

Exploring a new g/h separation models on the HAWC Observatory

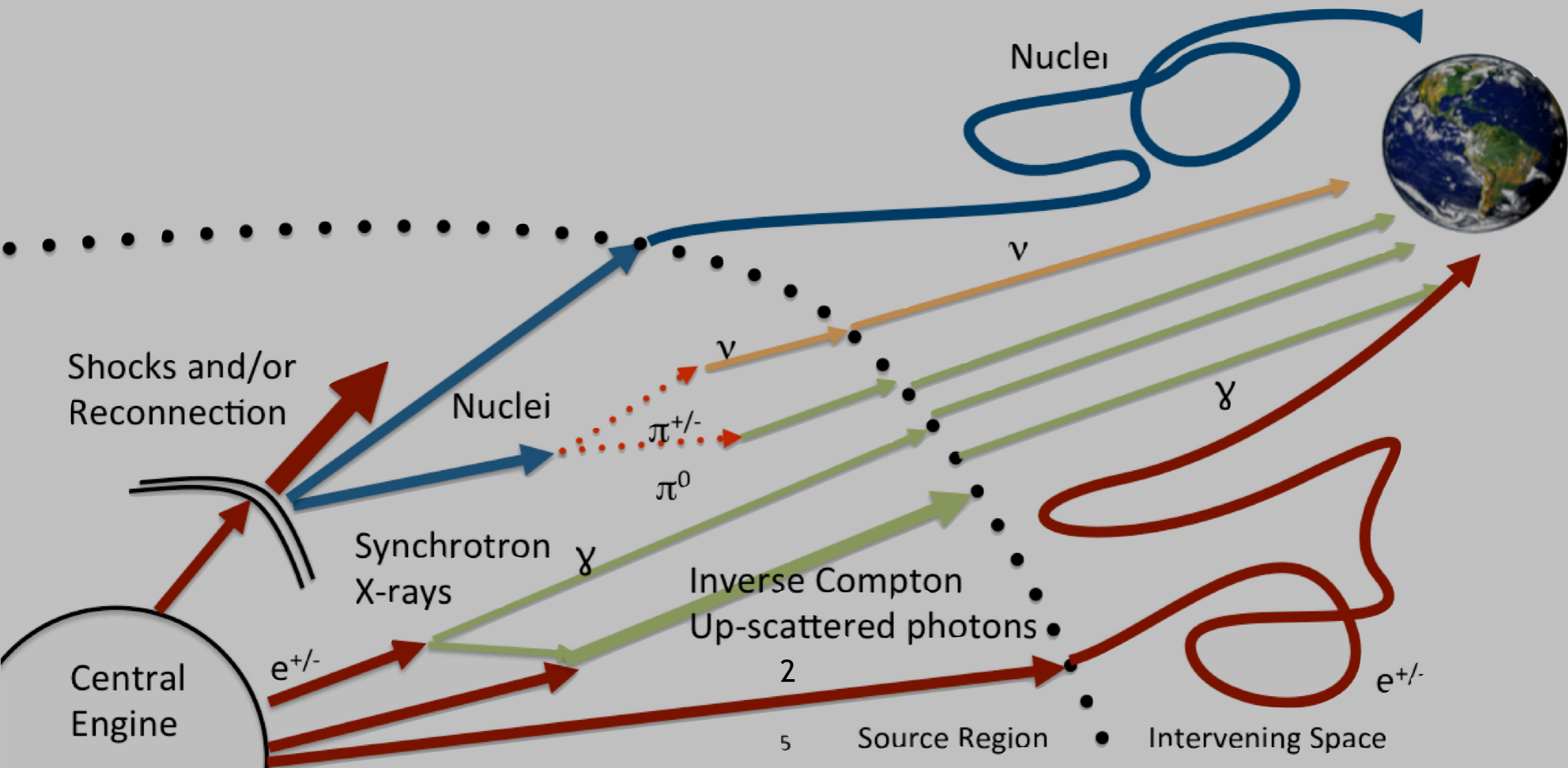
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INAOE-MSU-HKU

Meeting of the Cosmic Rays Section of the Mexican Physical Society

November 26th, 2019

Why study Gamma-rays?

Gamma-rays is not deflected by magnetic fields.



Pretz, J. (2015), Highlights from the High Altitude Water Cherenkov Observatory.

Gamma-ray Observatories

Wide-field
Continuous Operation



Fermi
AGILE
EGRET

TeV Sensitivity



HAWC
ARGO
Milagro



VERITAS
HESS
MAGIC
FACT

Pretz, J. (2015), Highlights from the High Altitude Water Cherenkov Observatory.



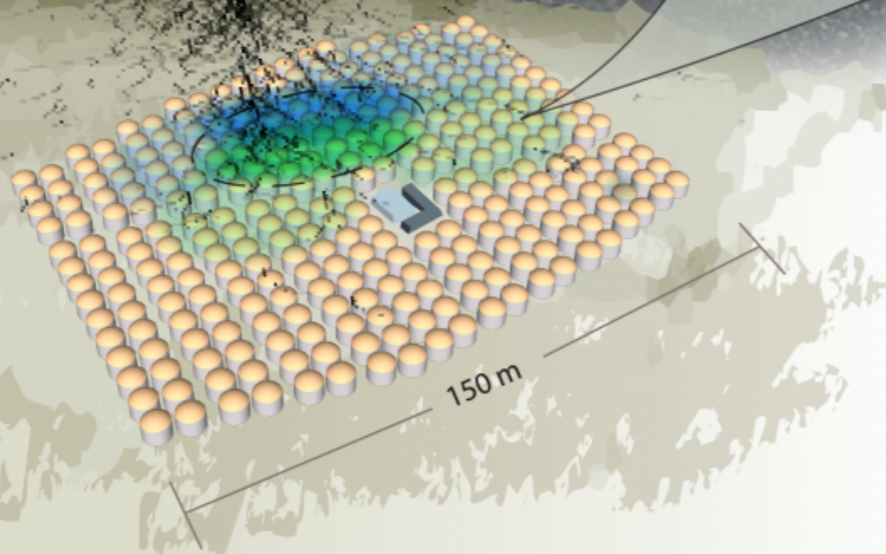
Mapping the Northern Sky in High-Energy Gamma Rays

HAWC Observatory

HAWC operates day and night, providing a large field of view for the observation of the highest energy gamma rays.



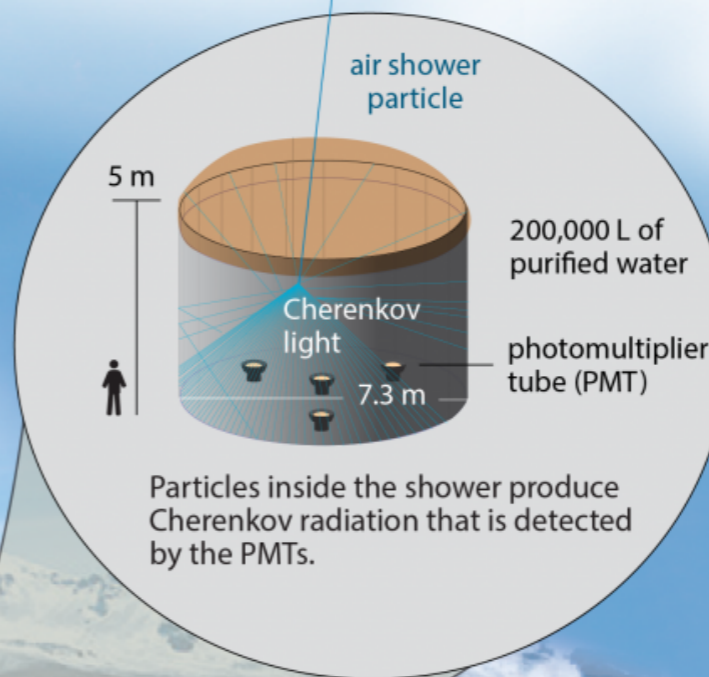
Pico de Orizaba (5,626 m)



HAWC is located at 4,100 m above sea level, covering an area of 20,000 m².

Water Cherenkov tank

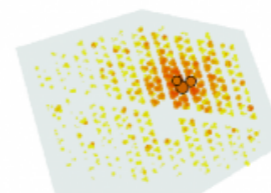
HAWC comprises an array of 300 tanks that record the particles created in gamma-ray and cosmic-ray showers.



Gamma rays vs cosmic rays

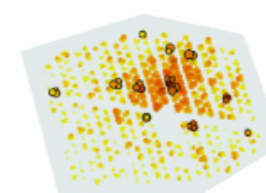
HAWC selects gamma rays from among a much more abundant background of cosmic rays.

gamma-ray shower



"hot" spots concentrate around the core

cosmic-ray shower



"hot" spots are more dispersed



Mapping the Northern Sky in High-Energy Gamma Rays

HAWC Observatory

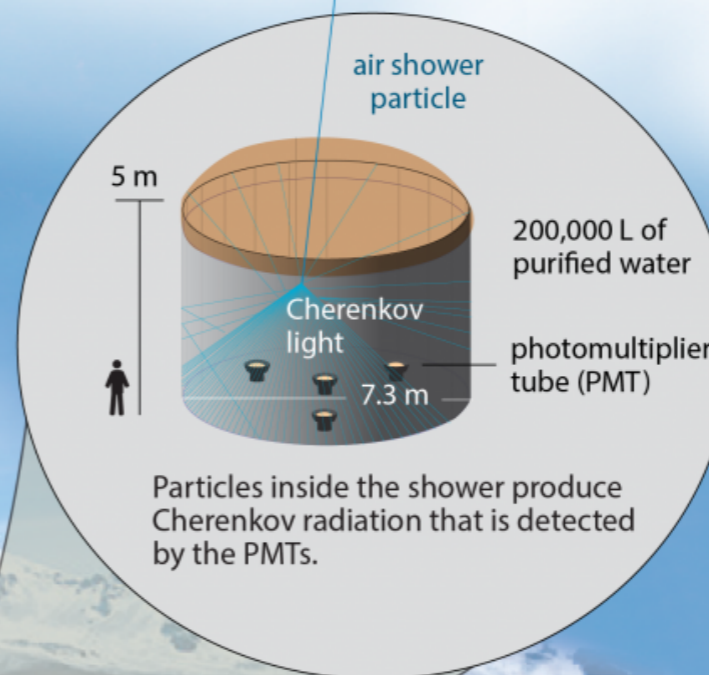
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Pico de Orizaba
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Water Cherenkov tank

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Gamma rays vs cosmic rays

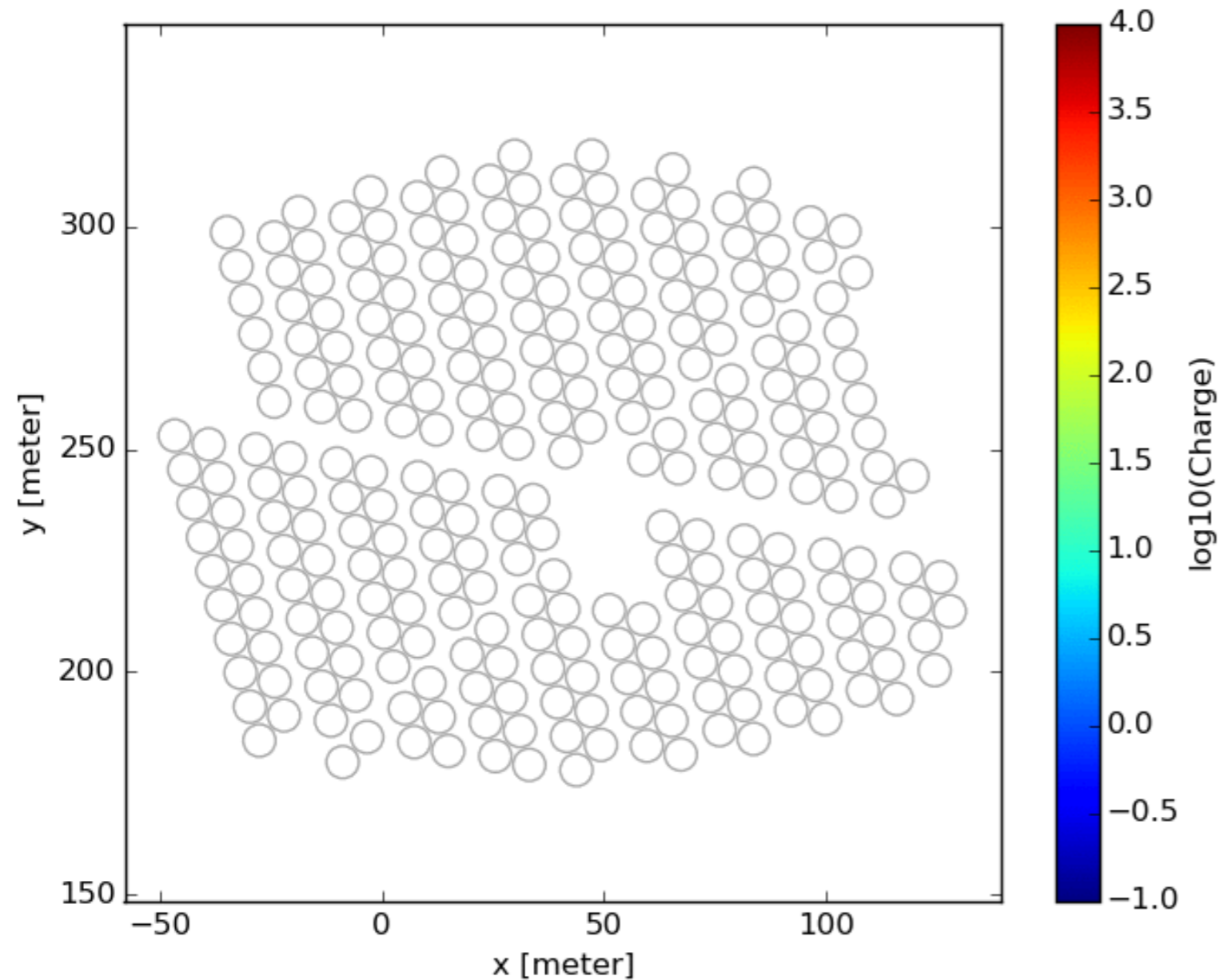
HAWC selects gamma rays from among a much more abundant background of cosmic rays.

HAWC is located at 4,100 m above sea level, covering an area of 20,000 m².

Main background: Hadronic cosmic ray

- ◆ Crab nebula: 400 photons/day
- ◆ Background: 15,000 cosmic ray/second

Event simulation detected by HAWC



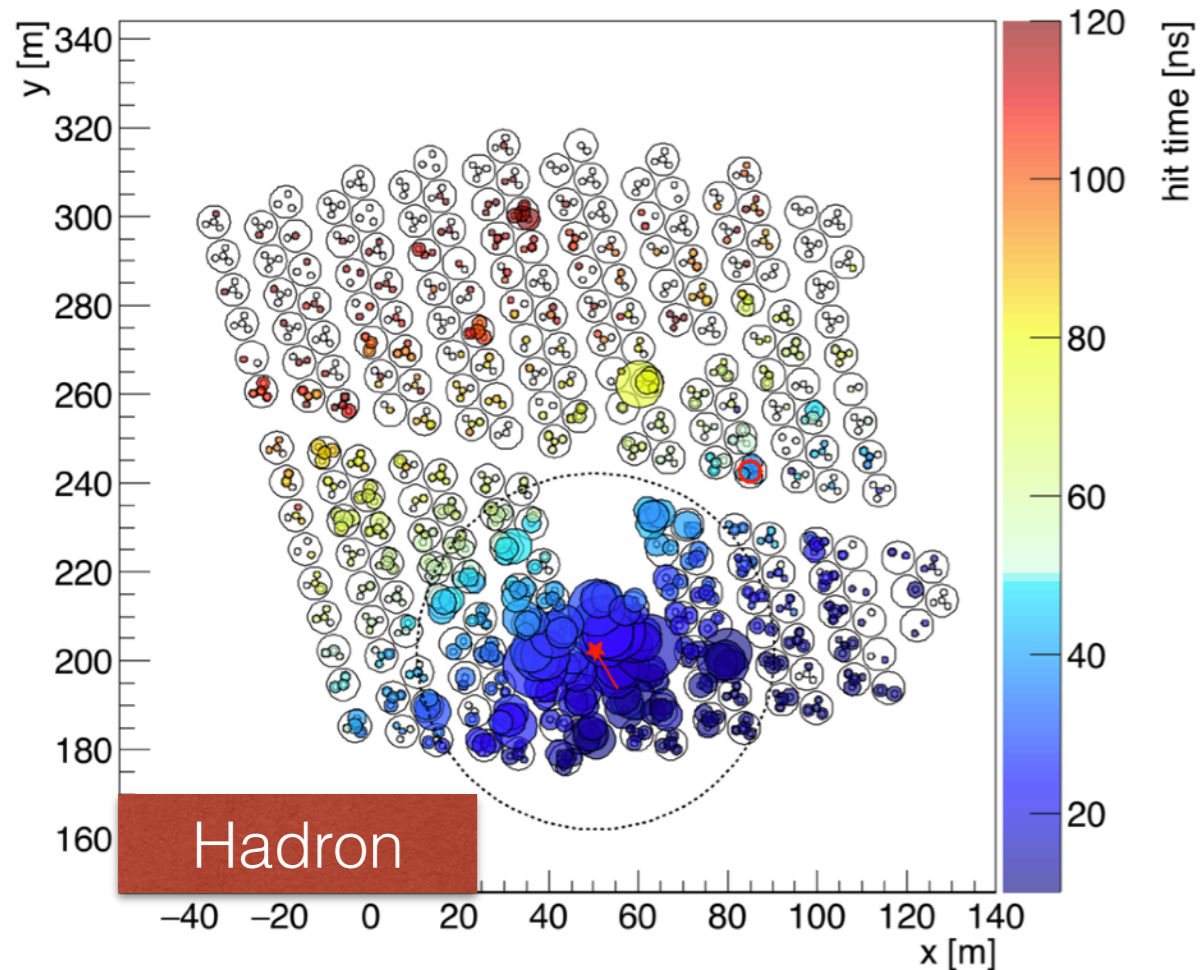
A. Timing information allows us determining where the particle comes.

B. Energy deposition in each PMT:

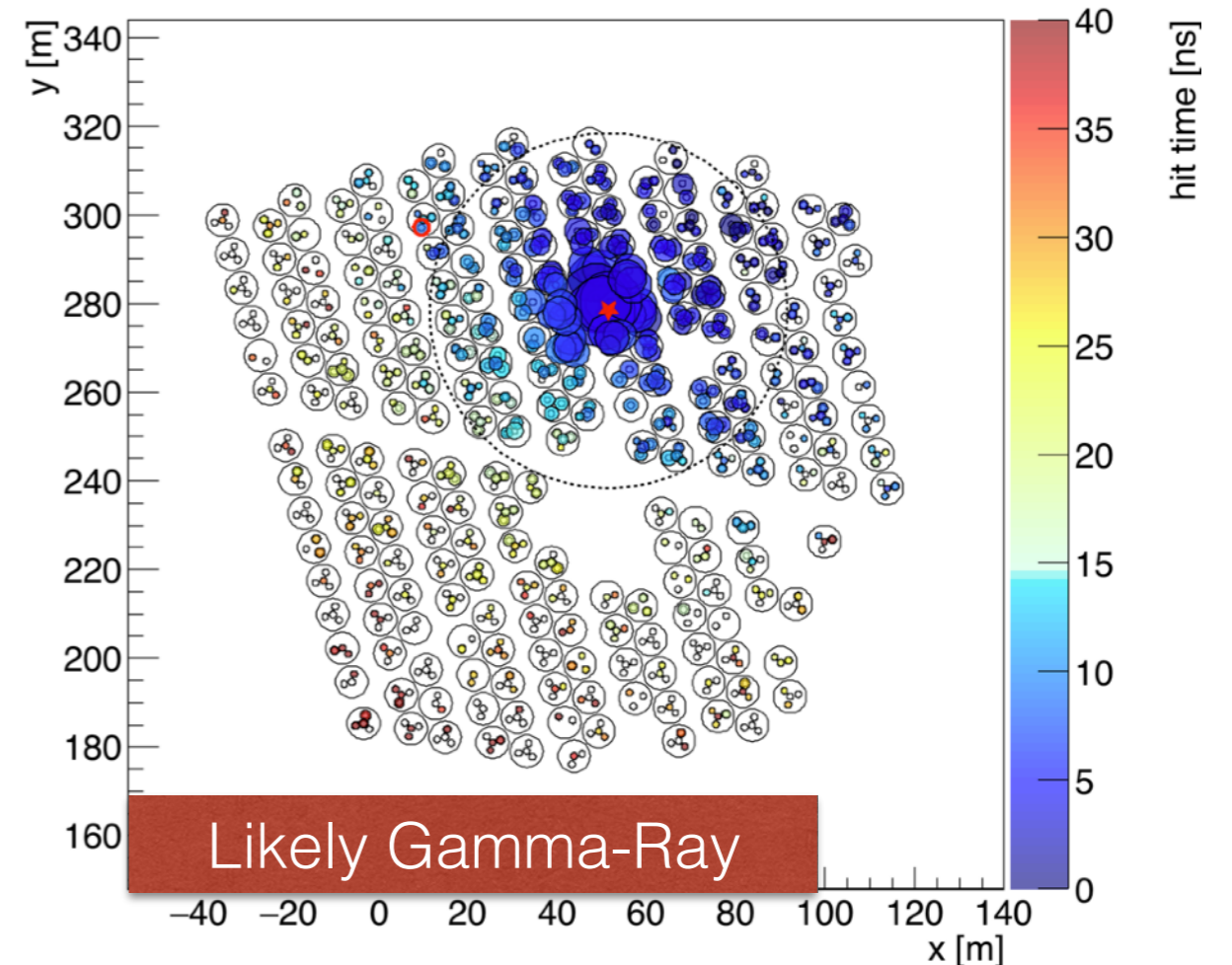
- Primary particle energy.
- The shower core.
- Gamma or Hadron?

Gamma Vs Hadron

Run 2118, TS 45004, Ev# 41, CXPE40= 55.7, Cmptrness= 10.7



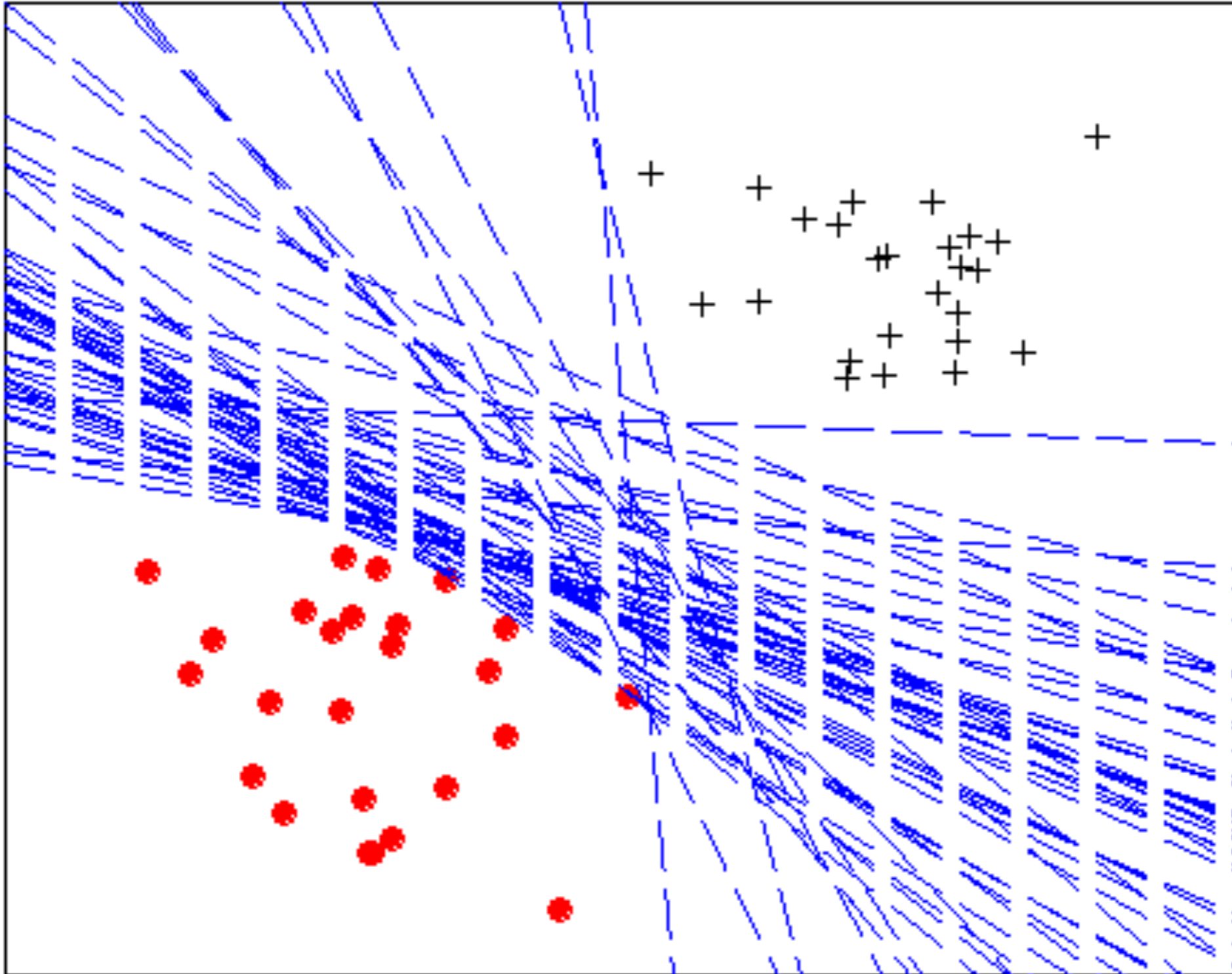
Run 2054, TS 584212, Ev# 226, CXPE40= 21.2, Cmptrness= 28.3



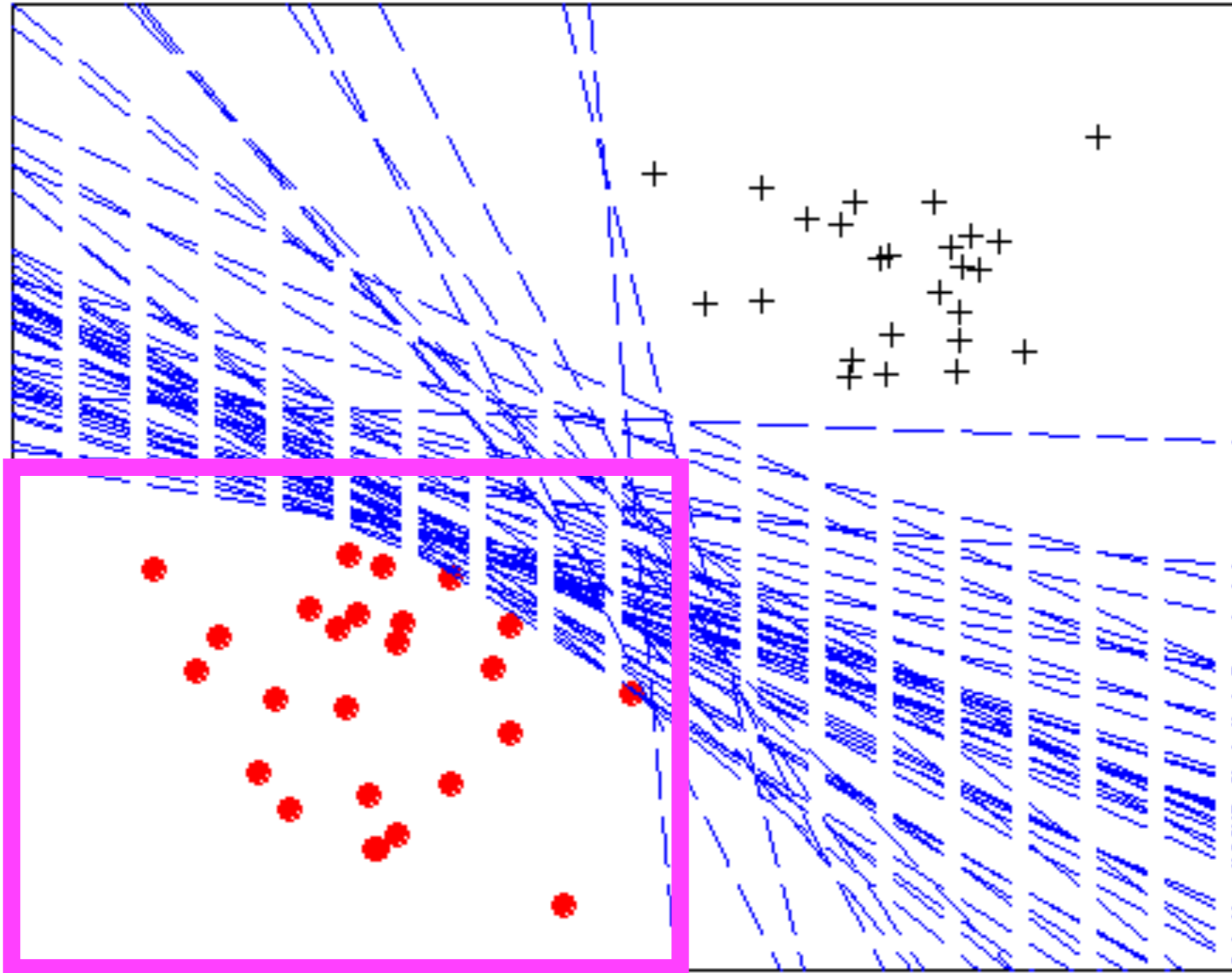
Task: Distinguishing between gammas and hadrons

<http://www.hawc-observatory.org/observatory/ghsep.php>

How recognize the particle?



Rectangle cut



Standard cuts

It is the official method in HAWC Observatory. It can describe as rectangle cut.

$$PINC \leq Cut_{PINC} \ \&\& \ LiC \leq Cut_{LiC}$$

Where

$$LiC = \text{Log}_{10}\left(\frac{CxPE40}{nHitSP20}\right)$$

$$PINCness = \frac{\sum_{i=0}^N \frac{(\log q_i - \log \bar{q}_i)^2}{\sigma_{\log q_i}^2}}{N}$$

Bins:

- 1. **fhit:** nHitSP20/nChAvail
- 2. **ebin:** logNNenergyV2

The fraction of the PMTs hit

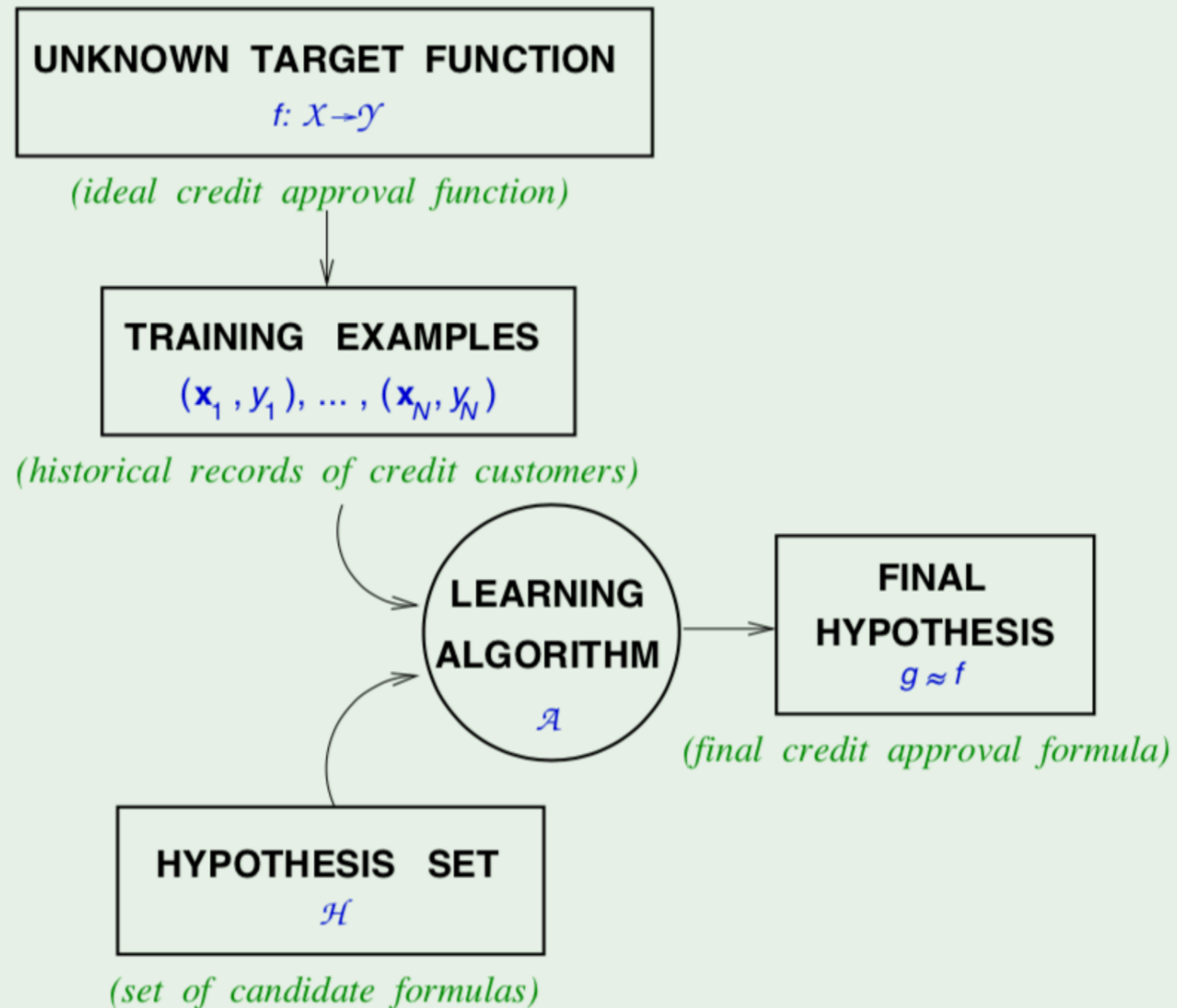
Energy estimator using a Neural Network

fhin	min fhin	max fhin
0	4.4%	6.7%
1	6.7%	10.5%
2	10.5%	16.2%
3	16.2%	24.7%
4	24.7%	35.6%
5	35.6%	48.5%
6	48.5%	61.8%
7	61.8%	74.0%
8	74.0%	84.0%
9	84.0%	100.0%

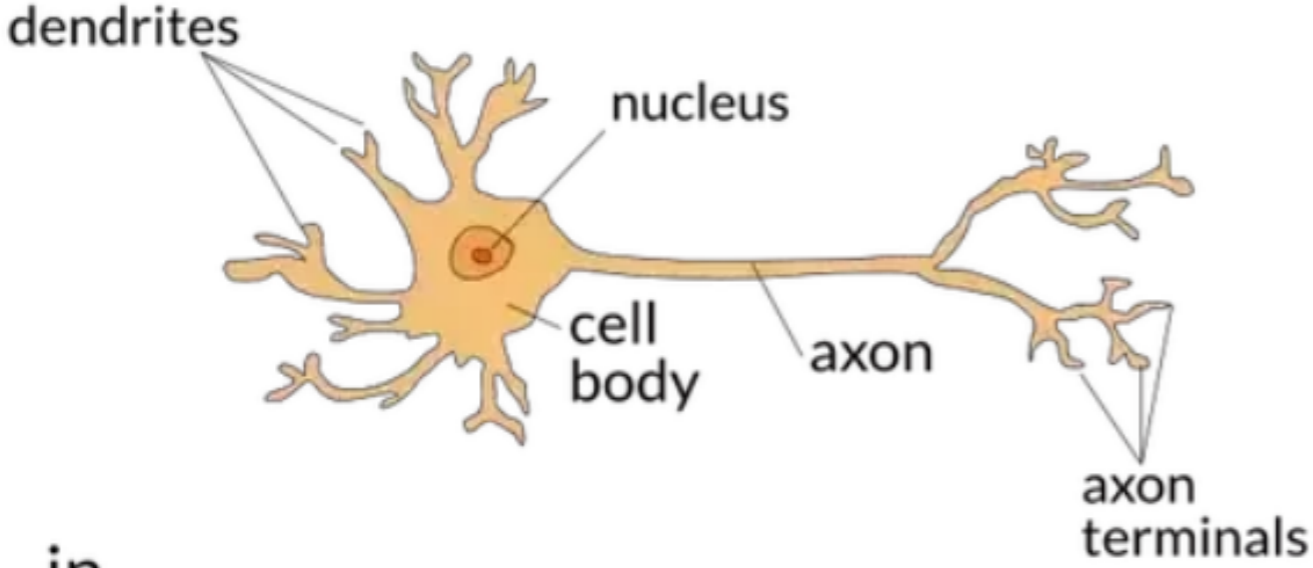
ebin	min ebin	max ebin	min ebin (GeV)	max bin (Gev)
0	2.50	2.75	316.23	562.34
1	2.75	3.00	562.34	1000.00
2	3.00	3.25	1000.00	1778.28
3	3.25	3.50	1778.28	3162.28
4	3.50	3.75	3162.28	5623.41
5	3.75	4.00	5623.41	10000.00
6	4.00	4.25	10000.00	17782.79
7	4.25	4.50	17782.79	31622.78
8	4.50	4.75	31622.78	56234.13
9	4.75	5.00	56234.13	100000.00
10	5.00	5.25	100000.00	177827.94
11	5.25	5.50	177827.94	316227.77

$$\hat{E} = \log_{10}(E / 1 \text{ GeV})$$

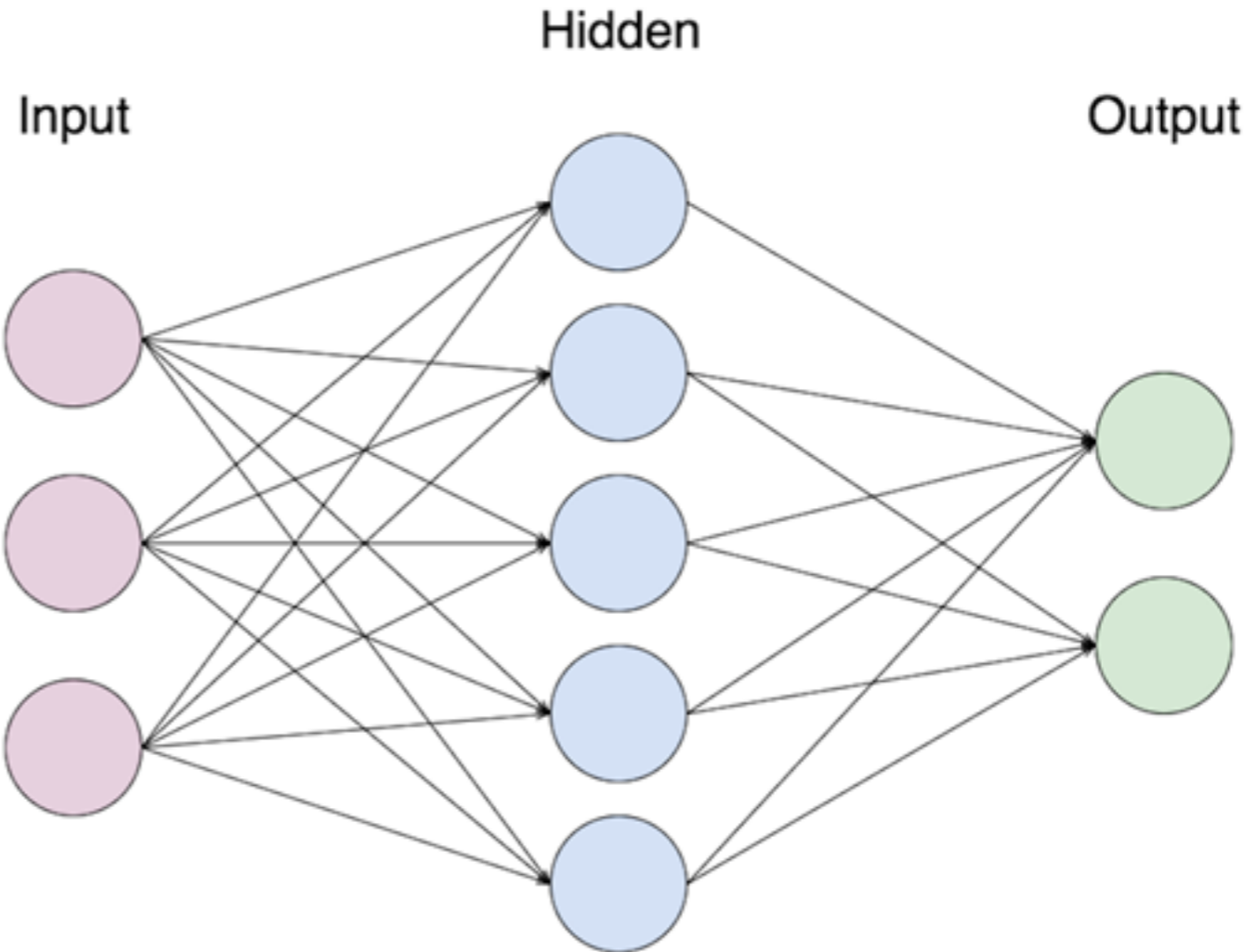
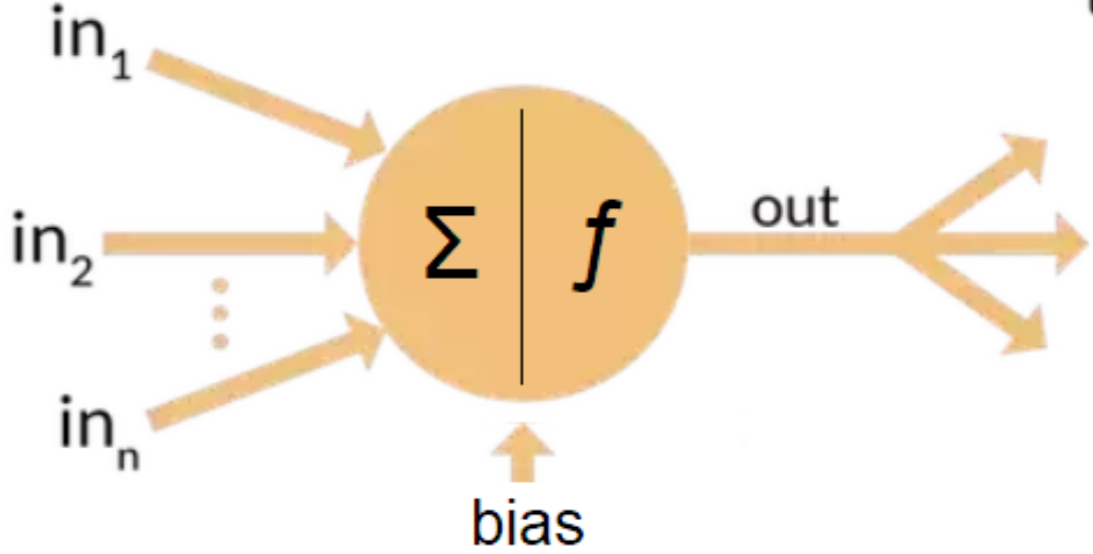
Learning from data



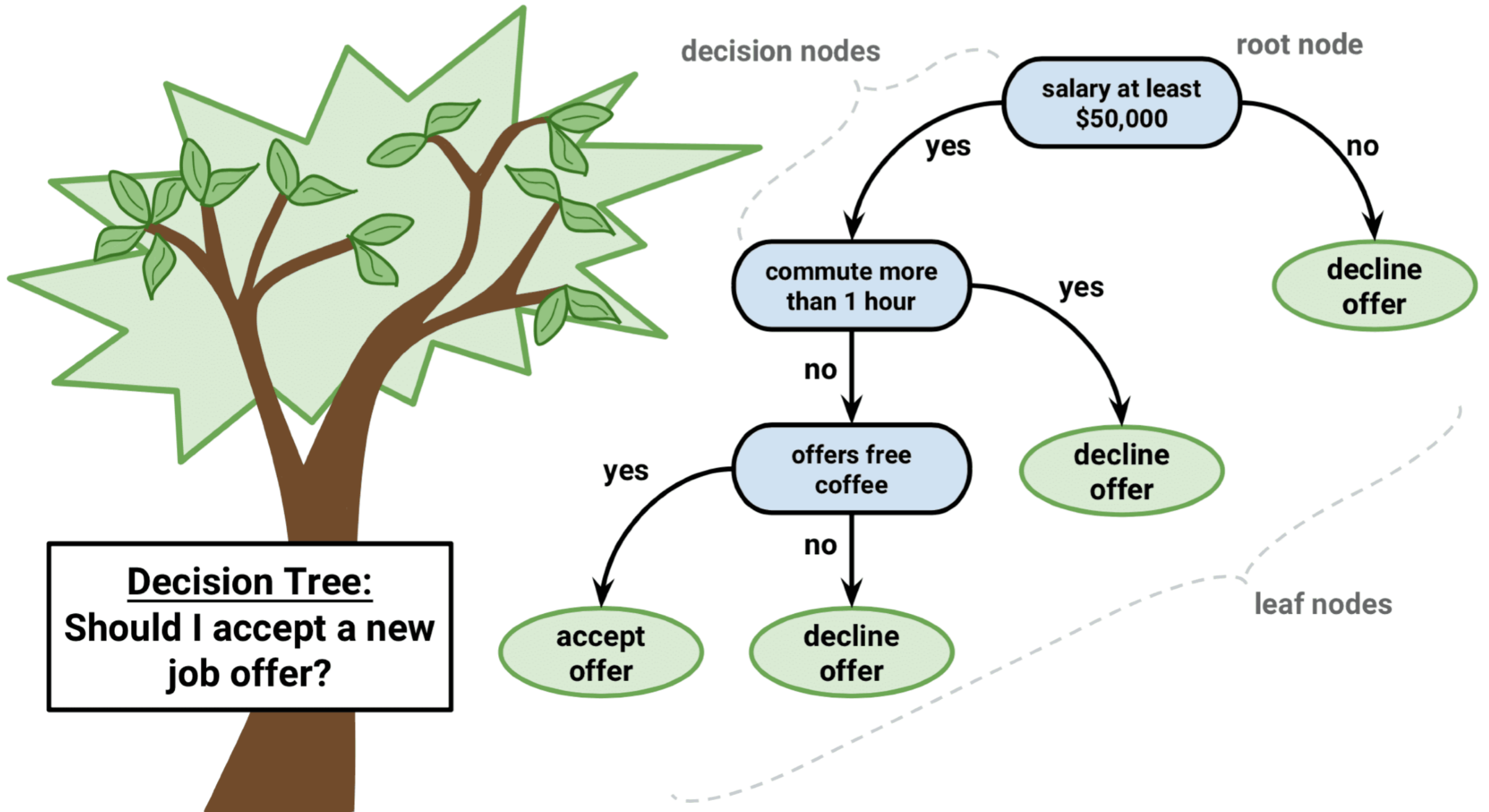
Neural Network



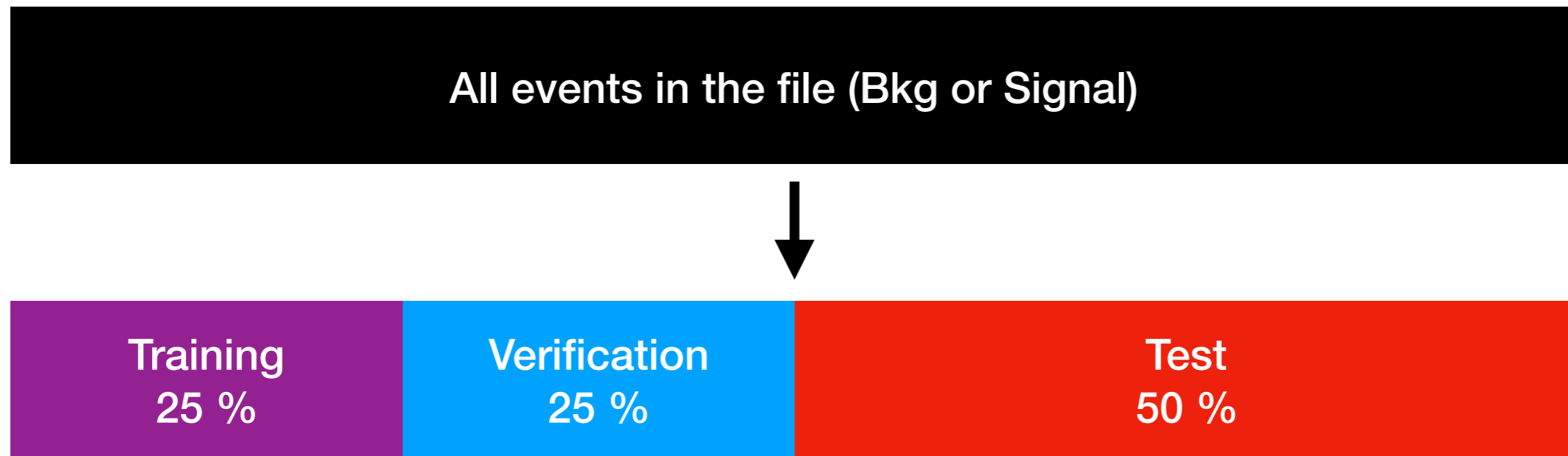
~



Boosted Decision Tree



Train a NN and BDT



A. Input parameters:

- $LIC = \log_{10}(CxPE40 / nHitSP20)$
- PINC
- logNNEnergyV2
- disMax
- LDFAmp
- LDFChi2
- $fbin = nHitSP20 / nChAvail$

MLT configuration:

Neural Network (NN):

B. **Architecture:** 7 : 10 : 10 : 1

C. Models trained with **TMVA**

Boosted Decision Tree (BDT):

B. **Model** with 500 tree

C. Models trained with python (**Xgboost** package)

Both Models

D. Don't use Physical weight

E. Models trained

- Low fbin: 0.044 to 0.162
- Medium fbin: 0.162 to 0.485
- High fbin: 0.485 to 1.000

F. Apply Quality cuts

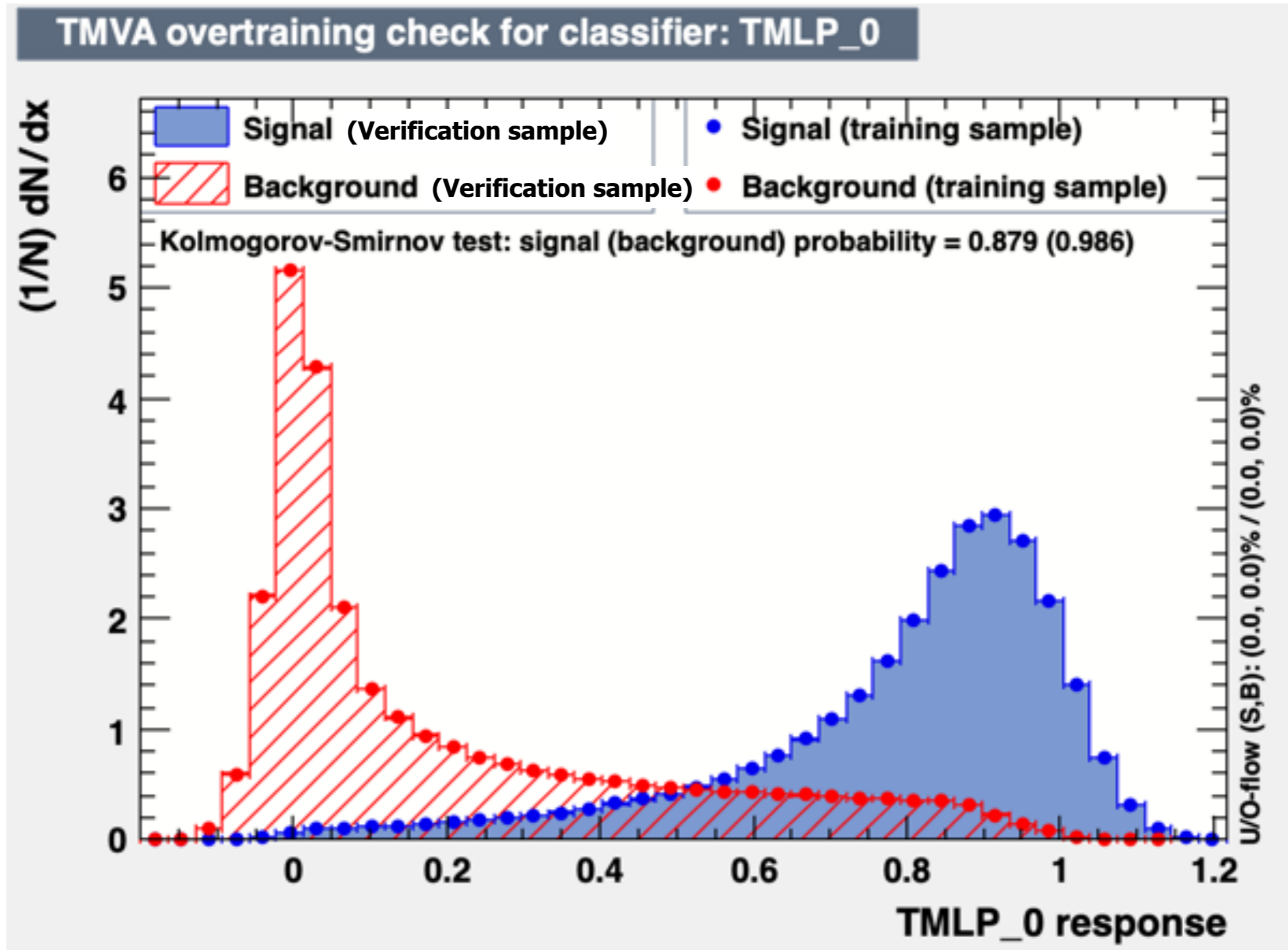
- `rec.angleFitStatus==0`
- `rec.coreFitStatus==0`

- `rec.nChTot >= 800`
- `rec.nChAvail > 0.9 * rec.nChTot`

G. Target

- `Signal = 1`
- `Background = 0`

After Training: NN model for low fbin

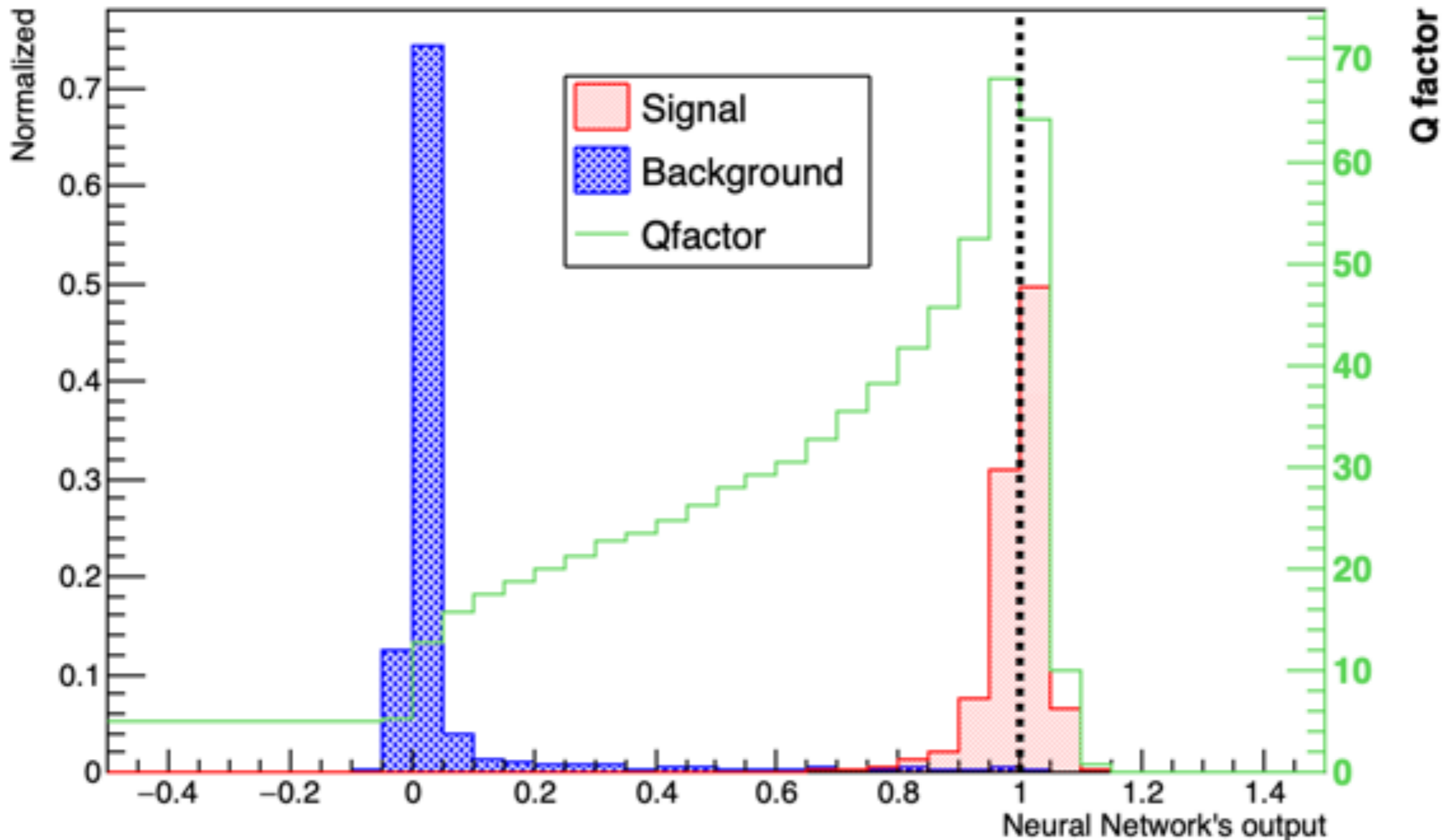


Find the cuts

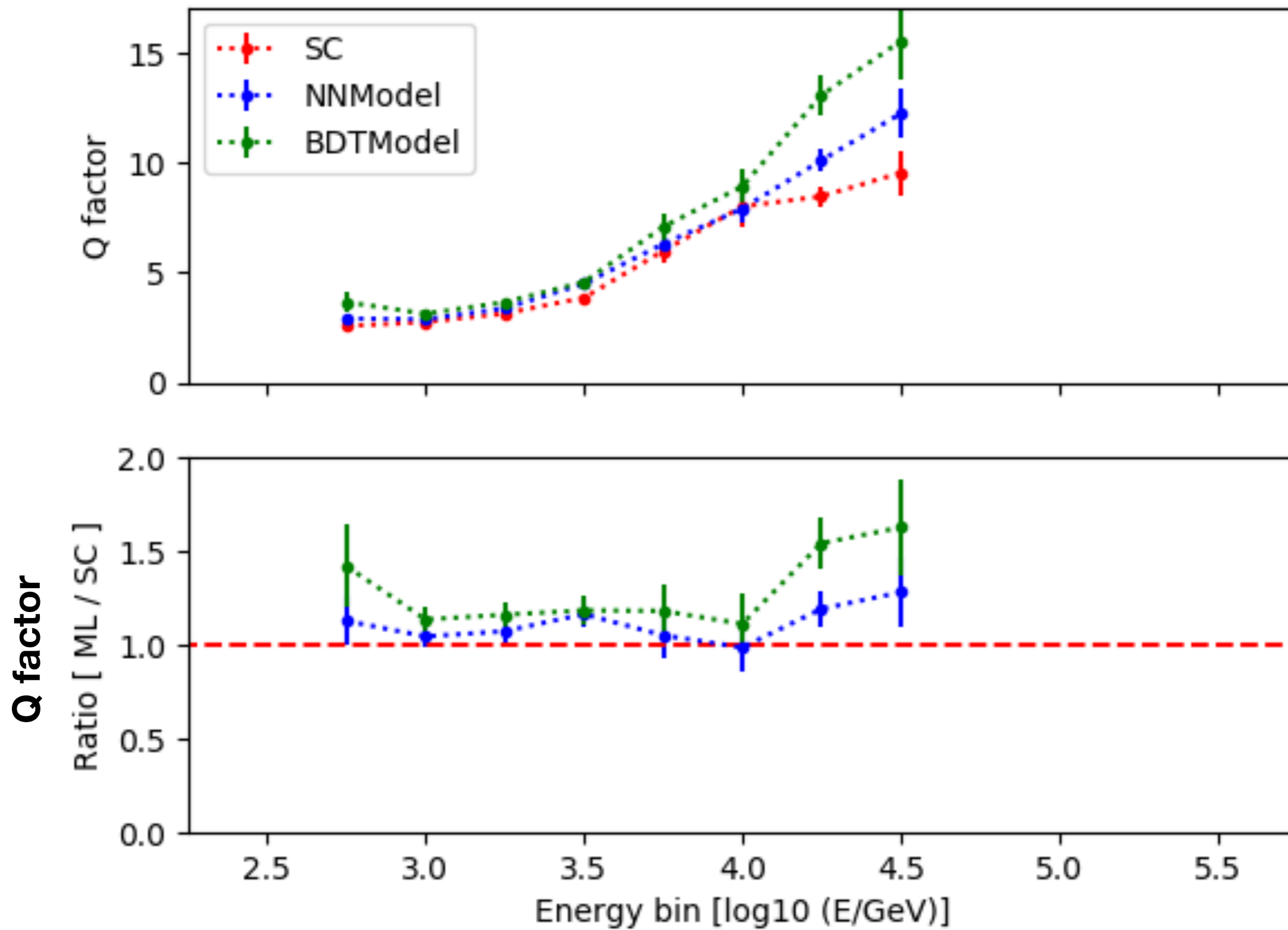
Conditions:

- Gamma efficiency $> 50\%$
- Hadron efficiency $> 0.1\%$

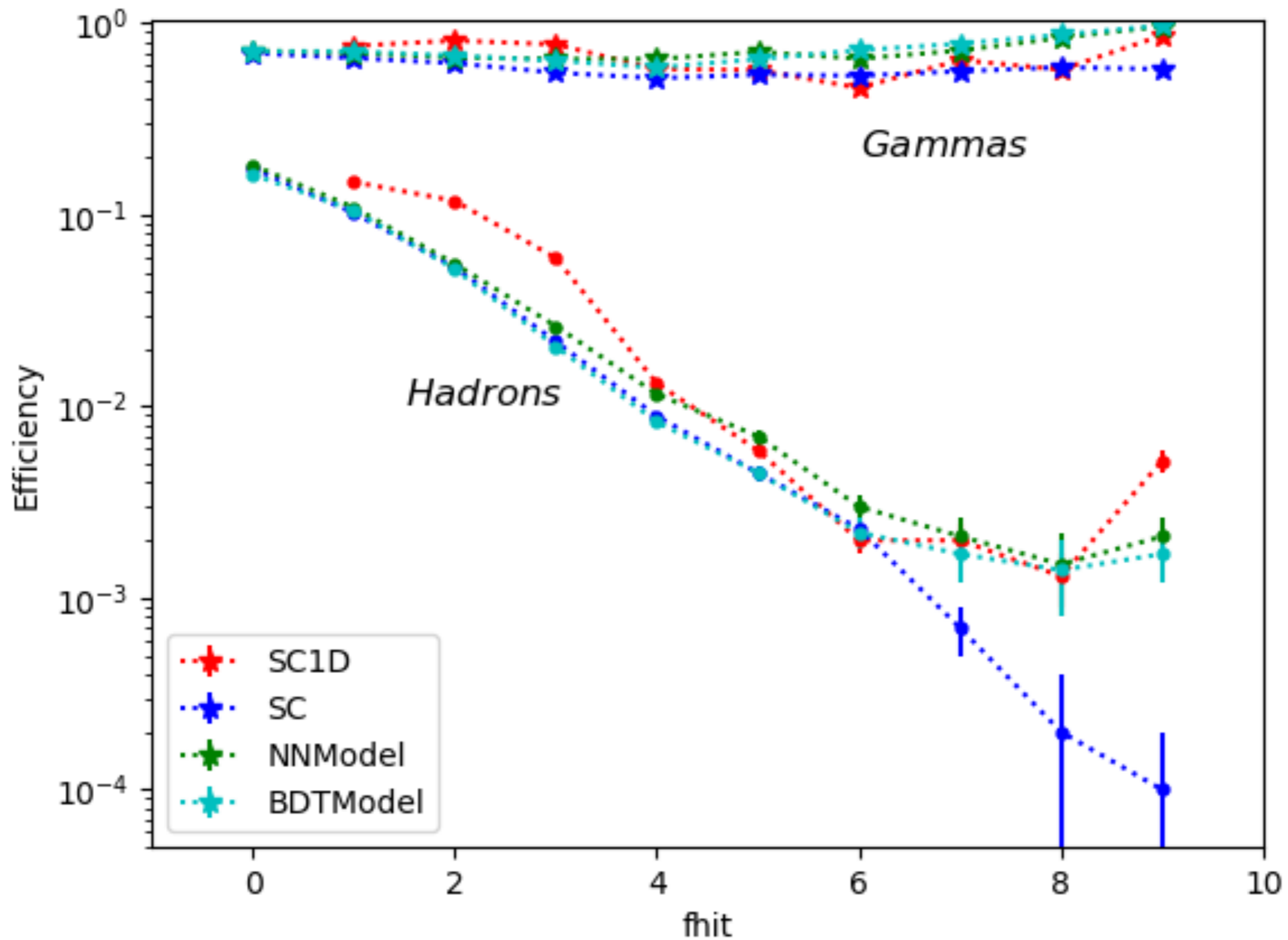
Example of fhit 7, ebin 3.75



Qfactor of fbin 3



MC Test



SC1D - <https://iopscience.iop.org/article/10.3847/1538-4357/aa7555>

Maps

Data used:

A. Period : from 2015/11/06 to 2017/12/20

B. Duration: ~837 days

1. Crab Nebula:

● RA : 83.6332

● DEC: 22.0145

2. Markarian 421:

● RA : 166.1138

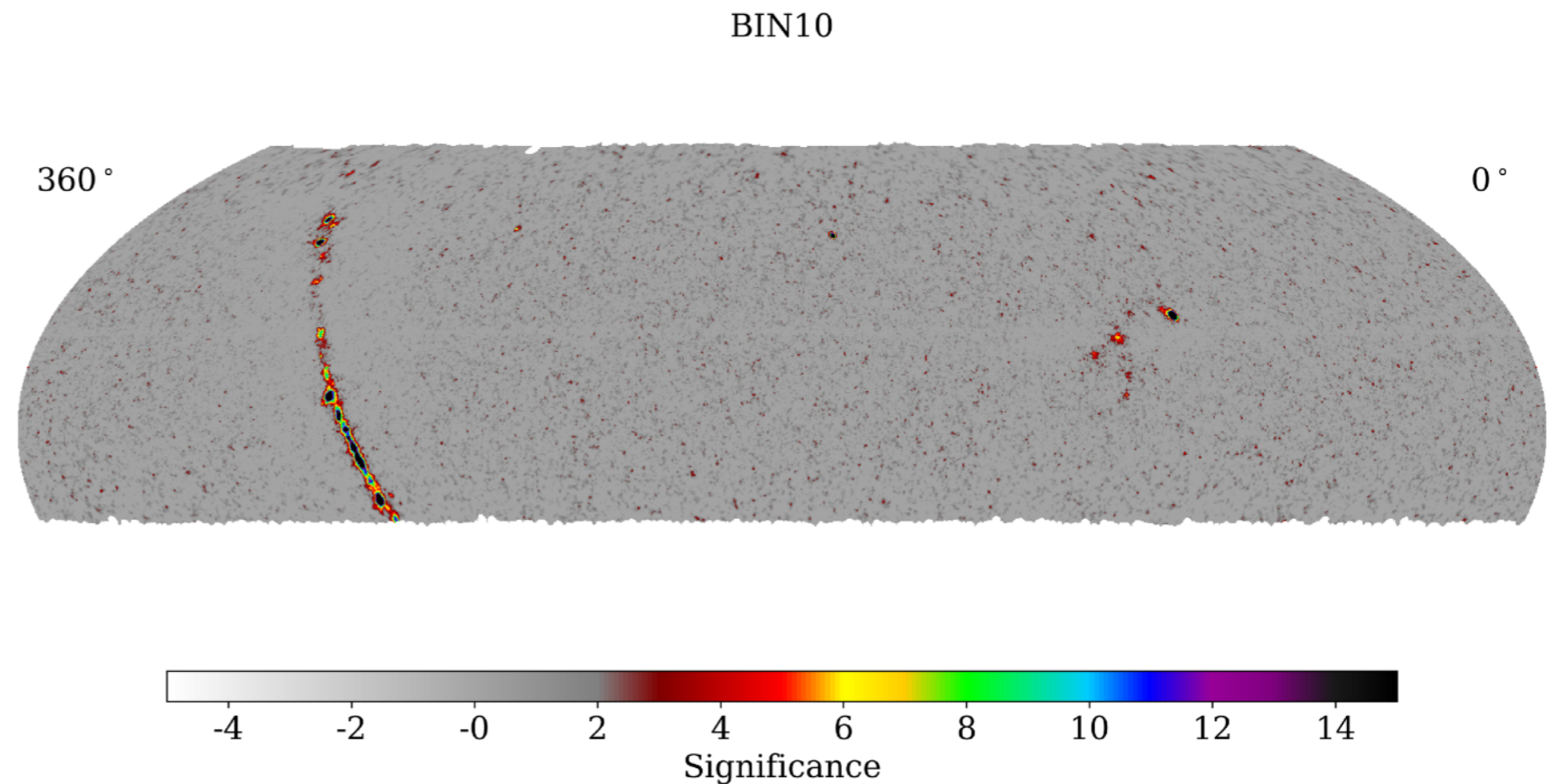
● DEC: 38.2088

3. Markarian 501:

● RA : 253.4675

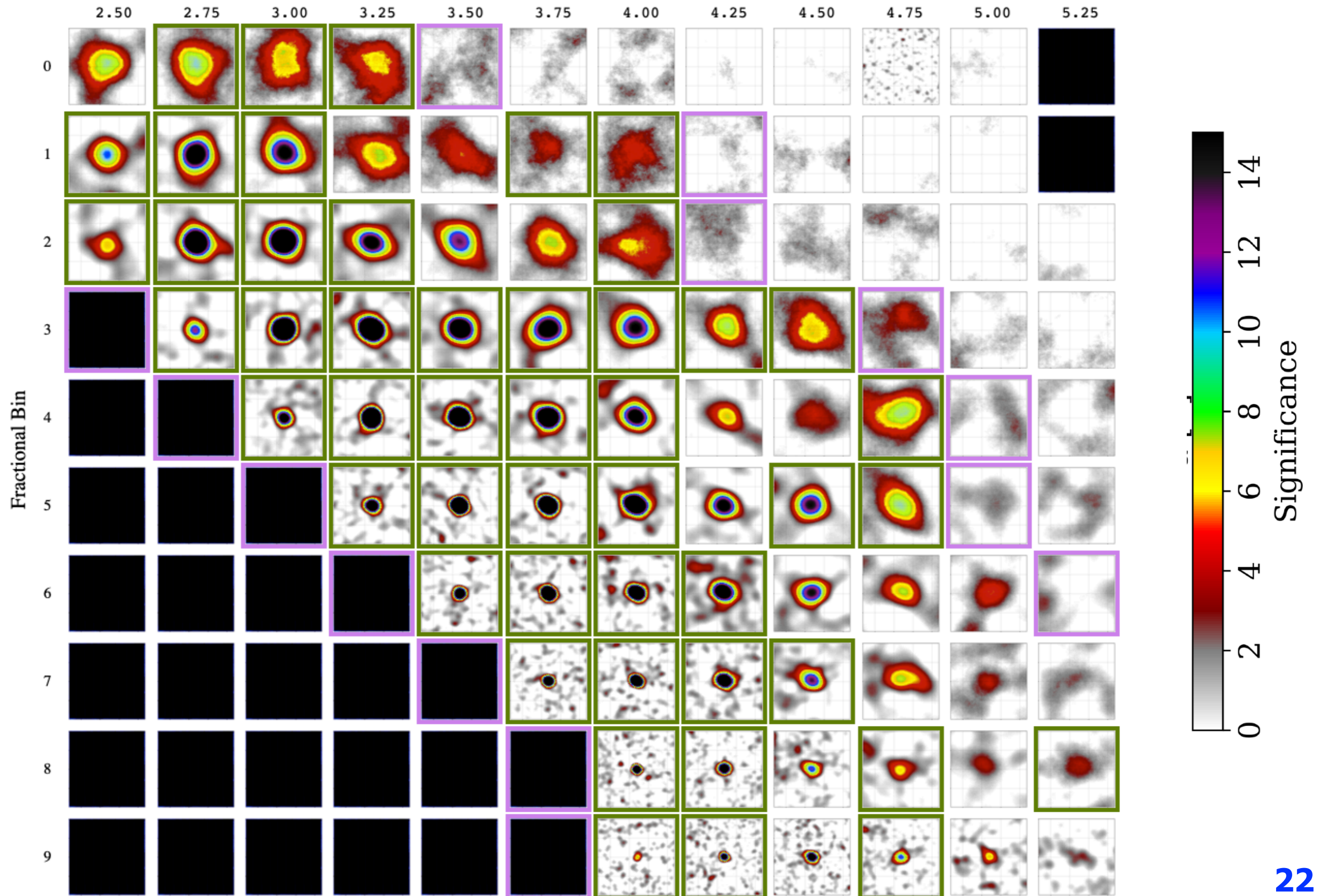
● DEC: 39.7604

4. List of 2nd HAWC Catalog (Use combine maps)

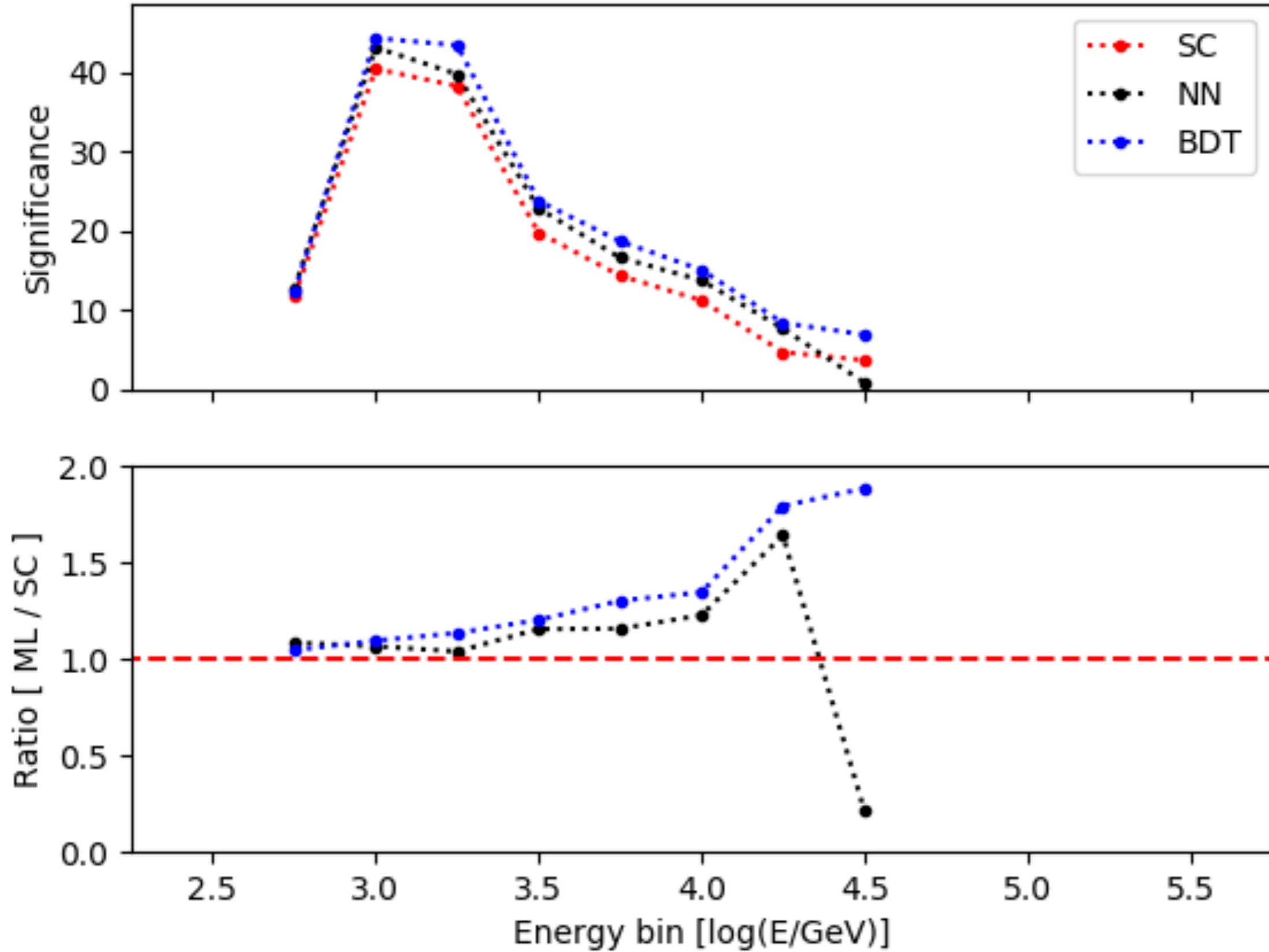


Crab Significance using BDT

Energy Bin lower bound ($\log(E/\text{GeV})$)



Significance at Crab position of fbin 3



Significance at the source position

fbin	Crab		Mrk 421		Mkr 501	
	NN / SC	BDT /SC	NN / SC	BDT /SC	NN / SC	BDT /SC
0	-3.1%	5.5%	2.0%	1.7%	-	-
1	-0.4%	2.4%	-6.2%	-2.2%	11.3%	21.1%
2	1.1%	5.1%	-4.2%	2.4%	6.2%	29.4%
3	6.0%	15.4%	3.8%	11.6%	-16.5%	-20.6%
4	9.5%	9.2%	11.6%	5.4%	19.9%	-14.0%
5	-2.2%	12.2%	1.9%	16.7%	14.4%	50.8%
6	-21.5%	7.3%	-3.6%	22.2%	-59.4%	14.1%
7	3.1%	5.5%	22.6%	15.9%	12.1%	29.5%
8	7.2%	6.4%	-10.1%	-51.3%	-13.5%	8.2%
9	9.2%	9.0%	26.1%	-60.9%	117.9%	81.3%
1-9	0.7%	9.6%	1.4%	9.0%	-4.0%	12.4%
0-9	0.7%	9.6%	1.2%	8.6%	-4.9%	11.9%

All source using official fhit

SOURCE	SC	NN	BDT	NN/SC	BDT/SC
J0534+220	155.74	156.87	170.69	0.73%	9.60%
J1104+381	35.26	35.96	38.63	1.99%	9.56%
J1825-134	31.32	34.06	35.76	8.75%	14.18%



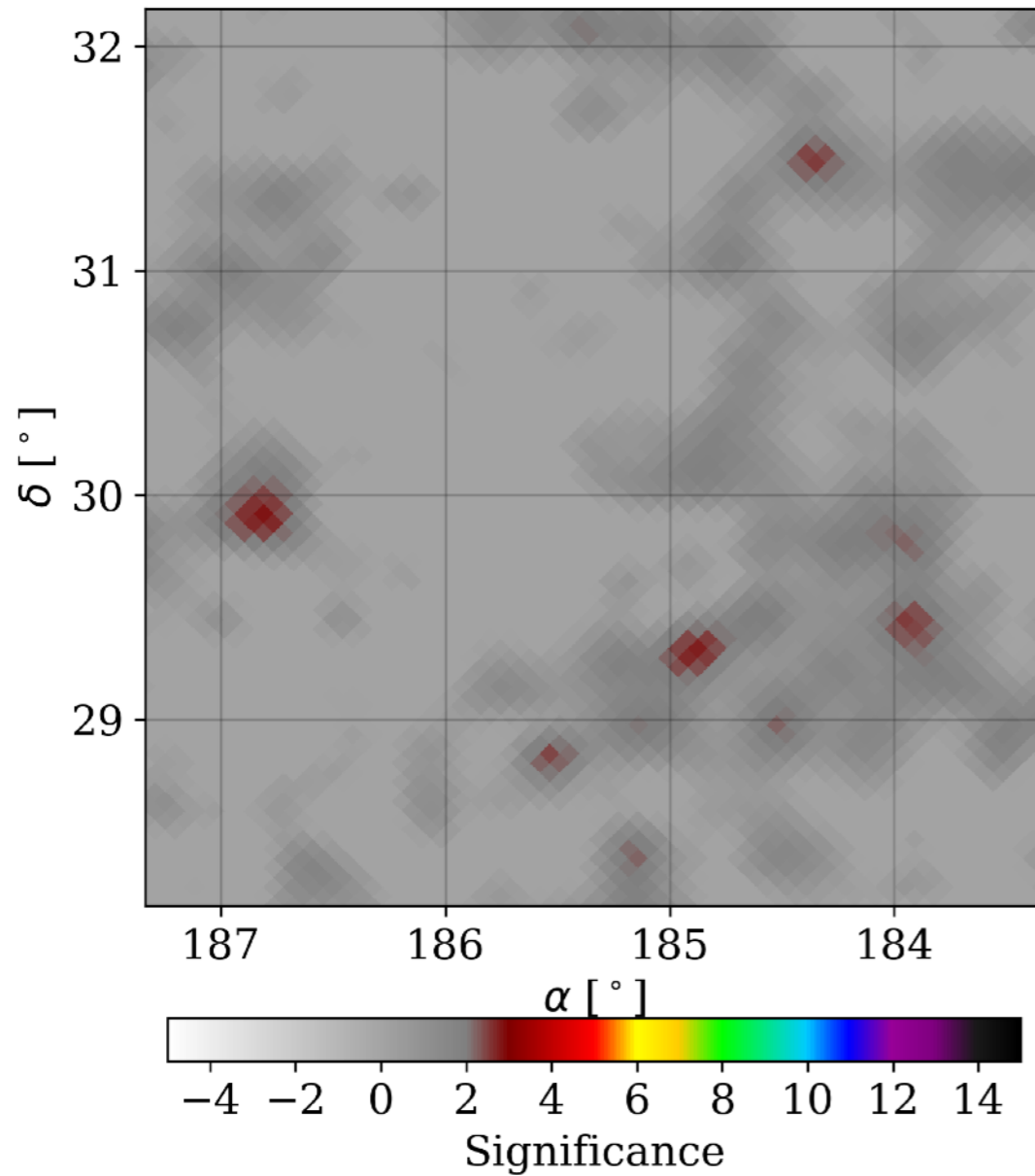
J0630+186	4.94	3.43	3.60	-30.57%	-27.13%
J2003+348	4.70	4.91	5.17	4.47%	10.00%
J1922+169	4.46	5.09	4.59	14.13%	2.91%
J1918+158	4.12	3.68	3.14	-10.68%	-23.79%



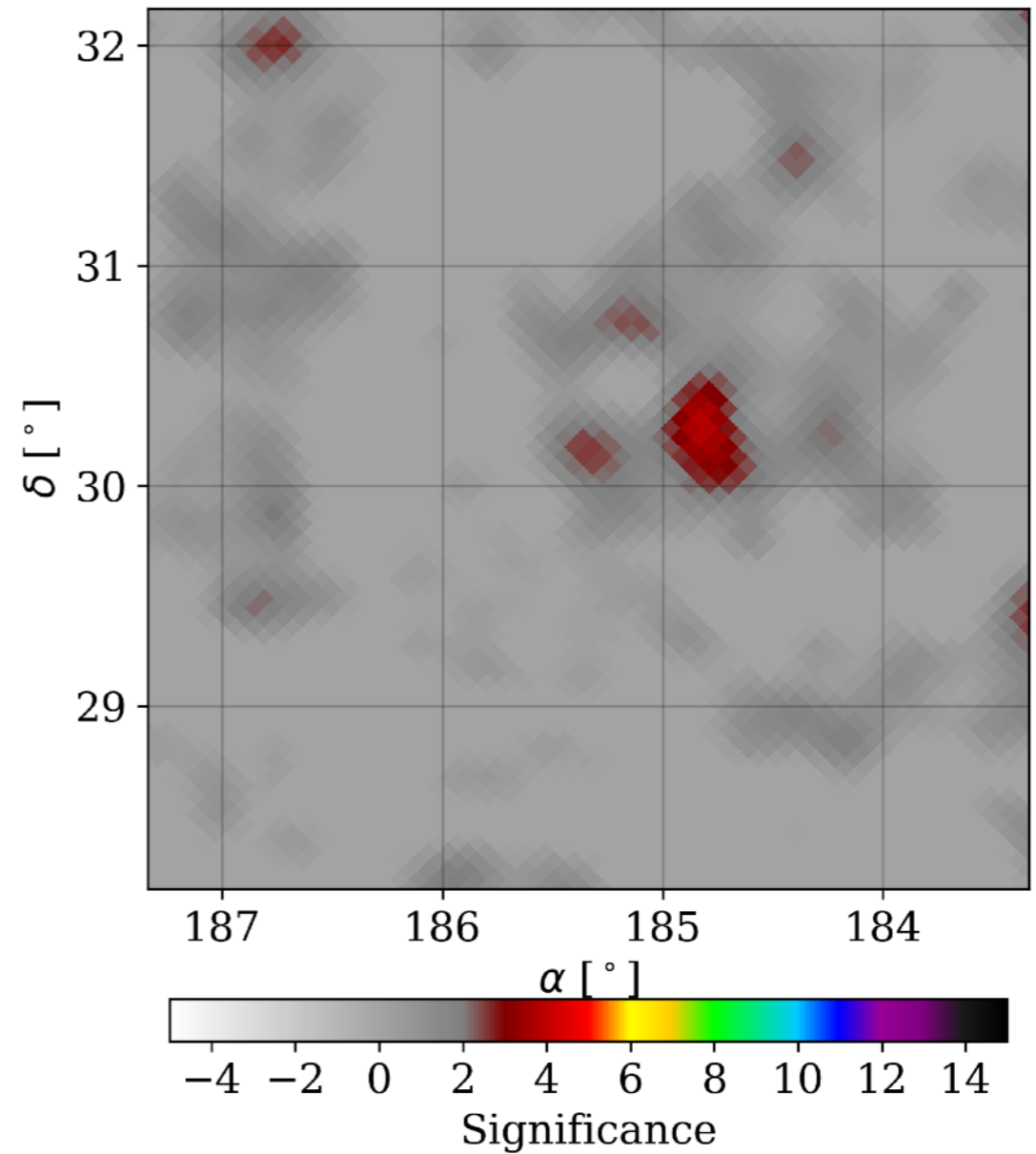
1ES_1215+303	2.20	3.89	4.02	76.82%	82.73%
J0709+108	1.96	3.16	2.90	61.22%	47.96%
PG_1218+304	1.95	3.89	4.02	99.49%	106.15%
1ES_2344+514	1.83	1.54	3.76	-15.85%	105.46%

PG 1218+304 using bin 1-9

SC

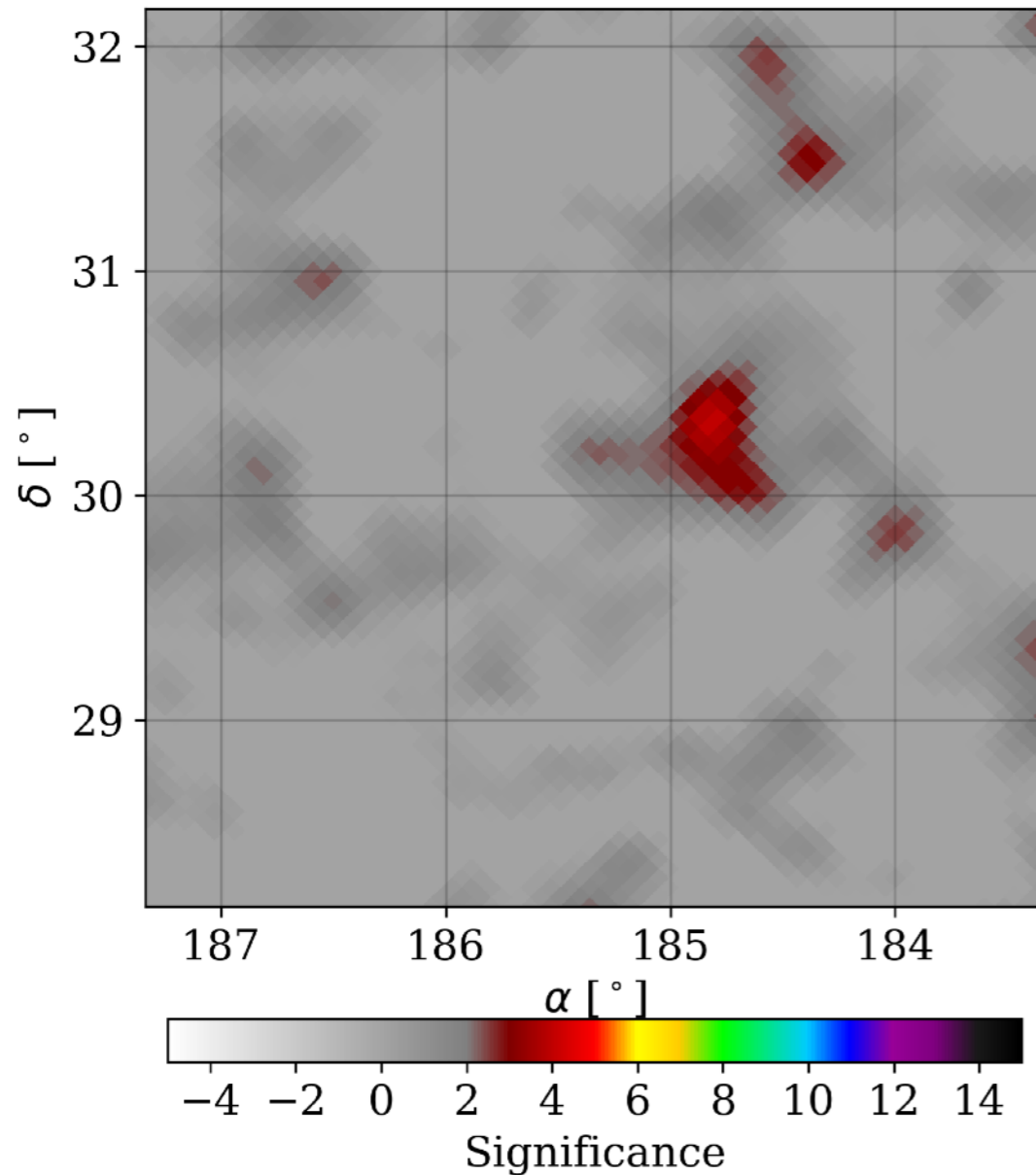


NN

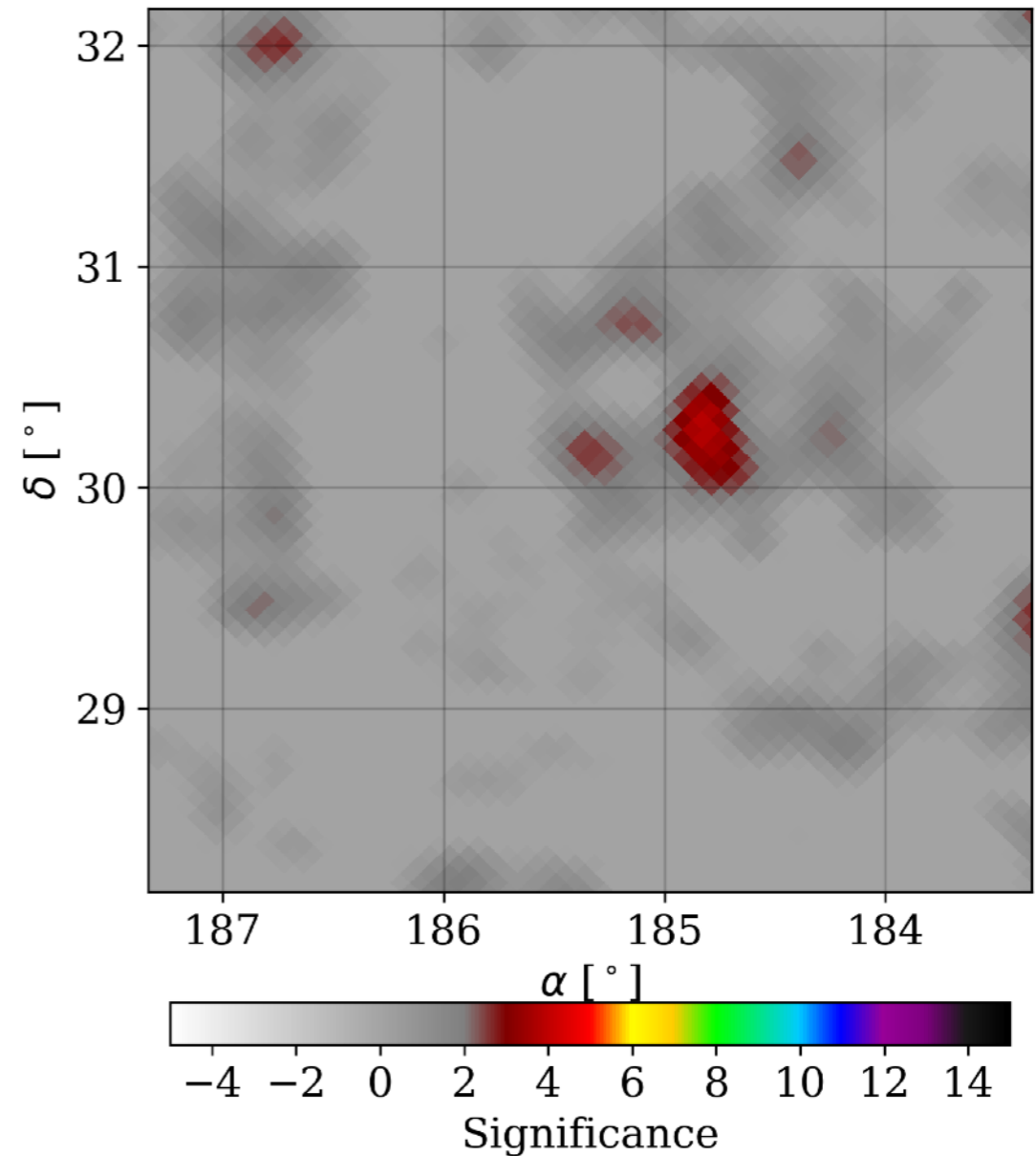


PG 1218+304 using bin 1-9

BDT

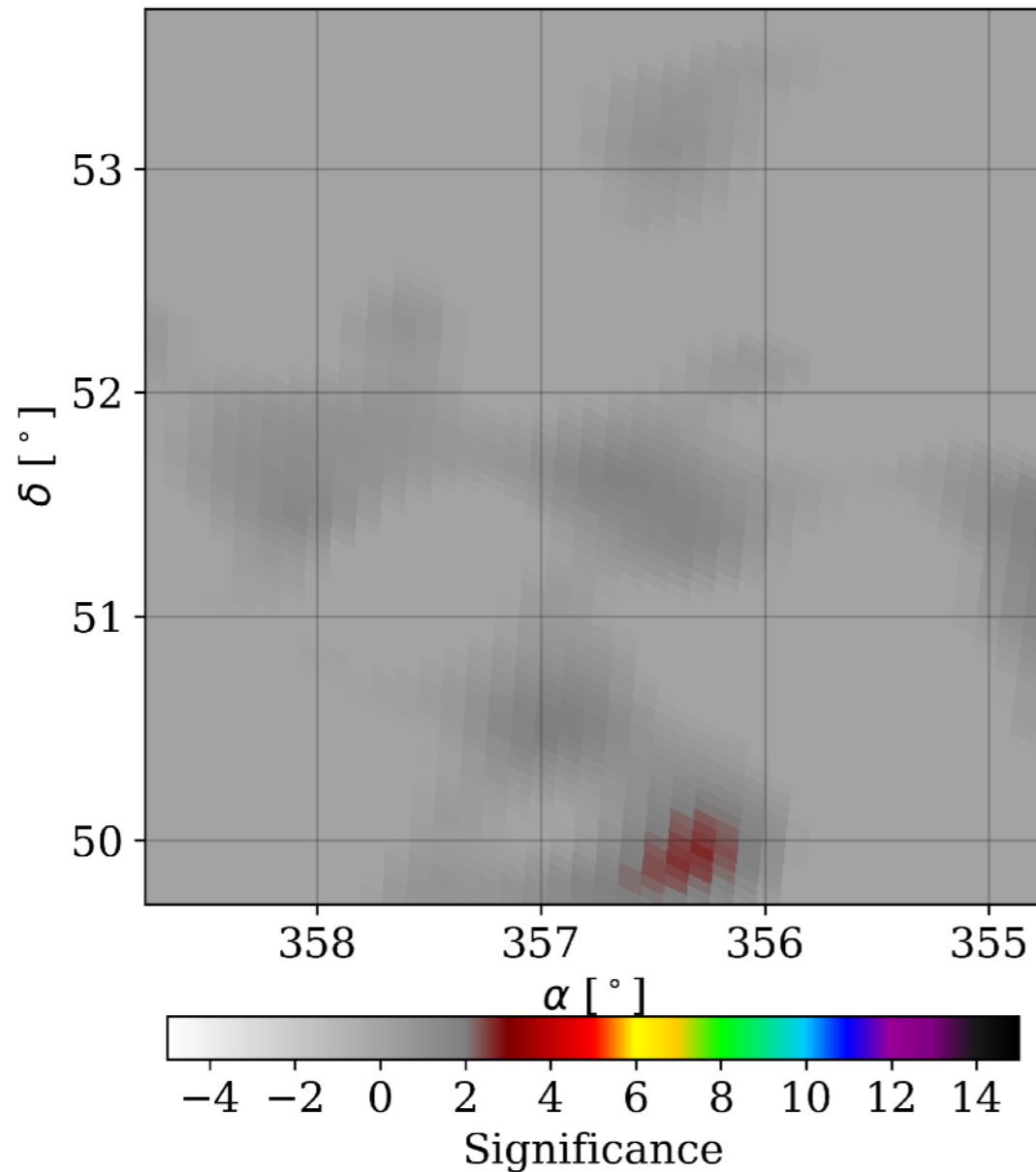


NN

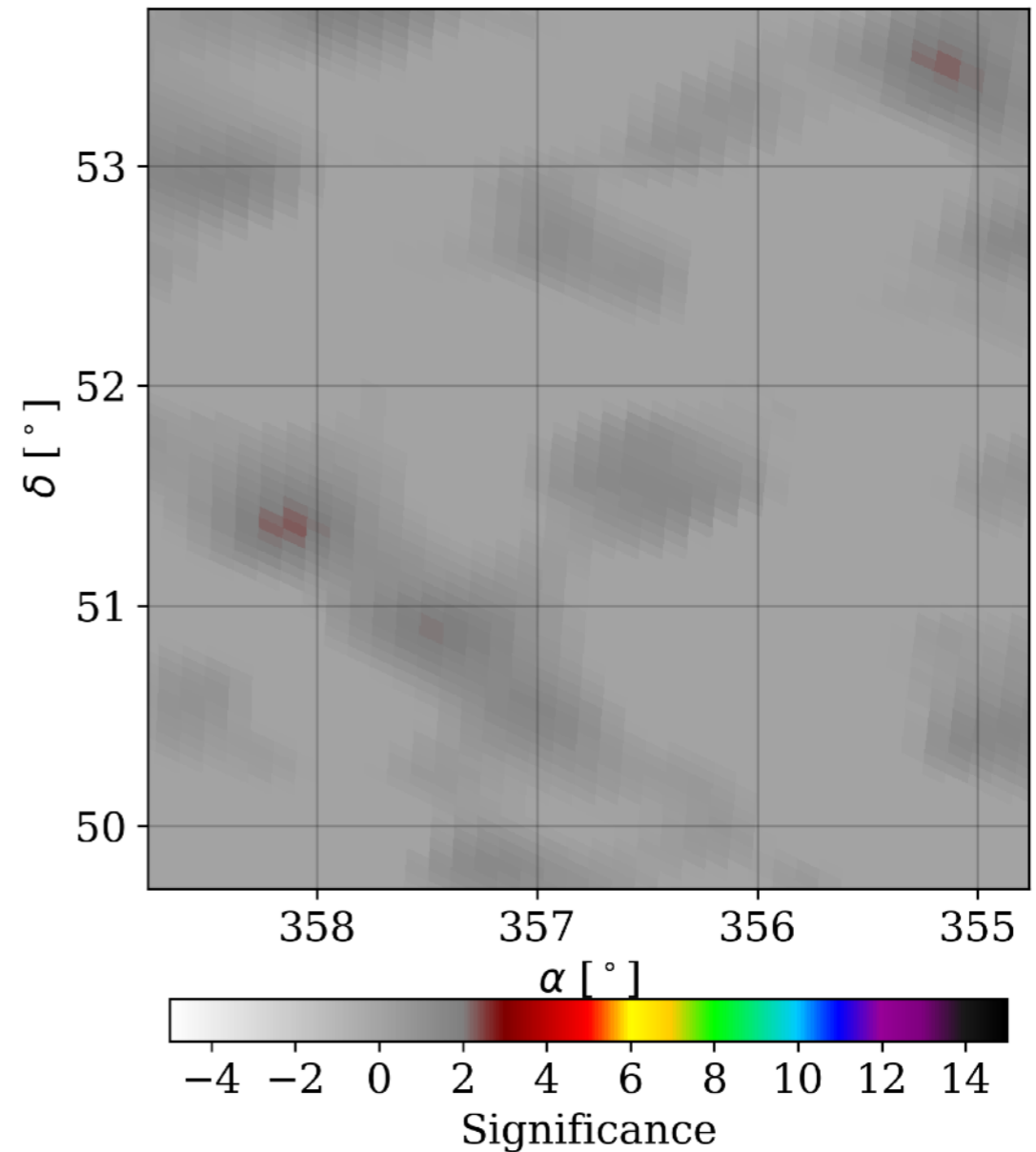


1ES_2344+514 using bin 1-9

SC

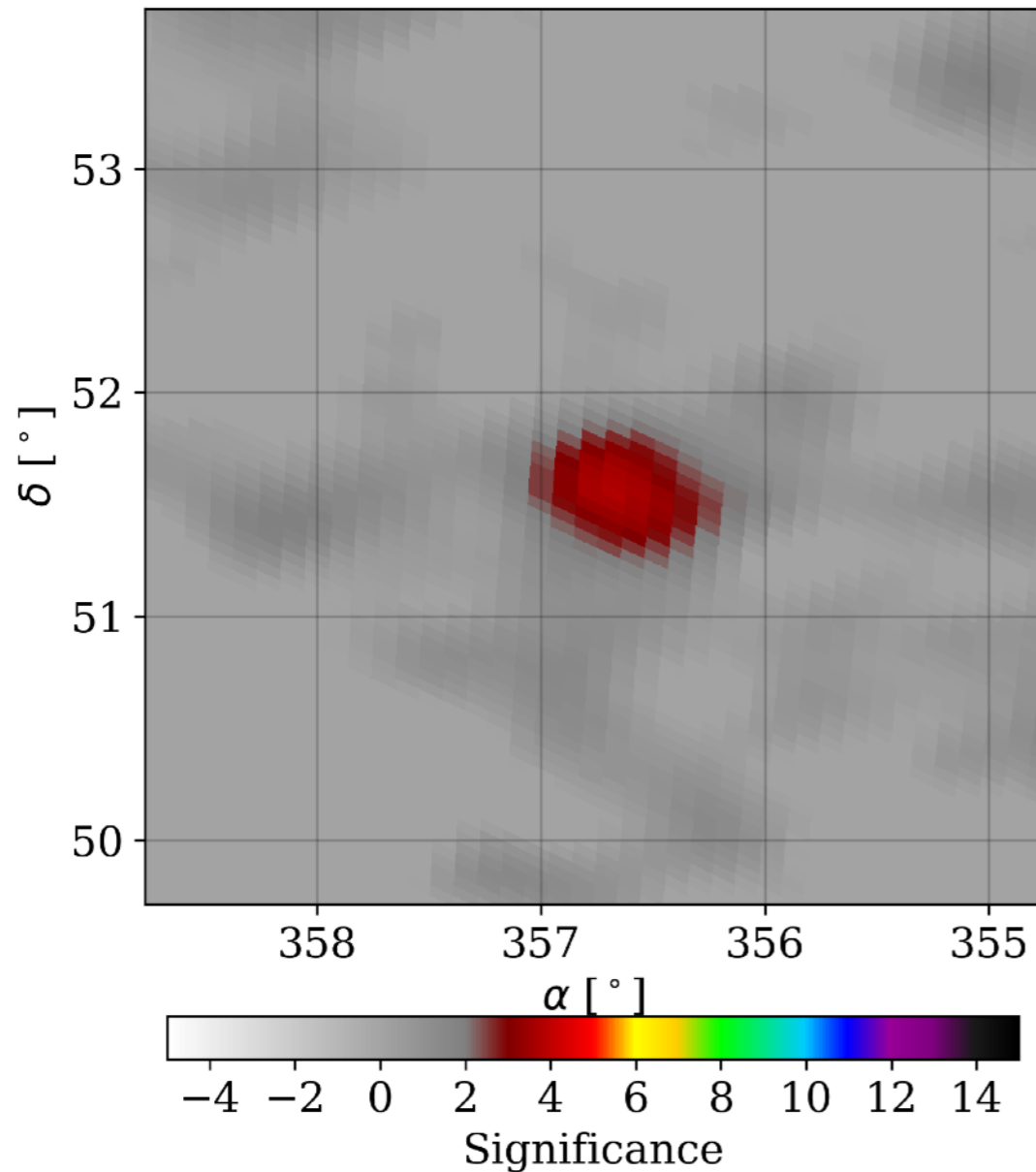


NN

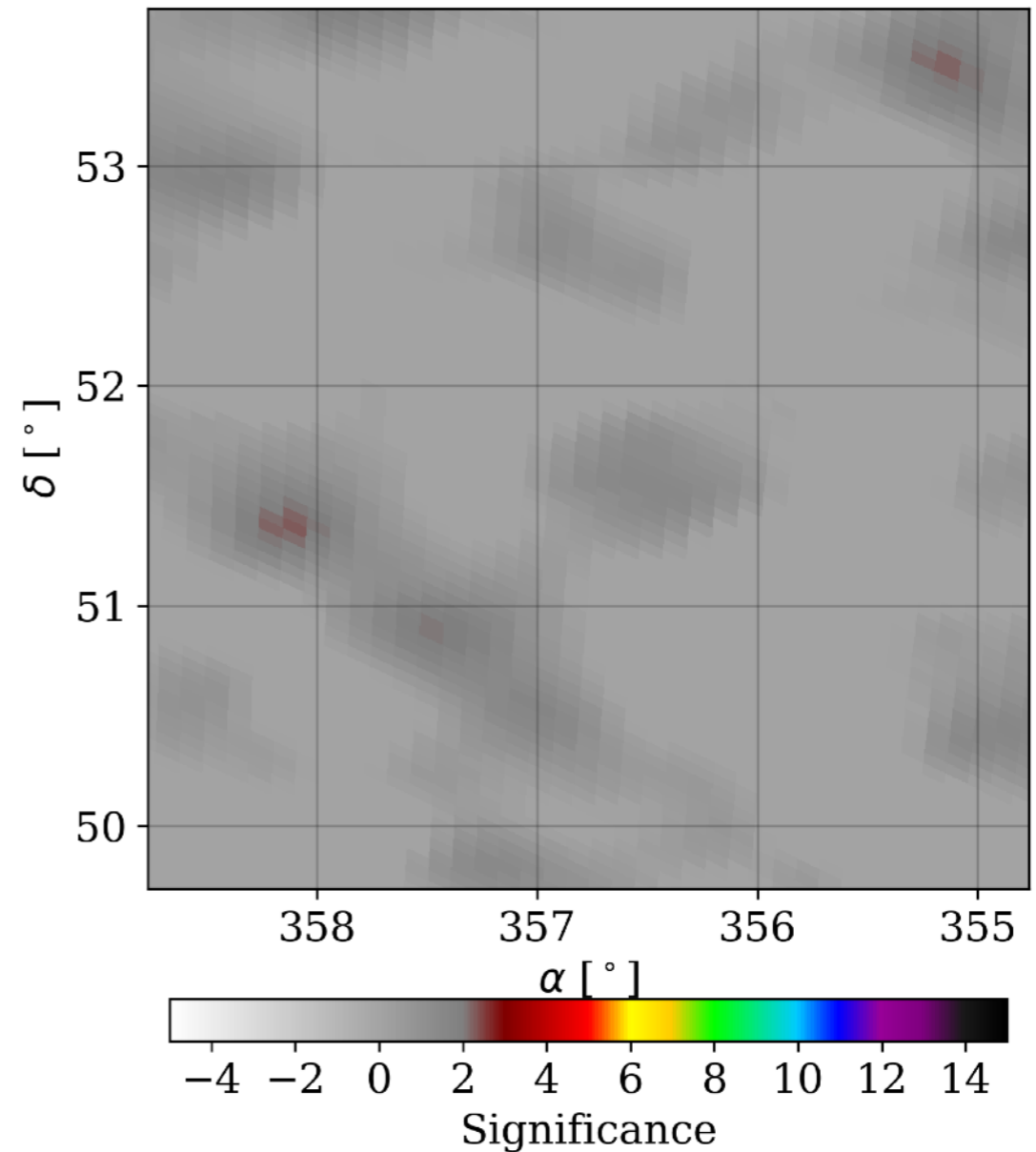


1ES_2344+514 using bin 1-9

BDT



NN



Summary

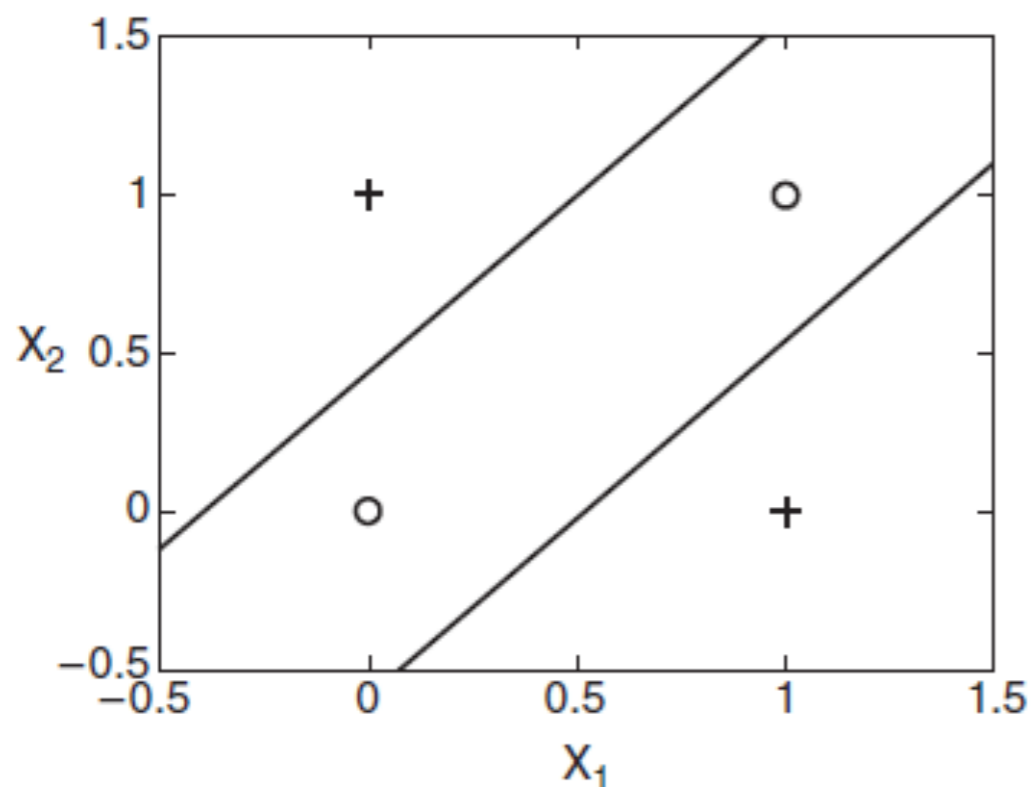
- A g/h separation model were built using the MLT.
- These MLT models where compare with the SC, and get successful results using MC data.
- The MLT has a good results using the Crab Nebula and Mrk 421.

Thanks

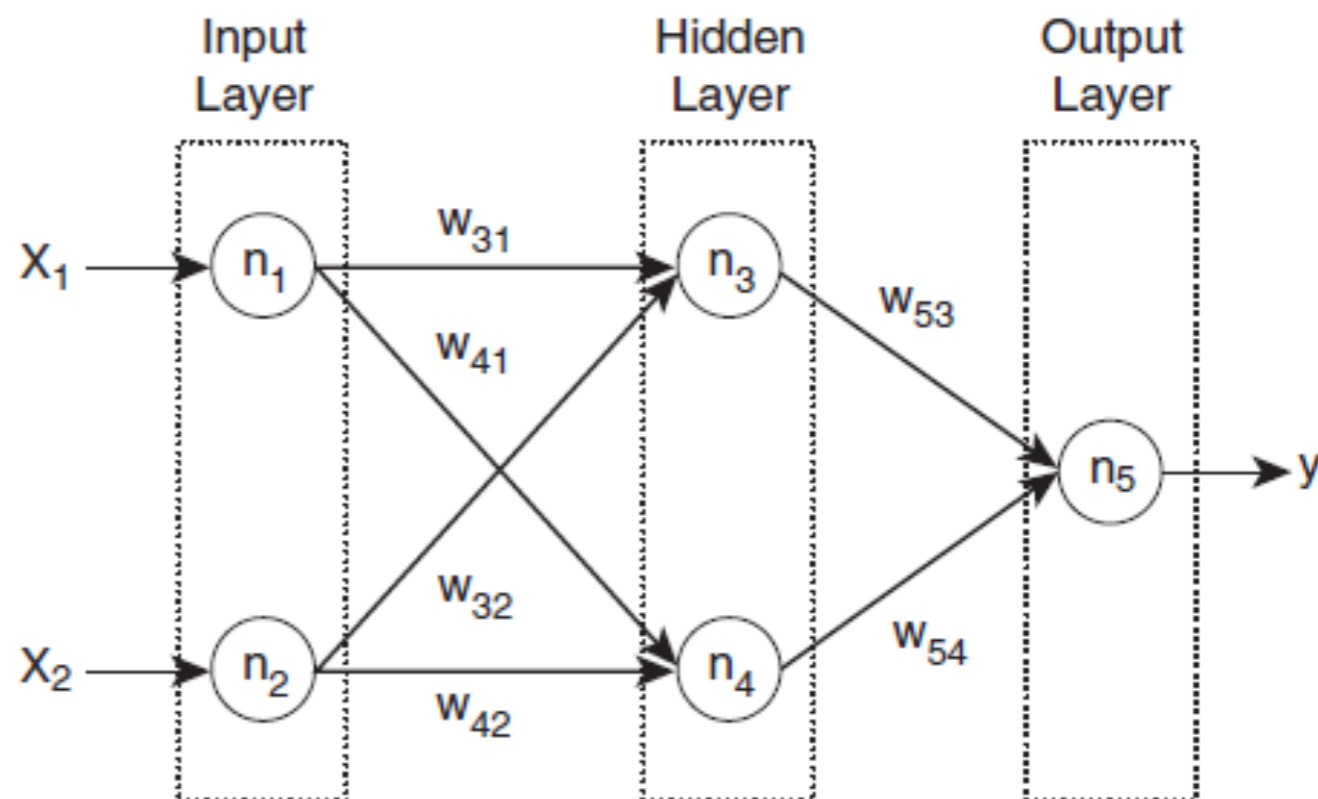
Backslide

Multilayer Neural Network

- It is a nonlinear classifier



(a) Decision boundary.

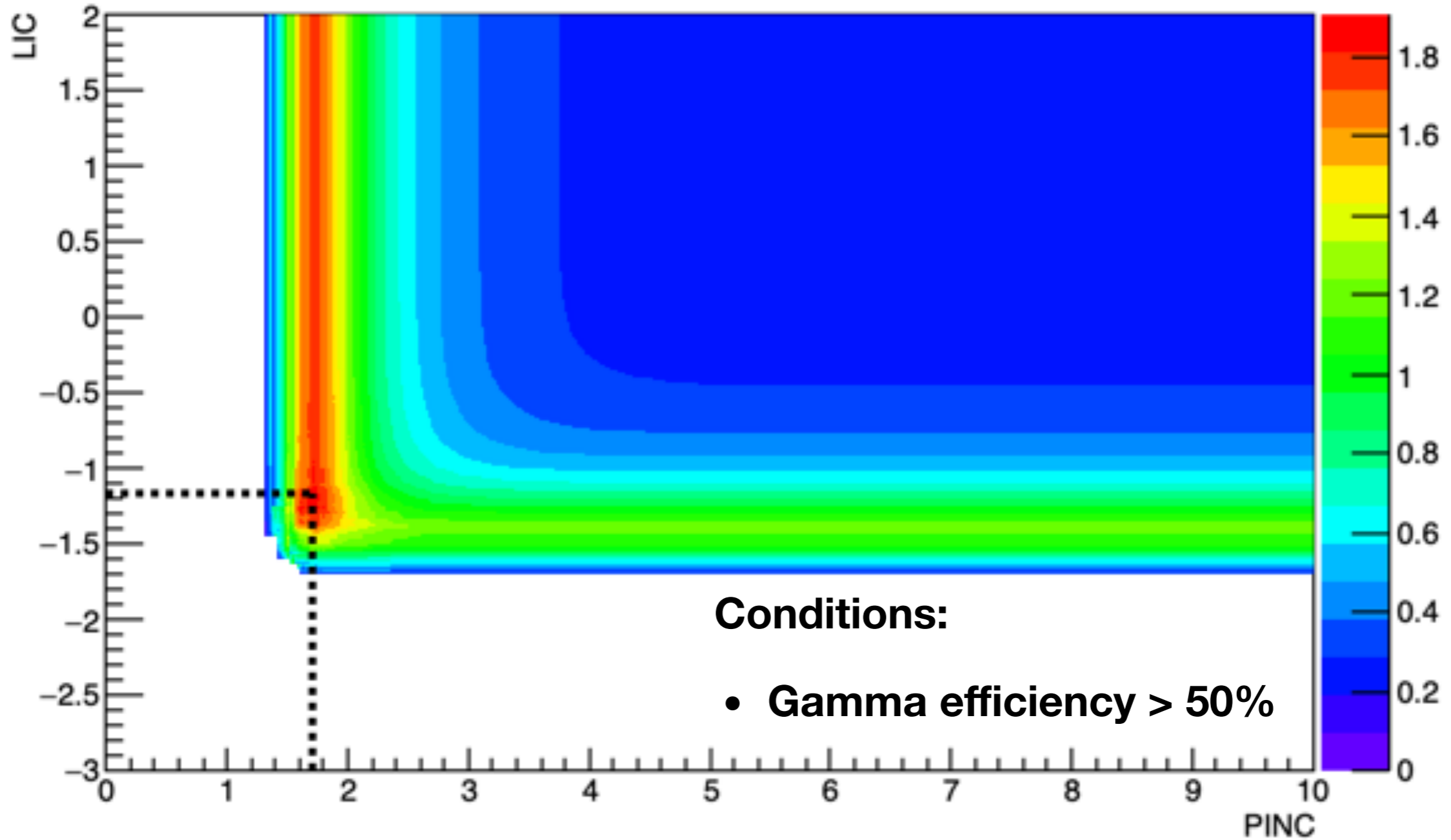


(b) Neural network topology.

Standard Cut

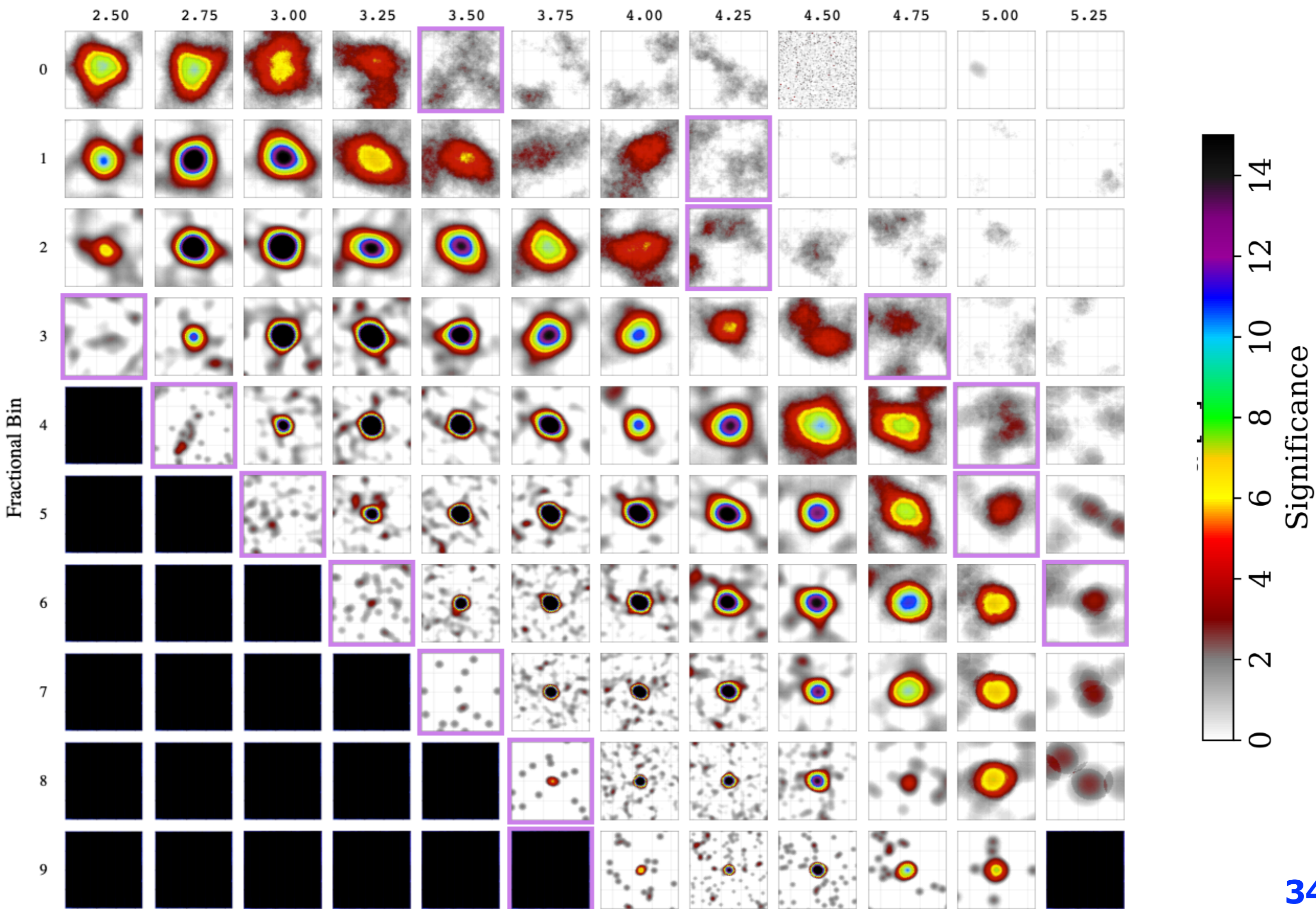
Example of fhit 7, ebin 3.75

Q factor



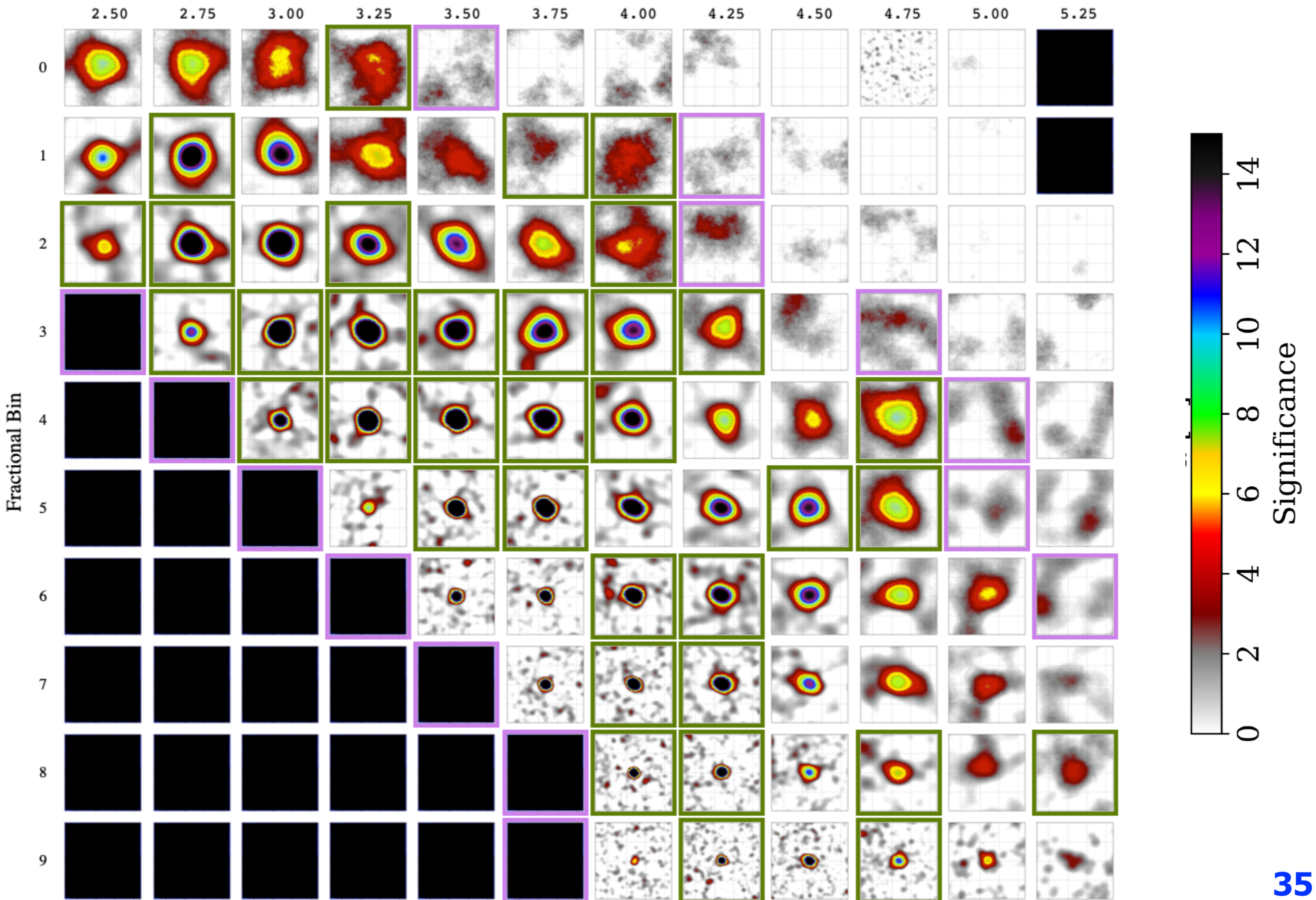
Crab Significance using SC

Energy Bin lower bound (log (E/GeV))



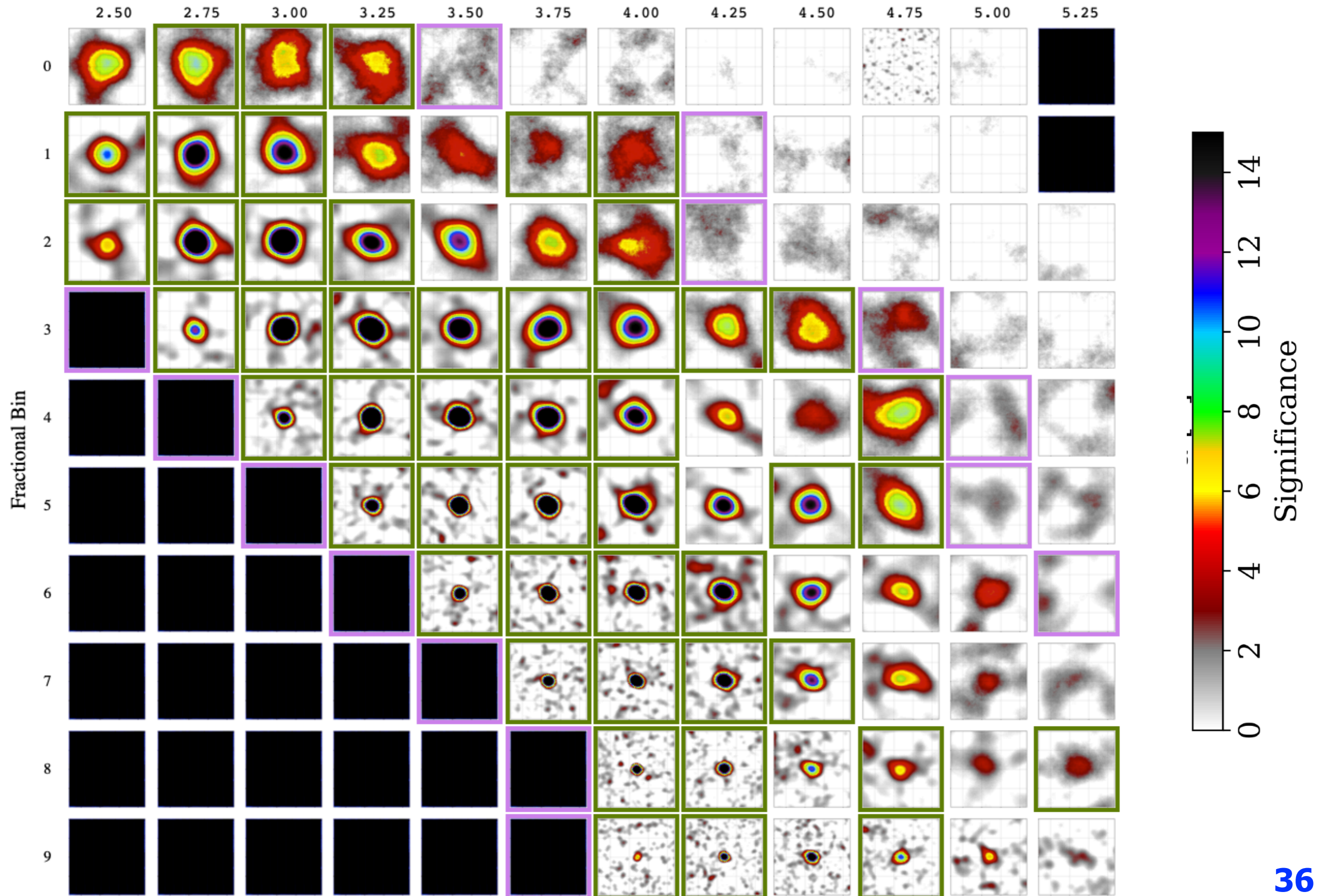
Crab Significance using NN

Energy Bin lower bound ($\log(E/\text{GeV})$)



Crab Significance using BDT

Energy Bin lower bound (log (E/GeV))



Combine ebin to get a fbin map

fbin	Crab			Mrk 421			Mkr 501		
	SC	NN	BDT	SC	NN	BDT	SC	NN	BDT
0	15.16	14.69	15.99	8.10	8.26	8.24	-	-	-
1	27.57	27.47	28.22	13.11	12.30	12.82	3.79	4.22	4.59
2	44.13	44.60	46.36	16.25	15.56	16.64	2.89	3.07	3.74
3	62.39	66.14	71.97	19.10	19.82	21.32	5.34	4.46	4.24
4	69.71	76.34	76.15	19.66	21.95	20.72	5.13	6.15	4.41
5	71.33	69.74	80.05	14.99	15.28	17.49	3.76	4.30	5.67
6	61.52	48.32	65.99	9.13	8.80	11.16	4.95	2.01	5.65
7	47.70	49.18	50.32	5.40	6.62	6.26	2.24	2.51	2.90
8	32.75	35.10	34.84	1.19	1.07	0.58	2.67	2.31	2.89
9	28.70	31.34	31.29	0.23	0.29	0.09	1.12	2.44	2.03
1-9	155.74	156.87	170.69	35.26	35.74	38.43	10.62	10.20	11.94
0-9	156.33	157.45	171.31	35.99	36.42	39.10	10.63	10.11	11.90