

PIERRE
AUGER
OBSERVATORY

Hybrid long term performance

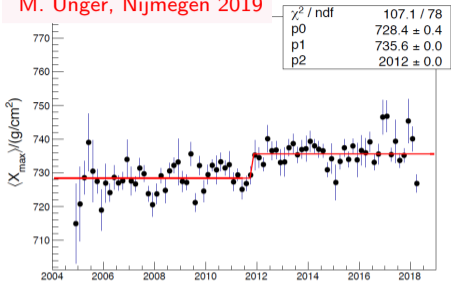
collection of plots

Alexey Yushkov

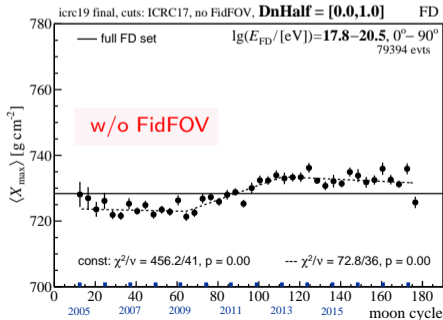
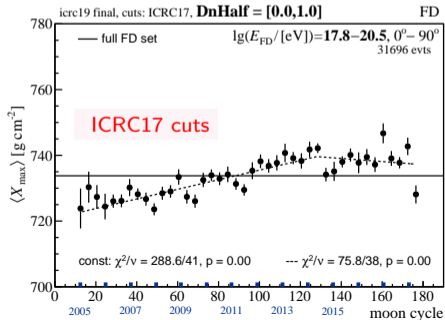
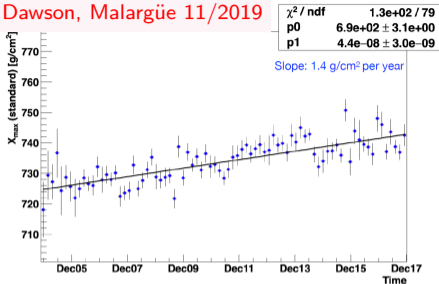
Fyzikální ústav AV ČR

$$X_{\max}, \lg(E/\text{eV}) = 17.8 - 20.5$$

M. Unger, Nijmegen 2019



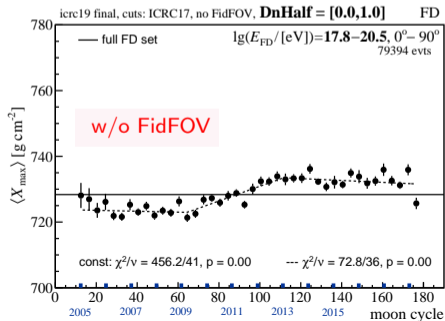
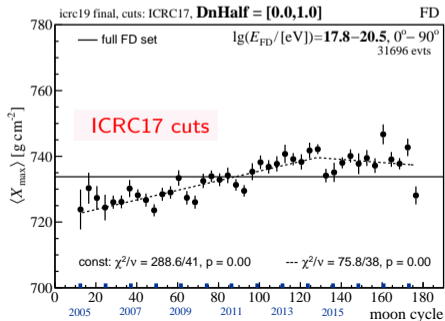
B. Dawson, Malargüe 11/2019



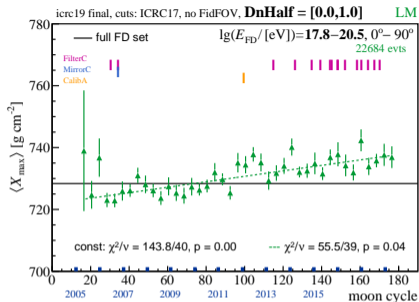
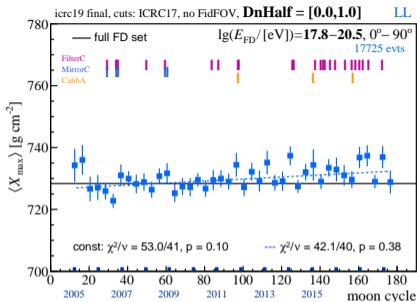
$$X_{\max}, \lg(E/\text{eV}) = 17.8 - 20.5$$

w/o FidFOV: $\langle X_{\max} \rangle$ is $\approx 5 \text{ g cm}^{-2}$ smaller, but similar trend

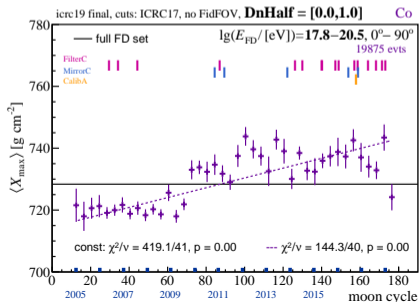
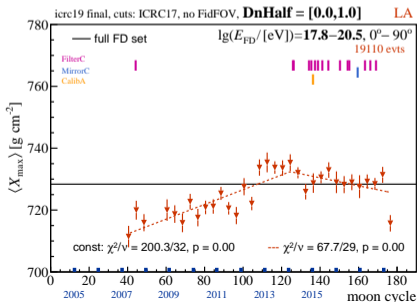
w/o FidFOV set is used further, 2.5 times more events



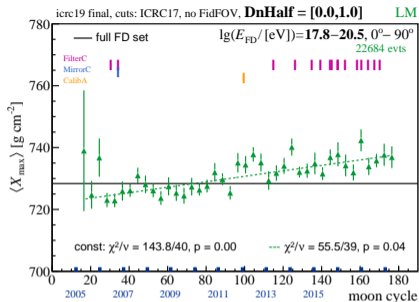
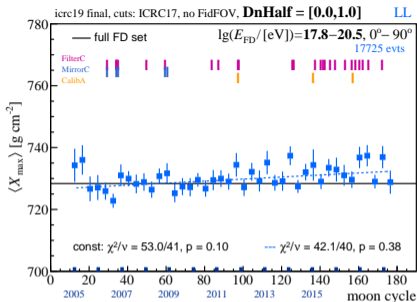
X_{\max} , individual FD sites



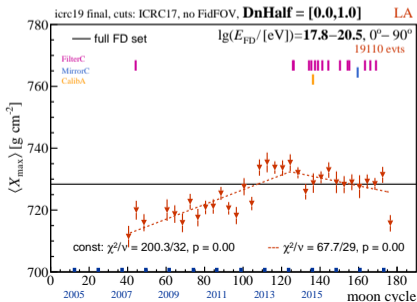
Marks:
Filter cleanings
Mirror cleanings
CaIA changes



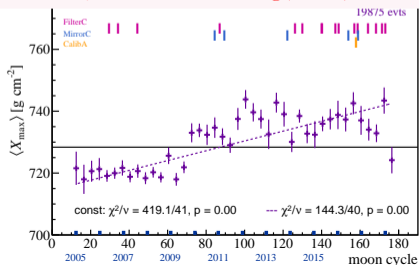
X_{\max} , individual FD sites



Marks:
Filter cleanings
Mirror cleanings
CaIA changes

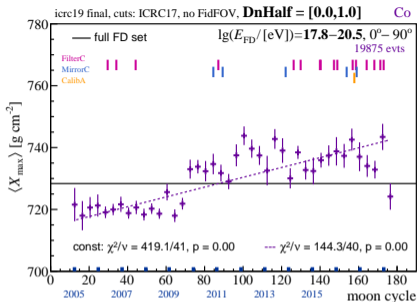


from 06/2010 events with $\lg(E/eV) = 17.8 - 18.1$ are in HeCo!

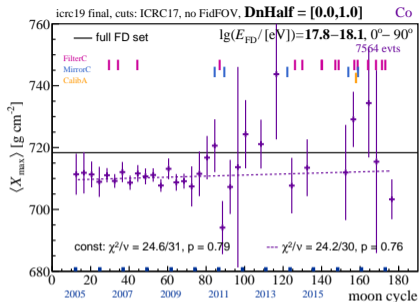


Co: individual energy bins

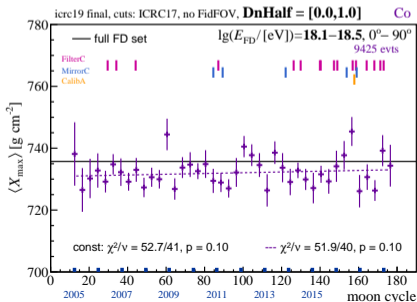
17.8 – 20.5



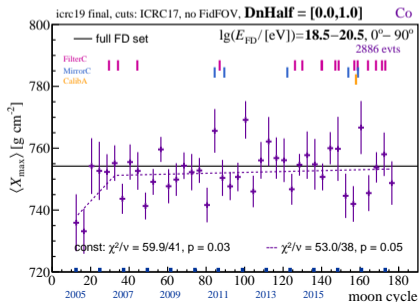
17.8 – 18.1



18.1 – 18.5

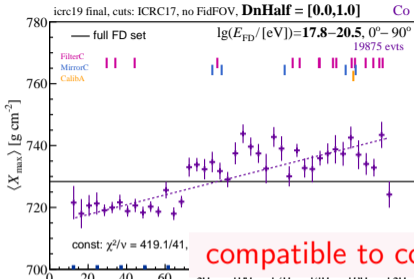


18.5 – 20.5

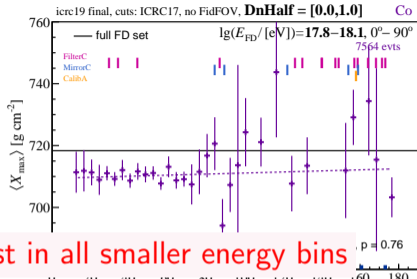


Co: individual energy bins

17.8 – 20.5



compatible to const in all smaller energy bins



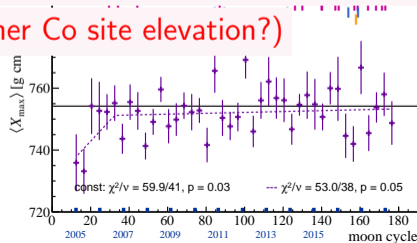
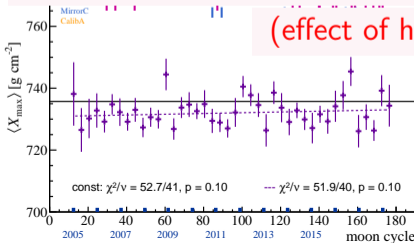
17.8 – 18.1

'jump' is just due to removal of events with $\lg(E/eV) = 17.8 - 18.1$ after 06/2010

at lower energies $\langle X_{\max} \rangle$ is systematically below the mean of the full FD

(effect of higher Co site elevation?)

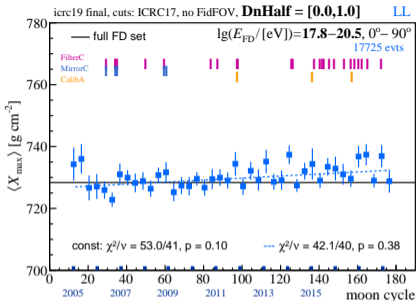
18.1 – 18.5



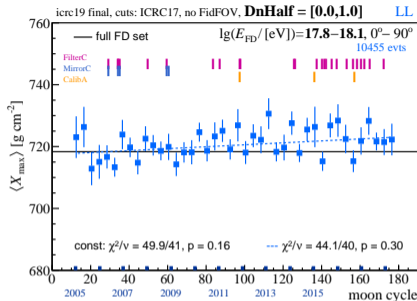
18.5 – 20.5

LL: individual energy bins

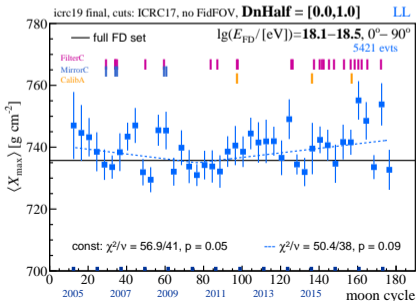
17.8 – 20.5



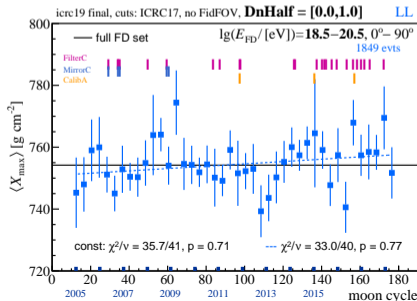
17.8 – 18.1



18.1 – 18.5

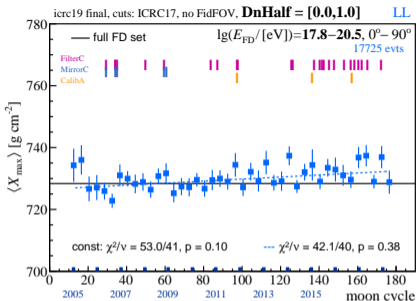


18.5 – 20.5

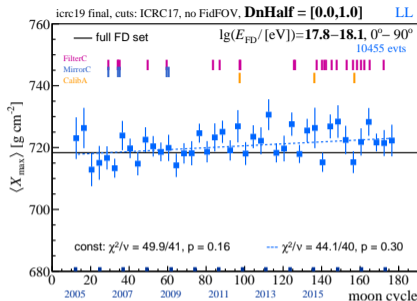


LL: individual energy bins

17.8 – 20.5

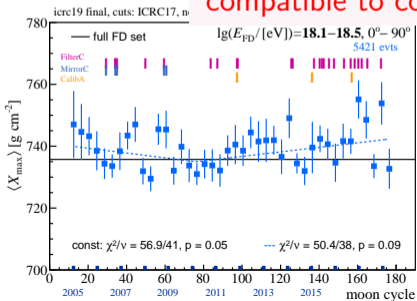


17.8 – 18.1

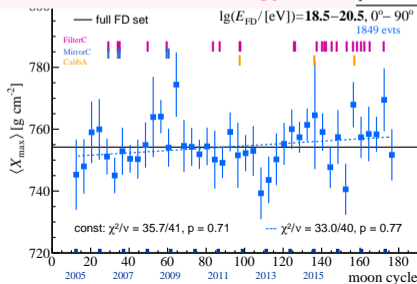


compatible to const in all smaller energy bins

18.1 – 18.5

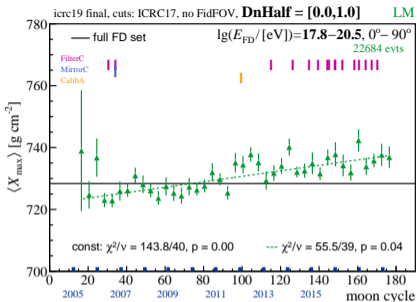


18.5 – 20.5

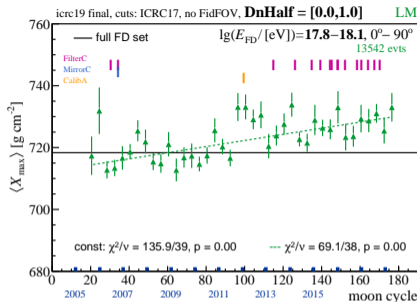


LM: individual energy bins

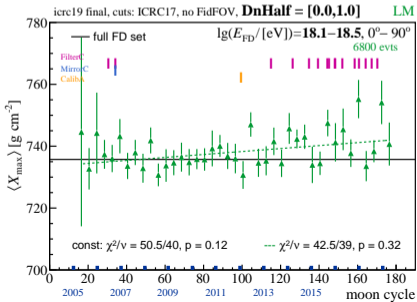
17.8 – 20.5



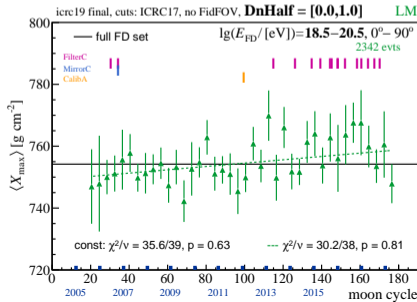
17.8 – 18.1



18.1 – 18.5

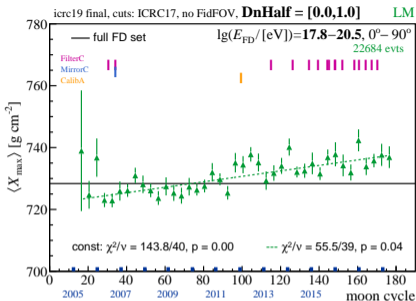


18.5 – 20.5

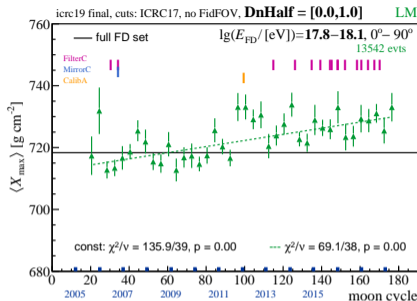


LM: individual energy bins

17.8 – 20.5

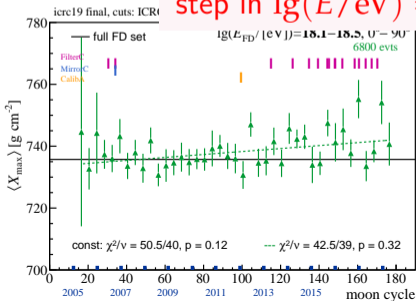


17.8 – 18.1

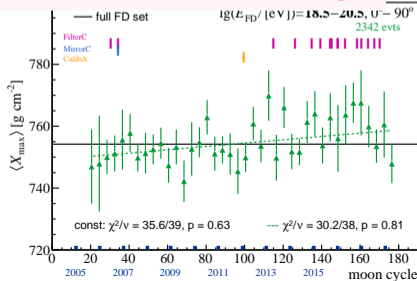


step in $\lg(E/eV) = 17.8 - 18.1$ at CalA change?

18.1 – 18.5

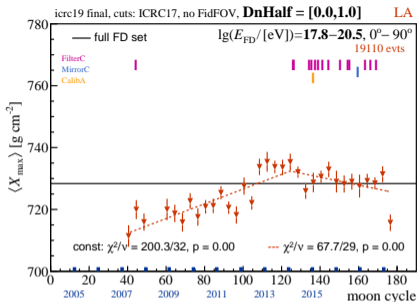


18.5 – 20.5

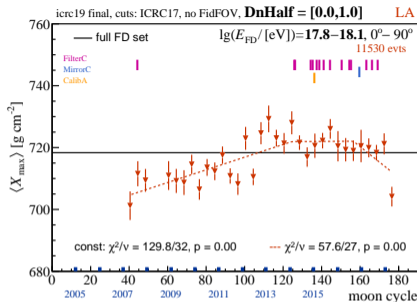


LA: individual energy bins

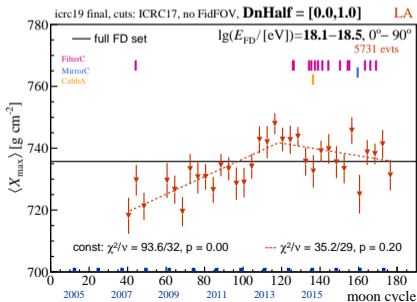
17.8 – 20.5



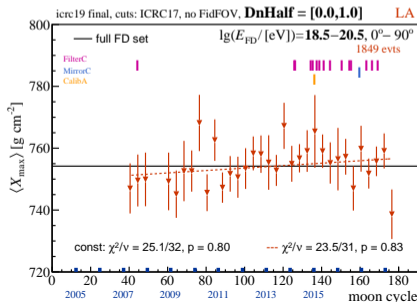
17.8 – 18.1



18.1 – 18.5

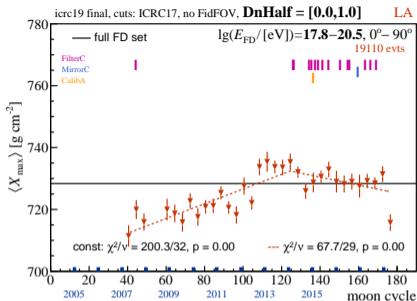


18.5 – 20.5

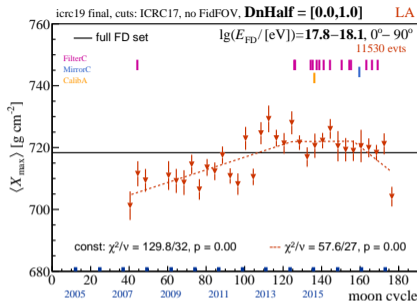


LA: individual energy bins

17.8 – 20.5

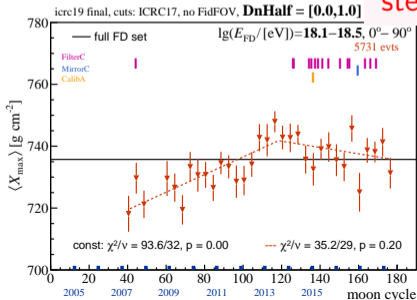


17.8 – 18.1

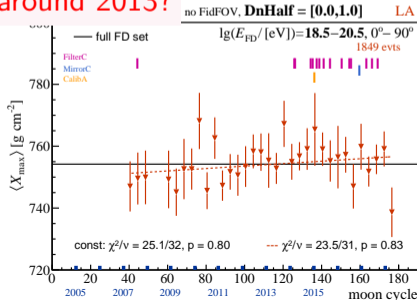


step around 2013?

18.1 – 18.5

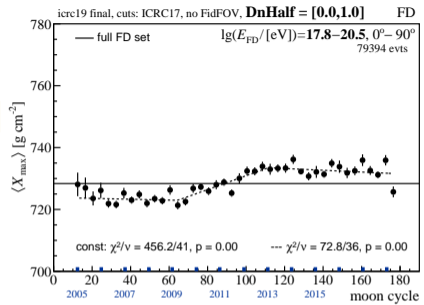


18.5 – 20.5

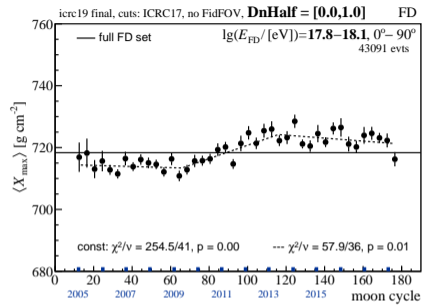


FD: individual energy bins

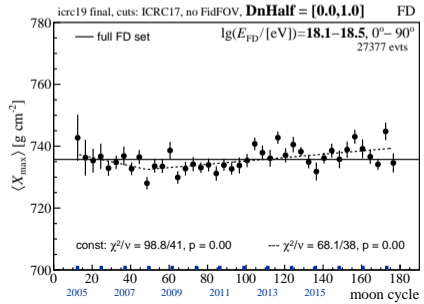
17.8 – 20.5



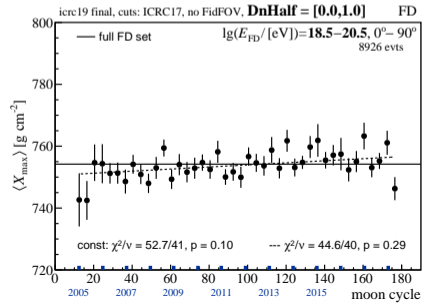
17.8 – 18.1



18.1 – 18.5

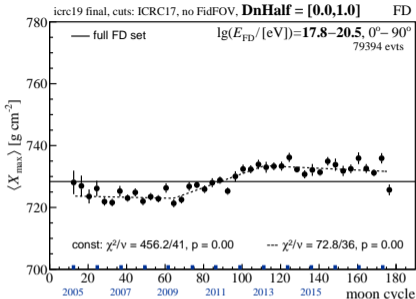


18.5 – 20.5

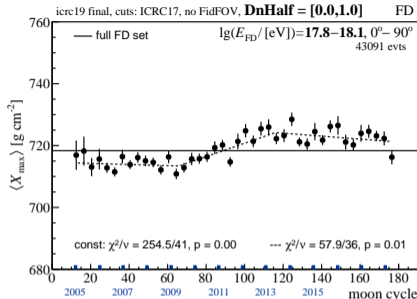


FD: individual energy bins

17.8 – 20.5

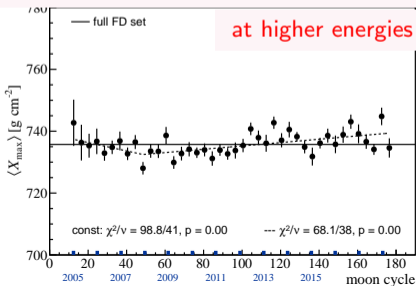


17.8 – 18.1



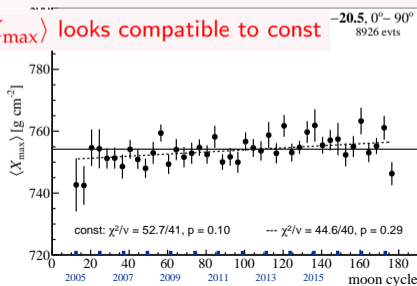
$\lg(E/eV) = 17.8 - 18.1$: step-like (?) LA & LM contributions, plus removal of Co events with lower $\langle X_{\max} \rangle$

18.1 – 18.5



at higher energies $\langle X_{\max} \rangle$ looks compatible to const

18.5 – 20.5



another one bites the dust

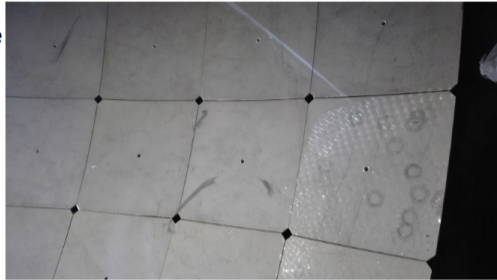


Mirror Cleaning at LL #4

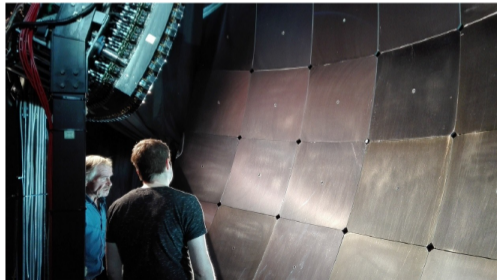


5th April 2019

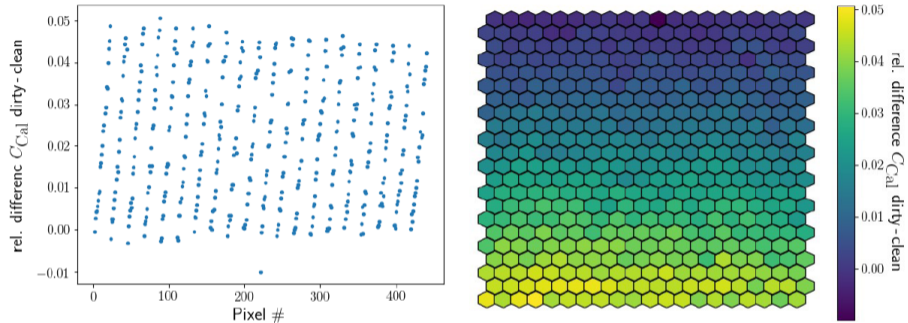
Before



After

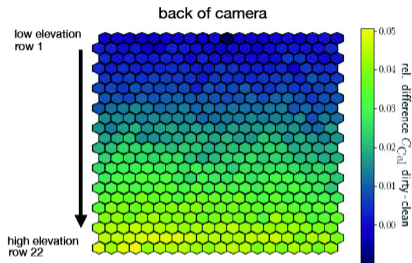
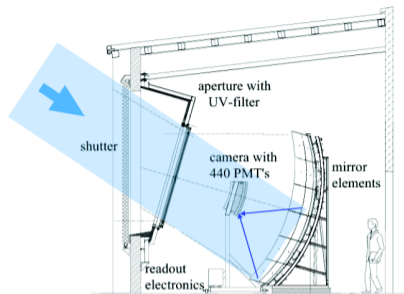


Mirror Cleaning



Up to 5% effect

Predominantly sensitive to absorption



To correct for the dust effect, impose a correction to the calibration constants, increasing their value (photons/ADC) with increasing elevation.

Imposing a universal calibration constant correction of 2% (0%) per year for row 22 (row 1) we correct the

- 1% per year drift in energy (Lorenzo, Phong)
- 1.4 g/cm² per year drift in X_{max}

Energy scale evolution

are events in cleaner (upper) mirror parts more stable?

X_{\max} evolution (though there might be none at higher energies)

are there subsets with different evolution? (systematics for Galactic plane anisotropy?)

Method

Separate events using fraction of charge deposited in top 11 camera rows

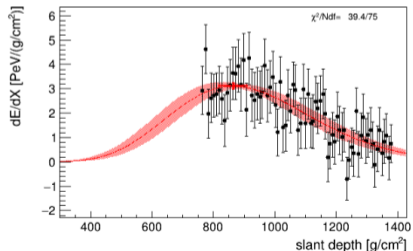
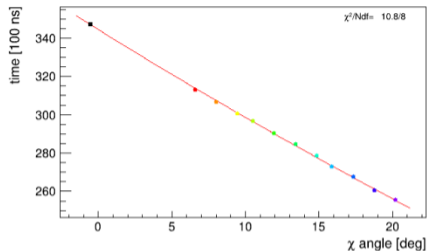
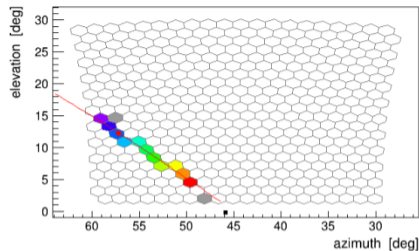
Camera top = Mirror top = Lower elevations

Variable **DnHalf: fraction of charge deposited in lower elevation**

3 classes of events with **DnHalf** $< 20\%$ (high), $40 - 60\%$ (cross), $> 80\%$ (low)

(using reconstructed pixels ePulseRecPix)

Events with $DnHalf = 1$ (fully in lower elevations)



Event Info | Pixels | PCGF

run 864, event 1056

time stamp: 827292608 s 130171826 ns
 Trigger: 'Physics - Int or L/R trigger', 'Shower Candidate'
 in Los Morados mirror 2 (in DAQ: 1 2 3 4 5 6)

geometry: hybrid, station 1005 (ToT), showerPlaneDistance = 504 m

$(\theta, \phi) = (53.1 \pm 1.1, 151.9 \pm 0.7)$ deg
 $(x, y) = (17.77 \pm 0.14, 10.00 \pm 0.11)$ km
 $R_p = 12.72 \pm 0.10$ km

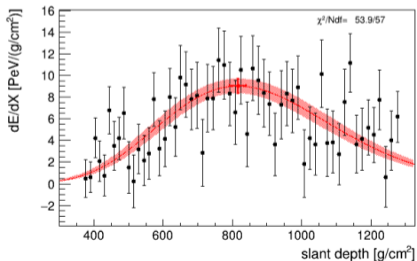
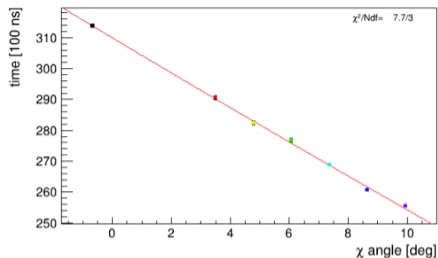
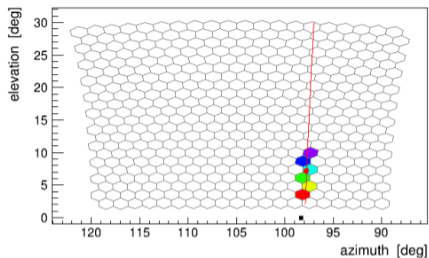
profile: 4-parameter Gaisser-Hillas (type: USP)

$E = (2.09 \pm 0.12 \pm 0.12) \times 10^{18}$ eV
 $X_{max} = 865 \pm 46$ g/cm²
 $(dE/dX)_{max} = 3.11 \pm 0.13$ PeV/(g/cm²)
 $(\lambda, X_0, L) = (58, -31, 229 \pm 8)$ g/cm², $R = 0.26 \pm 0.04$
 Cherenkov-fraction = 20%, $mva = 37$ deg.

databases:

Mie attenuation: measured ($h < 9.6$ km, VAOD at 3km: 0.04)
 LIDAR: no data ; CloudCam: max(ζ)=(0/0)% (elev>5.5**); CloudMap: max=0%
 molecular profile: GDAS; time correction: good

Events with $DnHalf = 1$ (fully in lower elevations)



Event Info | Pixels | PCGF

run 910, event 1001

time stamp: 801967288 s 772785674 ns
 Trigger: 'Physics - Int or L/R trigger', 'Shower Candidate'
 in Los Leones mirror 4 (in DAQ: 1 2 3 4 5 6)

geometry: hybrid, station 817 (ToT), showerPlaneDistance = 361 m

$(\theta, \phi) = (54.1 \pm 0.7, 66.8 \pm 2.1)$ deg
 $(x, y) = (-6.78 \pm 0.23, 0.9 \pm 0.23)$ km
 $R_0 = 17.78 \pm 0.35$ km

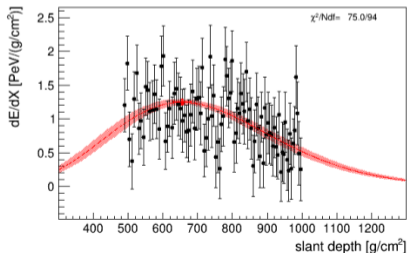
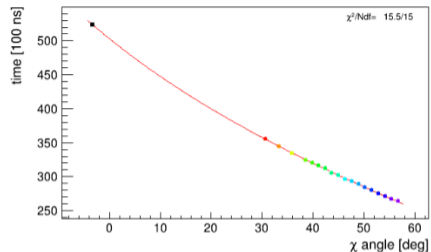
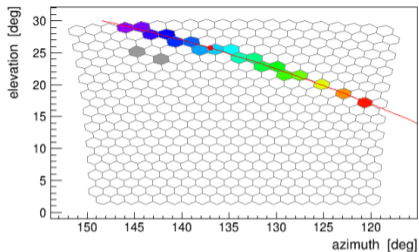
profile: 4-parameter Gaisser-Hillas (type: USP)

$E = (6.48 \pm 0.44 \pm 0.41) \times 10^{18}$ eV
 $X_{max} = 816 \pm 25$ g/cm²
 $(dE/dX)_{max} = 9.01 \pm 0.67$ PeV/(g/cm²)
 $(\lambda, X_0, L) = (62, -162, 247 \pm 10)$ g/cm², $R = 0.25 \pm 0.05$
 Cherenkov-fraction = 19%, $mva = 23$ deg.

databases:

Mie attenuation: measured ($h < 16.4$ km, VAOD at 3km: 0.00)
 LIDAR: no data ; CloudCam: $\max(\zeta) = (0/0)\%$ (elev > 5.5^{km}); CloudMap: no data
 molecular profile: GDAS; time correction: good

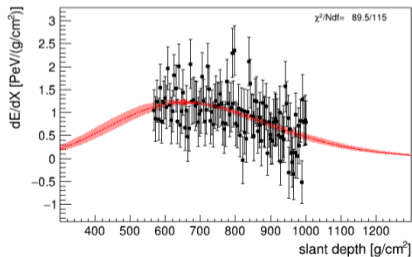
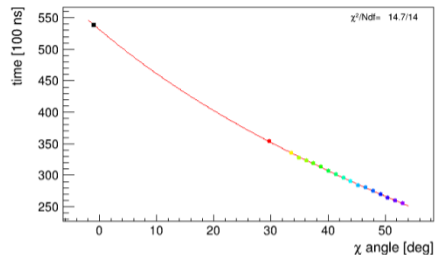
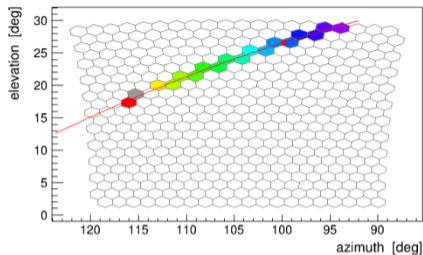
Events with $DnHalf = 0$ (fully in higher elevations)



Event Info | Pixels | PCGF

run 1469, event 21792
time stamp: 866089299 s 367026472 ns
Trigger: 'Physics - Int or L/R trigger', 'Shower Candidate'
in Los Morados mirror 5 (in DAQ: 1 2 3 4 5 6)
geometry: hybrid, station 726 (ToT), showerPlaneDistance = 603 m
 $(\theta, \phi) = (54.4 \pm 0.5, 247.2 \pm 1.2)$ deg
 $(x, y) = (12.57 \pm 0.07, -0.38 \pm 0.13)$ km
 $R_p = 10.04 \pm 0.04$ km
profile: 4-parameter Gaisser-Hillas (type: USP)
 $E = (8.73 \pm 0.46 \pm 0.44) \times 10^{17}$ eV
 $X_{max} = 655 \pm 25$ g/cm²
 $(dE/dX)_{max} = 1.25 \pm 0.04$ PeV/(g/cm²)
 $(\lambda, X_0, L) = (61, -236, 232 \pm 9)$ g/cm², $R = 0.26 \pm 0.05$
Cherenkov-fraction = 8%, $mva = 36$ deg.
databases:
 Mie attenuation: measured ($h < 8.4$ km, VAOD at 3km: 0.00)
 LIDAR: $h(\text{cloud}) = 100.0$ km, 0%; CloudCam: $\max(\square) = (0/0)\%$ (elev > 5.5^{ns}); CloudMap: no data
 molecular profile: GDAS; time correction: good

Events with $DnHalf = 0$ (fully in higher elevations)



Event Info | Pixels | PCGF

run 1519, event 5849

time stamp: 868583370 s 879036288 ns
 Trigger: 'Physics - Int or L/R trigger', 'Shower Candidate'
 in Los Morados mirror 4 (in DAQ: 1 2 4 5 6)

geometry: hybrid, station 541 (ToT), showerPlaneDistance = 338 m

(θ , ϕ) = (53.1 \pm 0.6, 97.7 \pm 1.4) deg
 (x, y) = (11.25 \pm 0.08, -8.56 \pm 0.16) km
 R_p = 10.92 \pm 0.05 km

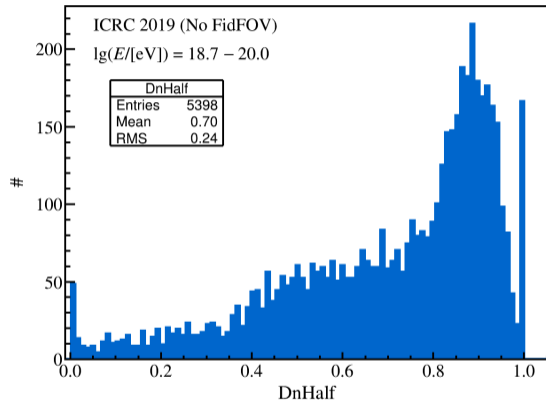
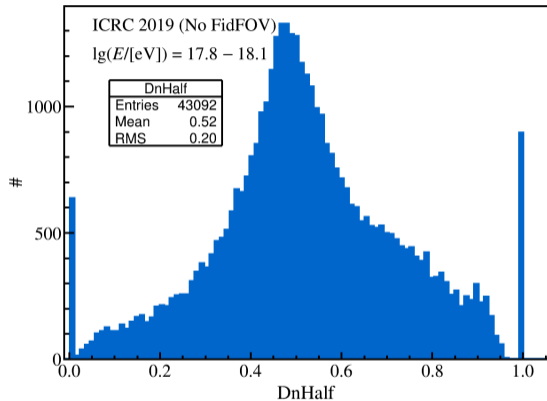
profile: 4-parameter Gaisser-Hillas (type: USP)

E = (8.23 \pm 0.53 \pm 0.42) $\times 10^{17}$ eV
 X_{max} = 655 \pm 31 g/cm²
 (dE/dX)_{max} = 1.22 \pm 0.05 PeV/(g/cm²)
 (λ , X_0 , L) = (57, -231, 225 \pm 10) g/cm², R = 0.25 \pm 0.05
 Cherenkov-fraction = 7%, mva=48 deg.

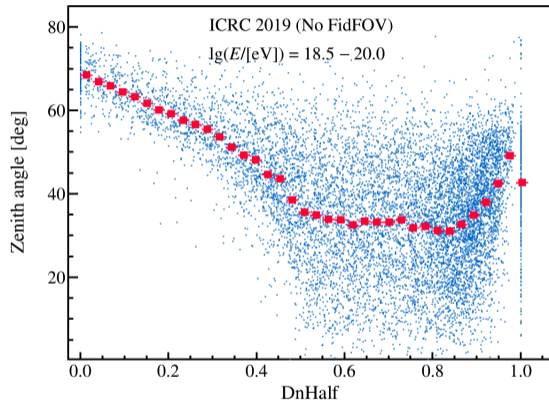
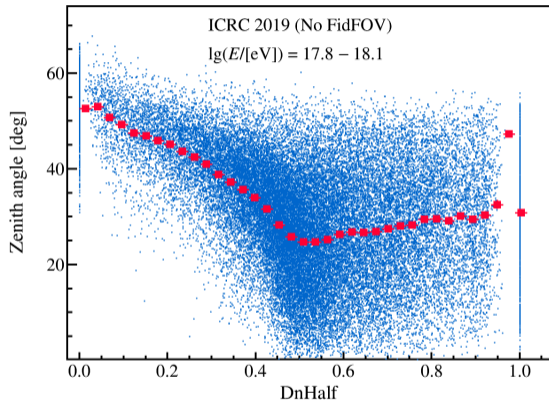
databases:

Mie attenuation: measured (h<10.1 km, VAOD at 3km: 0.00)
 LIDAR: h(cloud)=10.2 km, 79%; CloudCam: max(ζ)=(100/0)% (elev>5.5 $^{\circ}$); CloudMap: no data
 molecular profile: GDAS; time correction: good

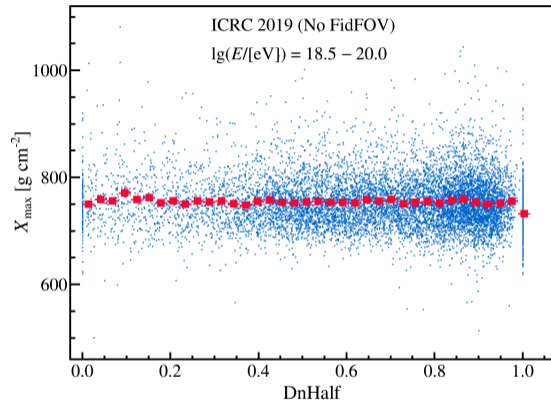
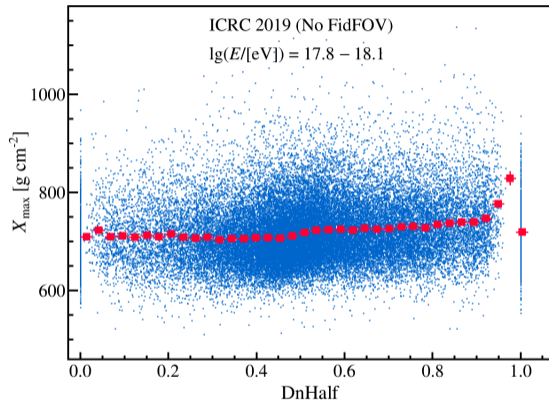
DnHalf distributions



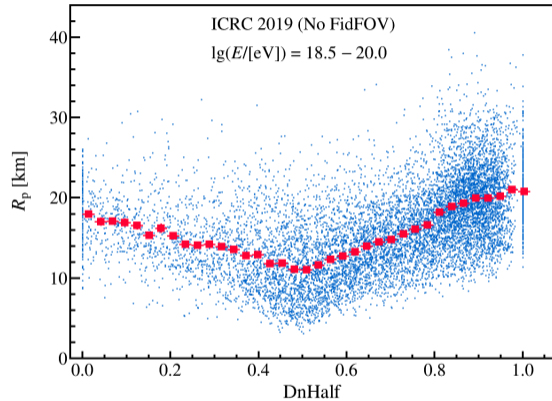
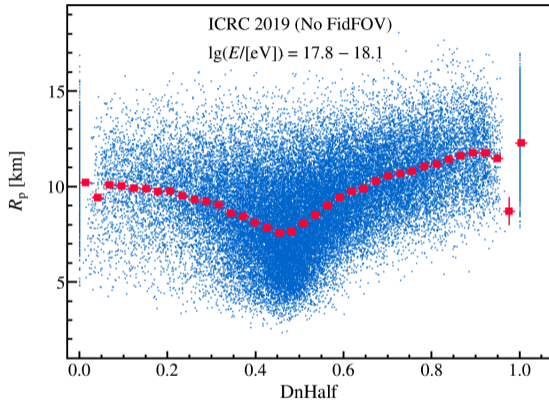
DnHalf distributions



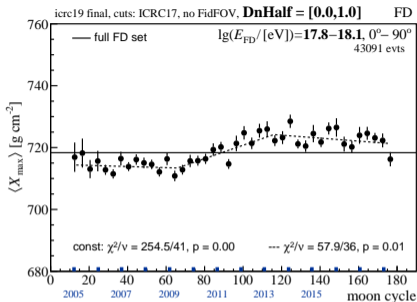
DnHalf distributions



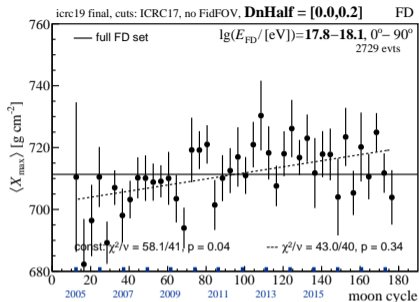
DnHalf distributions



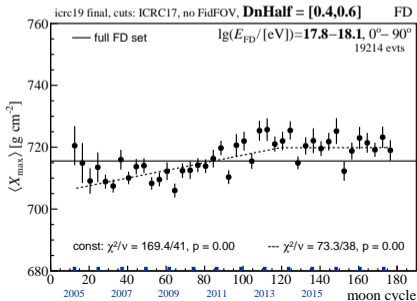
full



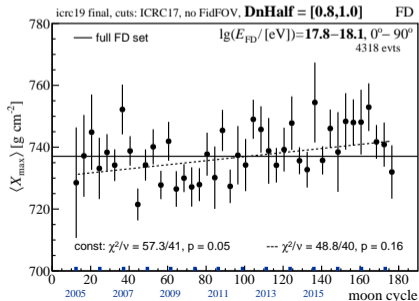
high



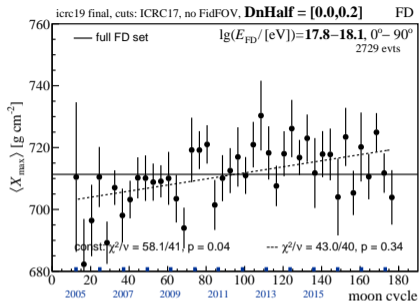
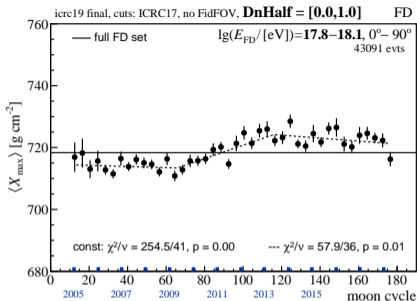
cross



low



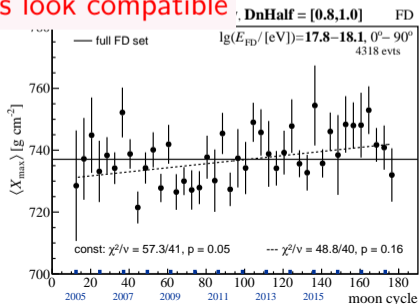
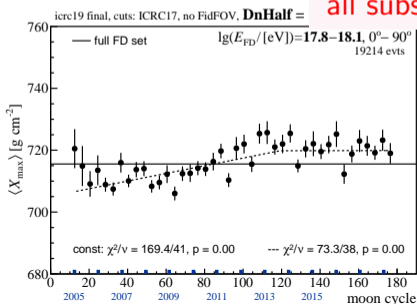
full



high

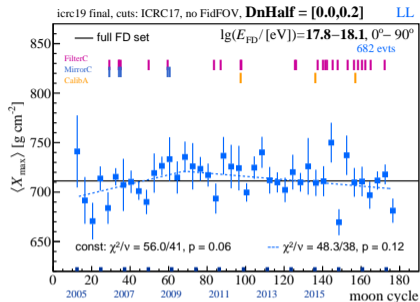
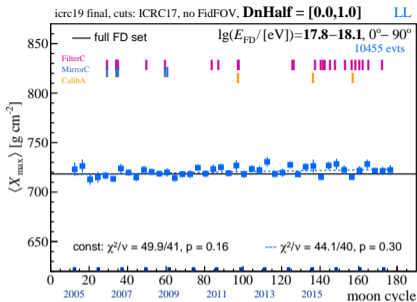
all subsets look compatible

cross



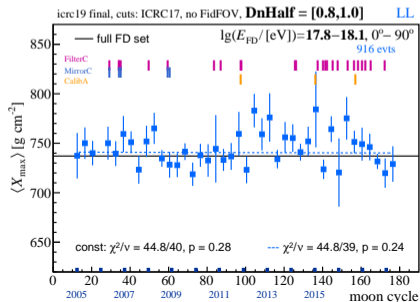
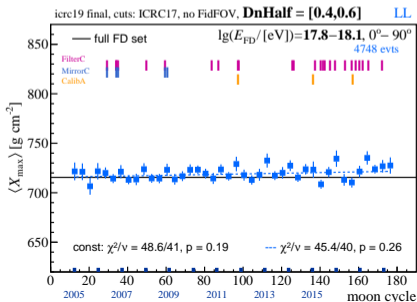
low

full



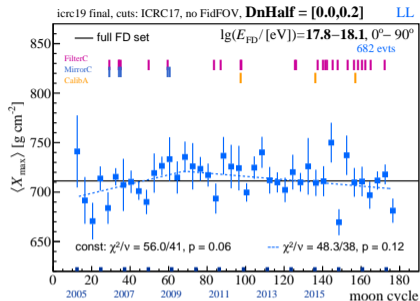
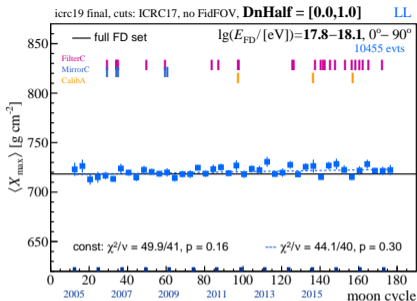
high

cross



low

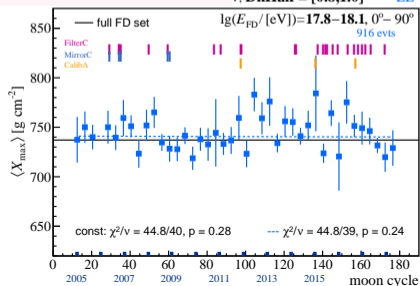
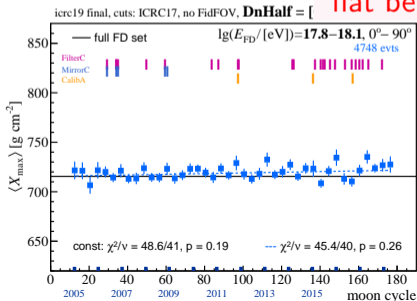
full



high

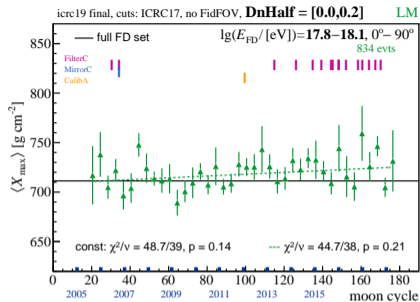
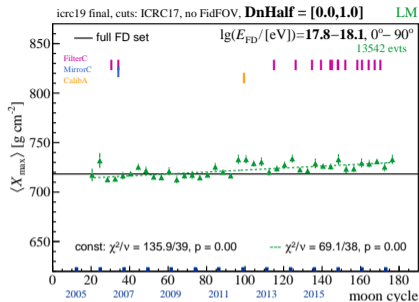
flat behavior in all subsets

cross



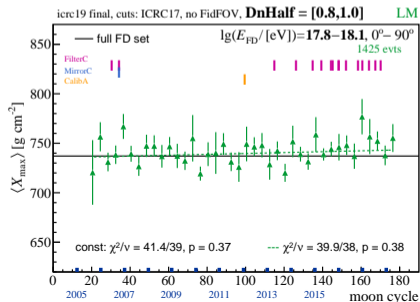
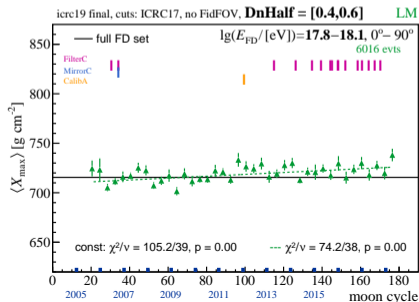
low

full



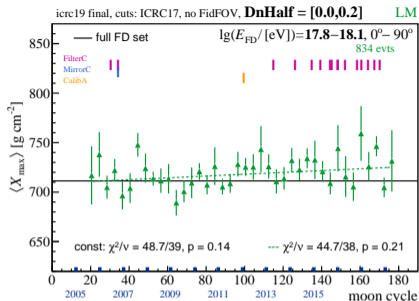
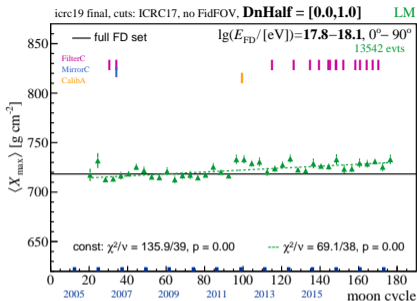
high

cross



low

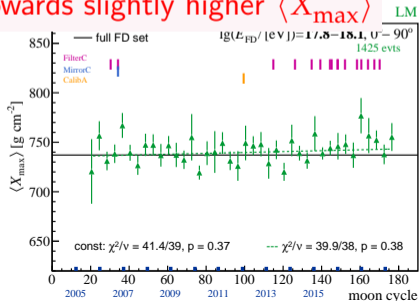
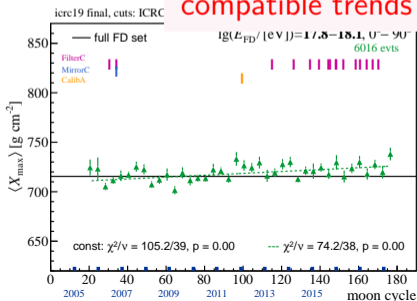
full



high

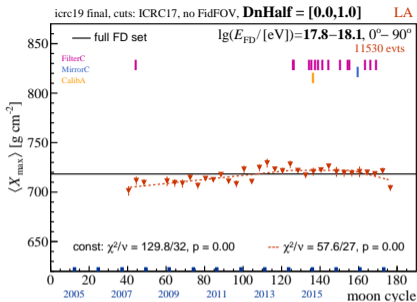
compatible trends towards slightly higher $\langle X_{max} \rangle$

cross

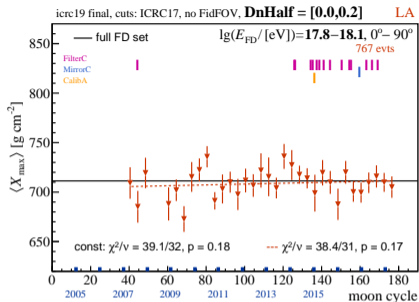


low

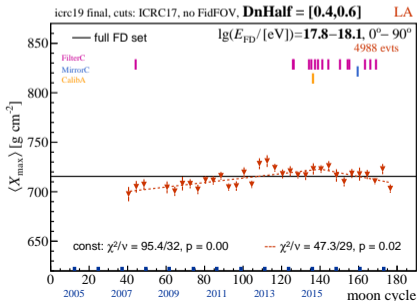
full



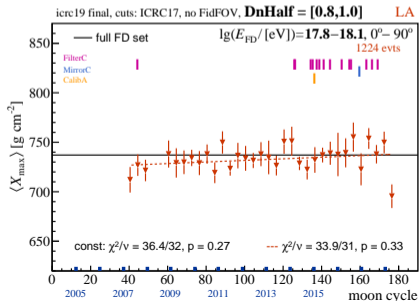
high



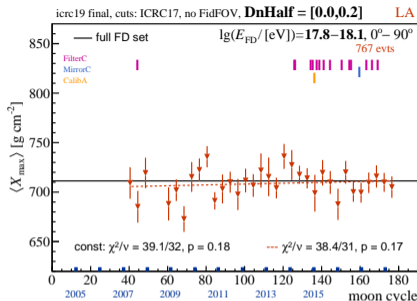
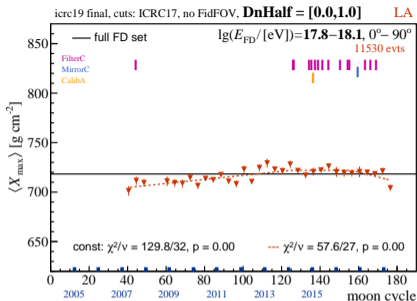
cross



low



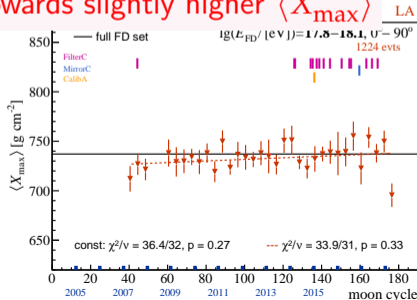
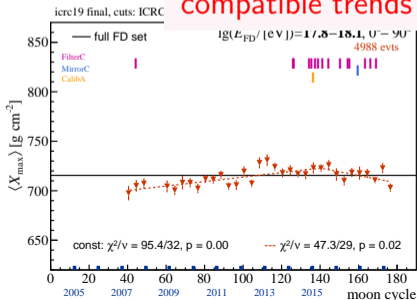
full



high

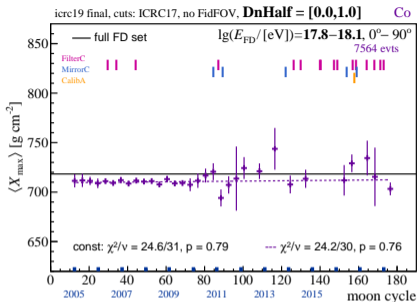
compatible trends towards slightly higher $\langle X_{max} \rangle$

cross

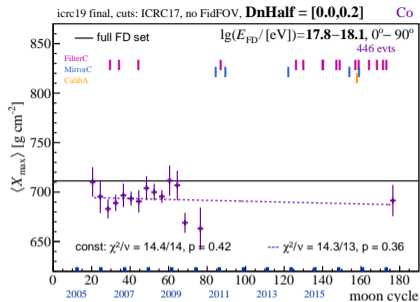


low

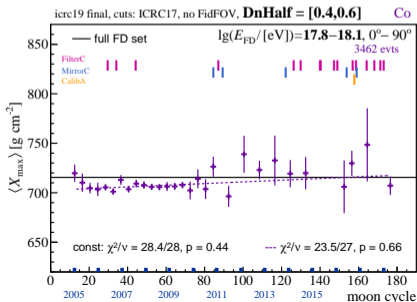
full



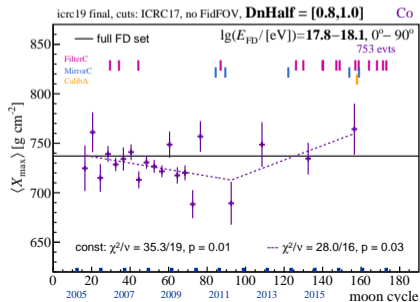
high



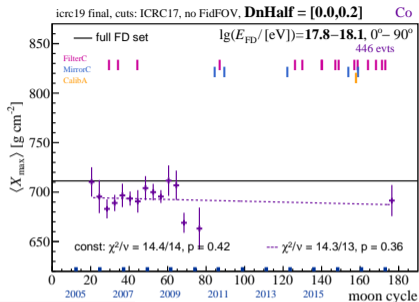
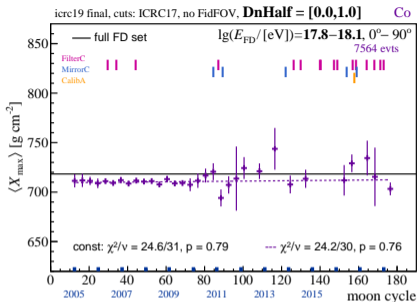
cross



low

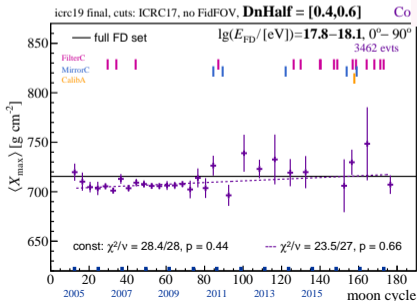


full

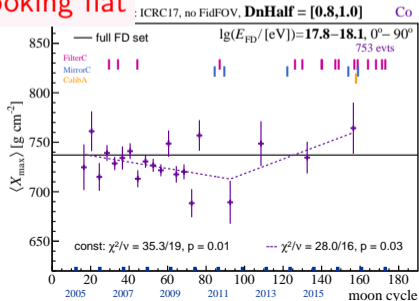


high

cross

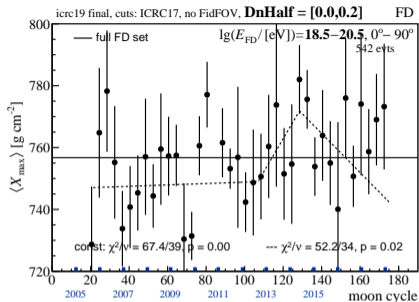
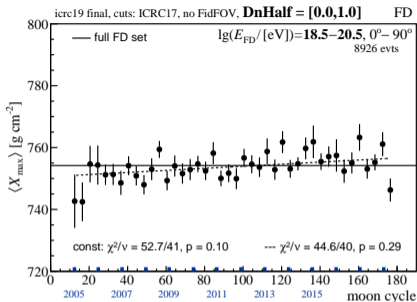


looking flat



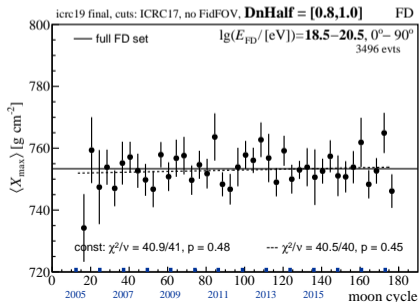
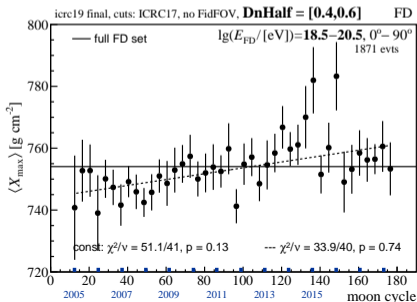
low

full



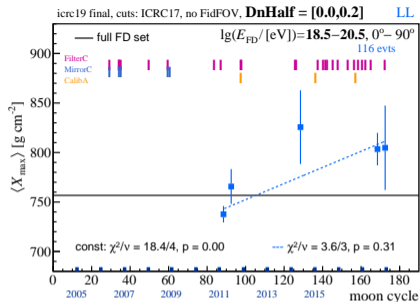
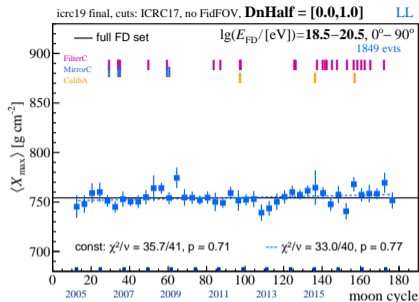
high

cross



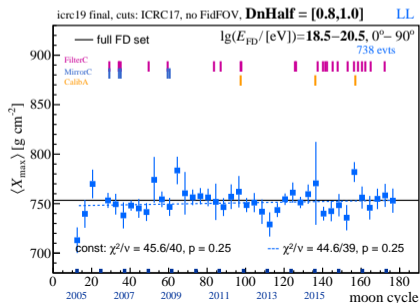
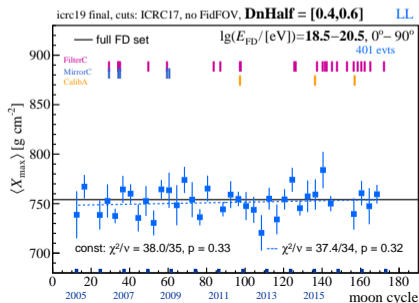
low

full



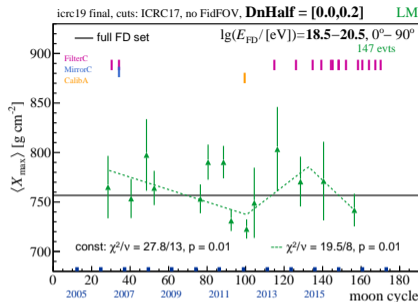
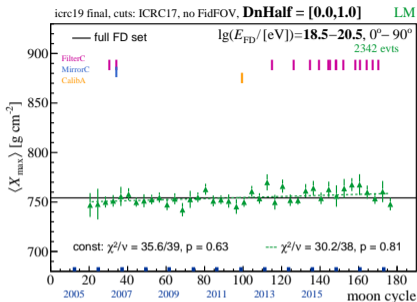
high

cross



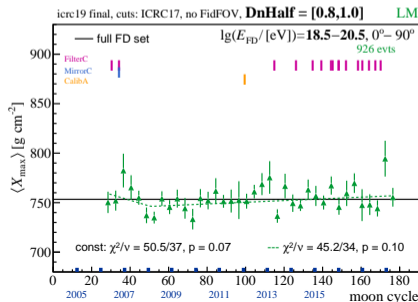
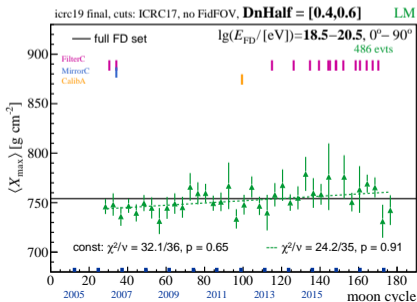
low

full



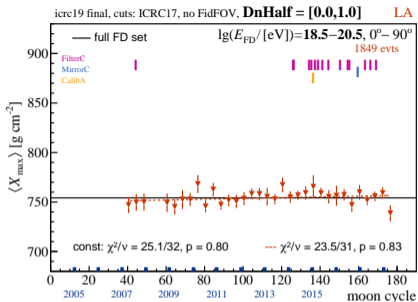
high

cross

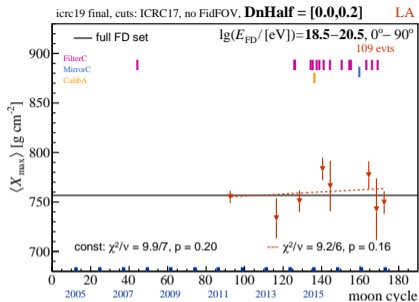


low

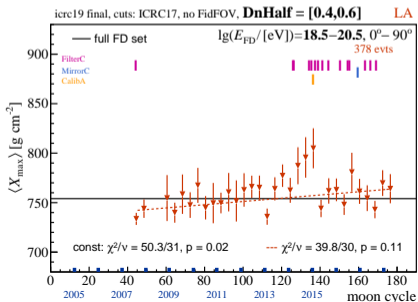
full



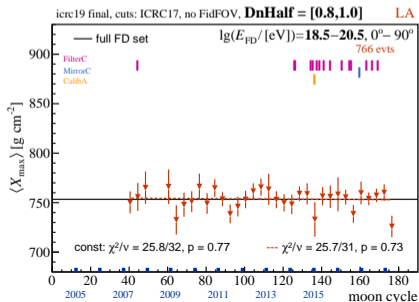
high



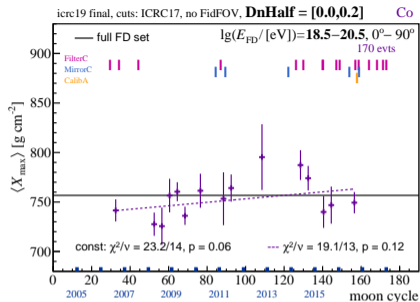
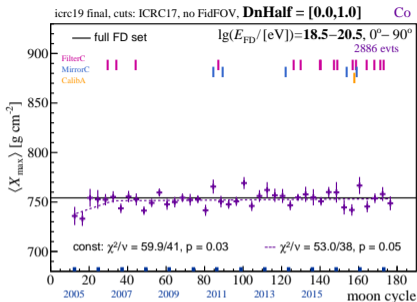
cross



low

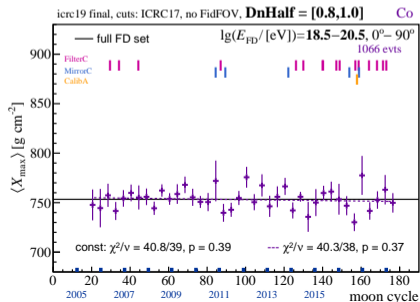
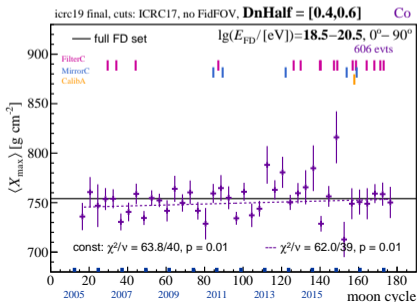


full



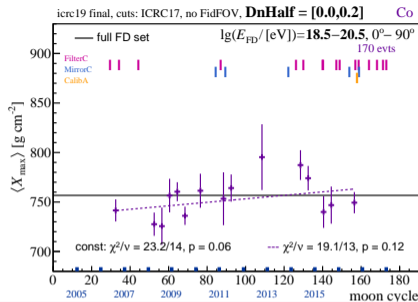
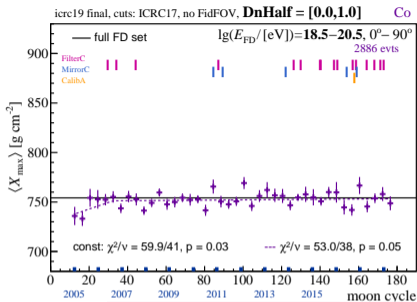
high

cross



low

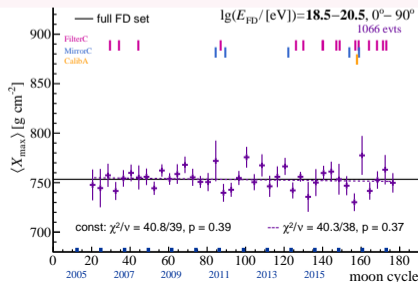
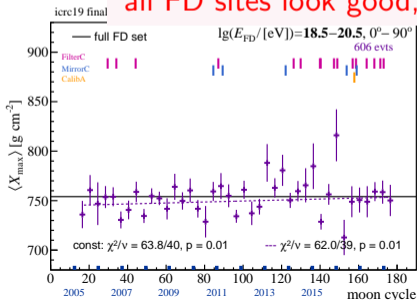
full



high

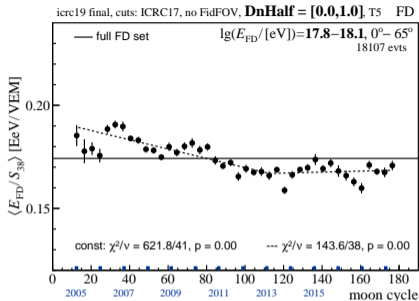
all FD sites look good, no troubles for Galactic anisotropy

cross

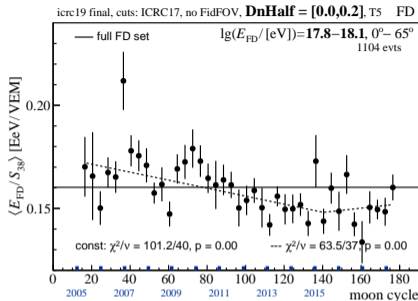


low

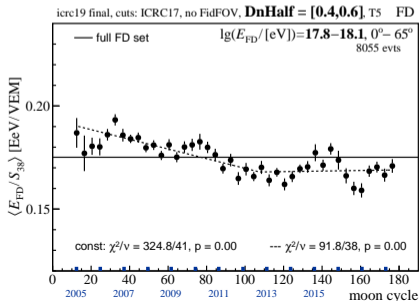
full



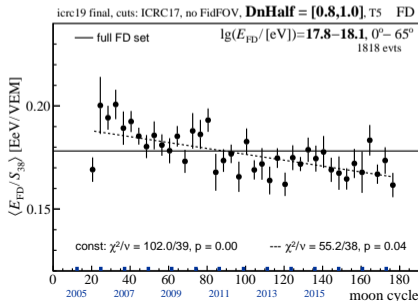
high



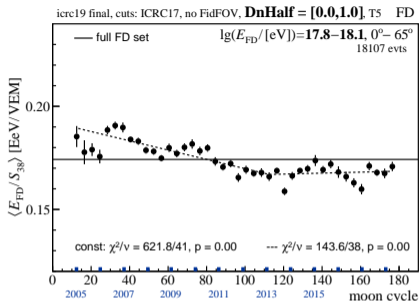
cross



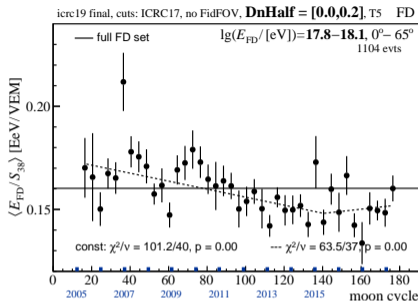
low



full

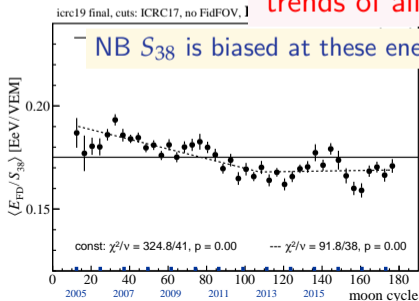


high

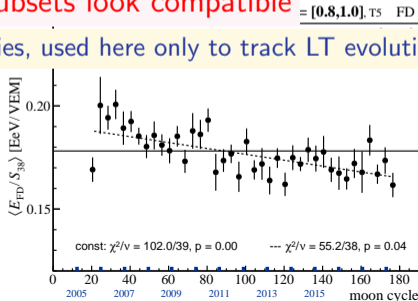


trends of all subsets look compatible

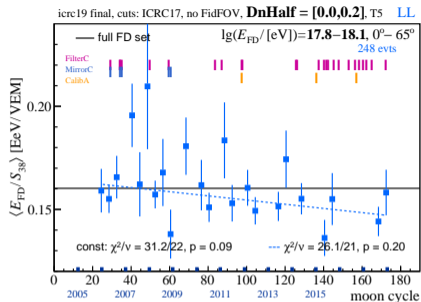
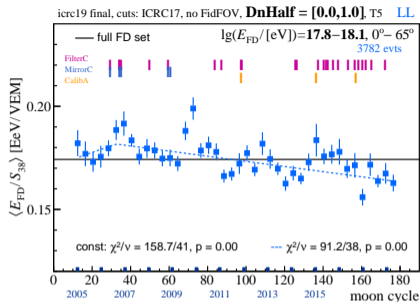
cross



low

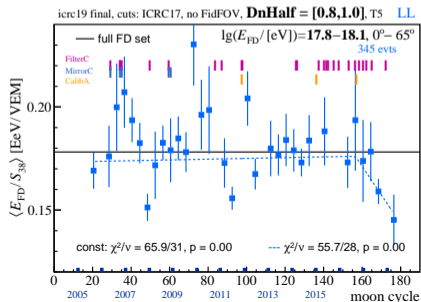
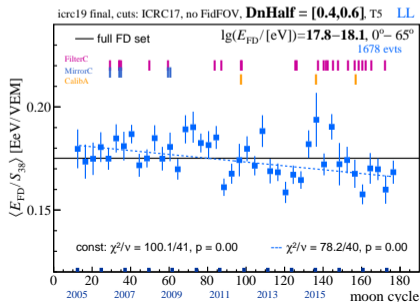


full



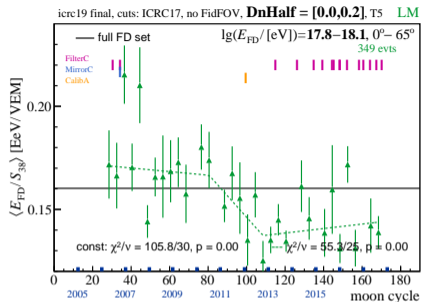
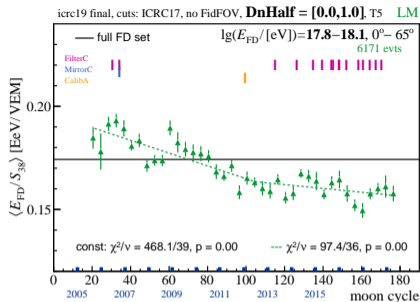
high

cross



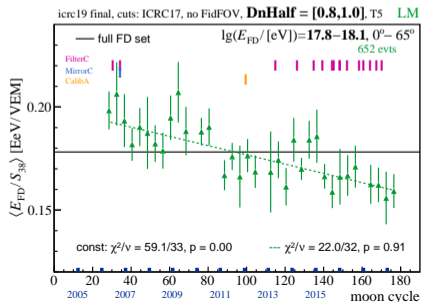
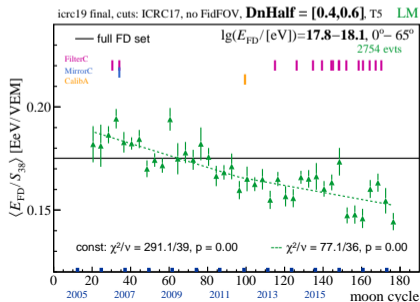
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full



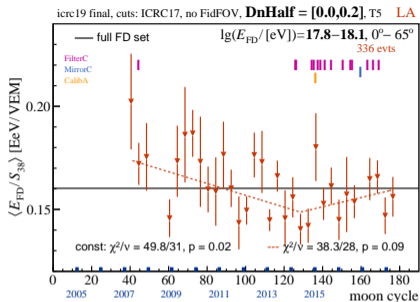
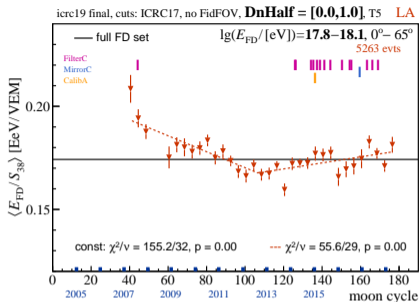
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cross



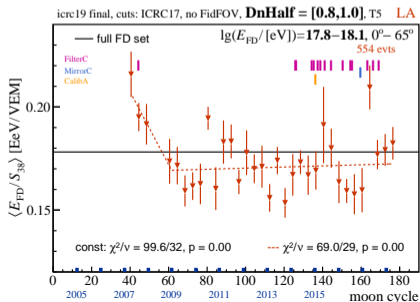
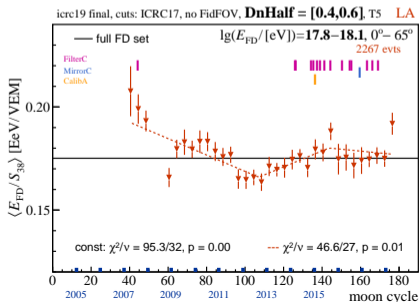
low

full



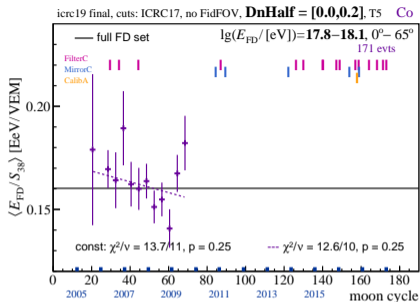
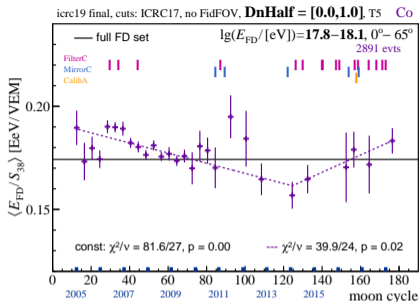
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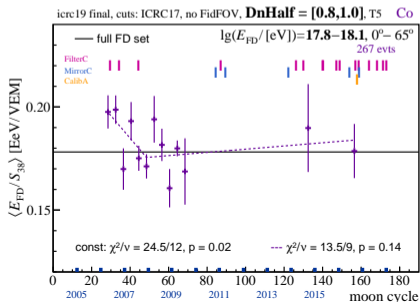
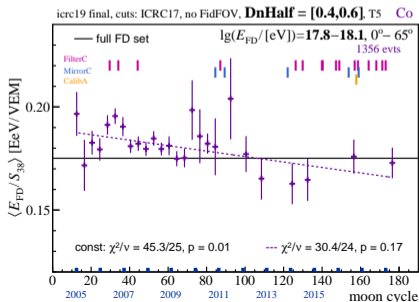
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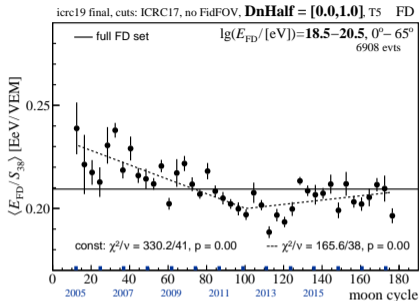
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cross

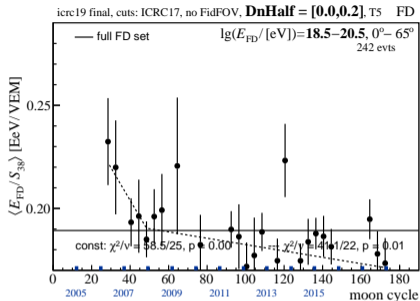


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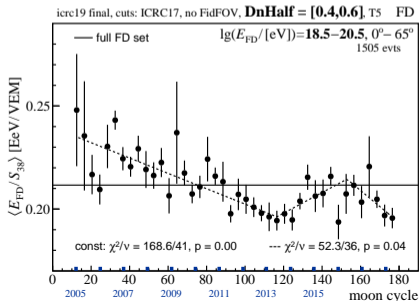
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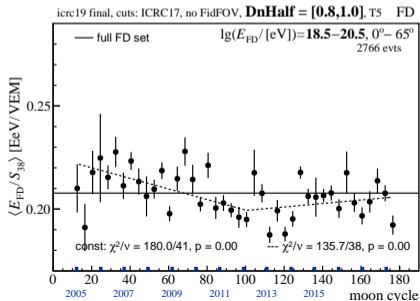
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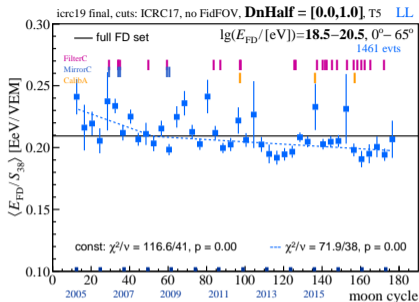
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low

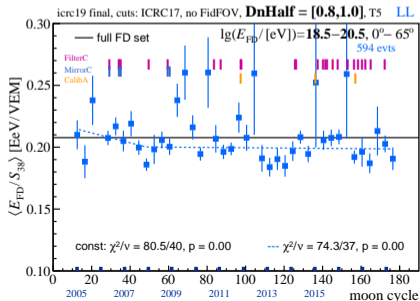
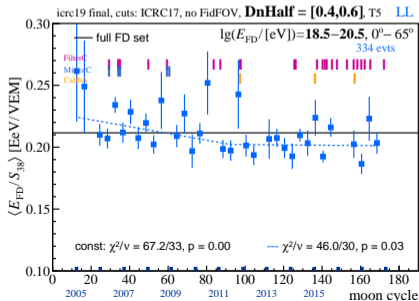


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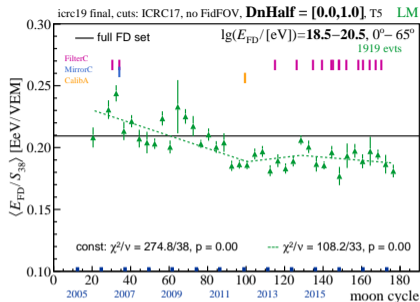
high

cross



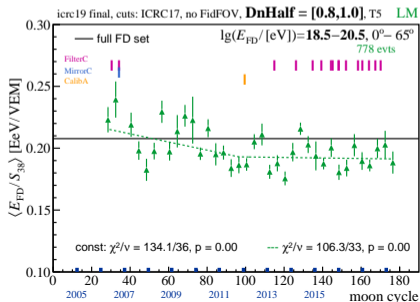
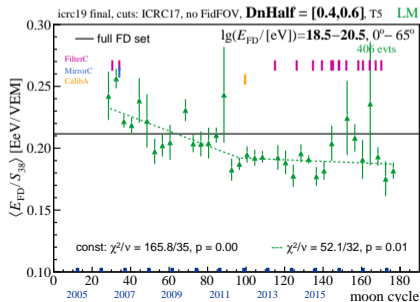
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full



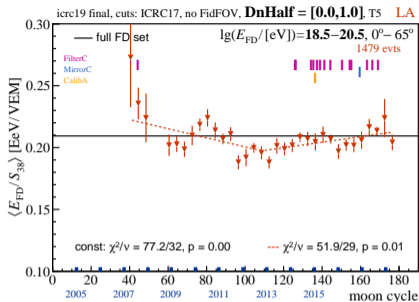
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cross



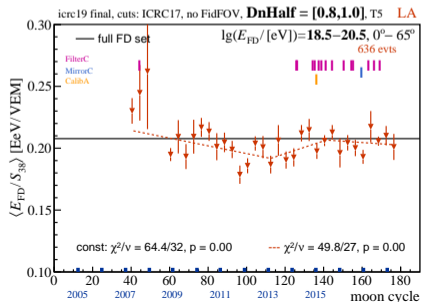
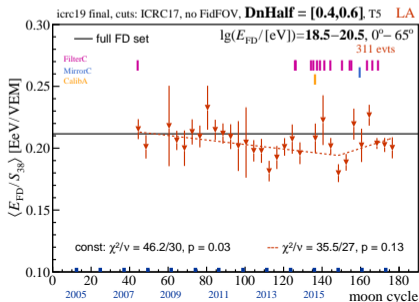
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full



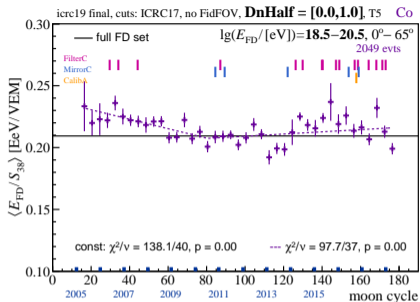
high

cross



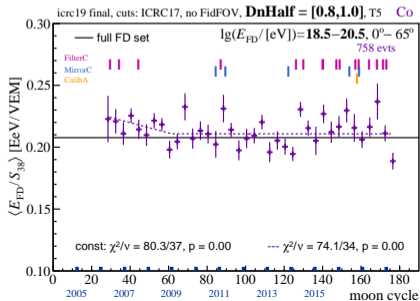
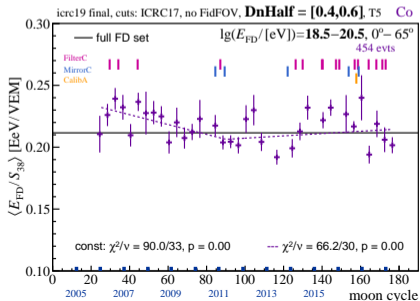
low

full



high

cross



low

$\langle X_{\max} \rangle$ evolution

stronger trends in $\lg(E/\text{eV}) = 17.8 - 18.1$ range

(possibly overlay of step-like changes at different FD sites at different times and Co removal)

at higher energies compatible to const (and also to small increase) for all FD sites/elevation subsets

Energy scale evolution

no indications of more stable ('cleaner') subsets

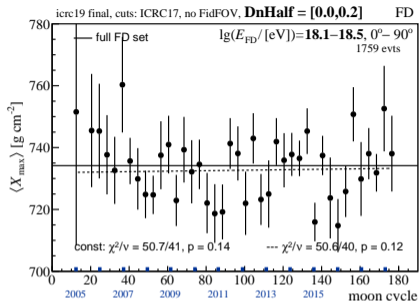
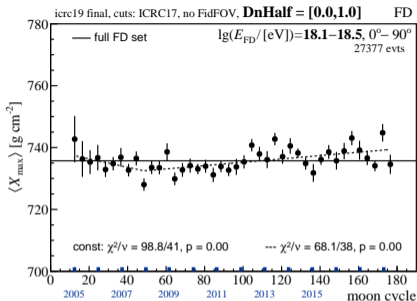
elevation subsets seem to have similar absolute scale

(comparison is applicable only to $\lg(E/\text{eV}) = 18.5 - 20.5$, for lower energies S_{38} is biased)

possibly the dust effects are too smooth (and mild) for such a coarse analysis

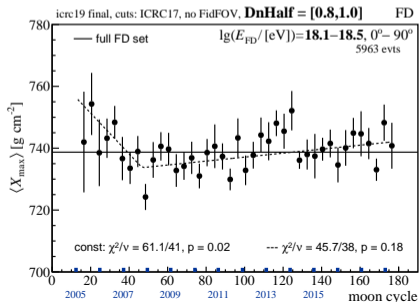
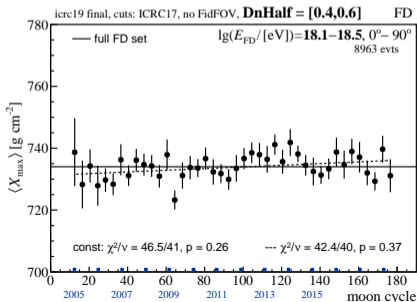
plots for $\lg(E/\text{eV}) = 18.1 - 18.5$

full



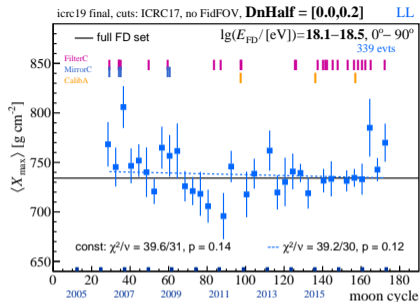
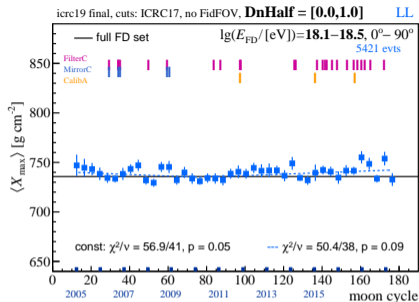
high

cross



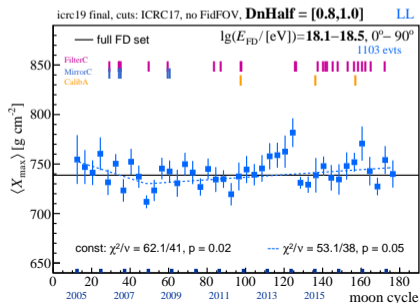
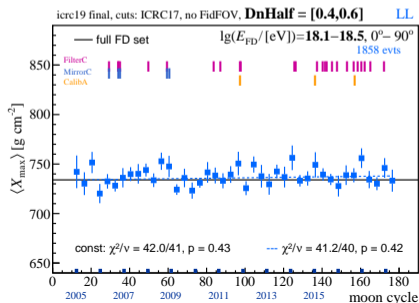
low

full



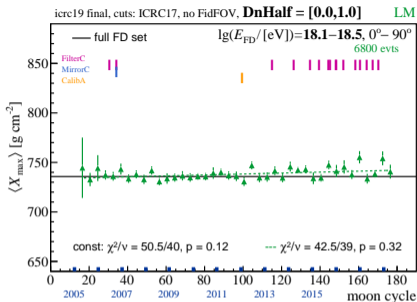
high

cross

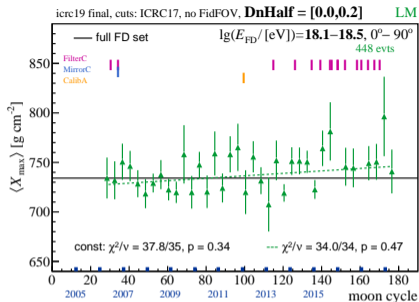


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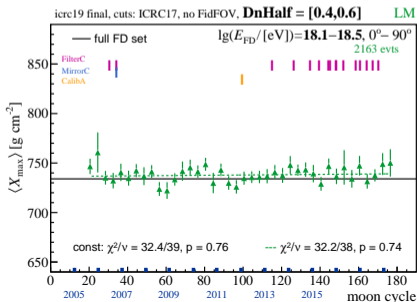
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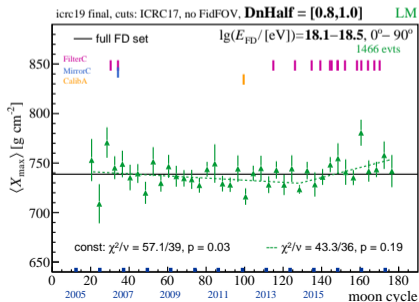
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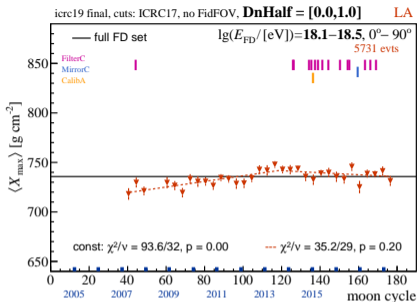
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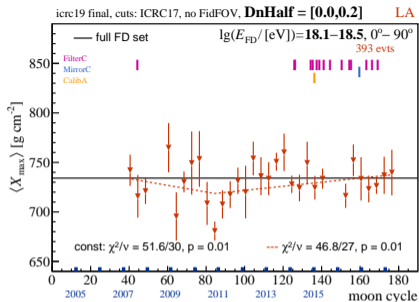
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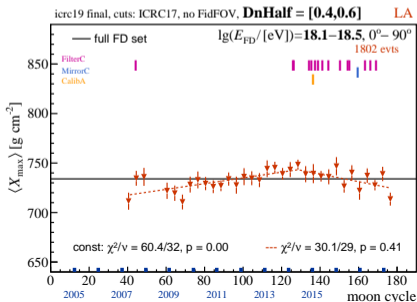
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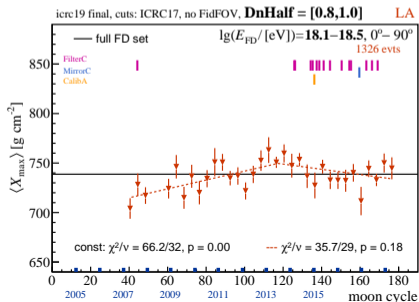
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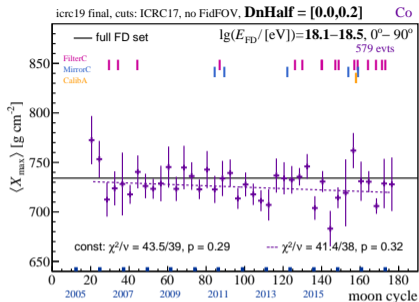
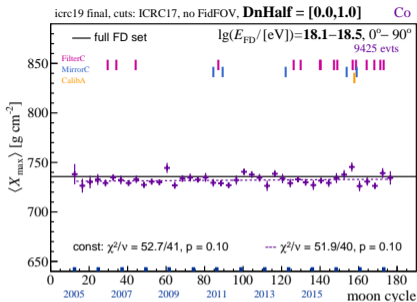
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low

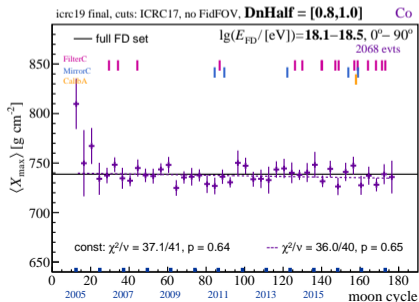
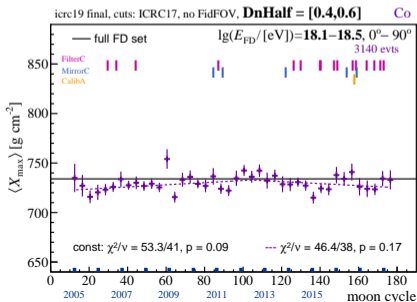


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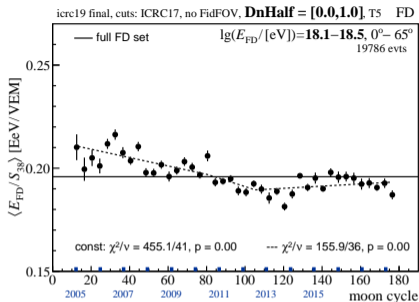
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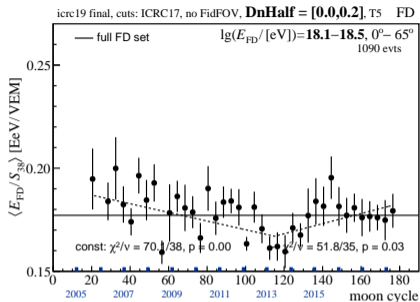


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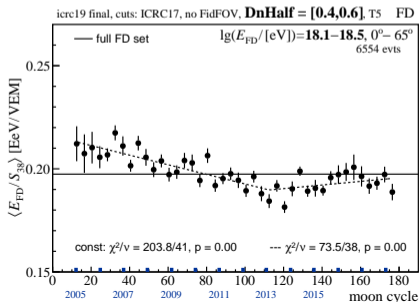
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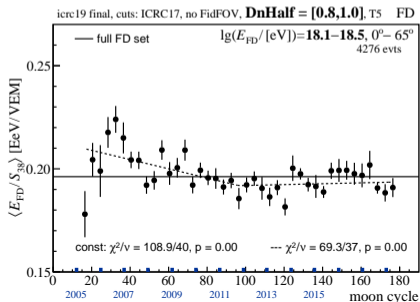
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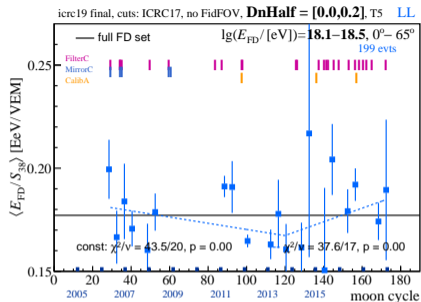
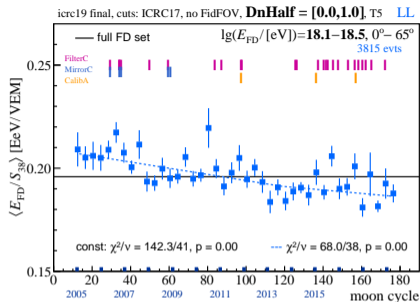
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low

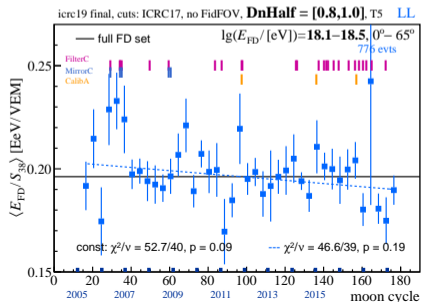
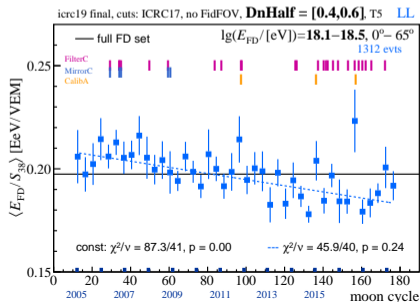


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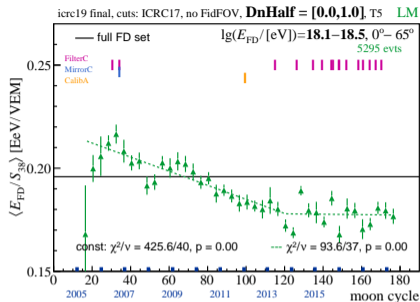
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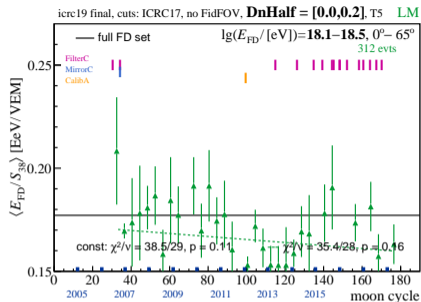


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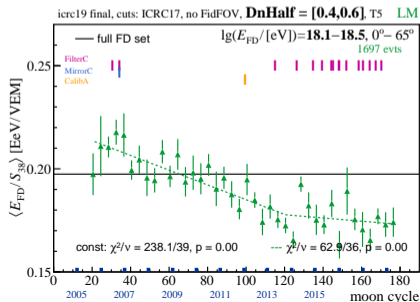
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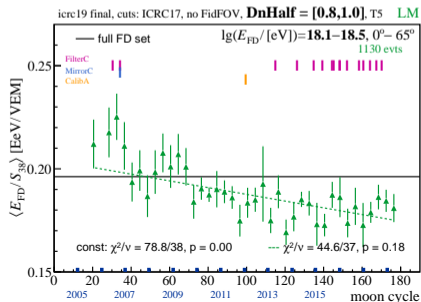
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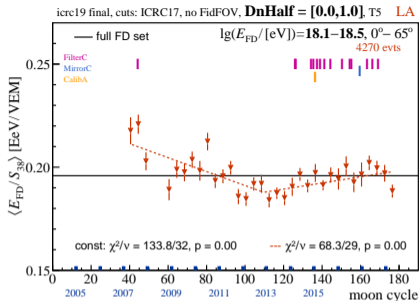
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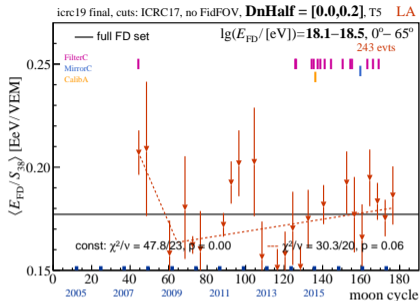
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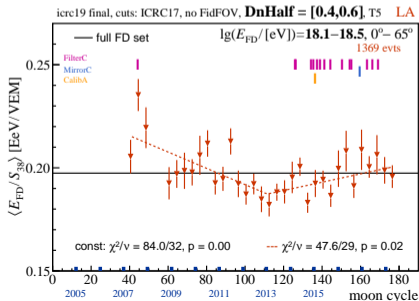
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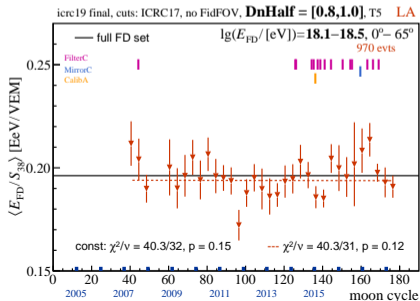
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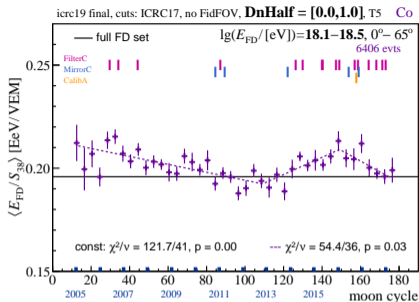
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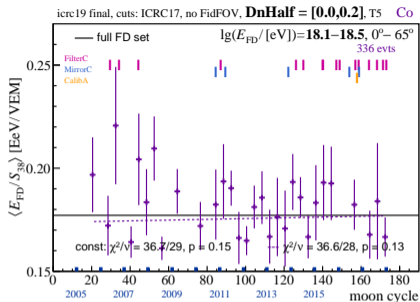
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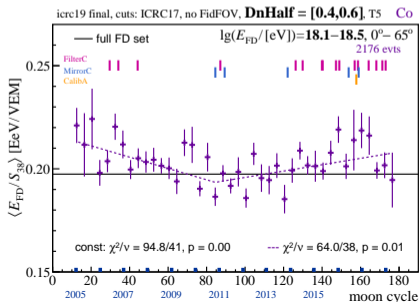
full



high



cross



low

