

# ICRC poster preparation

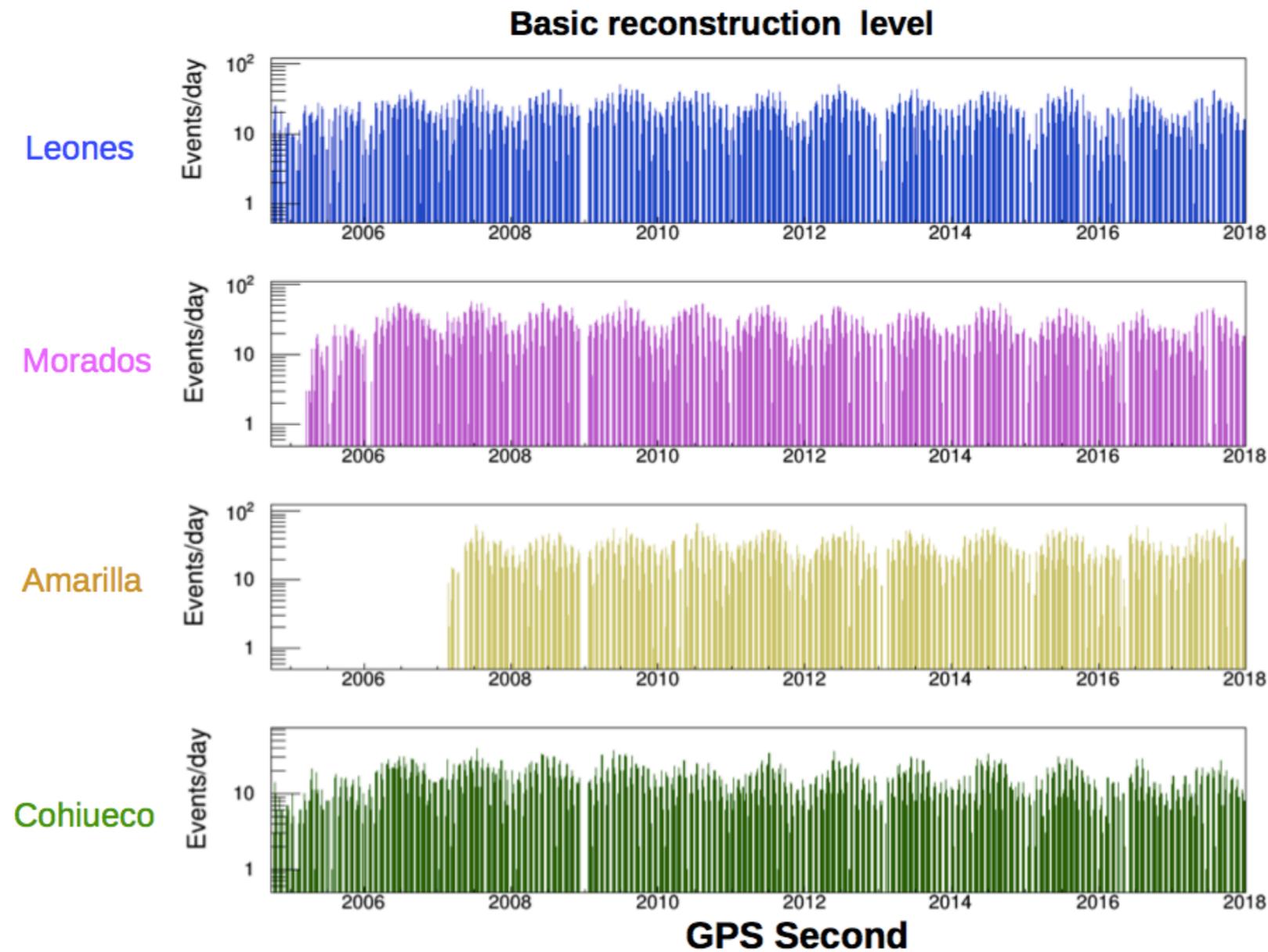
## abstract

Title: Long term performance of the Pierre Auger Observatory

The Surface Array Detector of the Pierre Auger Observatory is comprised of 1664 water Cherenkov detectors, and has been taking data for more than 14 years. We will present the performances and the time evolutions of the surface array, as well as predictions of the future performance. The data quality and long term behaviors of the fluorescence detectors, which are composed of 27 telescopes, will also be presented.

170119 Long term meeting  
Koun Choi (Université Libre de Bruxelles)

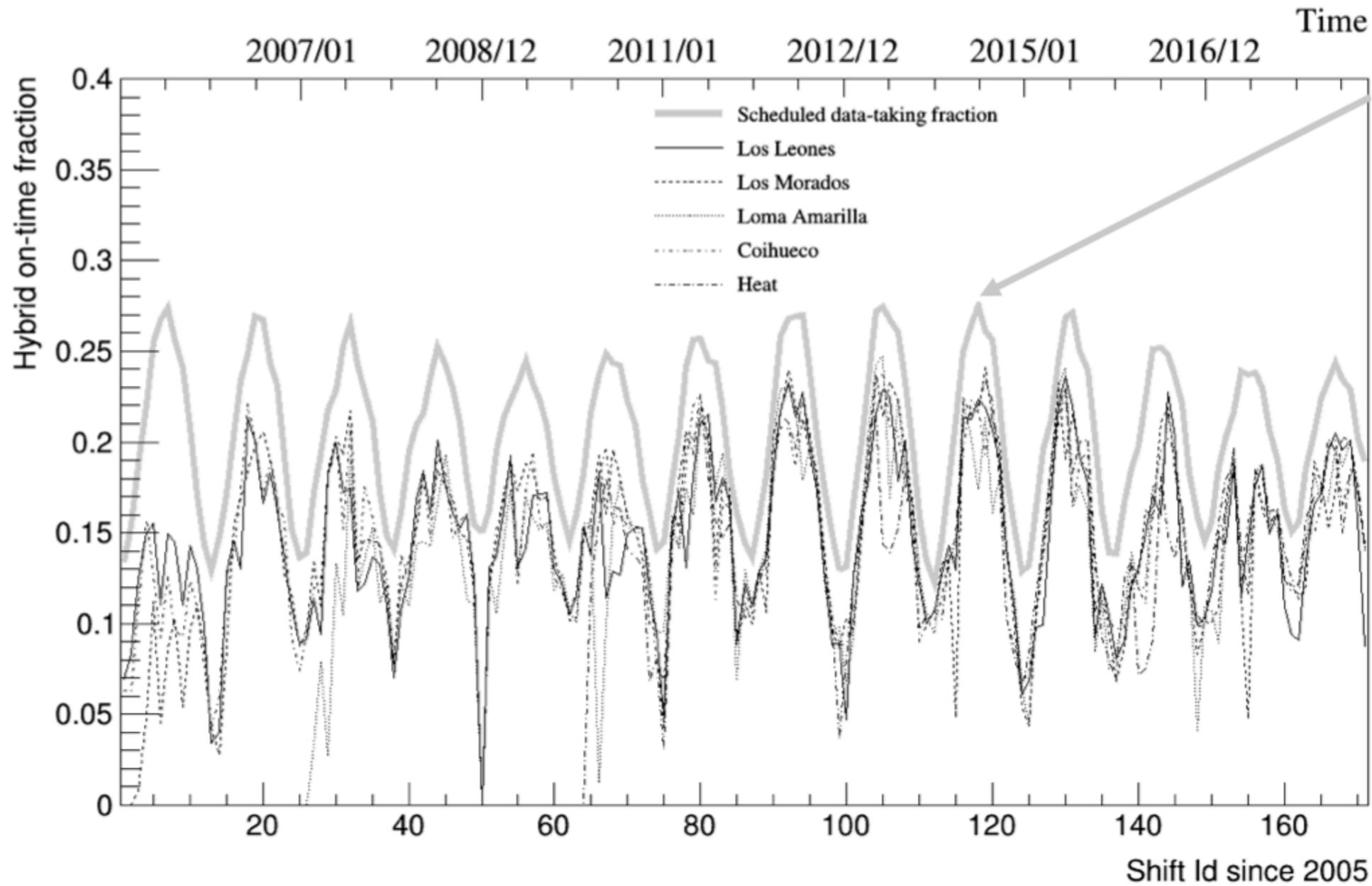
# 1. FD event rate



**raw event rate**  
**to be updated**  
**(F.Salamida, L.Perrone)**

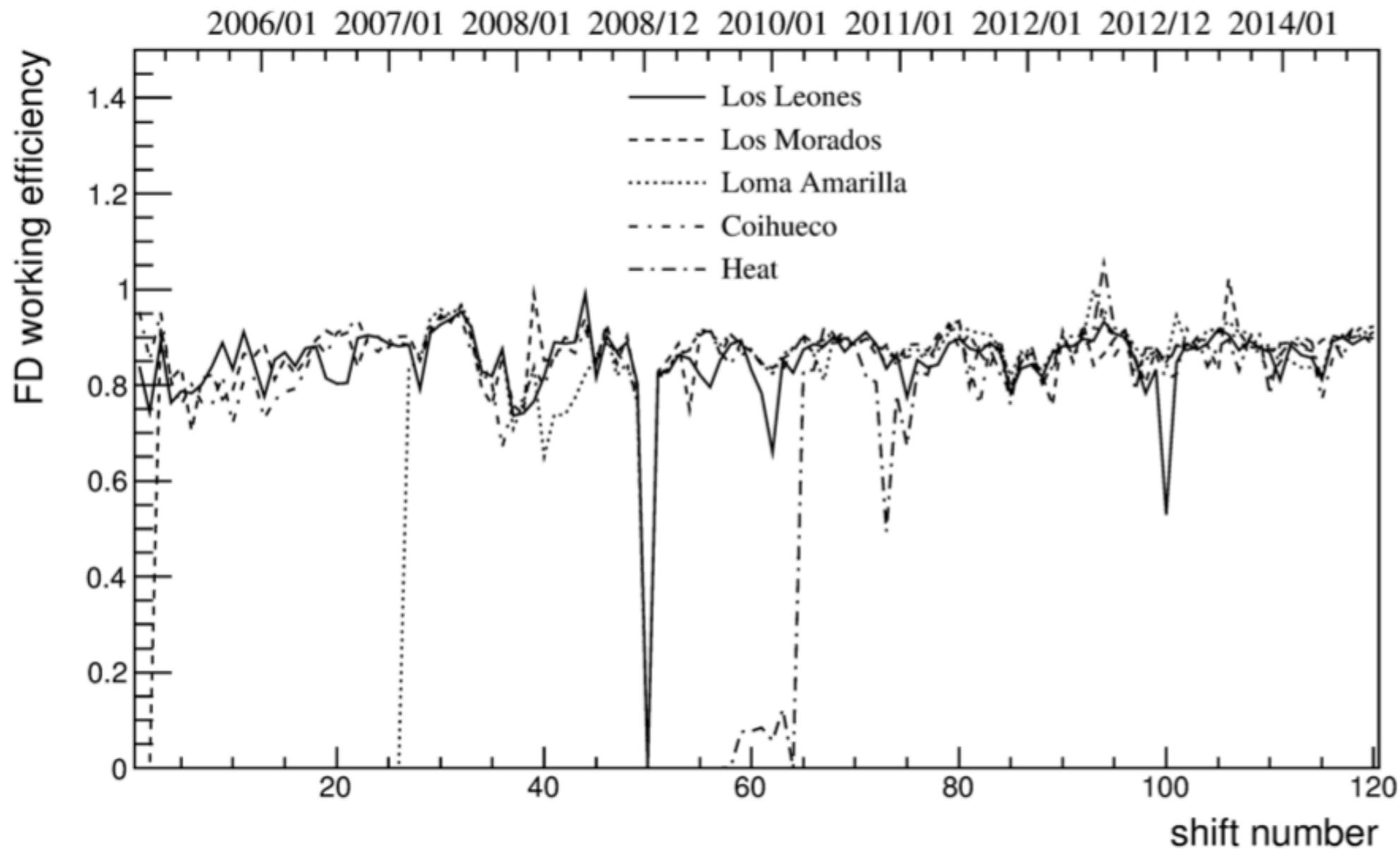
**event rate after X**  
**max/spectrum**  
**selection(to**  
**show efficiency)**  
**to be produced**

# 2. FD uptime



to be updated  
(F.Salamida,  
L.Perrone)

# 3. FD Efficiency



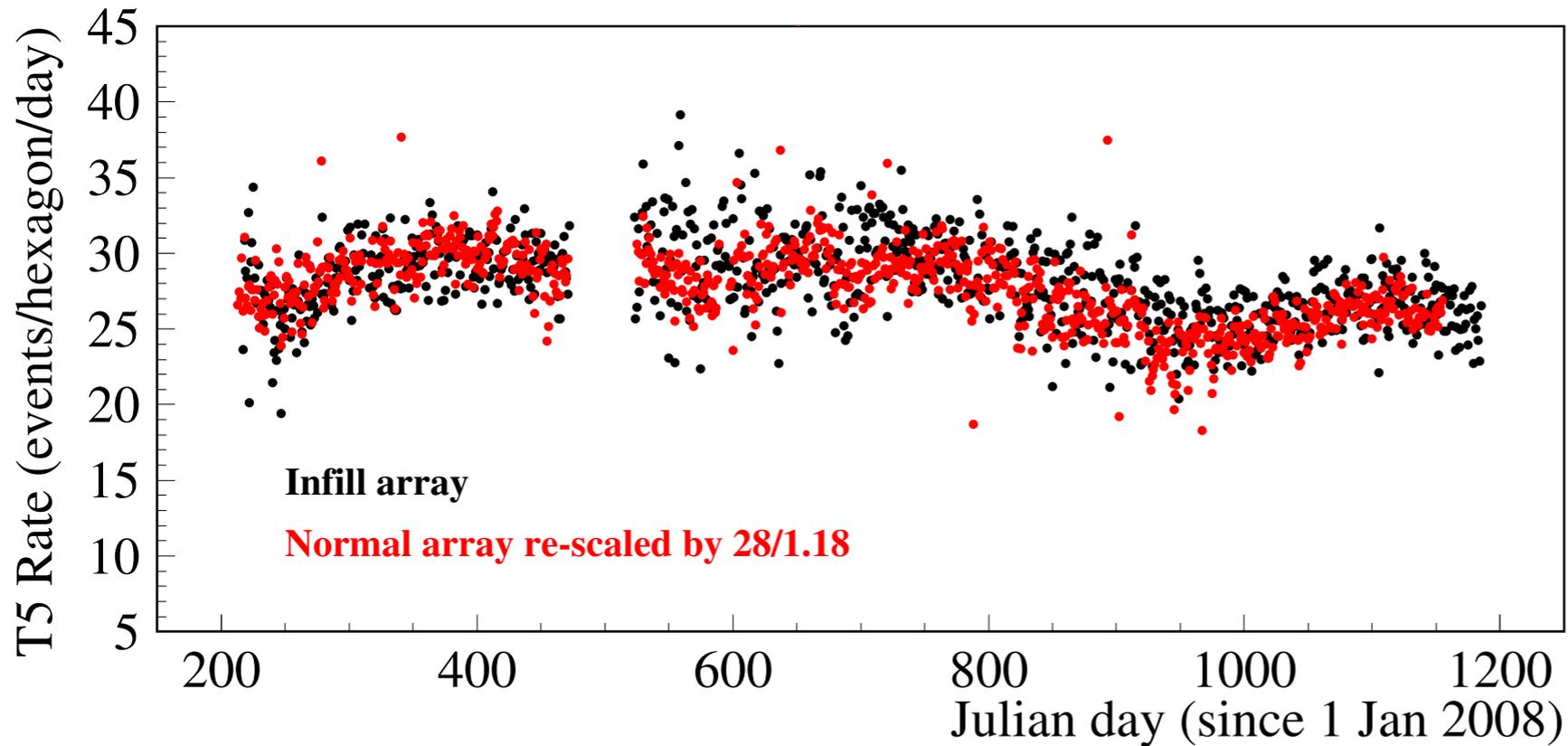
to be updated  
(2015 NIM paper)

# 4. FD/SD cumulative # of events

to be done

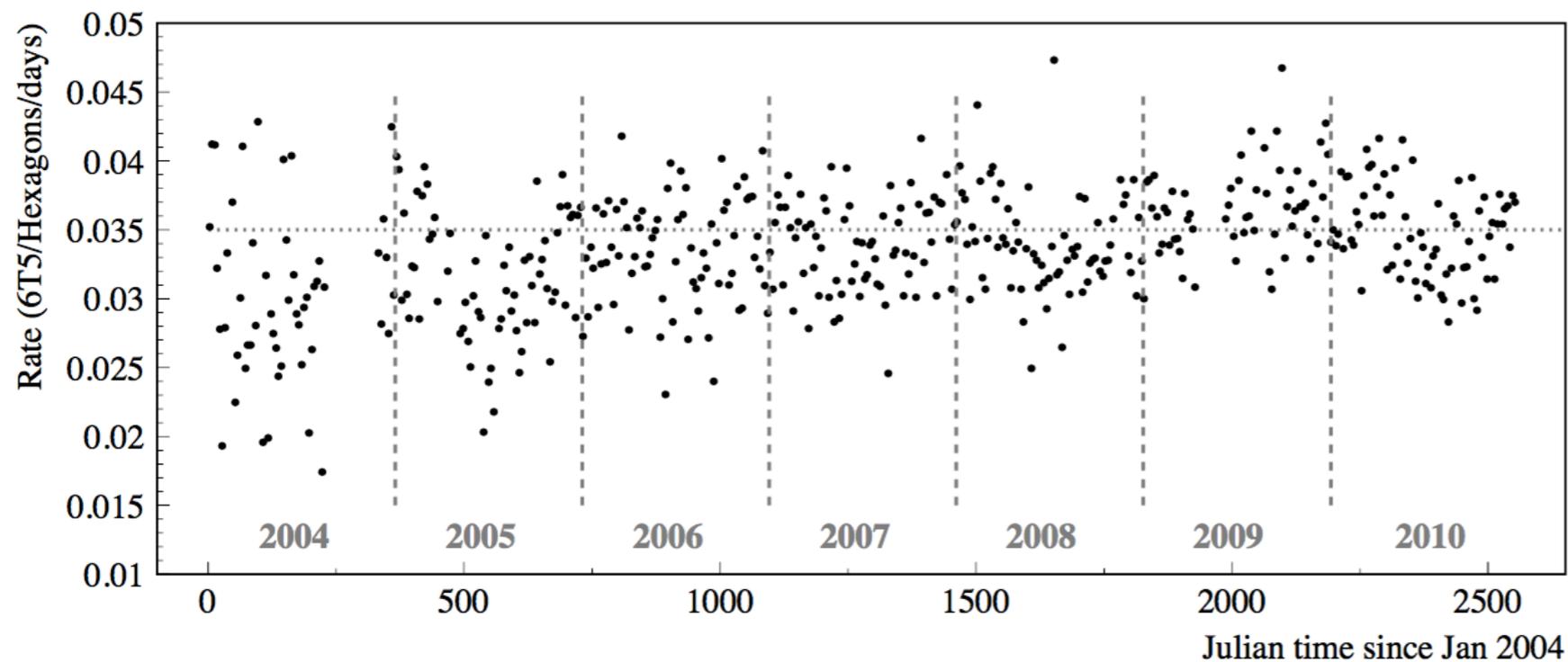


# 5. SD event rate



**750/1500 m arrays  
(all events)**

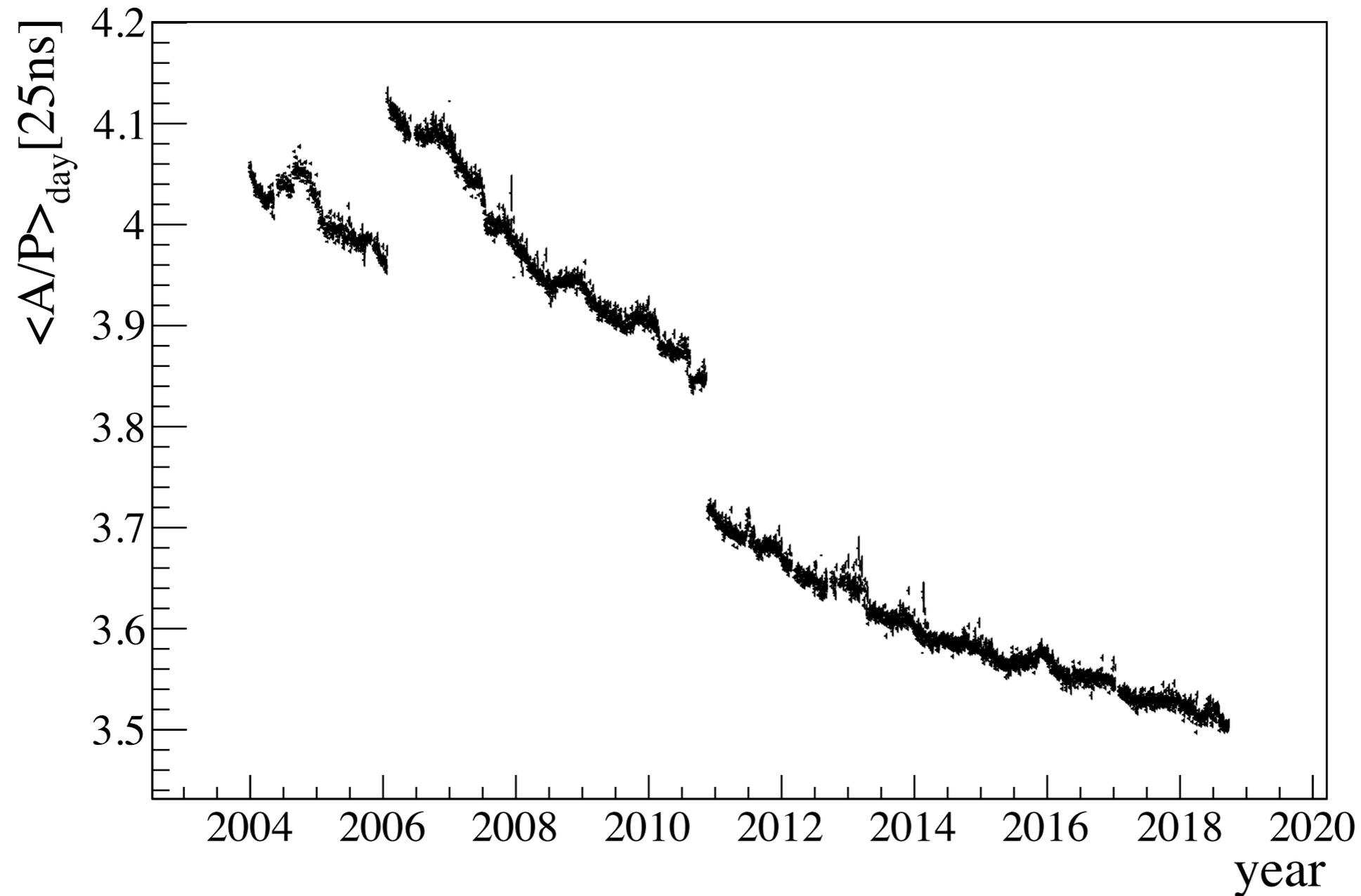
**to be updated  
(I. Maris, R.Sato,  
C.Bonifazi, 2011  
ICRC)**



**750/1500 m arrays  
(above the threshold)**

**to be updated  
R.Sato, 2011 ICRC)**

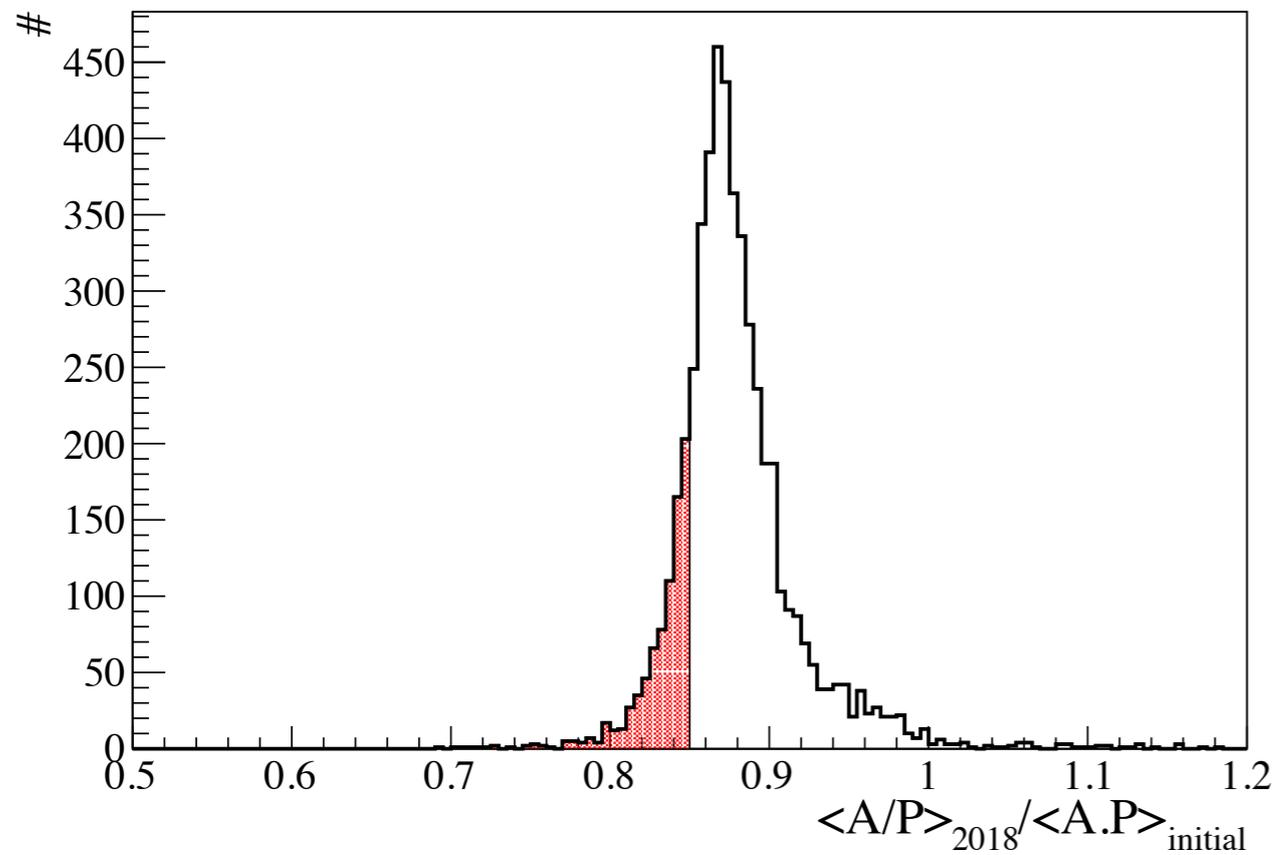
# 6. an example A/P profile



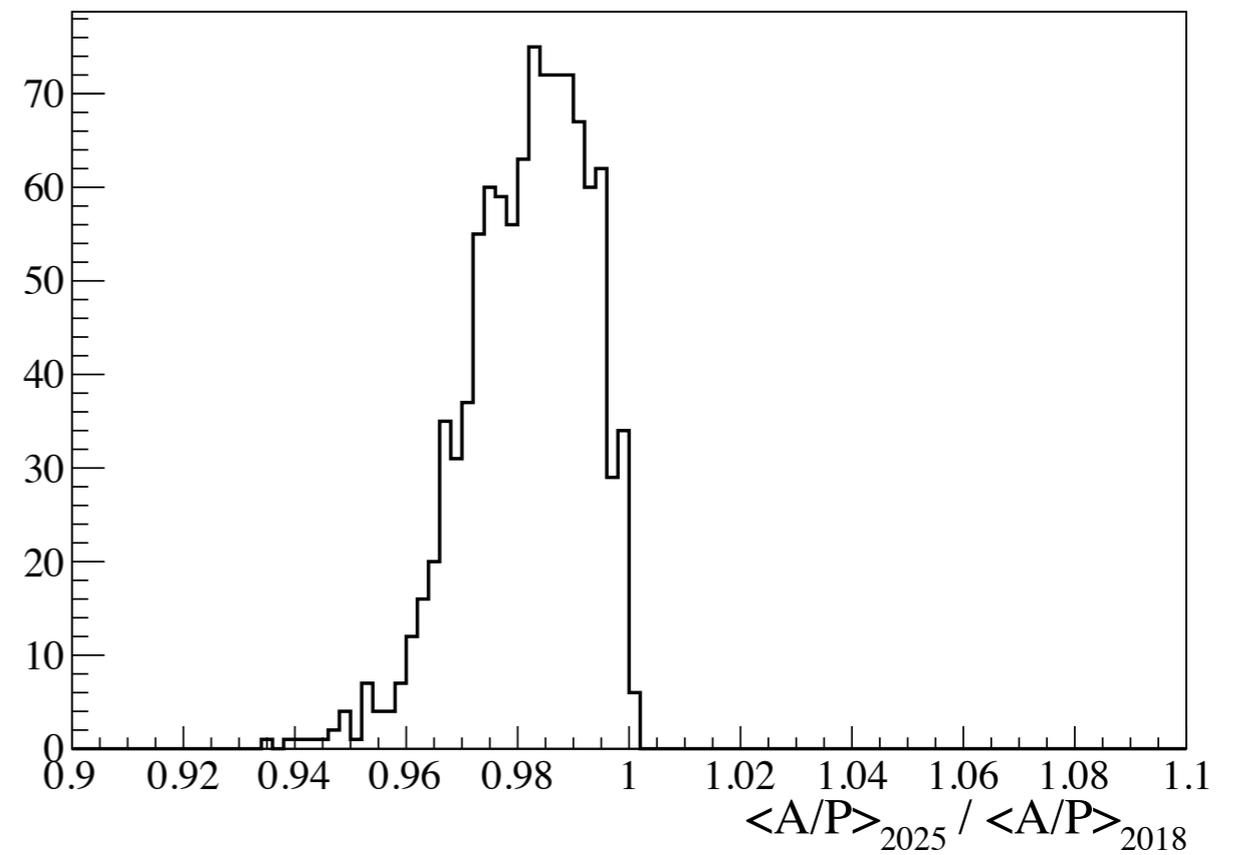
**to be updated**  
**(profile without**  
**jumps)**

# 6. A/P ratio expected in 2025

current status



future expectation



# Brainstorming

- mean A/P of air showers
- mean  $X_{\max}$  vs time
- weighted(for the flux) mean of energy of events(FD/SD) vs time above certain threshold
- SD uptime
- battery/solar panel lifetime
- attenuation parameters(CIC) vs time
- ...