

Quantum baryon number fluctuations in subsystems of a hot relativistic fermionic gas

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Motivation:

The Joint Institute for Computational Fundamental Science

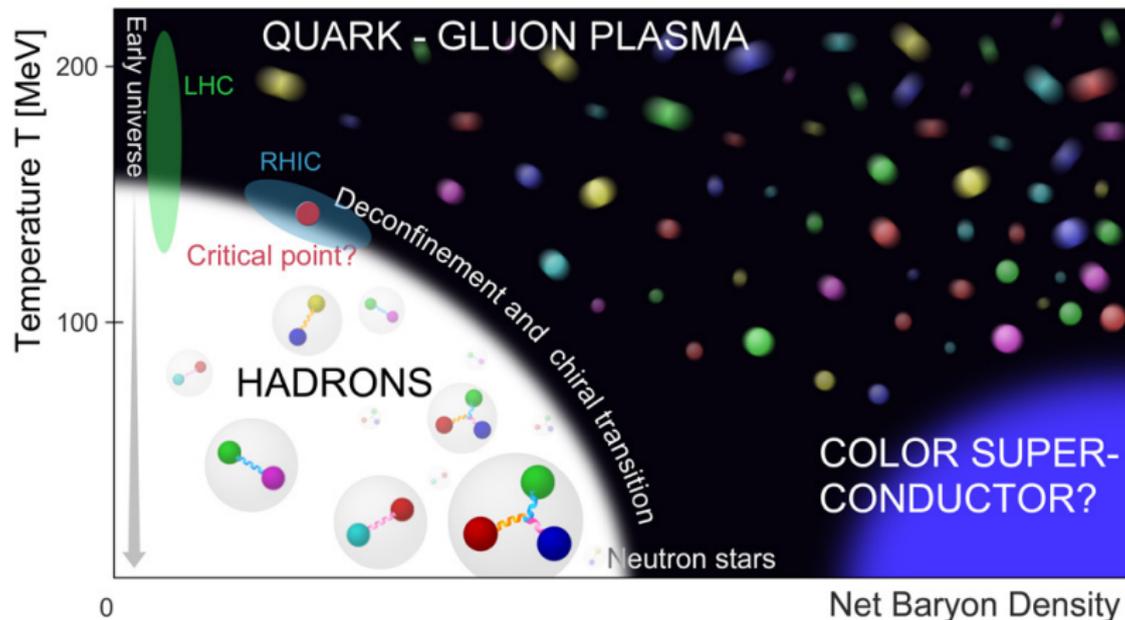


Figure: Theoretical diagram based on Lattice QCD simulations depicting expected quark-gluon phase transition.

Motivation:

Fluctuations of various physical quantities play a very important role in all fields of physics, as they reveal the information about

- possible phase transitions
- formation of structures in the Early Universe
- dissipative phenomena

Most common fluctuations we deal with are those arising from quantum uncertainty relation or those present in thermodynamic systems.

Motivation:

- We discuss **fluctuations of the baryon number density** in a hot and dense relativistic gas of fermions.
- Our analysis is relevant for **relativistic heavy-ion physics, in particular**, in the context of the beam energy scan (BES).
- Hunt for the conjectured **critical endpoint in the QCD phase diagram** has triggered vast theoretical and experimental studies of many fluctuation observables.
- Study of fluctuations of baryon number might provide an **excellent opportunity to study the critical phenomena**.

Basic concepts and definitions:

- We consider the fluctuation of the baryon number in the subsystem S_a of the thermodynamic system S_V described by the grand canonical ensemble characterized by the temperature (T) and the baryon chemical potential (μ).
- The volume V of the larger system S_V is larger than the characteristic volume of the subsystem S_a .
- We derive a compact formula that defines quantum fluctuations of the baryon number operator in subsystems of a hot and dense relativistic gas.
- Then we apply this formula to get physical insights into situations expected in relativistic heavy-ion collisions.

Basic concepts and definitions:

A **quantum field operator** for spin- $\frac{1}{2}$ particle has the standard form:

$$\psi(t, \mathbf{x}) = \sum_r \int \frac{d^3k}{(2\pi)^3 \sqrt{2\omega_{\mathbf{k}}}} \left(U_r(\mathbf{k}) a_r(\mathbf{k}) e^{-ik \cdot x} + V_r(\mathbf{k}) b_r^\dagger(\mathbf{k}) e^{ik \cdot x} \right),$$

where $a_r(\mathbf{k})$ and $b_r^\dagger(\mathbf{k})$ are annihilation and creation operators for particles and antiparticles, respectively, satisfying the canonical commutation relations $\{a_r(\mathbf{k}), a_s^\dagger(\mathbf{k}')\} = (2\pi)^3 \delta_{rs} \delta^{(3)}(\mathbf{k} - \mathbf{k}')$ and $\{b_r(\mathbf{k}), b_s^\dagger(\mathbf{k}')\} = (2\pi)^3 \delta_{rs} \delta^{(3)}(\mathbf{k} - \mathbf{k}')$, whereas $\omega_{\mathbf{k}} = \sqrt{\mathbf{k}^2 + m^2}$ is the energy of a particle.

Basic concepts and definitions:

To perform **thermal averaging**, it is sufficient to know the expectation values of the products of two and four creation and/or annihilation operators

$$\begin{aligned}\langle a_r^\dagger(\mathbf{k}) a_s(\mathbf{k}') \rangle &= (2\pi)^3 \delta_{rs} \delta^{(3)}(\mathbf{k} - \mathbf{k}') f(\omega_{\mathbf{k}}), \\ \langle a_r^\dagger(\mathbf{k}) a_s^\dagger(\mathbf{k}') a_{r'}(\mathbf{p}) a_{s'}(\mathbf{p}') \rangle &= (2\pi)^6 \left(\delta_{rs'} \delta_{r's} \delta^{(3)}(\mathbf{k} - \mathbf{p}') \delta^{(3)}(\mathbf{k}' - \mathbf{p}) \right. \\ &\quad \left. - \delta_{rr'} \delta_{ss'} \delta^{(3)}(\mathbf{k} - \mathbf{p}) \delta^{(3)}(\mathbf{k}' - \mathbf{p}') \right) f(\omega_{\mathbf{k}}) f(\omega_{\mathbf{k}'}).\end{aligned}$$

Here $f(\omega_{\mathbf{k}}) = 1/(\exp(\beta(\omega_{\mathbf{k}} - \mu)) + 1)$ is the Fermi–Dirac distribution function for particles. For antiparticles, the Fermi–Dirac distribution function differs by the sign of the baryon chemical potential μ , i.e. $\bar{f}(\omega_{\mathbf{k}}) = 1/(\exp(\beta(\omega_{\mathbf{k}} + \mu)) + 1)$.

Basic concepts and definitions:

We define the baryon number density operator \hat{J}_a^0 , associated with the conserved baryon current in a subsystem S_a using a smooth Gaussian profile placed at the origin of the coordinate system

$$\hat{J}_a^0 = \frac{1}{(a\sqrt{\pi})^3} \int d^3\mathbf{x} \hat{J}^0(\mathbf{x}) \exp\left(-\frac{\mathbf{x}^2}{a^2}\right)$$

where $\hat{J}^0 = \psi^\dagger \psi$.

To determine the baryon number fluctuation of the subsystem S_a , we consider the variance

$$\sigma^2(a, m, T, \mu) = \langle : \hat{J}_a^0 :: \hat{J}_a^0 : \rangle - \langle : \hat{J}_a^0 : \rangle^2$$

and the normalized standard deviation as

$$\sigma_n(a, m, T, \mu) = \frac{(\langle : \hat{J}_a^0 :: \hat{J}_a^0 : \rangle - \langle : \hat{J}_a^0 : \rangle^2)^{1/2}}{\langle : \hat{J}_a^0 : \rangle}$$

Mean value for baryon number density operator:

Using the thermal averaging of two creation and/or annihilation operators, the thermal expectation value of $:\hat{J}_a^0:$ has the form

$$\langle : \hat{J}_a^0 : \rangle = 2 \int dK \left[f(\omega_{\mathbf{k}}) - \bar{f}(\omega_{\mathbf{k}}) \right]$$

This expression agrees with the standard kinetic-theory definition, with the factor of 2 accounting for the spin degeneracy.

Quantum fluctuation expression:

$$\begin{aligned}\sigma^2(a, m, T, \mu) &= \langle : \hat{J}_a^0 :: \hat{J}_a^0 : \rangle - \langle : \hat{J}_a^0 : \rangle^2 \\ &= \int \frac{dK}{\omega_{\mathbf{k}}} \frac{dK'}{\omega_{\mathbf{k}'}} (\omega_{\mathbf{k}} \omega_{\mathbf{k}'} + \mathbf{k} \cdot \mathbf{k}' + m^2) e^{-\frac{a^2}{2}(\mathbf{k}-\mathbf{k}')^2} \times \\ &\quad [f(\omega_{\mathbf{k}})(1 - f(\omega_{\mathbf{k}'})) + \bar{f}(\omega_{\mathbf{k}})(1 - \bar{f}(\omega_{\mathbf{k}'}))] \\ &- \int \frac{dK}{\omega_{\mathbf{k}}} \frac{dK'}{\omega_{\mathbf{k}'}} (\omega_{\mathbf{k}} \omega_{\mathbf{k}'} + \mathbf{k} \cdot \mathbf{k}' - m^2) e^{-\frac{a^2}{2}(\mathbf{k}+\mathbf{k}')^2} \times \\ &\quad [f(\omega_{\mathbf{k}})(1 - \bar{f}(\omega_{\mathbf{k}'})) + \bar{f}(\omega_{\mathbf{k}})(1 - f(\omega_{\mathbf{k}'}))]\end{aligned}$$

Variation of normalized baryon density fluctuation:

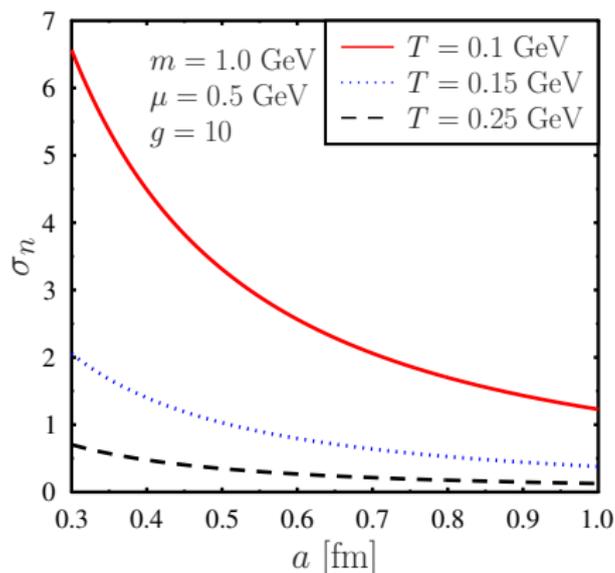


Figure: Variation of normalized fluctuation σ_n in the subsystem S_a with the scale a for different values of the temperature T and fixed particle mass $m = 1.0$ GeV and baryon chemical potential $\mu = 0.5$ GeV.

Variation of normalized baryon density fluctuation:

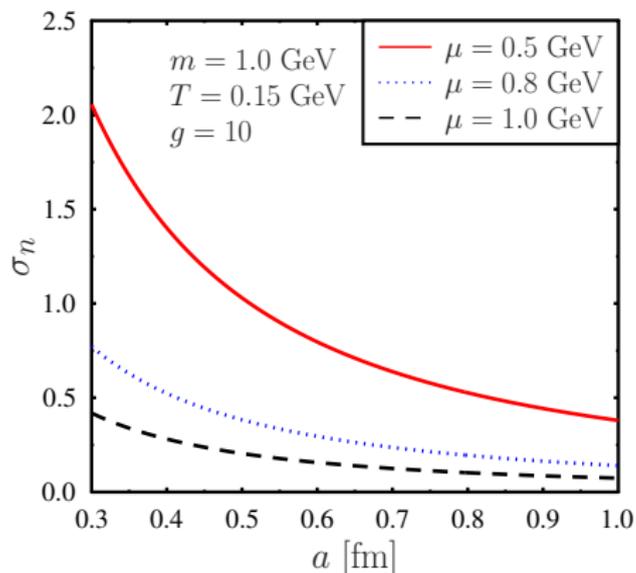


Figure: Variation of normalized fluctuation σ_n in the subsystem S_a with the scale a for different values of the baryon chemical potential μ and fixed particle mass $m = 1.0 \text{ GeV}$ and temperature $T = 0.15 \text{ GeV}$.

Variation of normalized baryon density fluctuation:

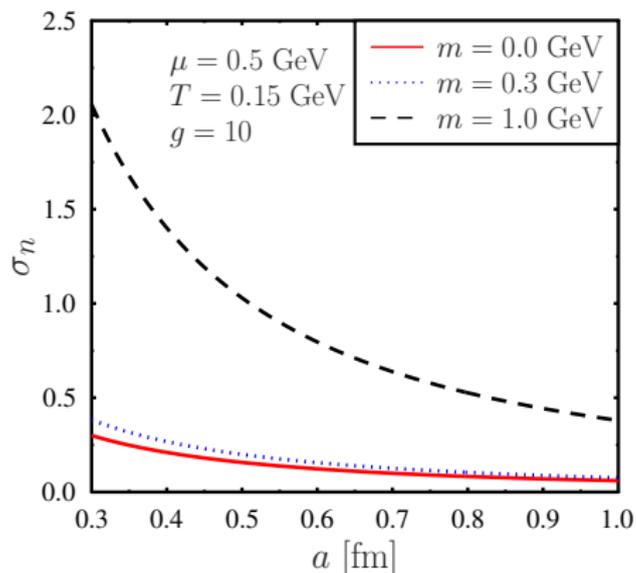


Figure: Variation of normalized fluctuation σ_n in the subsystem S_a with the scale a for different values of the particle mass and fixed temperature $T = 0.15 \text{ GeV}$ and baryon chemical potential $\mu = 0.5 \text{ GeV}$.

Variation of normalized baryon fluctuation:

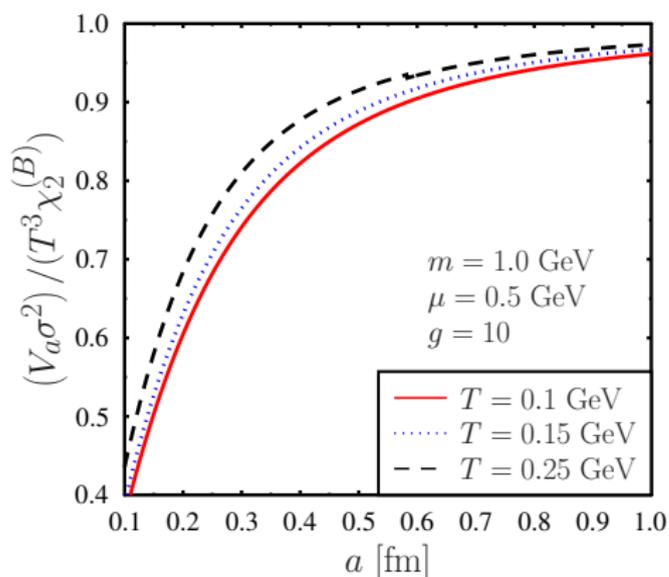


Figure: Variation of normalized fluctuation $V_a \sigma^2 / (T^3 \chi_2^{(B)})$ for different values of temperature (T) but with fixed baryon chemical potential (μ) and particle mass (m).

Variation of normalized baryon fluctuation:

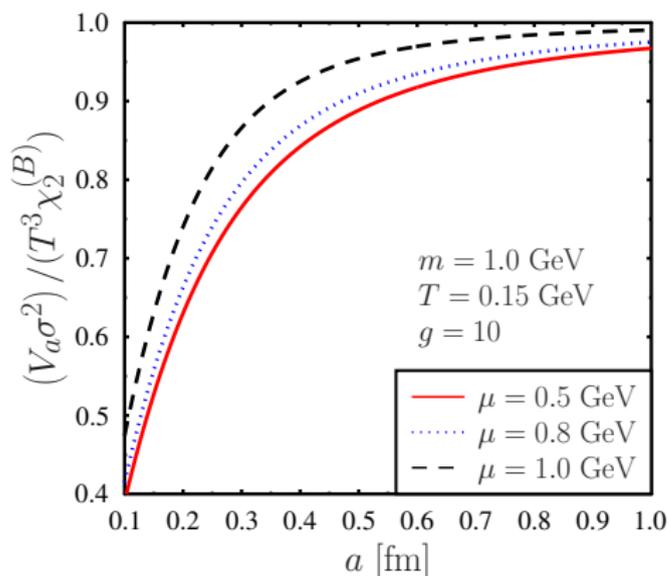


Figure: Variation of normalized fluctuation $V_a \sigma^2 / (T^3 \chi_2^{(B)})$ for different values of baryon chemical potential (μ) but with fixed temperature (T) and particle mass (m).

Variation of normalized baryon fluctuation:

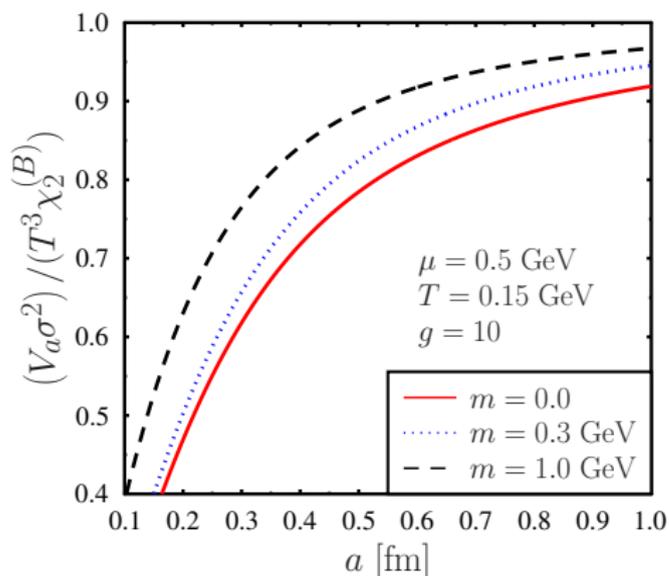
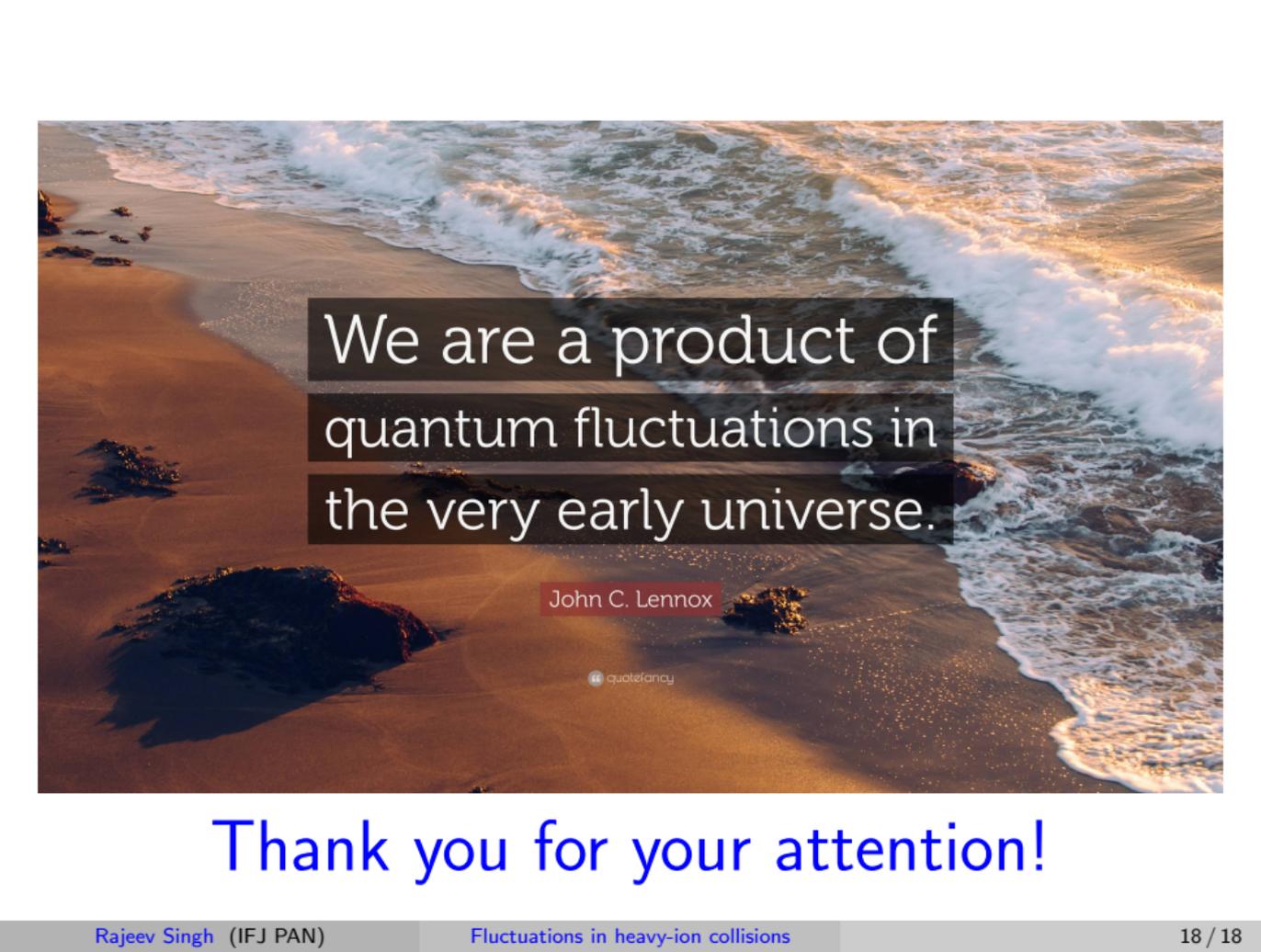


Figure: Variation of normalized fluctuation $V_a \sigma^2 / (T^3 \chi_2^{(B)})$ for different values of particle mass (m) but with fixed temperature (T) and baryon chemical potential (μ).

Summary:

- We have analyzed quantum baryon-number fluctuations in subsystems of a hot and dense relativistic gas of fermions
- And found that they diverge for small system sizes.
- Our results agree with the results known from statistical physics for sufficiently large system size a .
- In this way, we have delivered a useful formula that accounts for both statistical and quantum features of the fluctuations.
- The numerical results obtained here can be useful to interpret and shed new light on the heavy-ion experimental data.



We are a product of
quantum fluctuations in
the very early universe.

John C. Lennox

quote fancy

Thank you for your attention!