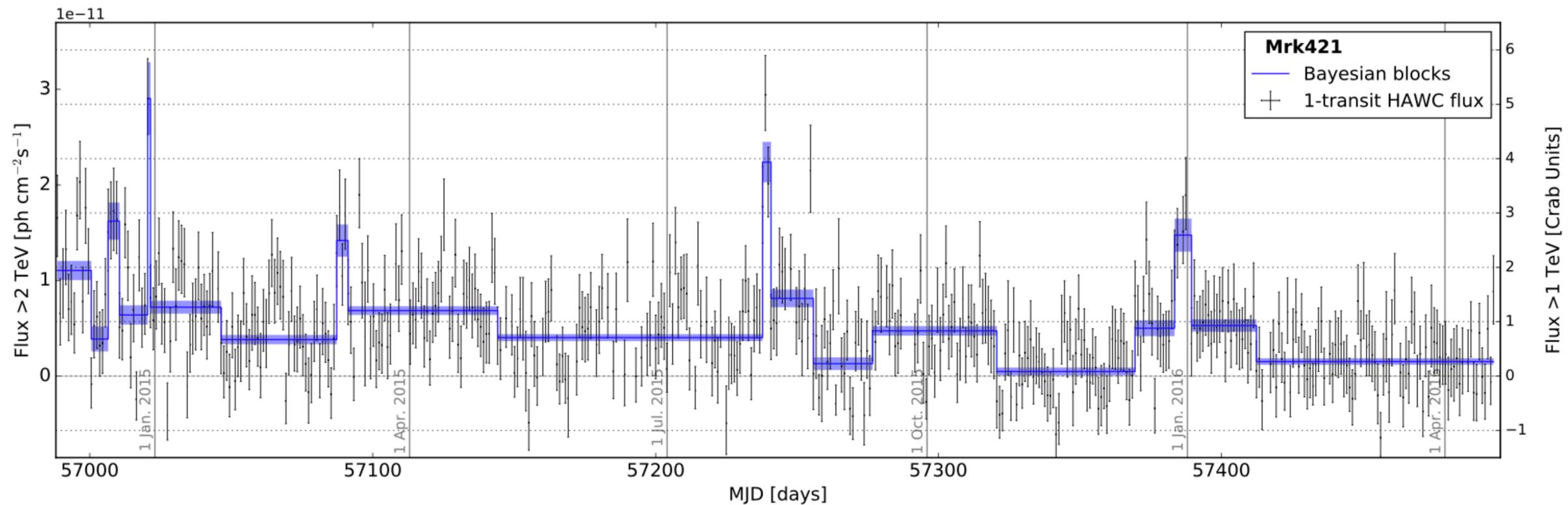


Figure 4. Flux light curve for Mrk 421 with sidereal-day sampling for 471 transits between 2014 November 26 and 2016 April 20. The integrated fluxes are derived from fitting F_i in $dN/dE = F_i (E/(1 \text{ TeV}))^{-2.2} \exp(-E/(5 \text{ TeV}))$ and converted to Crab Units via dividing by the HAWC measurement of the average Crab Nebula gamma-ray flux. The blue lines show the distinct flux states between change points identified via the Bayesian blocks analysis with a 5% false positive probability.

A. U. Abeysekara (2017) DOI: 10.3847/1538-4357/aa729e

Radio Galaxy (RG)

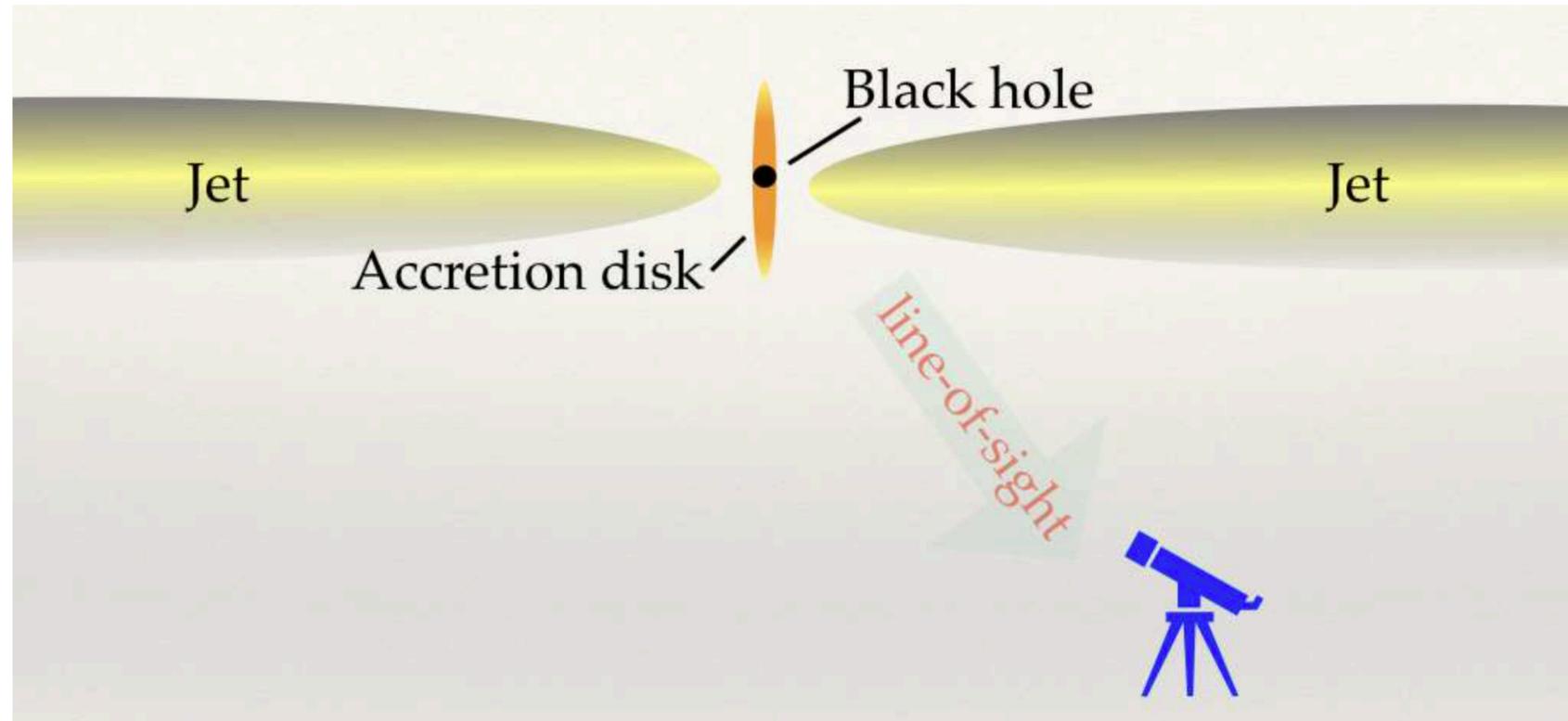


Figure 1. Sketch representing the mis-aligned radio emitting bipolar jets of a radio galaxy (not to scale). Jets typically extend up to a few hundreds of kilo-parsec to mega-parsec scales.

B. Rrani (2018) [arXiv:1811.00567](https://arxiv.org/abs/1811.00567)

- RG is a class of Active Galactic Nuclei (AGN).
- Host a relativistic jet misaligned with the line of sight.
- Probe the physics of VHE emission process.

Radio Galaxies at TeV

Table 1. 3LAC MAGN: (i) Source name; (ii) 3FGL name; (iii) right ascension; (iv) declination; (v) redshift; (vi) radio classification; (vii) detected TeV counterpart. 3C 120 has been added to the list since it is a confirmed γ -ray source, and also Tol1326-379 and PKS1718-379 recently associated with 3FGL sources.

Source	3FGL Name	RA (J2000)	Dec (J2000)	z	FR type	TeV source
3C 78	3FGLJ0308.6+0408	03 08 26.22	+04 06 39.3	0.02865	FR I	no
IC310	3FGLJ0316.6+4119	03 16 42.97	+41 19 29.61	0.019	FR I	yes
NGC1275	3FGLJ0319.8+4130	03 19 48.16	+41 30 42.1	0.0175	FR I	yes
ForA(lobes)	3FGLJ0322.5-3721	03 21 37.75	-37 12 49.1	0.00587	FR I	no
4C+39.12	3FGLJ0334.2+3915	03 34 18.41	+39 21 24.4	0.02059	FR I	no
Pictor A	3FGLJ0519.2-4542	05 19 49.72	-45 46 43.85	0.03506	FR II	no
PKS0625-35	3FGLJ0627.0-3529	06 27 06.72	-35 29 15.33	0.05494	FR I	yes
3C 189	3FGLJ0758.7+3747	07 58 28.1	+37 47 11.8	0.04284	FR I	no
3C 221	3FGLJ0934.1+3933	09 35 06.63	+39 42 06.7	–	FR II/SSRQ	no
3C 264	3FGLJ1145.1+1935	11 45 05.0	+19 36 22.74	0.02172	FR I	yes
M87	3FGLJ1230.9+1224	2 30 49.42	+12 23 28.04	0.00428	FR I	yes
CenA(core)	3FGLJ1325.4-4301	13 25 27.61	-43 01 08.8	0.0018	FR I	yes
3C 303	3FGLJ1442.6+5156	14 43 02.76	+52 01 37.29	0.14119	FR II	no
NGC6251	3FGLJ1630.6+8232	16 32 31.96	+82 32 16.39	0.02	FR I	no
3C 111	3FGLJ0418.5+3813c	04 18 21.27	+38 01 35.8	0.0485	FR II	no
CenB	3FGLJ1346.6-6027	13 46 49.04	-60 24 29.35	0.01292	FR I	no
TXS0348+013	3FGLJ0351.1+0128	03 50 57.36	+01 31 04.91	1.12	FR II/SSRQ	no
3C 207	3FGLJ0840.8+1315	08 40 47.58	+13 12 23.56	0.681	FR II/SSRQ	no
PKS1203+04	3FGLJ1205.4+0412	12 06 19.92	+04 06 12.04	0.536	FR II/SSRQ	no
3C 275.1	3FGLJ1244.1+1615	12 43 57.64	+16 22 53.39	–	FR II/SSRQ	no
3C 380	3FGLJ1829.6+4844	18 29 31.78	+48 44 46.16	0.695	FR II/SSRQ	no
3C 120	–	04 33 11.1	+05 21 16	0.033	FR I	no
Tol1326-379	3FGL1330.0-3818	13 29 19.2	-38 14 18	0.028	FR 0	no
PKS1718-649	3FGLJ1728.0-6446	17 23 41.0	-65 00 37	0.014	CSO	no

RG detected with TeV

counterpart:

1. IC310
2. NGC1275
3. PKS0623-35
4. 3C 264
5. M87
6. CenA(core)

Radio Galaxies at TeV

Table 1. Radio galaxies detected at TeV energies

Source	Type	Redshift (Distance in Mpc)	M_{BH} (M_{\odot})	L_{VHE} (erg s^{-1})
Centaurus A	FR1	0.00183 (3.7) [23]	5×10^7 [24]	10^{40}
M87	FR1	0.0044 (16) [25]	6×10^9 [26]	10^{41}
3C 84	FR1	0.0177 (71) [27]	$(3-8) \times 10^8$ [28,29]	10^{45}
IC 310	FR1	0.0189 (80) [30]	$(1-7) \times 10^8$ [24,31]	10^{44}
3C 264	FR1	0.0217 (95) [32]	2.6×10^8 [33]	6×10^{43}
PKS 0625-35	FR1/BL Lac	0.05488 (220) [34]	3×10^9 [24]	5×10^{41}

B. Rani (2018) arXiv:1811.00567

The RG are:

- The closest Extragalactic objects.
- Harbors a Super Massive Black Hole $\sim 10^{(8-9)}$ solar masses.
- Luminosities up to 10^{45} ergs s^{-1} .
- Classified as Fanaroff and Riley type I (FR1).

Radio Galaxies at TeV

“Like Blazars, TeV radio galaxies exhibit variability on multiple timescale” B. Rani (2018).

There are two state: flaring and quiescent.

A. M87 was the first TeV detected radio galaxy.

Three active states:

1. 2005 flare is reported by H.E.S.S.
2. 2008 flare is reported by MAGIC
3. 2010 flare is reported by VERITAS

B. NGC1275

Two active states:

4. October 2016
5. Dember 2016

C. 3C 264

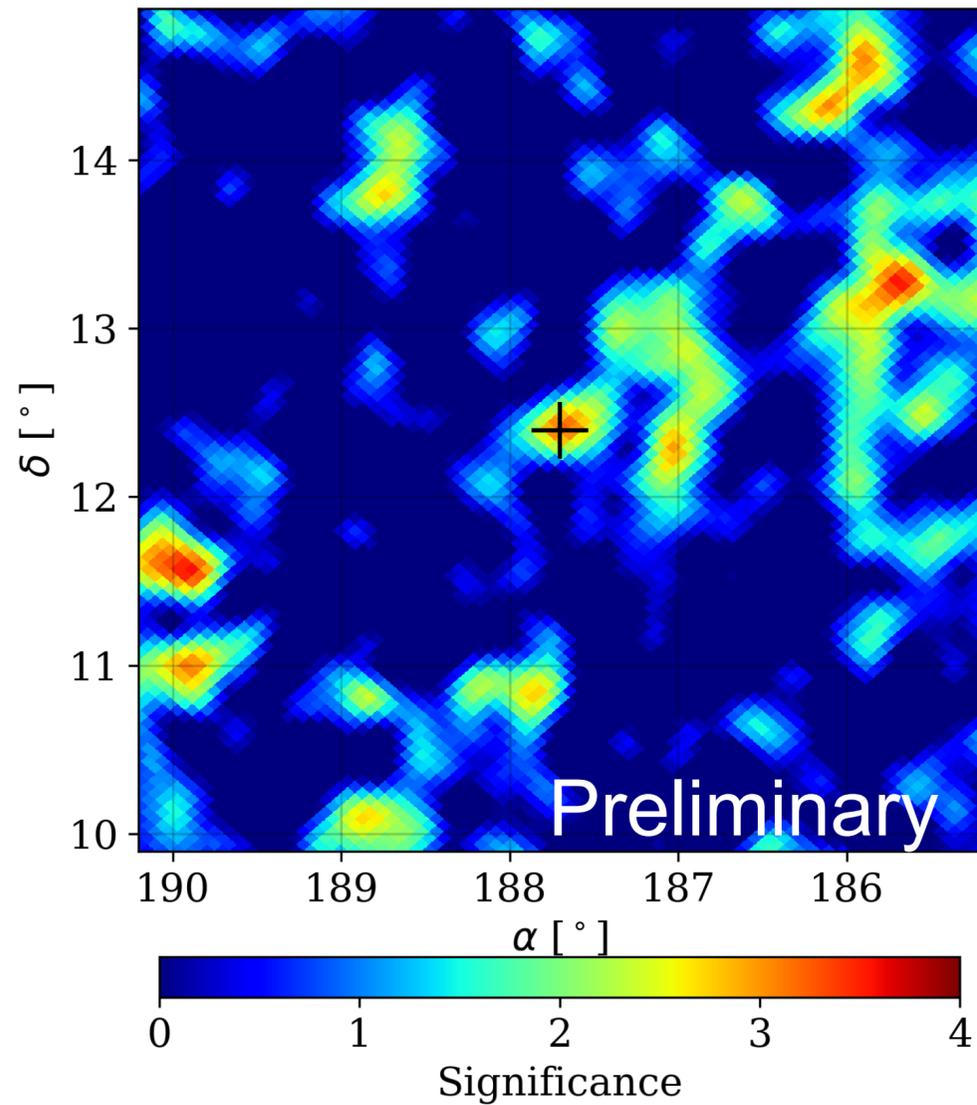
Figures

Using 1523 days

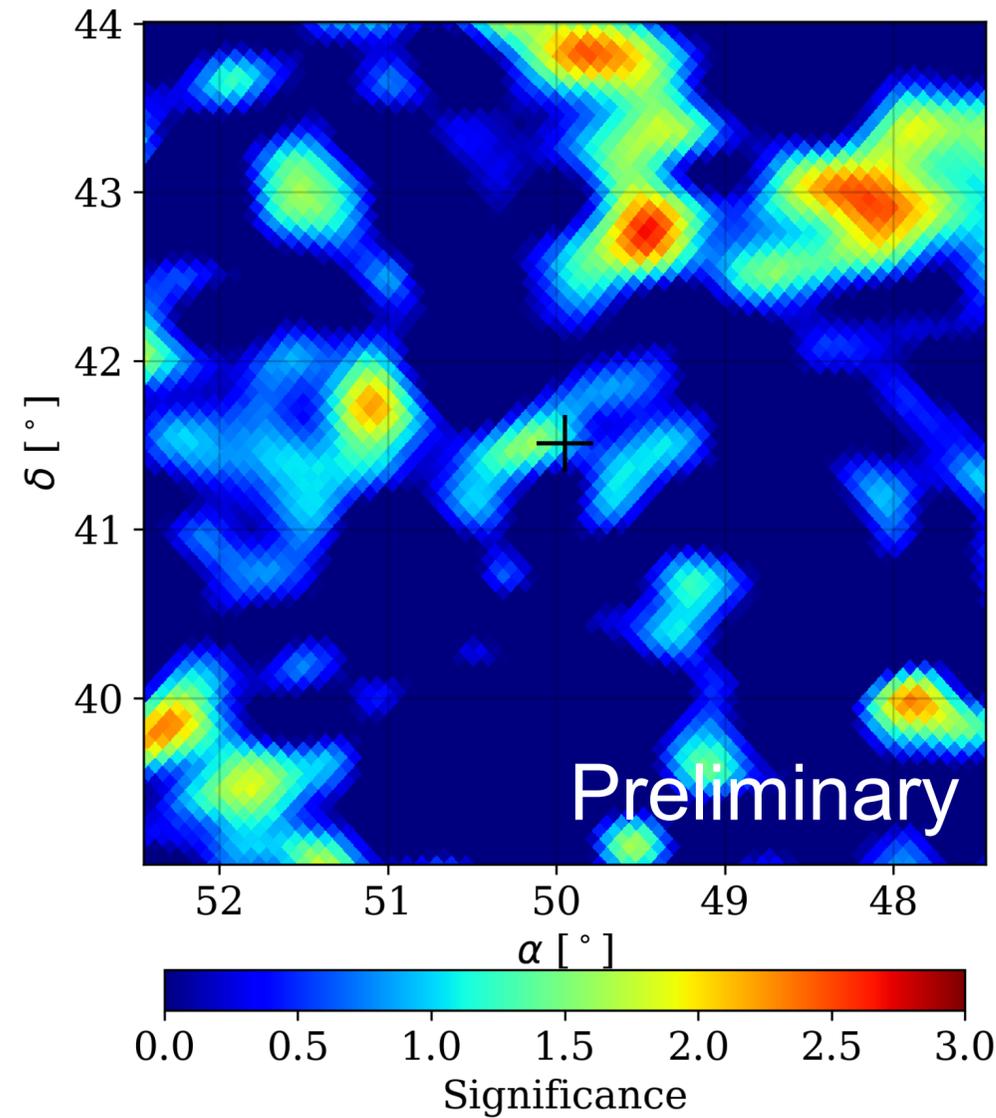
M87

NGC1275

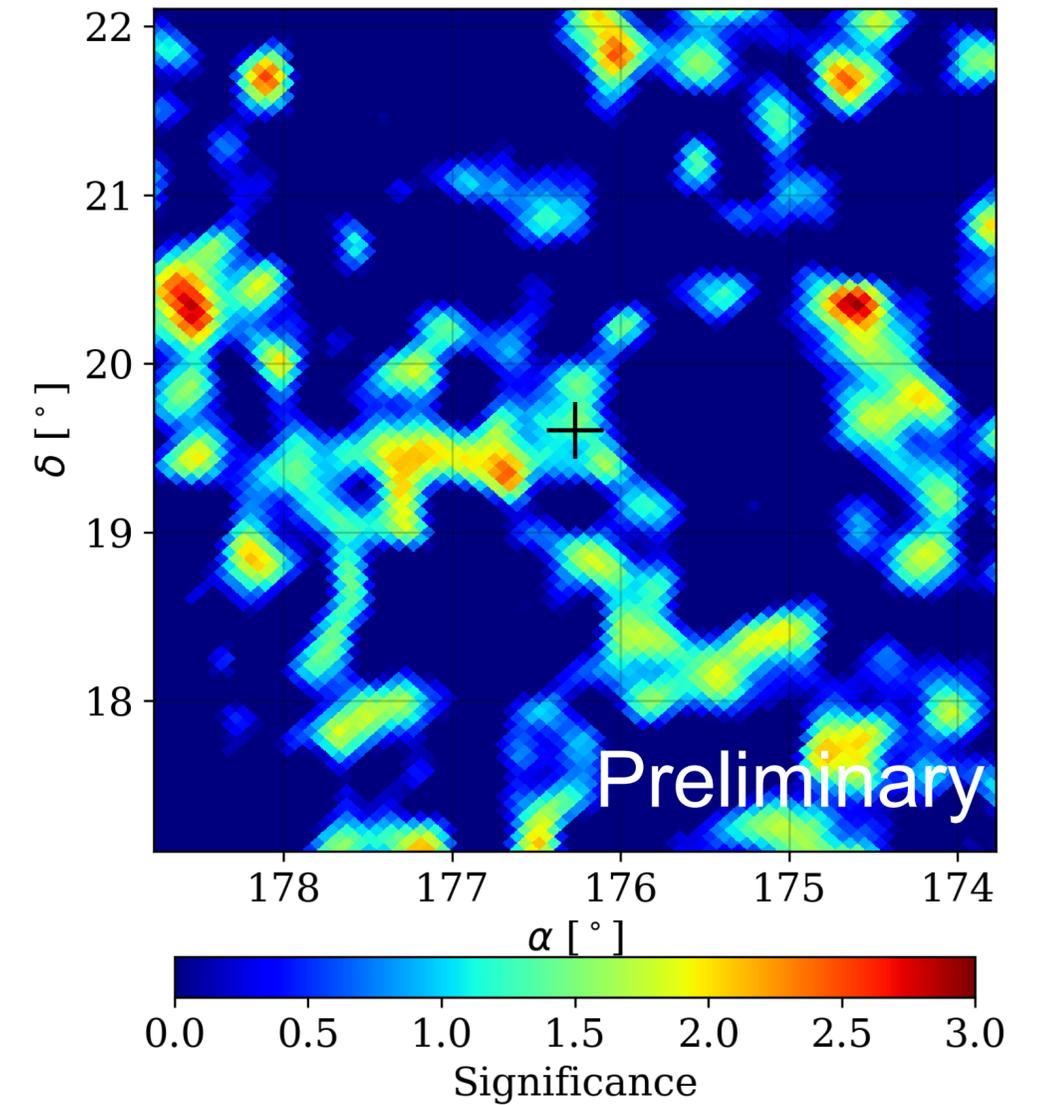
3C 264



At source: 3.20



At source: 1.15



At source: 1.47

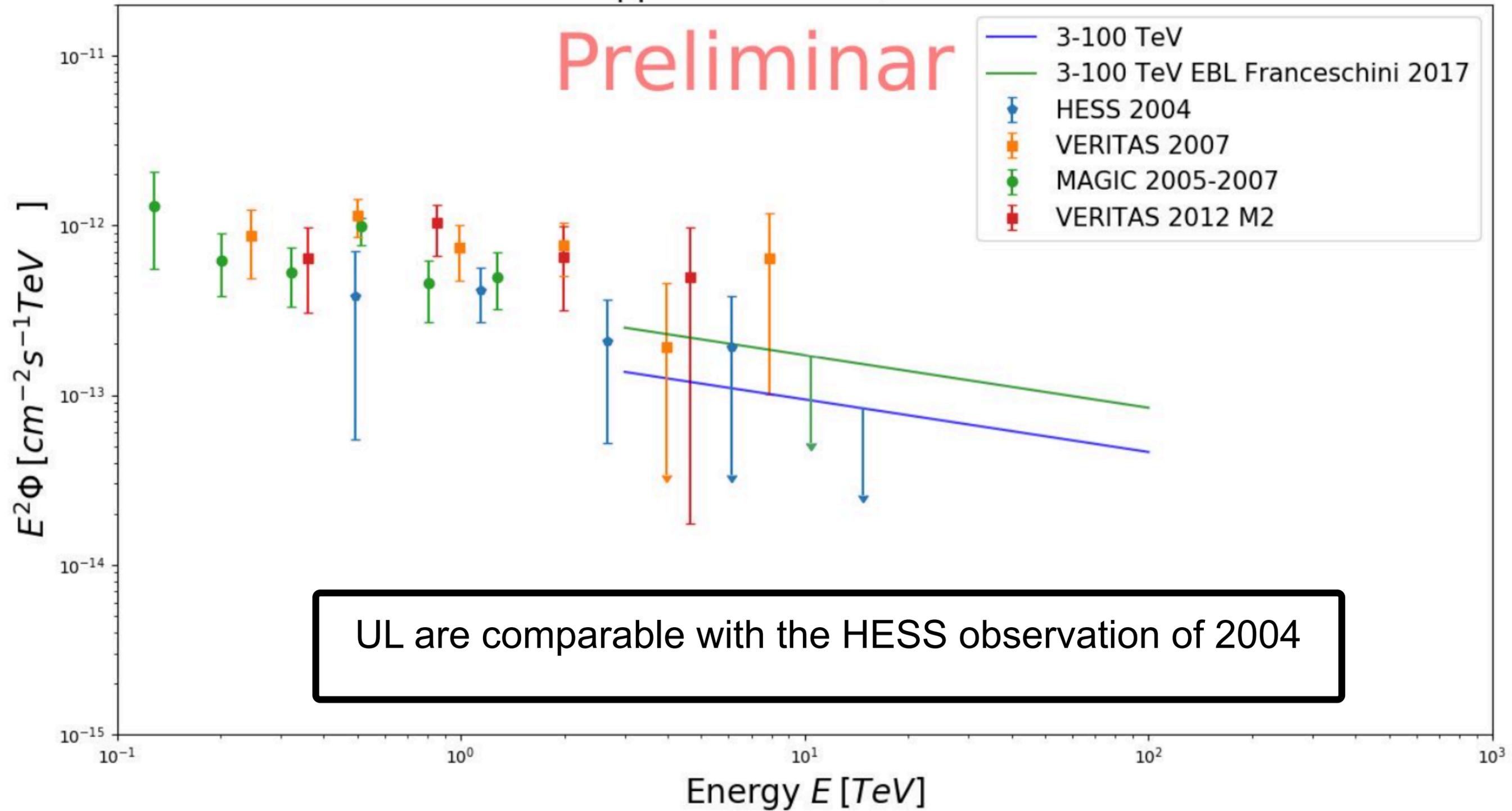
Upper Limit

Source	UL with EBL (TeV⁻¹cm⁻²s⁻¹)	UL without EBL (TeV⁻¹cm⁻²s⁻¹)	Sigma
M87	2.63E-13	2.63E-13	2.77
NGC1275	1.21E-12	1.05E-11	0.82
3C 264	1.89E-13	2.56E-13	-1.36

EBL Model = Franceschini 2008

Upper Limits M87, SI 2.31

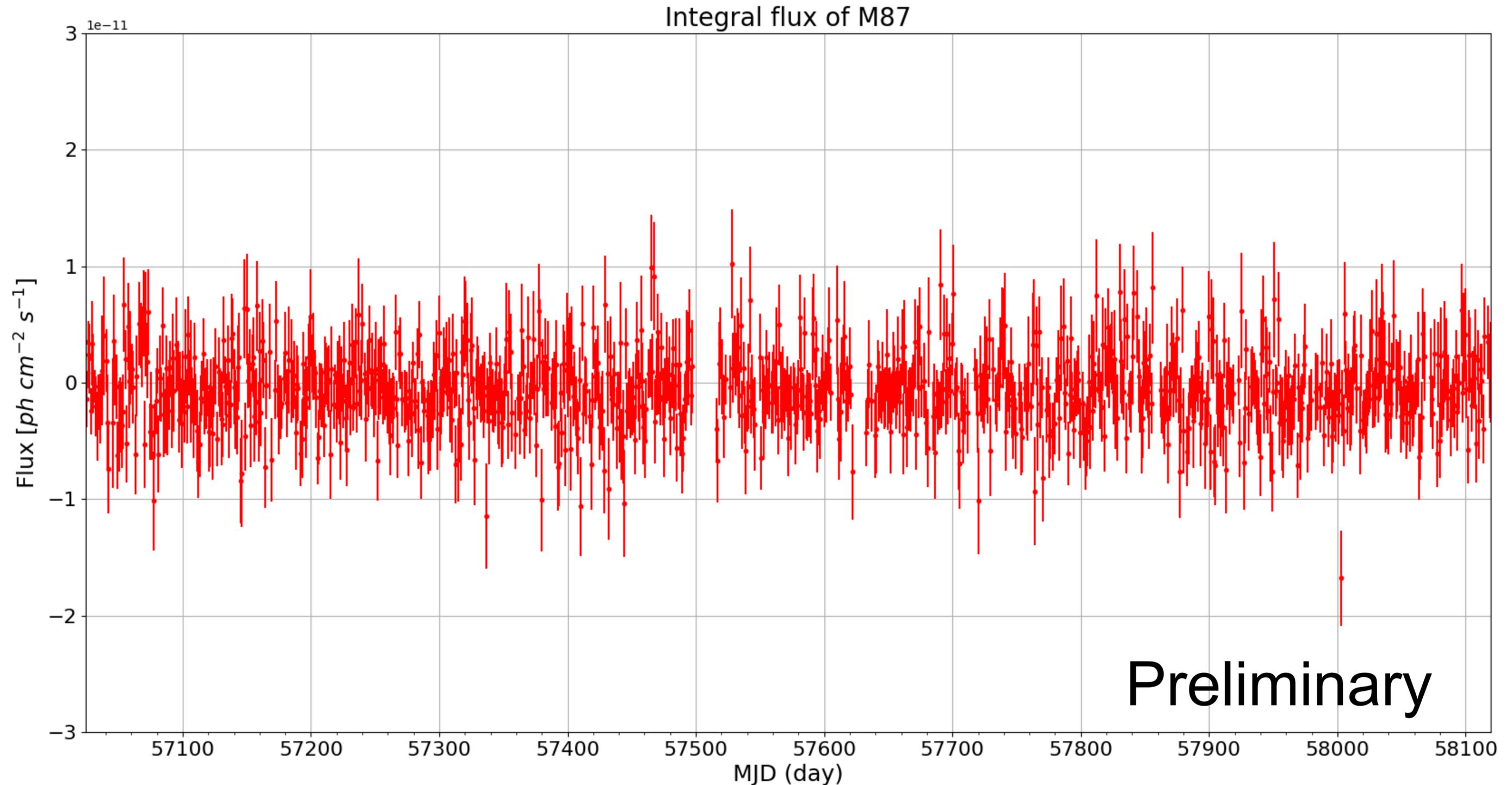
Preliminary



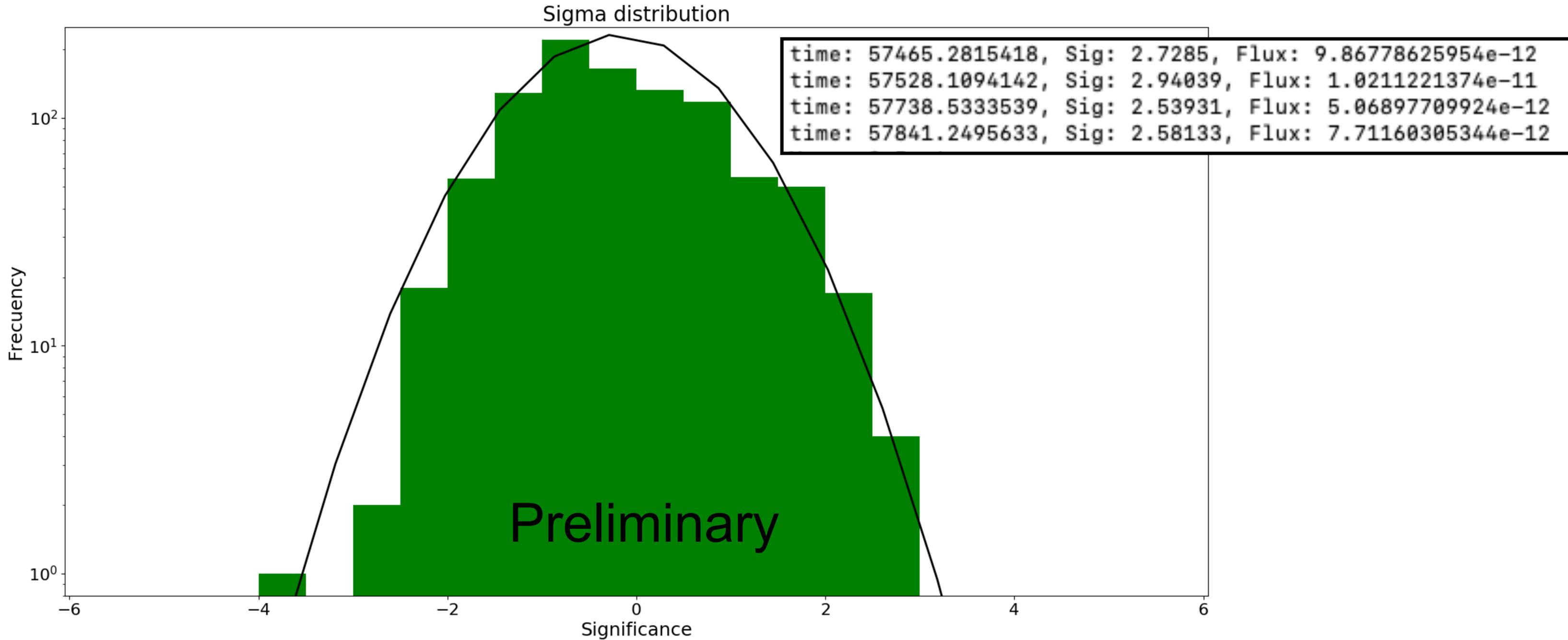
D. Avila (2019) arXiv:1909.00075 & thesis

M87 light curve

Simple Power Law
Index = 2.31
Redshift = 0.0044
Energy pivot = 1 TeV



M87 light curve

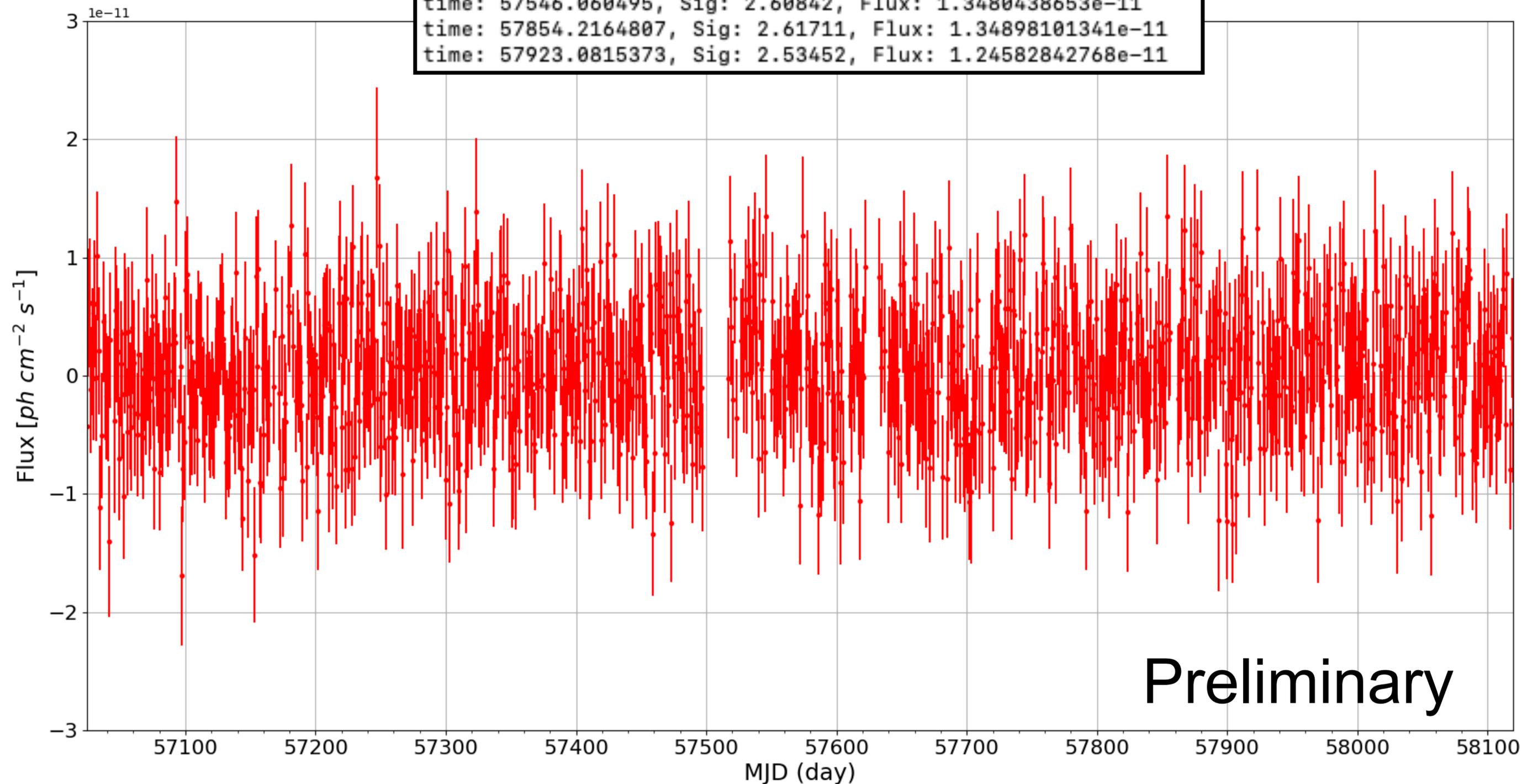


Fit results: $\mu = -0.19$, $\text{std} = 1.02$

NGC 1275 light curve

CutOff Power Law
Index = 3.0
Redshift = 0.017559
Energy pivot = 1 TeV
Energy cut = 500 GeV

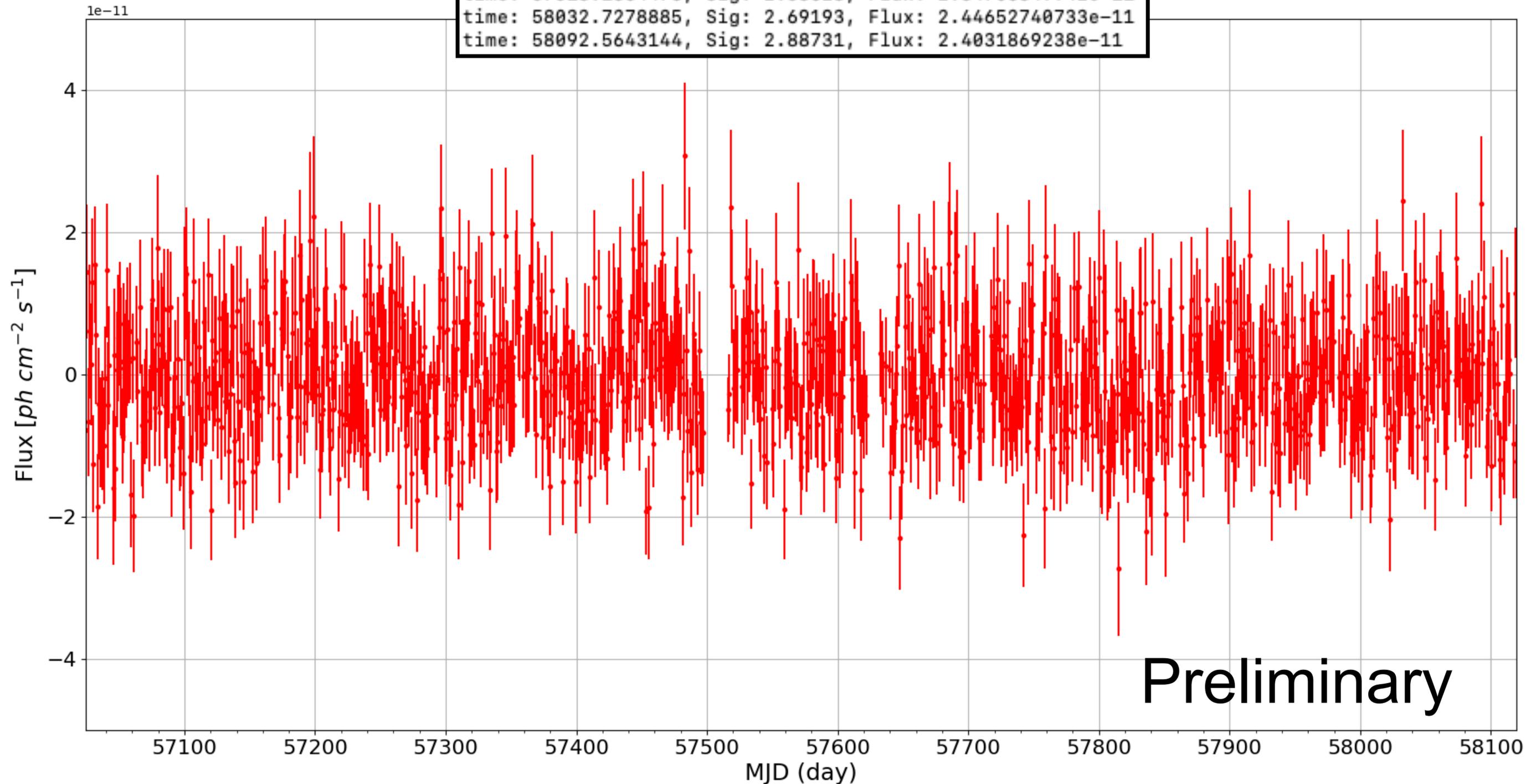
time: 57093.2727599, Sig: 2.75227, Flux: 1.47457091373e-11
time: 57404.4481087, Sig: 2.50094, Flux: 1.24323394368e-11
time: 57546.060495, Sig: 2.60842, Flux: 1.3480438653e-11
time: 57854.2164807, Sig: 2.61711, Flux: 1.34898101341e-11
time: 57923.0815373, Sig: 2.53452, Flux: 1.24582842768e-11



3C 264 light curve

CutOff Power Law
Index = 2.3
Redshift = 0.022
Energy pivot = 1 TeV
Energy cut = 10 TeV

time: 57296.7433623,	Sig: 2.94752,	Flux: 2.33722909339e-11
time: 57483.2325958,	Sig: 3.32255,	Flux: 3.07200992702e-11
time: 57518.1364476,	Sig: 2.53628,	Flux: 2.34706349942e-11
time: 58032.7278885,	Sig: 2.69193,	Flux: 2.44652740733e-11
time: 58092.5643144,	Sig: 2.88731,	Flux: 2.4031869238e-11



Summary

- The significance of the RG are less than 5 sigmas.
- The UL were computed for three radio galaxies.
- The LC were obtained for the three radio galaxies.

Thanks!