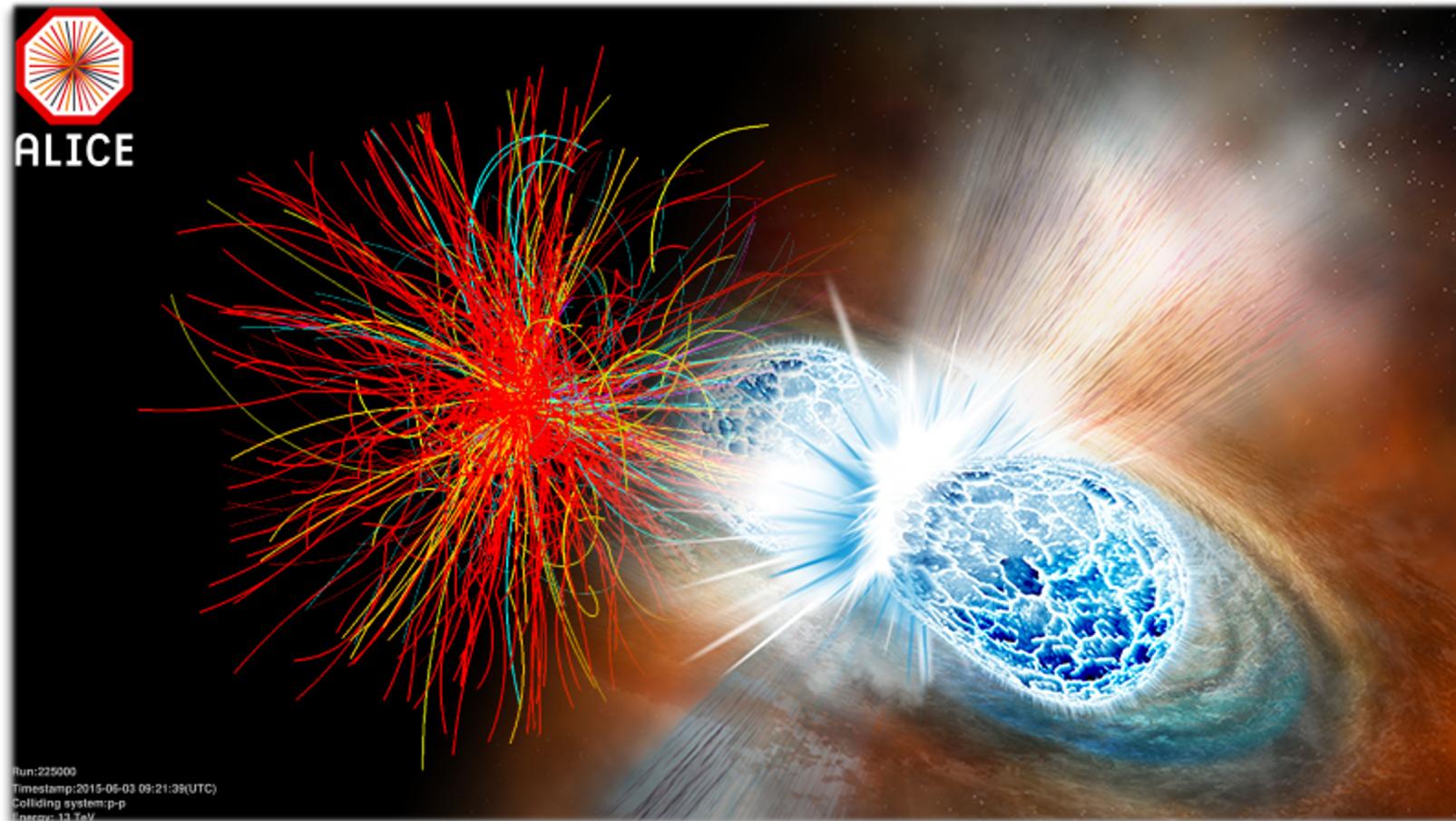




ALICE



ALICE unveils strong interaction
among stable and unstable hadrons.

SFB 1258
Neutrinos
Dark Matter
Messengers

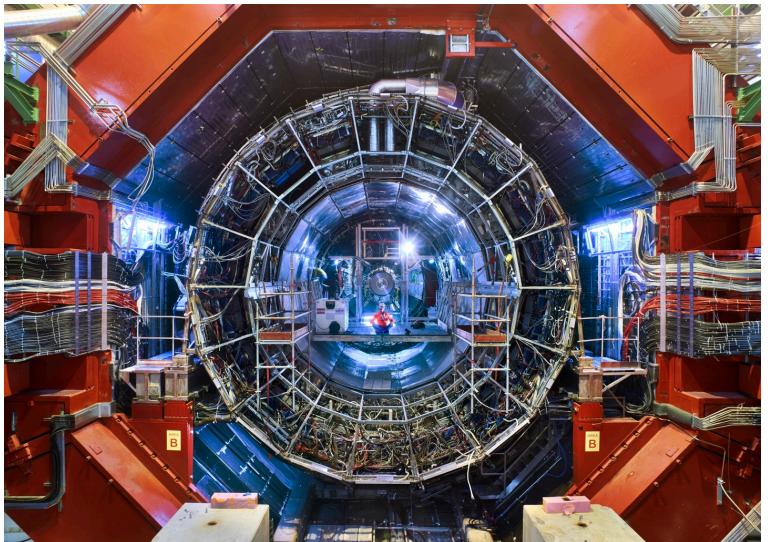


Laura Fabbietti, Technische Universität München

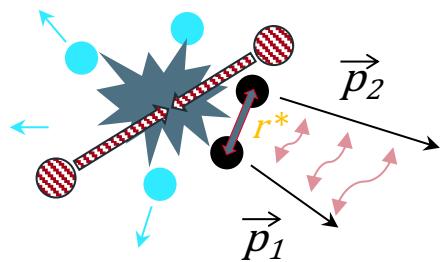


Outline

ALICE at the LHC

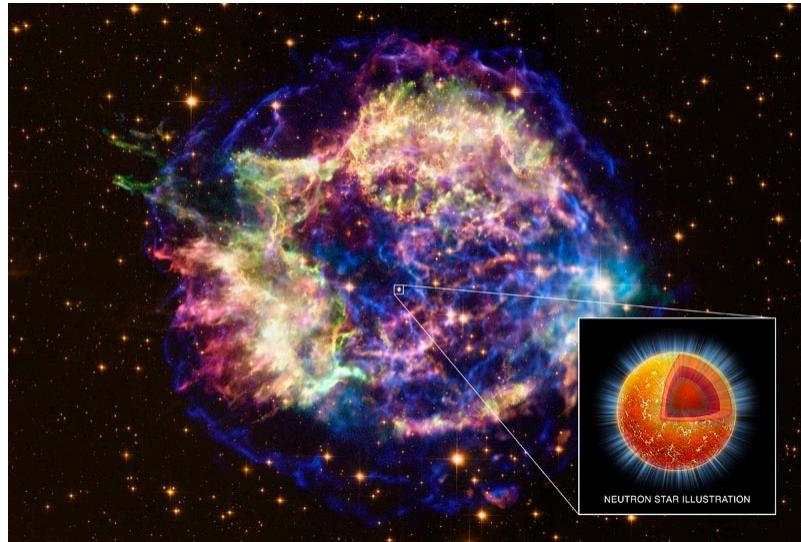


Femtoscopy



Hyperon-nucleon
Hyperon-Hyperon
Hyperon-NN...

Hyperon Stars



Hyperon appearance in neutron stars?

Dimensions

$R \sim 10 - 15 \text{ km}$

$M \sim 1.5 - 2.2 M_{\odot}$

Outer Crust

Ions, electron gas,
Neutrons

Inner Core

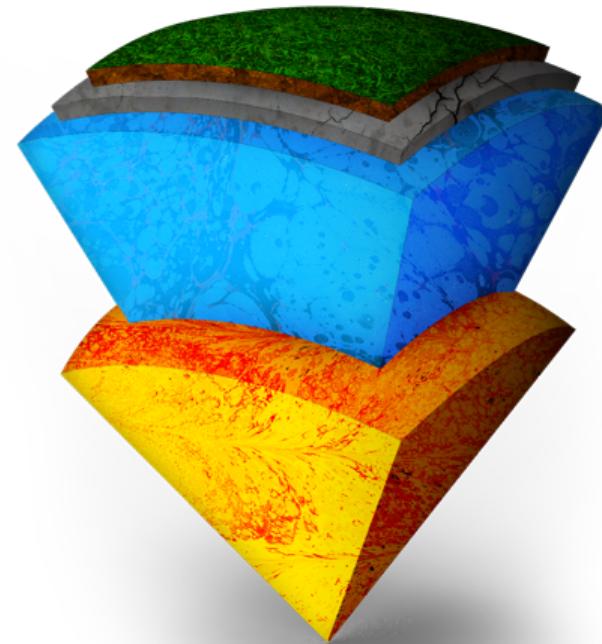
Neutrons?

Protons?

Hyperons?

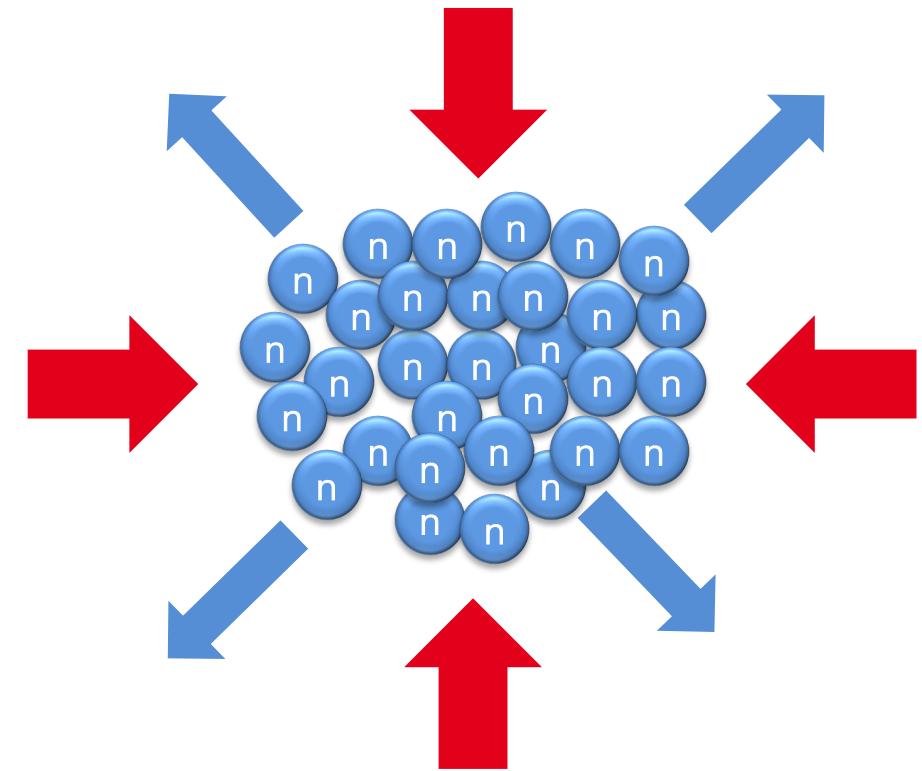
Kaon condensate?

Quark Matter?

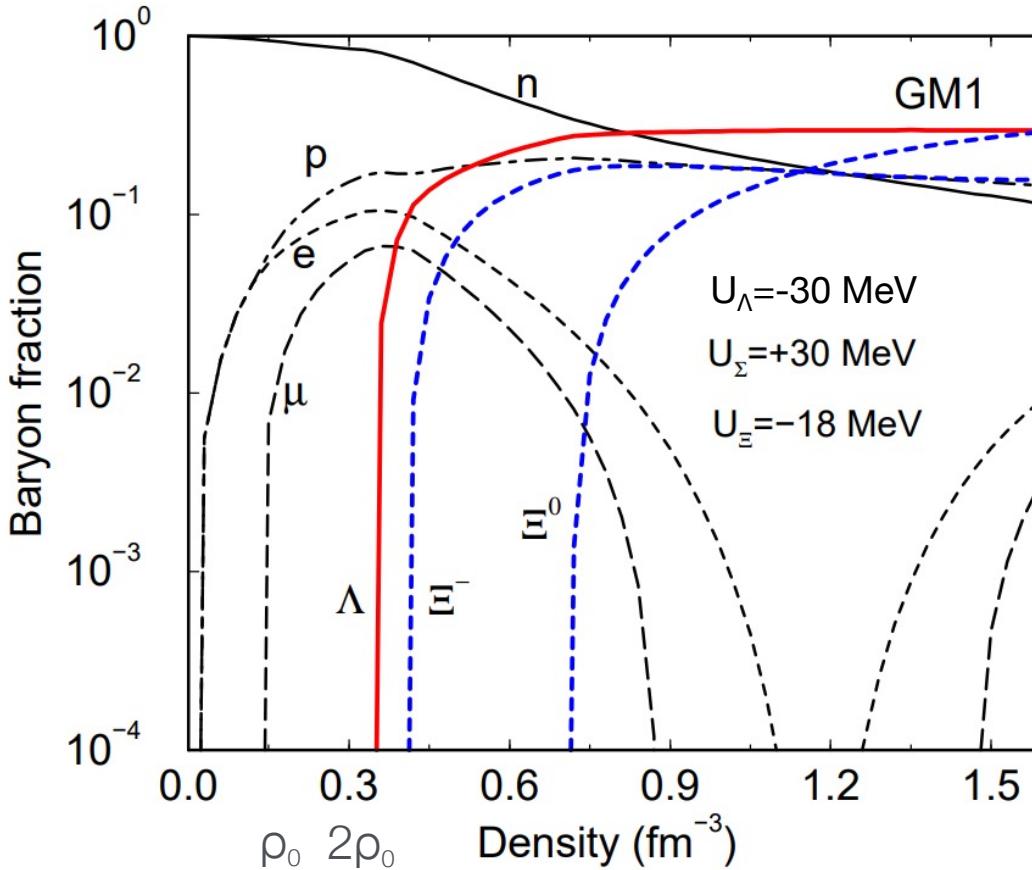


Neutron Stars: very dense, compact objects

- What is the EoS?
 - What are the constituents to consider?
 - How do they interact?



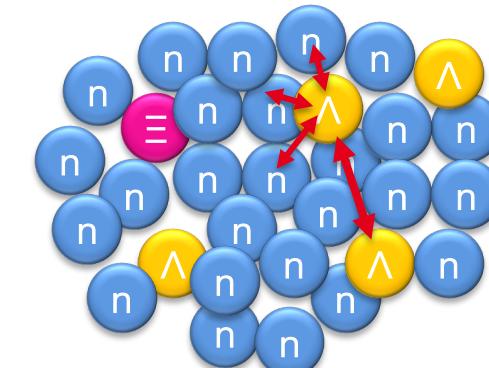
Hyperon appearance in neutron stars?



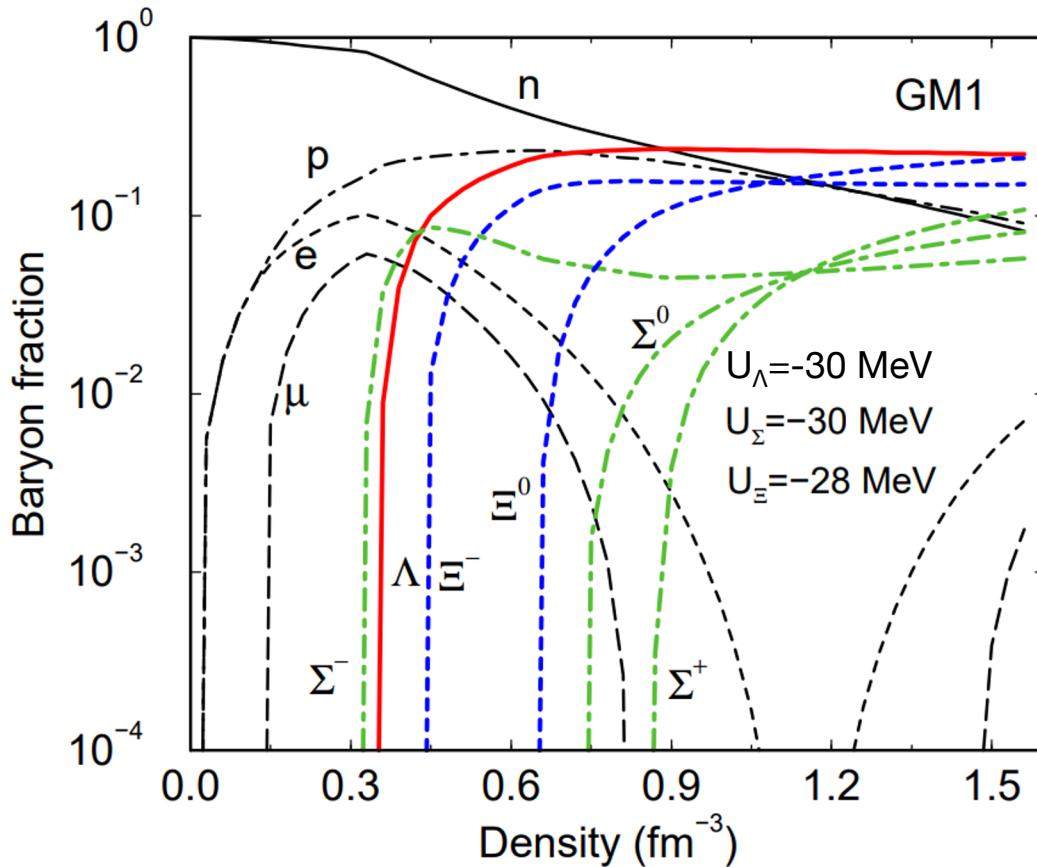
J. Schaffner-Bielich, Nucl. Phys. A 835 (2010) 279

Neutron Stars: very dense, compact objects

- At finite densities **hyperon** production becomes energetically favorable \Rightarrow softening of EoS
 - Appearance as a function of ρ depends on the Y interaction with medium and Y interactions among themselves
 - Naïve introduction of Λ incompatible with astrophysical observations



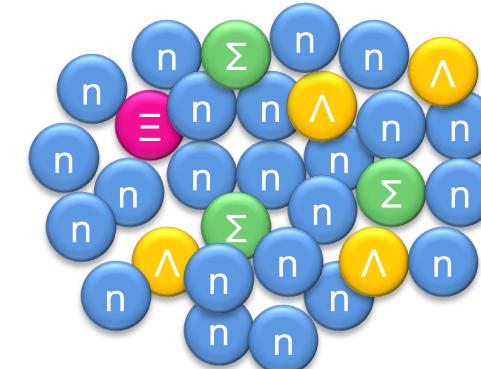
Hyperon appearance in neutron stars?



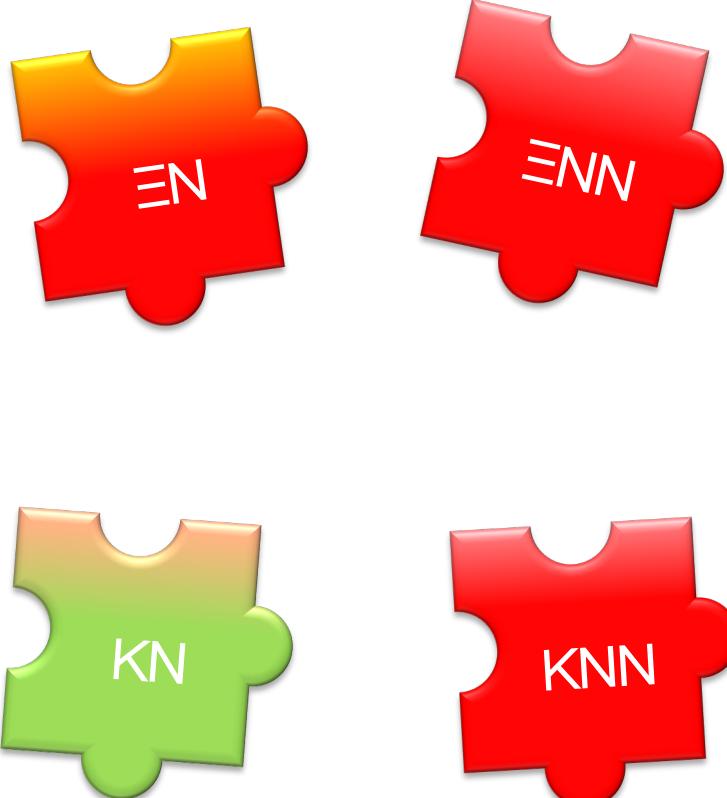
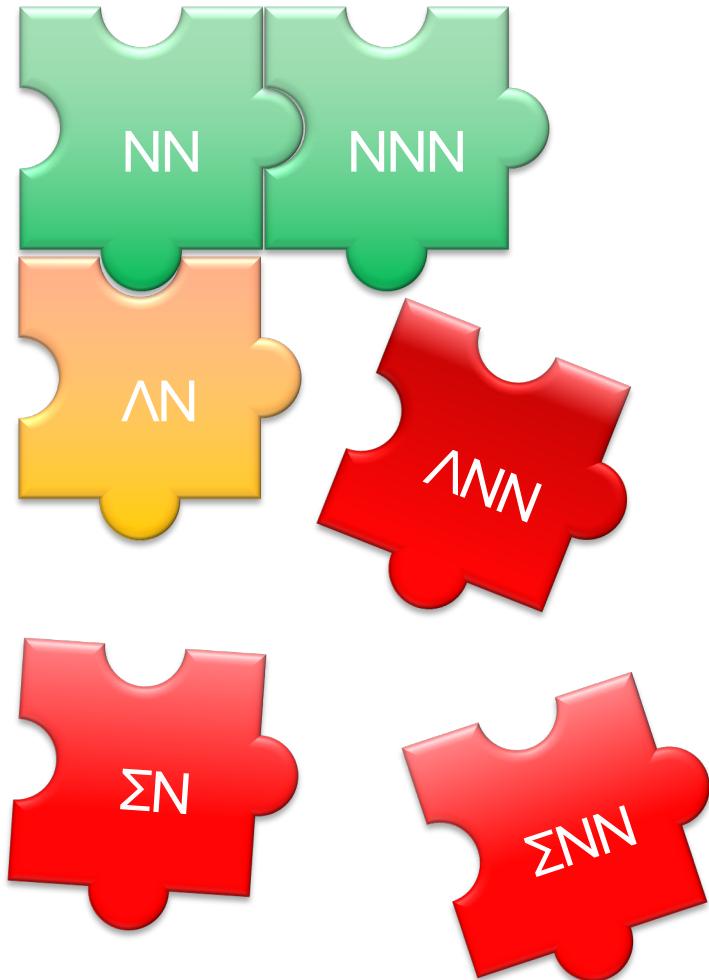
J. Schaffner-Bielich, Nucl. Phys. A 835 (2010) 279

Neutron Stars: very dense, compact objects

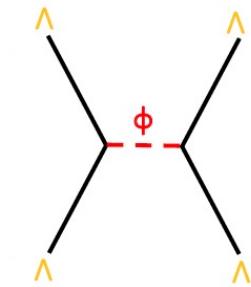
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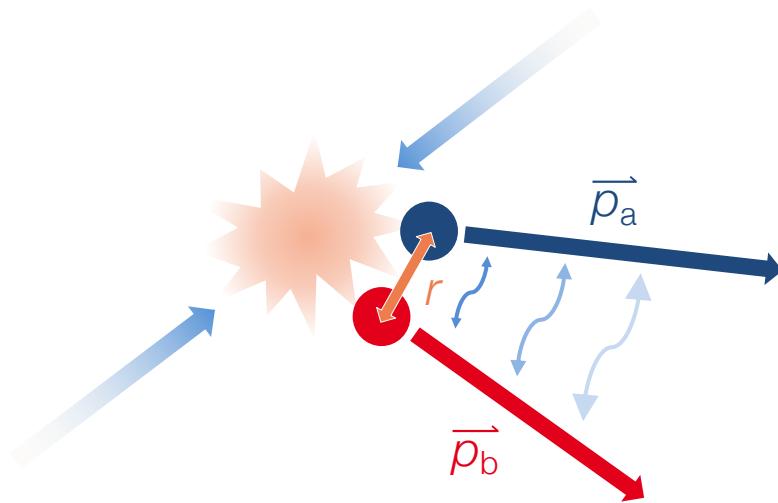
The building blocks for the EoS of hyperon stars: two- and three-body interactions



Vector meson exchange to describe $\Lambda\Lambda$ interaction



Femtoscopy with small sources

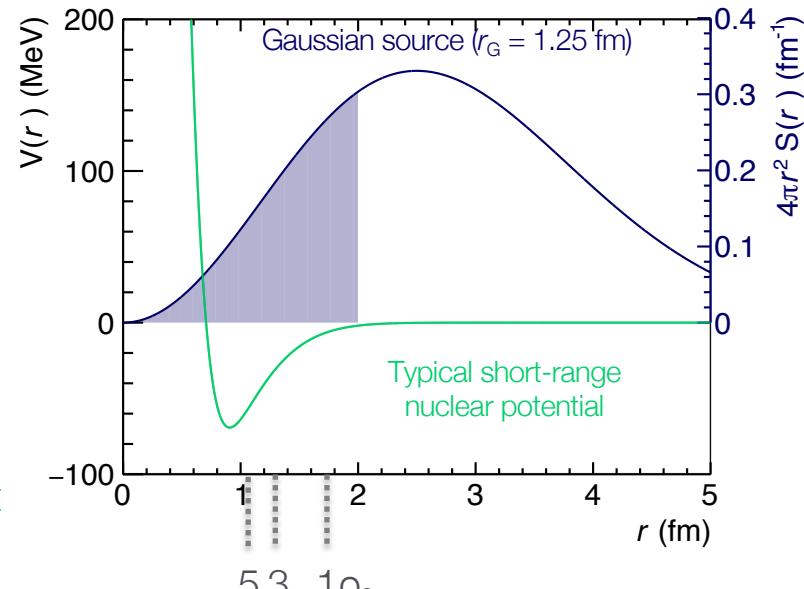


$$C(k^*) = \int S(r) |\psi(\vec{k}^*, \vec{r})|^2 d^3r = \zeta(k^*) \cdot \frac{N_{same}(k^*)}{N_{mixed}(k^*)}$$

Emission source

Two-particle
wave function

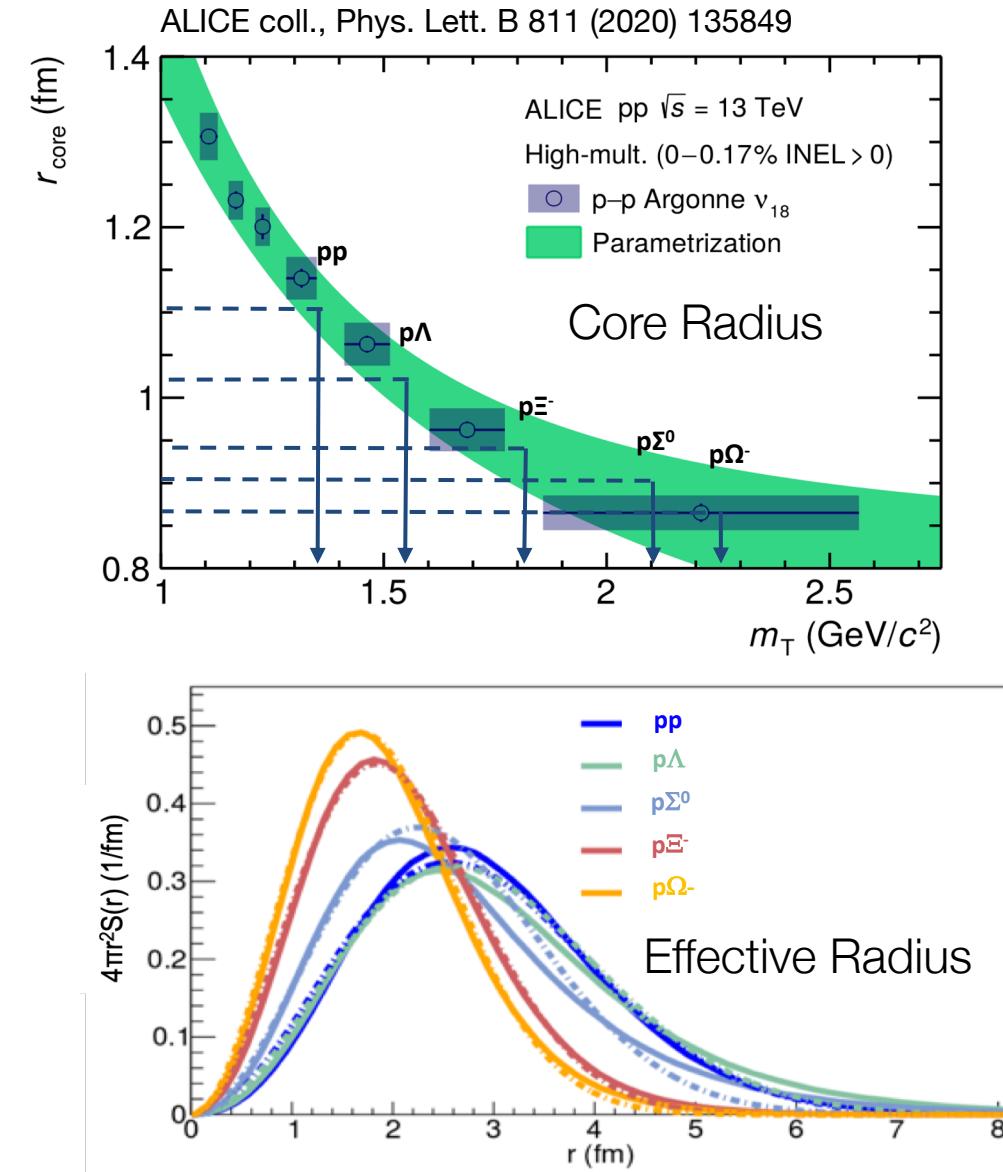
- > 1 Attract
- = 1 No Inter.
- < 1 Repulsive



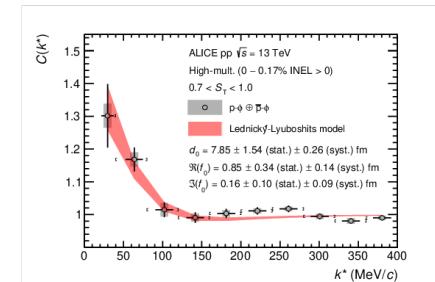
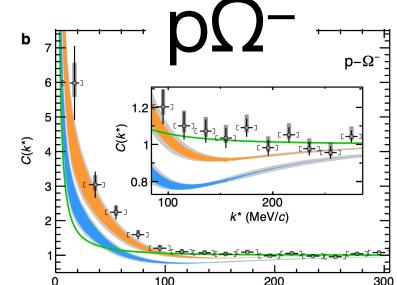
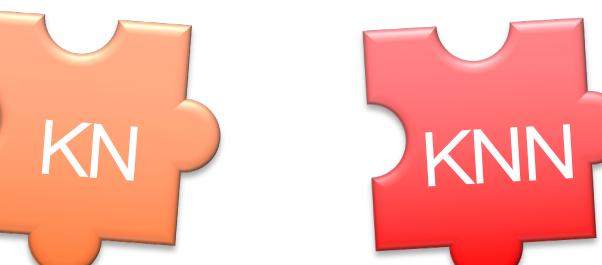
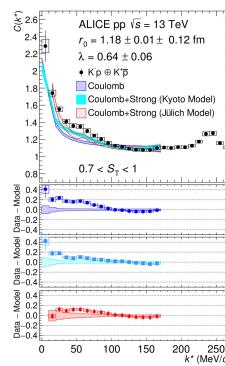
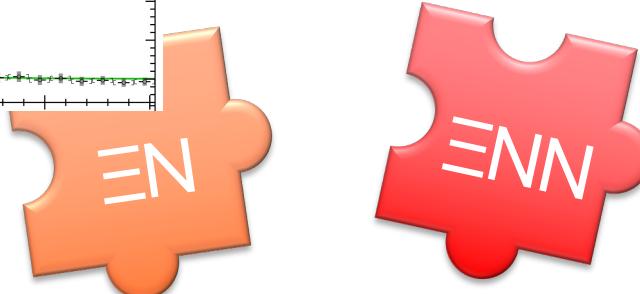
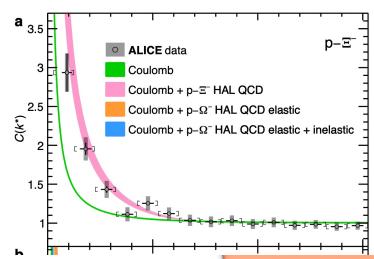
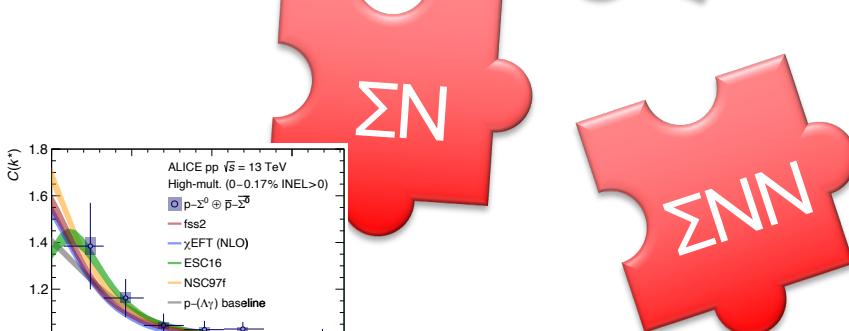
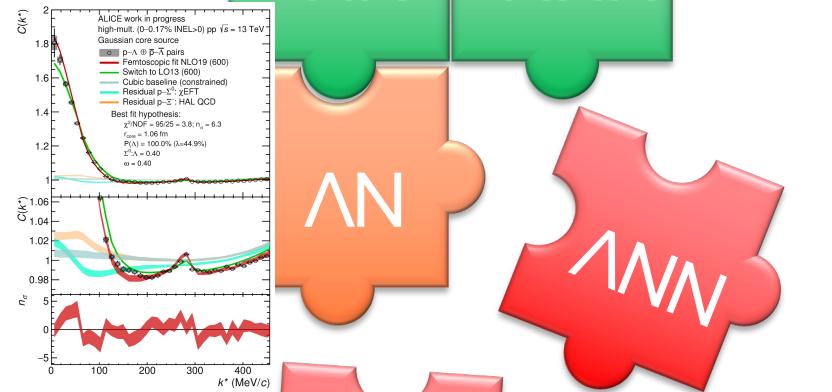
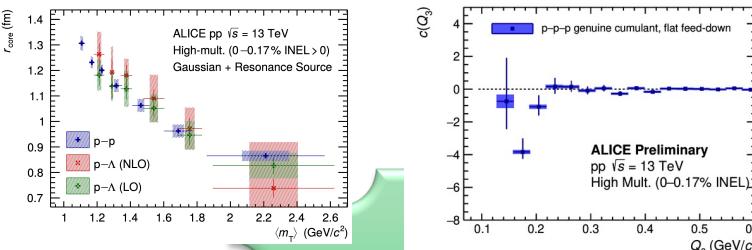
- Small particle-emitting source created in pp and p–Pb collisions at the LHC
 - Essential ingredient for detailed studies of the strong interaction
 - D. Mihaylov *et al.*, *Eur. Phys. J. C*78 (2018) 394
 - Small interparticle distance → doorway to studying large densities

Small Sources: Collective Effects and Strong Resonances

- Modeling of the emitting source in pp collisions with a common core + halo of strongly decaying resonances for each pair
- Core constrained with femtoscopic studies for p-p pairs in pp HM 13 TeV
- Resonance contributions evaluated from statistical models and kinematics from transport models
- Determination of the source for each pair at the corresponding $\langle m_T \rangle \Rightarrow$ direct access to the interaction signal of the correlation function

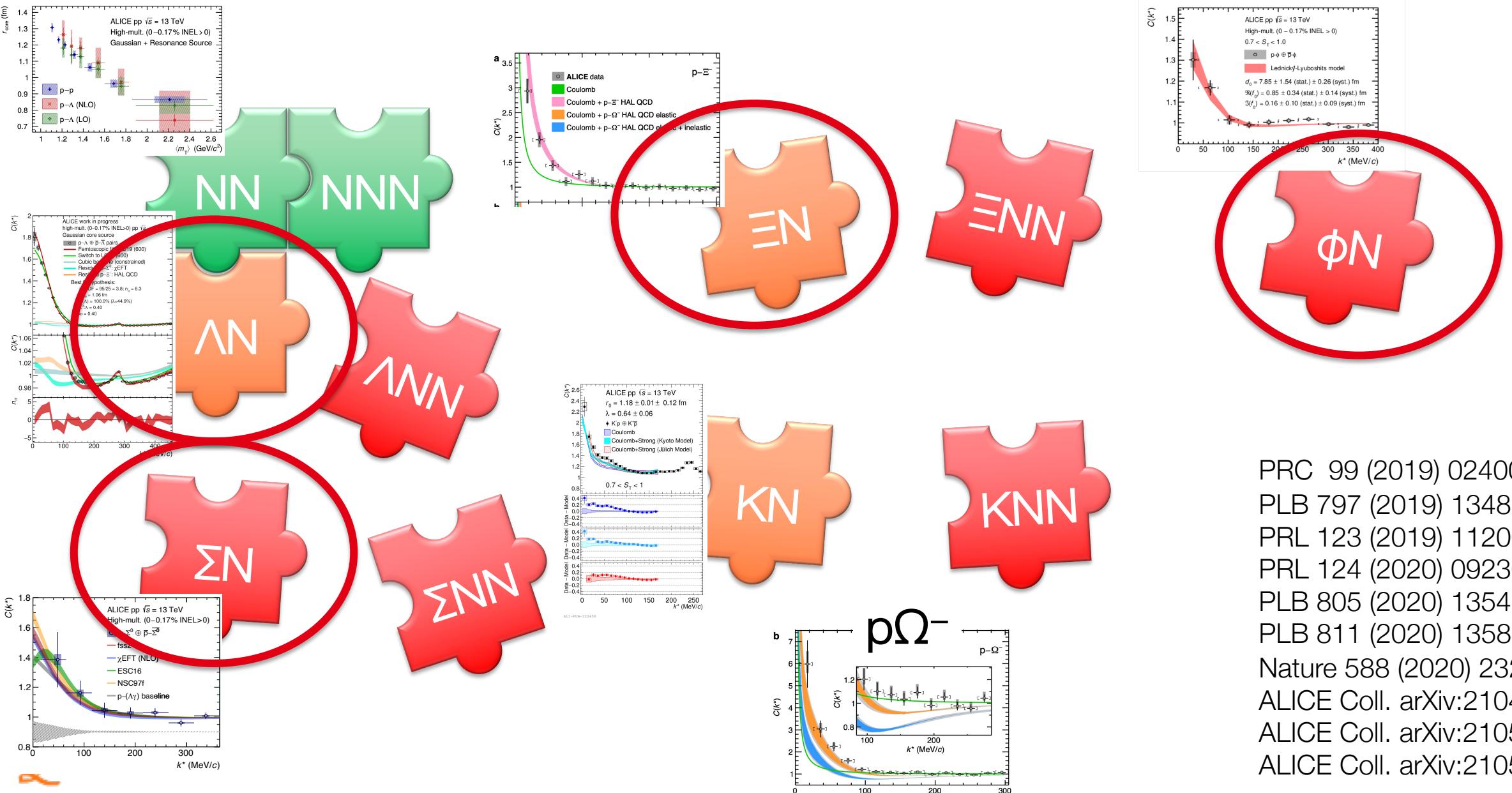


Harvest of LHC Run 1 & 2



- PRC 99 (2019) 024001
PLB 797 (2019) 134822
PRL 123 (2019) 112002
PRL 124 (2020) 09230
PLB 805 (2020) 135419
PLB 811 (2020) 135849
Nature 588 (2020) 232-238
ALICE Coll. arXiv:2104.04427
ALICE Coll. arXiv:2105.05578
ALICE Coll. arXiv:2105.05190

Harvest of LHC Run 1 & 2 – What we will see today

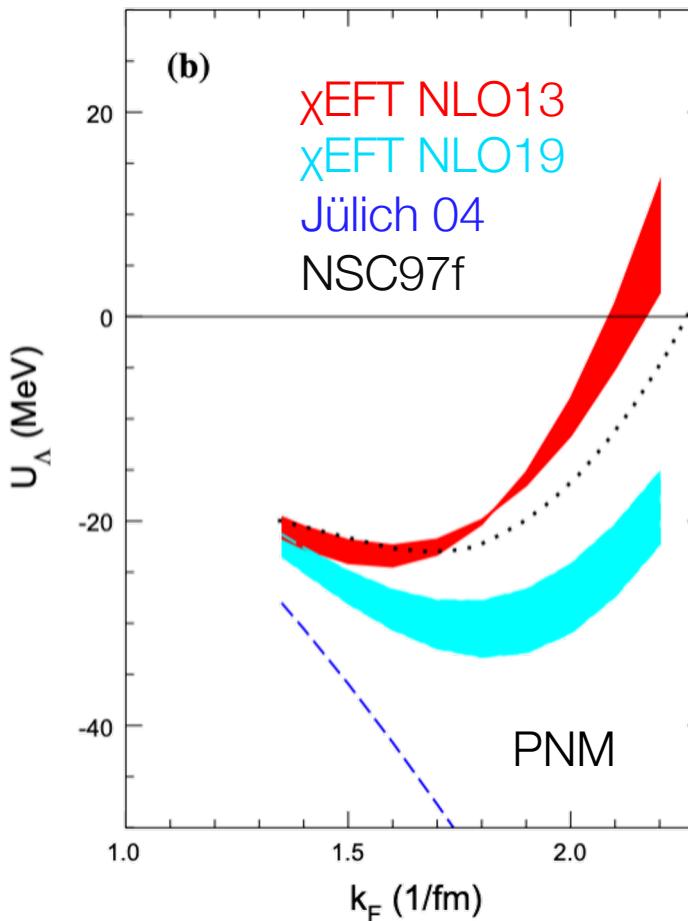
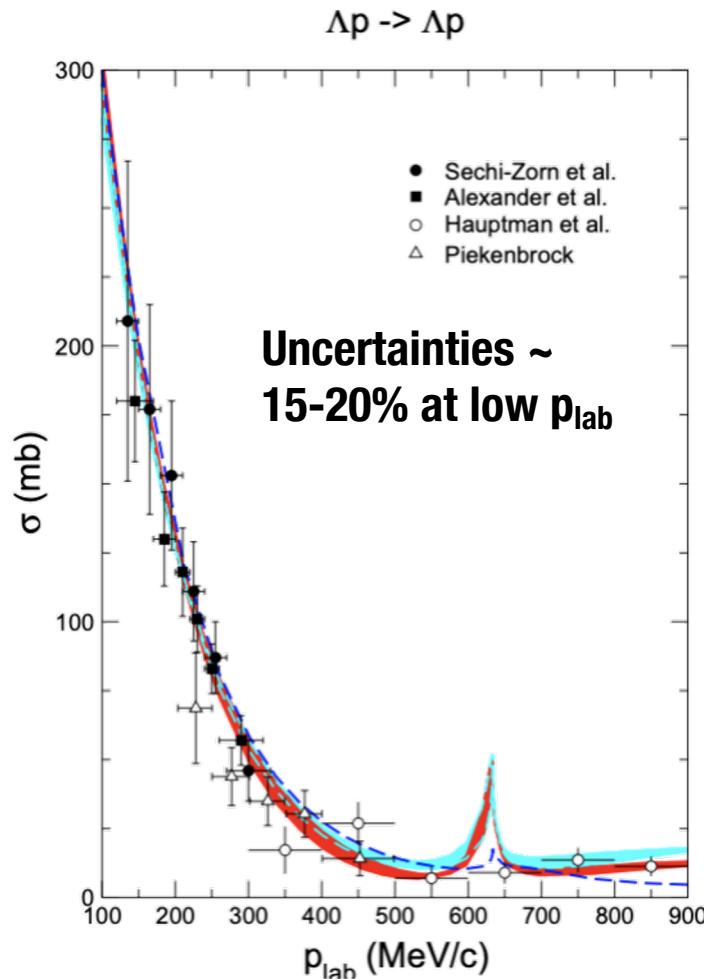


- PRC 99 (2019) 024001
 PLB 797 (2019) 134822
 PRL 123 (2019) 112002
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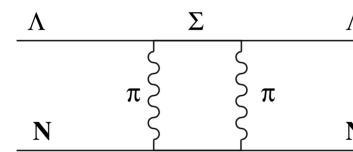
$|S|=1$: $p\Lambda$ Interaction and role of the ΣN coupling



NLO13: J.Haidenbauer, N.Kaiser et al., NPA 915, 24 (2013)
 NLO19: J.Haidenbauer, U. Meißner, Eur.Phys.J.A 56 (2020)



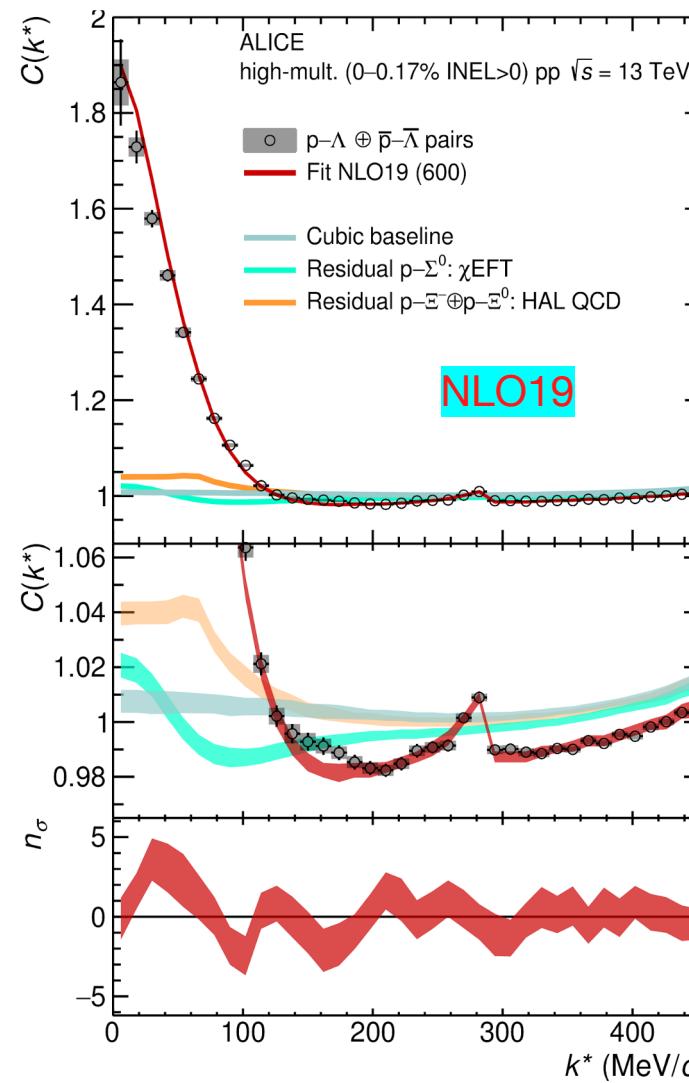
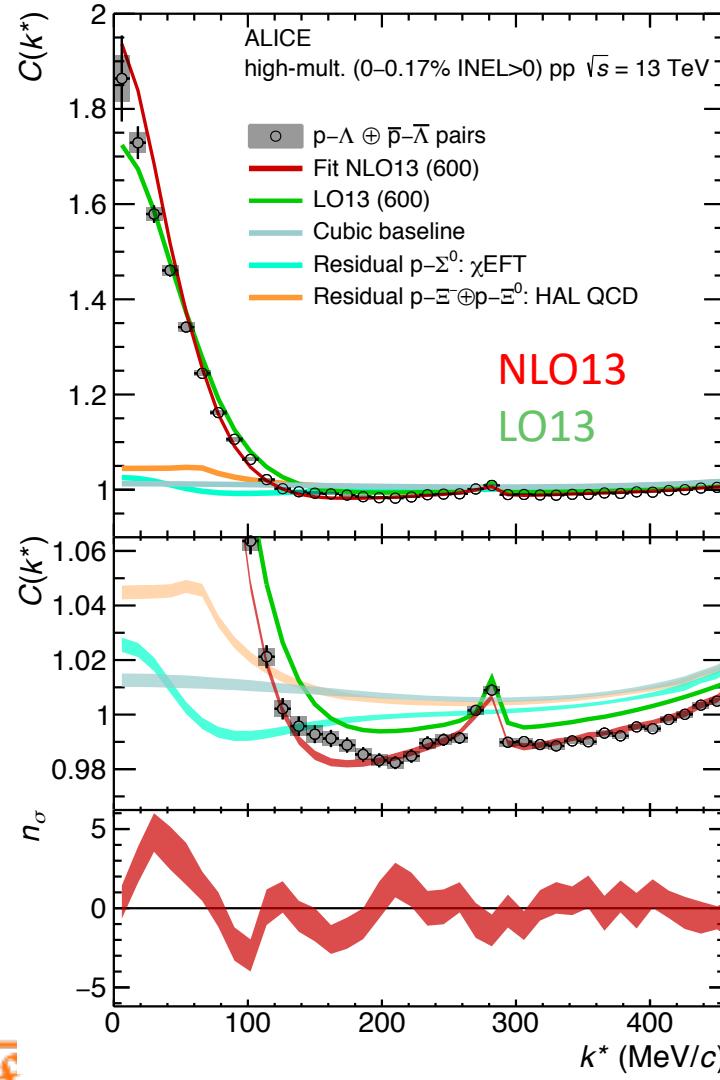
- $p\Lambda$ interaction:
 - low-statistics scattering data and hypernuclei, not available at low momenta ($p_{lab} > 100$ MeV/c)
 - 2-body coupling to ΣN is experimentally not (yet) measured
- ΣN coupling strength deeply affects the behaviour of Λ at finite density
- Relevance for EoS in NS (“hyperon puzzle”) and for connection to role of ΛNN three-body interaction
- Updated NLO19 with weaker coupling strength in $\Lambda N - N\Sigma$ leading to different Λ properties in nuclear matter





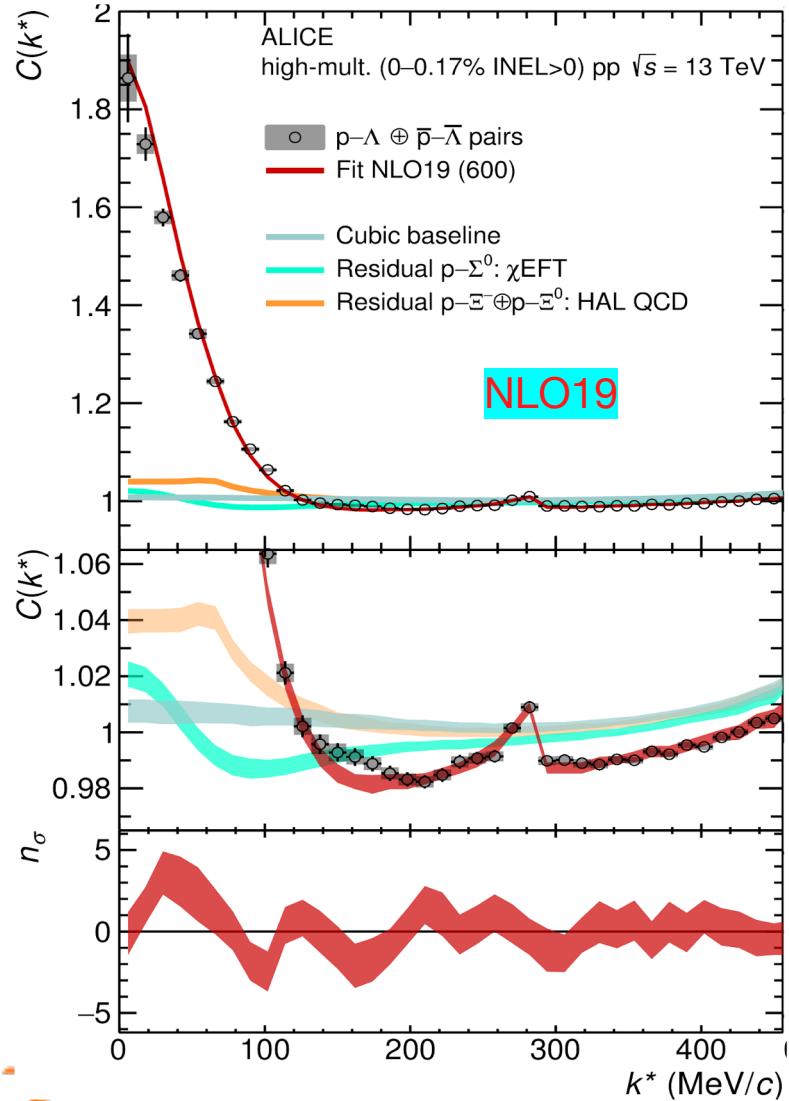
$|S|=1 : \Lambda p$ Interaction and access to the ΣN coupling

ALICE Coll. arXiv:2104.04427 submitted to PRL



- Extension of the kinematic range
- **Factor 20 improved precision of the measurement**
- **First experimental evidence ΣN opening in 2 body channel**
- Sensitivity to different strength of ΣN coupling (NLO13 vs NLO19)
- Sensitivity to residual $p\Sigma^0$ contribution \Rightarrow shallow interaction
- **NLO19 favoured \Rightarrow very attractive $U_\Lambda \Rightarrow$ too soft EoS....**

$|S|=1 : \Lambda p$ Interaction and access to the ΣN coupling



$p-\Sigma^0 (\rightarrow)$ $p-\Lambda (\downarrow)$	χ EFT	Flat
LO13-600	4.7 (5.7)	7.7 (10.8)
NLO13-500	5.3 (7.4)	3.3 (3.7)
NLO13-550	3.7 (4.5)	2.5 (2.5)
NLO13-600	4.4 (4.5)	2.9 (3.3)
NLO13-650	4.1 (4.0)	3.1 (3.8)
NLO19-500	4.0 (4.8)	2.8 (2.8)
NLO19-550	3.4 (3.5)	2.3 (2.7)
NLO19-600	3.2 (3.1)	2.3 (2.9)
NLO19-650	3.2 (3.2)	2.6 (3.4)

- Extension of the kinematic range
- **Factor 20 improved precision of the measurement**
- **First experimental evidence ΣN opening in 2 body channel**
- Sensitivity to different strength of ΣN coupling (NLO13 vs NLO19)
- Sensitivity to residual $p\Sigma^0$ contribution \Rightarrow shallow interaction
- **NLO19 favoured \Rightarrow very attractive $U_\Lambda \Rightarrow$ too soft EoS....**



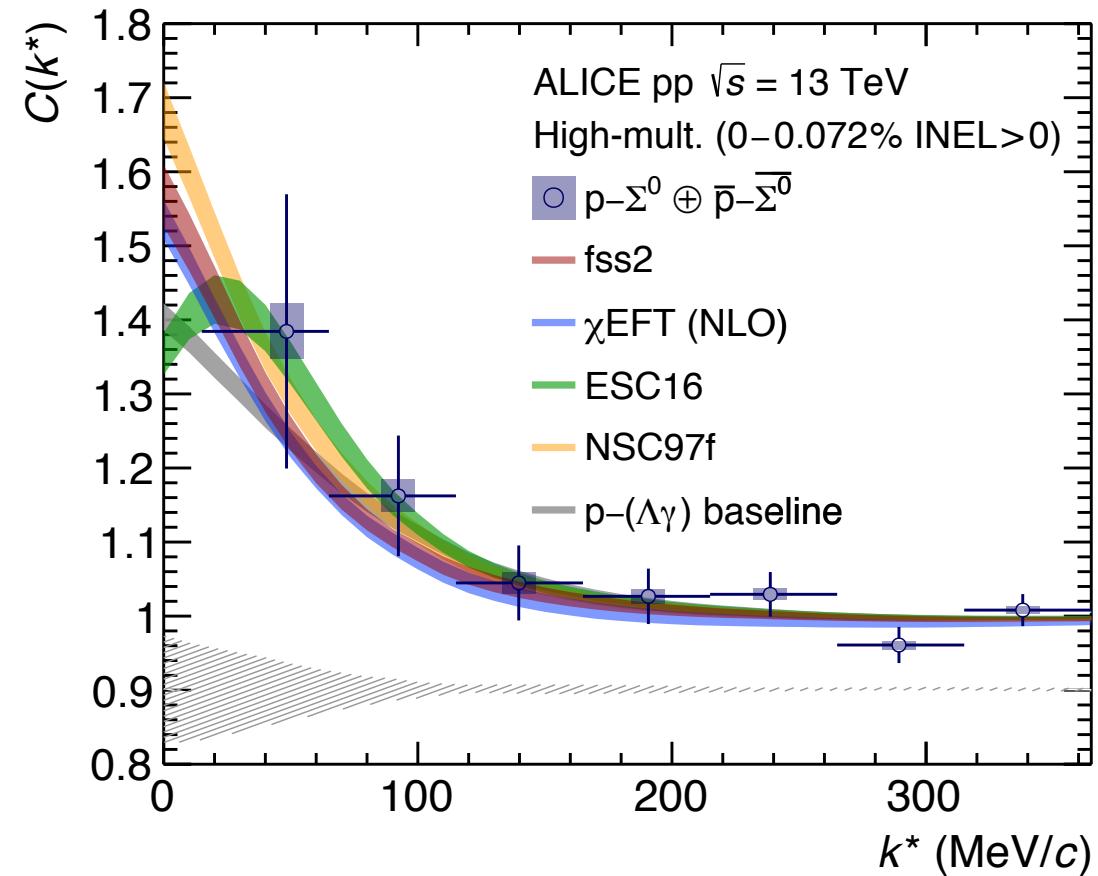
$|S|=1$: First measurement of the $p\Sigma^0$ interaction

- Experimental data on hypernuclei too scarce for a final conclusion: attractive or repulsive interaction?
- Very challenging measurement via the difficult electromagnetic decay $\Sigma^0 \rightarrow \Lambda \gamma$
- **Correlation function is above the background \Rightarrow pointing to a very shallow attractive interaction**

Model	$p-(\Lambda\gamma)$ baseline	fss2	χ EFT	NSC97f	ESC16
n_σ ($k^* < 150$ MeV/c)	0.2–0.8	0.2–0.9	0.3–1.0	0.2–0.6	0.1–0.5

- Relevant for dense neutron matter is the interaction with neutrons and the interaction of $\Sigma^{+,-}$!
 - Disentangle the different isospin contributions
- **Larger statistics in Run3 and Run4 will definitely increase the precision and constraints on the Σ -N interaction**

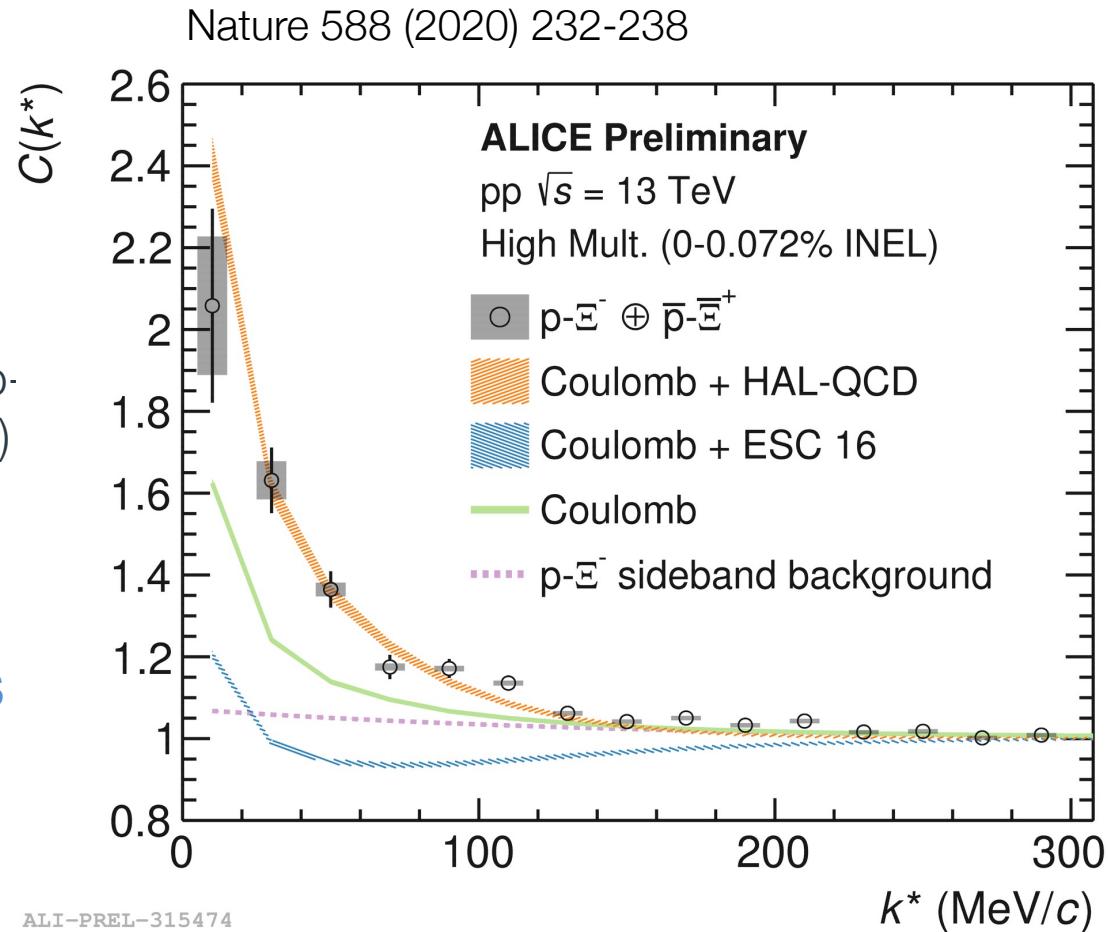
ALICE coll., PLB 805 (2020) 135419



$|S|=2$ sector: $p\bar{\Xi}^-$ -interaction and first test of LQCD

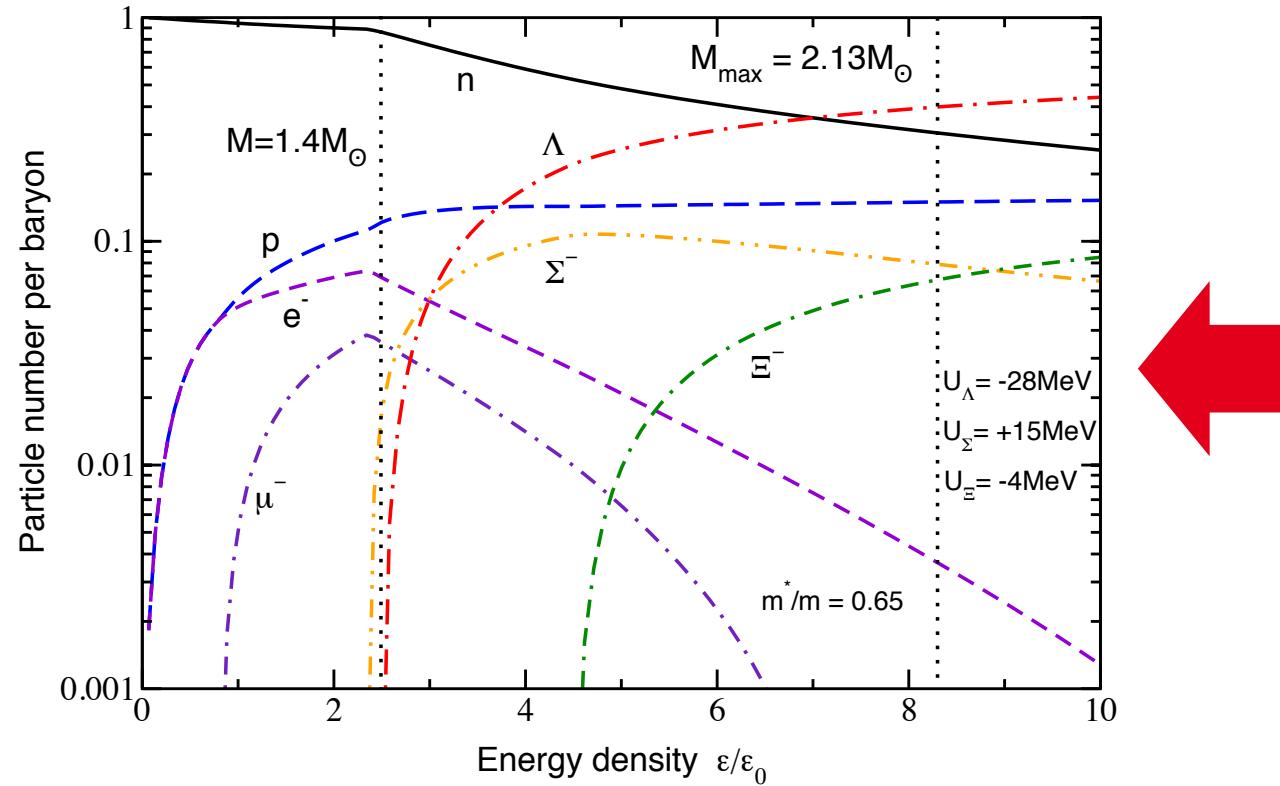


- Lattice QCD potentials from HAL QCD collaboration available \Rightarrow femtoscopy unique tool to study this system
- Observation of the strong interaction beyond Coulomb
- Agreement with LQCD calculations confirmed in pp and p-Pb colliding systems (Phys. Rev. Lett 123, (2019) 112002)
- **At finite density HAL QCD potentials predict in PNM a slightly repulsive $U_{\Xi} \sim +6$ MeV (HAL QCD Coll., PoS INPC2016 (2016) 277)**
 - pushing production of Ξ to higher densities \Rightarrow stiffer EoS

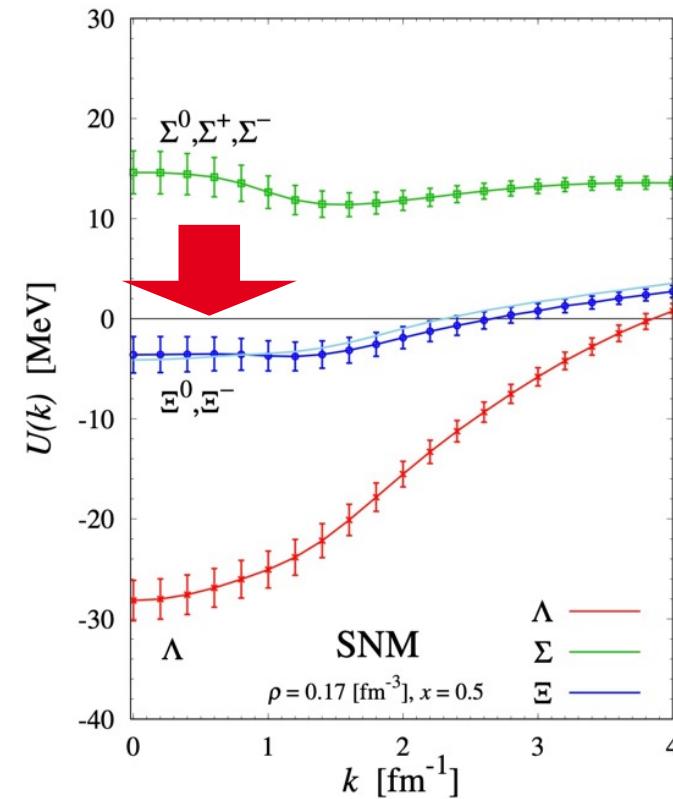


Single particle potentials from lattice QCD

Courtesy J. Schaffner-Bielich 2020



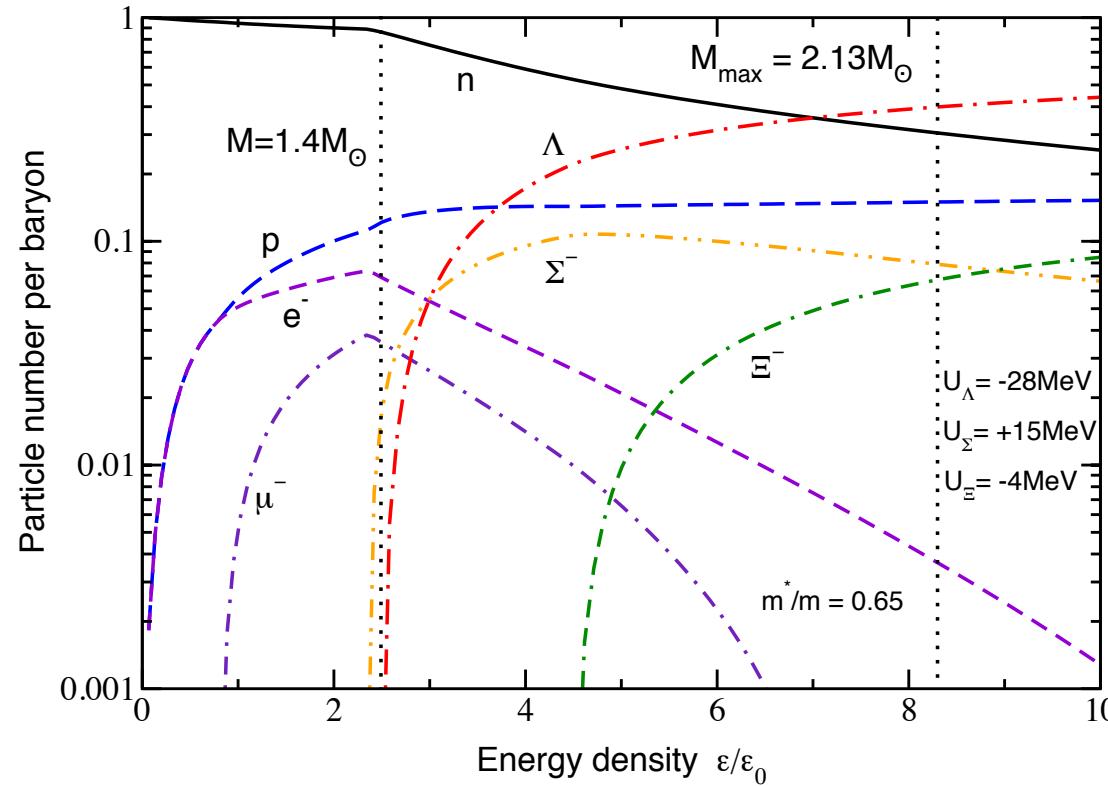
Attractive $p\Xi^-$ interaction lead to slightly attractive single particle potential in symmetric nuclear matter (SNM) and slight repulsion in neutron rich matter. Ξ^- appears at larger densities in NS



Inoue, T. Strange nuclear physics from QCD on lattice. AIP Conference Proceedings 2130, 020002 (2019)

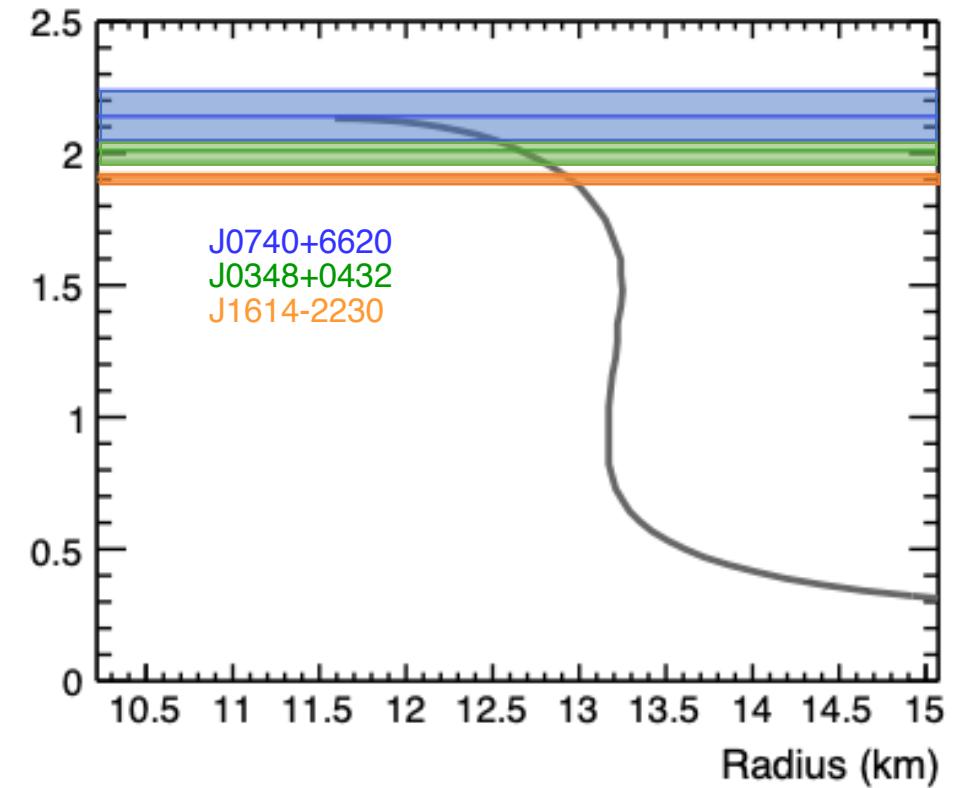
Equation of state for neutron stars

Courtesy J. Schaffner-Bielich 2020



Mass (M_\odot)

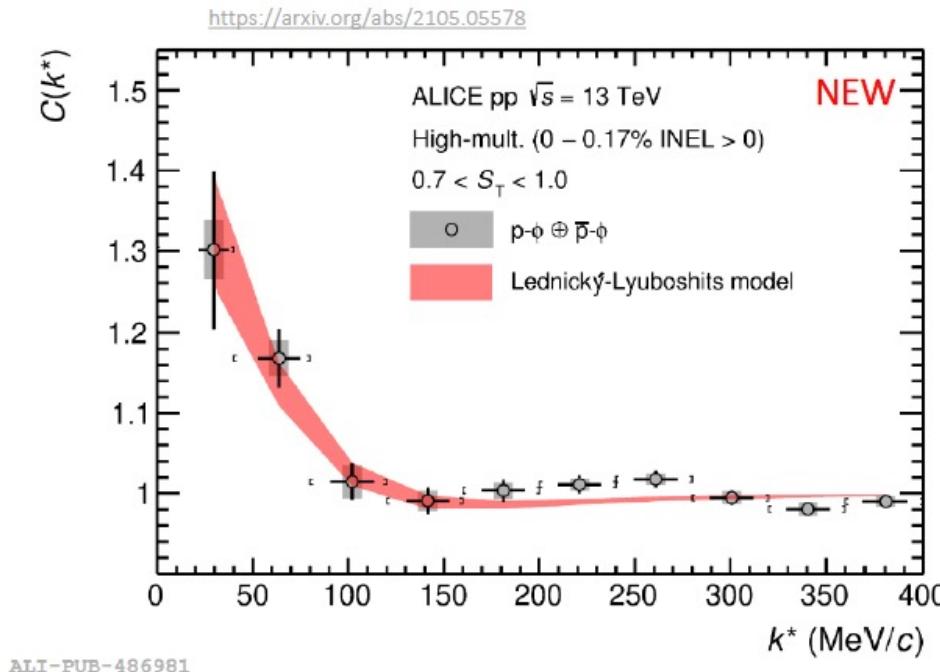
LF, V. Mantovani Sarti, O. Vazquez Doce arXiv:2012.09806



The resulting equation of state for neutron star is stiffer and the observation of 2 solar masses is matched.

This is not the end of the story...

Proton- ϕ Interaction



Extracted genuine correlation function
After the subtraction of combinatorics, mini-jets and unfolding
according to the genuine p- ϕ contribution.

1) Lednicky-Lyuboshits Model

$$C(k^*) = \sum_S \rho_S \left[\frac{1}{2} \left| \frac{f(k^*)}{r_0} \right|^2 \left(1 - \frac{d_0}{2\sqrt{\pi}r_0} \right) + \frac{2\Re f(k^*)}{\sqrt{\pi}r_0} F_1(2k^*r_0) - \frac{\Im f(k^*)}{r_0} F_2(2k^*r_0) \right]$$

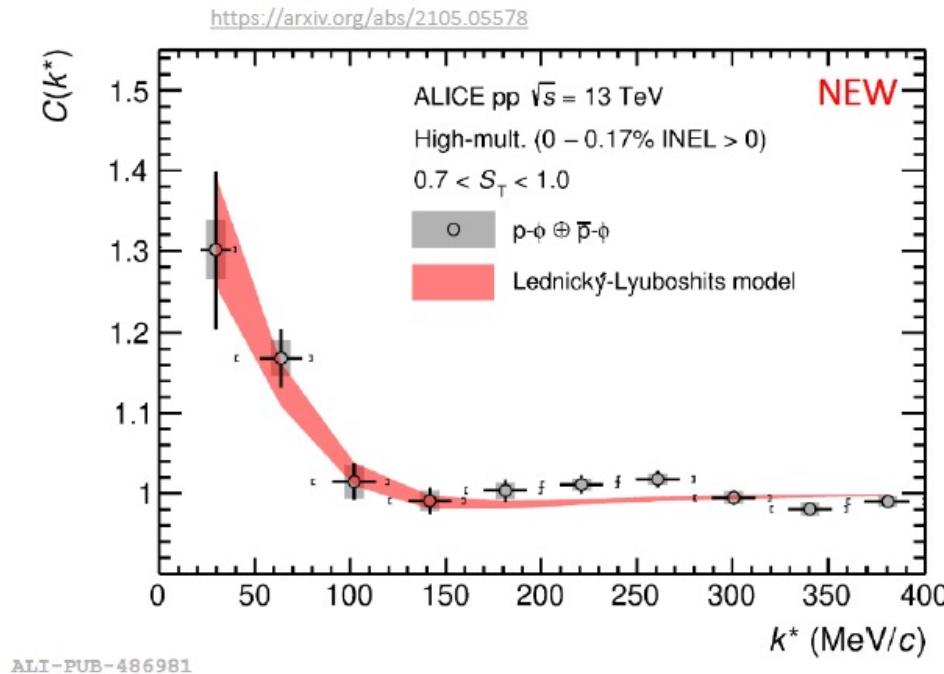
$$d_0 = 7.85 \pm 1.54(\text{stat.}) \pm 0.26(\text{syst.}) \text{ fm}$$

$$\Re(f_o) = 0.85 \pm 0.34(\text{stat.}) \pm 0.14(\text{syst.}) \text{ fm}$$

$$\Im(f_o) = 0.16 \pm 0.10(\text{stat.}) \pm 0.09(\text{syst.})$$

Large range
Dominant elastic interaction!

Proton- ϕ Interaction



Extracted genuine correlation function
After the subtraction of combinatorics, mini-jets and unfolding
according to the genuine p- ϕ contribution.

2) Yukawa-type potential with real parameters

$$V(r) = -A \frac{e^{\alpha r}}{r}$$

Correlation obtained with CATS

Fit Results:

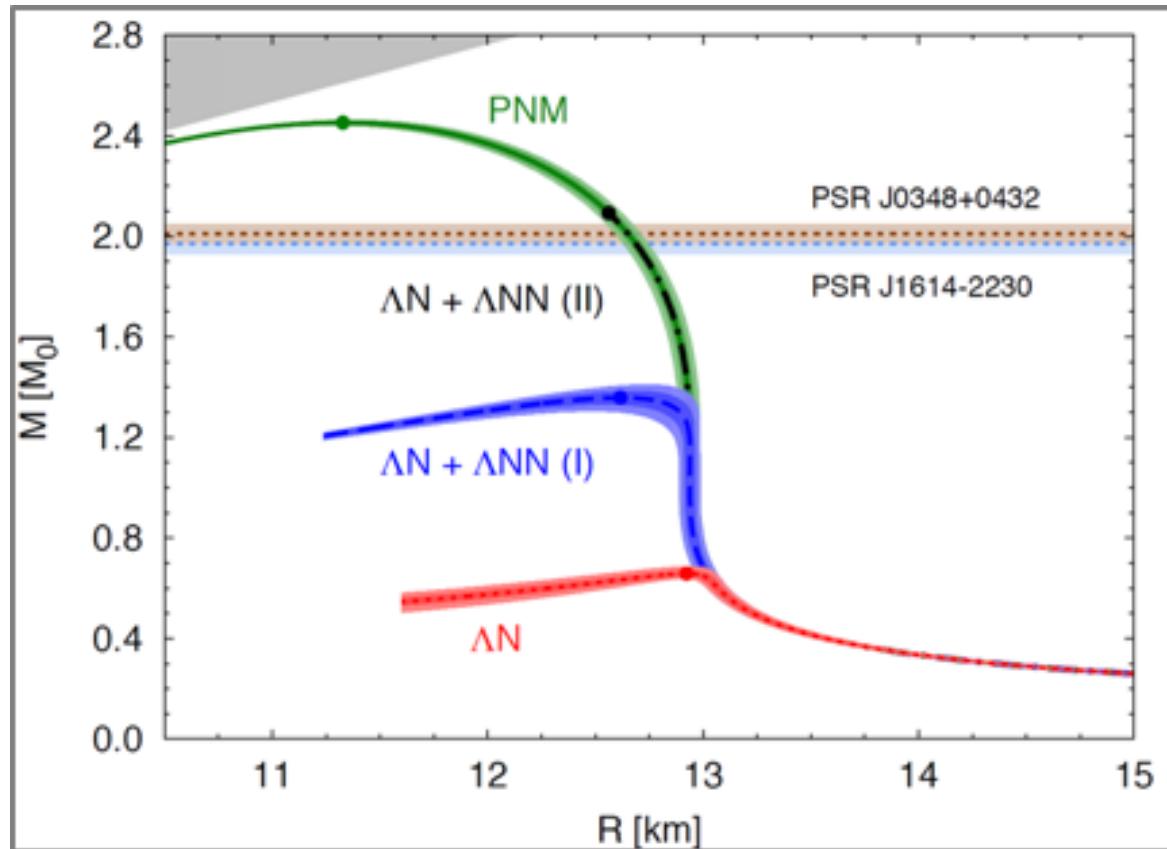
$$A = 0.021 \pm 0.009(\text{stat.}) \pm 0.006(\text{syst.})$$

$$\alpha = 65.9 \pm 38.0(\text{stat.}) \pm 17.5(\text{syst.})$$

Extraction of the N- ϕ coupling constant as \sqrt{A}

$$g_{\phi N} = 0.14 \pm 0.03(\text{stat.}) \pm 0.02(\text{syst.})$$

Effect of the 3-body Forces



Repulsive 3-body forces can contribute to stiffen the hyperon stars equation of state

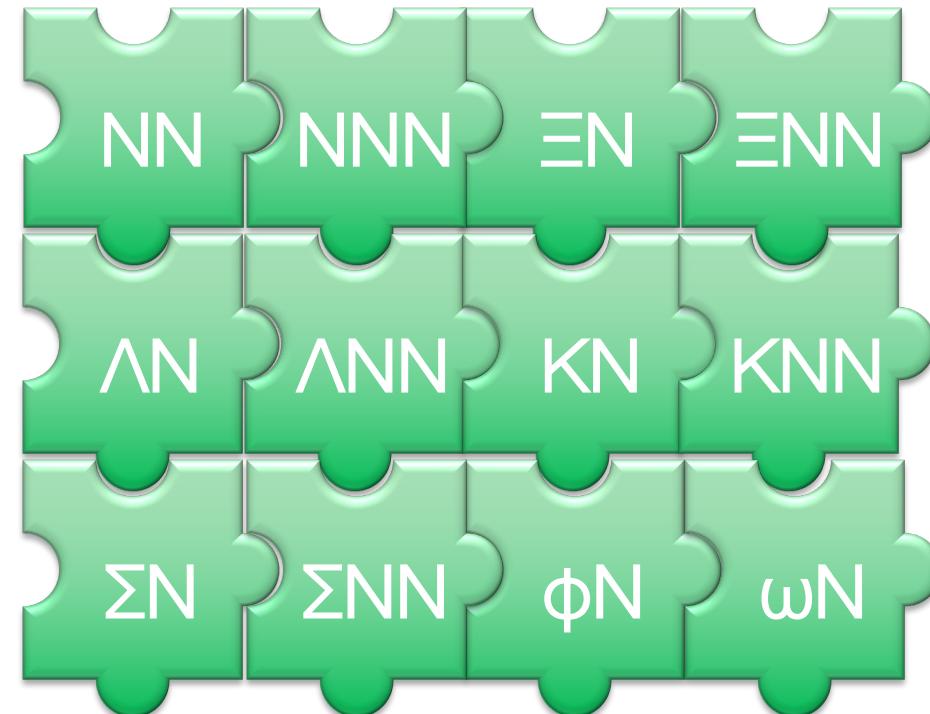
Prediction from Quantum Montecarlo adjust repulsive 3-body forces to hypernuclei binding energies

Current χ EFT calculations predict an overall attractive 3-body contribution and a repulsive 2-body core

(S. Petschauer et al Theory. Front. Phys. 8 (2020) 12)
(D.Logoteta et al Eur. Phys. J. A (2019) 55: 207)

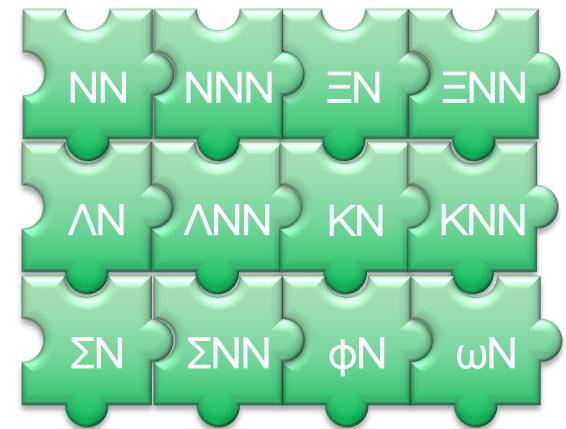
Need a direct measurement!
Talk by Dr. Raffaele Del Grande Tue 19:00

... towards the solution of the puzzle



Summary

- Femtoscopy as a tool to provide unprecedented constraints on hadron-hadron interactions
- Most precise data on several YN interactions providing constraints for hadron-hadron potentials
- First measurement of the proton- ϕ interaction
- More precision studies within reach with the large data samples collected in Run 3 & 4
 - Direct measurements of three-body interactions for the first time
 - Eventually solve the hyperon puzzle...
 - Probing QCD bound states



Thank you!