

# Effective theory approach to partly neutrinos.

## Theory and application

F del Águila, S. Bar Shalom, A. Soni, J.W.

# The main issue

$\nu$  oscillations suggest the presence of  $\nu_R = N$ :

$$\mathcal{L}_{\nu SM} \equiv \mathcal{L}_{SM} + \left( \bar{N}_a M_{ab} N_b^c / 2 - \bar{L}_i \tilde{\phi} Y_{ia} N_a + \text{H.c.} \right)$$

If  $M$  does not allow a conserved fermion #

$\Rightarrow N \sim \text{Majorana} \Rightarrow \text{LNV signals}$

$$m_\nu = -m_D M^{-1} m_D^T, \quad m_D = \langle \phi \rangle Y = v Y / \sqrt{2}$$

Two examples giving  $m_\nu \sim 0.01$  eV:

- $M \sim 100$  GeV,  $m_D \sim m_{\text{electron}}/10$  ( $Y \sim 10^{-7}$ )
- $M \sim 10^{15}$  GeV,  $m_D \sim m_W$  ( $Y \sim 1$ )

But the N **decouple** in both cases:

When  $M \sim 10^{15}$  GeV because they are very heavy

When  $M \sim 100$  GeV because they couple weakly:

$$\mathcal{L}_{V-A}^W = -(g/\sqrt{8})U_{\ell N}\overline{N^c}\gamma^\mu(1 - \gamma_5)\ell W_\mu^+ + \text{H.c.}$$

$$U_{\ell N} \sim \sqrt{m_\nu/M} \sim Y \sim 10^{-7}$$

At least as far as  $\mathcal{L}_{\nu \text{ SM}}$  is concerned

# Masses & mixings

$$\mathcal{L}_{\nu \text{ mass}} = \frac{1}{2} \bar{N} M N^c - \bar{\nu}_L m N + \bar{\nu}_L \mu \nu_L^c + \text{H.c.}$$

Absorbed the effects of  
 $(\bar{N} N^c) \langle \phi^\dagger \phi \rangle / \Lambda$

Dirac mass:  
 $m = \langle \phi \rangle Y = v Y$

From  $(\bar{L} \tilde{\phi})(\phi^\dagger L^c)$   
 Naturally small:  $\mu \sim v^2/\Lambda$

Simplest hierarchy :  $M \gg m = vY \gg \mu = \frac{v^2}{\Lambda}$

masses :  $m_{\text{light}} = \mu + (\mathbf{Im}m)M^{-1}(\mathbf{Im}m)^T - (\mathbf{Re}m)M^{-1}(\mathbf{Re}m)^T + \dots$   
 $m_{\text{heavy}} = M + \dots$

$$\nu_L - N \text{ mixing} : \sim \frac{m}{M} \lesssim \sqrt{\frac{m_{\text{light}}}{M}}$$

# Newer physics

Observable N effects likely imply more new physics

I will assume the newer physics has scale  $\Lambda$  ...

$$\Lambda > M$$

$$1 \text{ TeV} > M > 100 \text{ GeV}$$

... and that the scale  $\Lambda$  is not directly probed

⇒ use an effective  $\mathcal{L}$  to describe the N – SM interactions

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\nu SM} + \sum_{n=5}^{\infty} \Lambda^{4-n} \sum_i \alpha_i \mathcal{O}_i^{(n)}$$

- $\mathcal{O}$ : gauge-invariant built of N and SM fields
- The  $\alpha_i$  cannot be calculated (unless the underlying physics is known) ...  
... but they can be bound by naturality
- Largest effects from tree-level generated operators

## Dimension 5 terms

$L$ : left-handed lepton isodoublets  
 $Q$ : left-handed quark isodoublets  
 $e$ : right-handed charged lepton isosinglets  
 $u, d$ : right-handed charged quark isosinglets  
 $\phi$ : scalar isodoublet

$$(\bar{L}\tilde{\phi})(\phi^\dagger L^c)$$

$v_L$  Majorana mass  $\sim v^2/\Lambda$   
+ H interactions

$$(\bar{N}N^c)(\phi^\dagger\phi)$$

$v_R$  Majorana mass  $\sim v^2/\Lambda$   
+ H interactions

## Dimension 6 terms

$$\mathcal{O}_{LN\phi} = (\phi^\dagger \phi)(\bar{L}N\tilde{\phi})$$

$$\mathcal{O}_{NN\phi} = i(\phi^\dagger D_\mu \phi)(\bar{N}\gamma^\mu N)$$

$$\mathcal{O}_{Ne\phi} = i(\phi^T \varepsilon D_\mu \phi)(\bar{N}\gamma^\mu e)$$

$$\mathcal{O}_{duNe} = (\bar{d}\gamma^\mu u)(\bar{N}\gamma^\mu e)$$

$$\mathcal{O}_{fNN} = (\bar{f}\gamma_\mu f)(\bar{N}\gamma^\mu N)$$

$$\mathcal{O}_{LNLe} = (\bar{L}N)\varepsilon(\bar{L}e)$$

$$\mathcal{O}_{LNQd} = (\bar{L}N)\varepsilon(\bar{Q}d)$$

$$\mathcal{O}_{QuNL} = (\bar{Q}u)(\bar{N}L)$$

$$\mathcal{O}_{QNLd} = (\bar{Q}N)\varepsilon(\bar{L}d)$$

$$\mathcal{O}_{LN} = |\bar{L}N|^2$$

$$\mathcal{O}_{QN} = |\bar{Q}N|^2$$

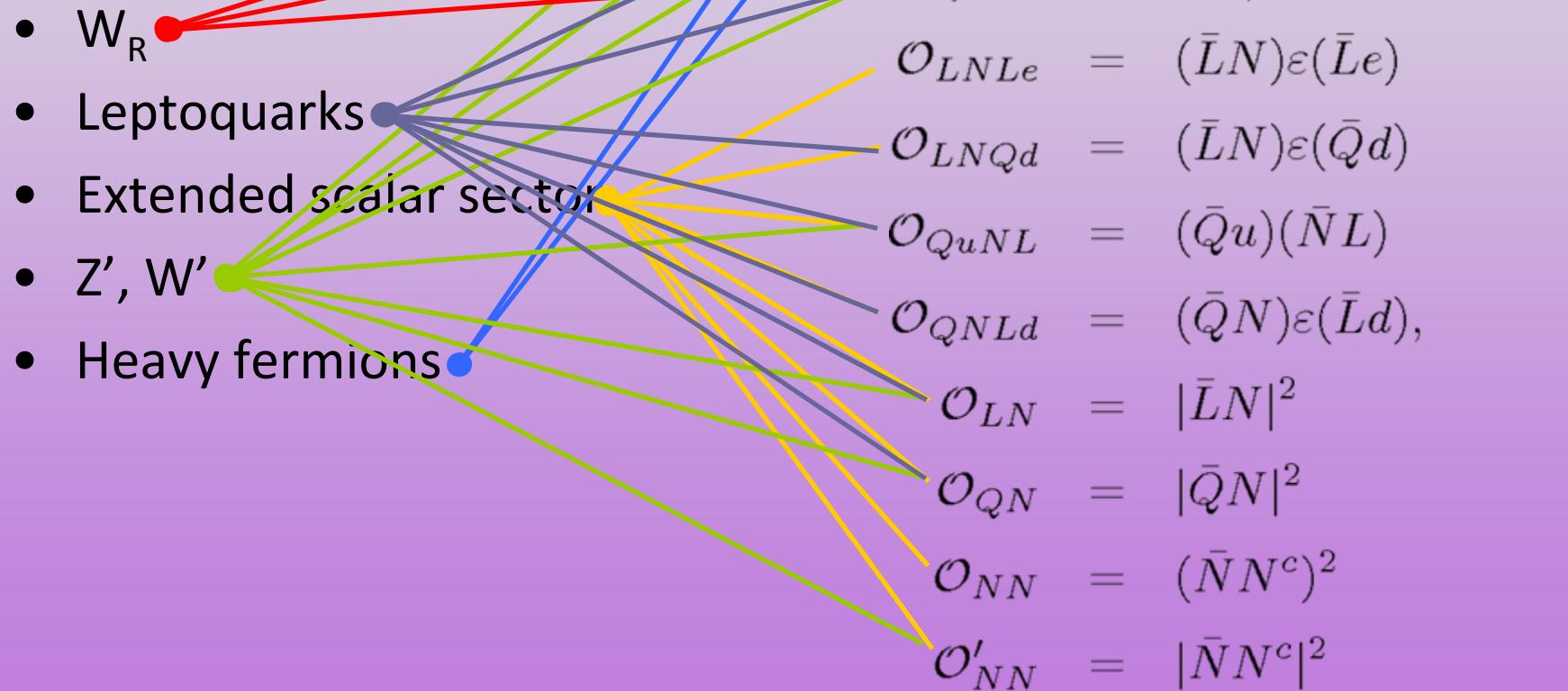
$$\mathcal{O}_{NN} = (\bar{N}N^c)^2$$

$$\mathcal{O}'_{NN} = |\bar{N}N^c|^2$$

$L$ : left-handed lepton isodoublets  
 $Q$ : left-handed quark isodoublets  
 $e$ : right-handed charged lepton isosinglets  
 $u, d$ : right-handed charged quark isosinglets  
 $\phi$ : scalar isodoublet

+ B-violating operators + loop generated operators

Generated by



# Application: $pp, p\bar{p} \rightarrow \ell^+ \ell^+ j j$

## Lagrangian

$$\mathcal{L}_{eff}^N = \Lambda^{-2} \left[ -\sqrt{2} v m_w \alpha_{wl} \overline{N^c} \gamma^\mu e_L W_\mu^+ \right.$$

From  $\mathcal{L}_v^{\text{SM}}$   
Small coupling  $\sim 10^{-7}$

$$-\sqrt{2} v m_w \alpha_{wr} \bar{N} \gamma^\mu e_R W_\mu^+$$

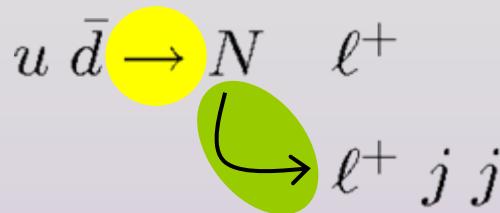
$$+ \alpha_v (\bar{d}_R \gamma^\mu u_R) (\bar{N} \gamma_\mu e_R)$$

From  $i(\phi^T \varepsilon D_\mu \phi)(\bar{N} \gamma^\mu e)$

$$+ \alpha_{s1} (\bar{u}_R d_L) (\bar{e}_L N)$$

$$- \alpha_{s2} (\bar{u}_L d_R) (\bar{e}_L N)$$

$$\left. + \alpha_{s3} (\bar{u}_L N) (\bar{e}_L d_R) + \text{H.c.} \right]$$



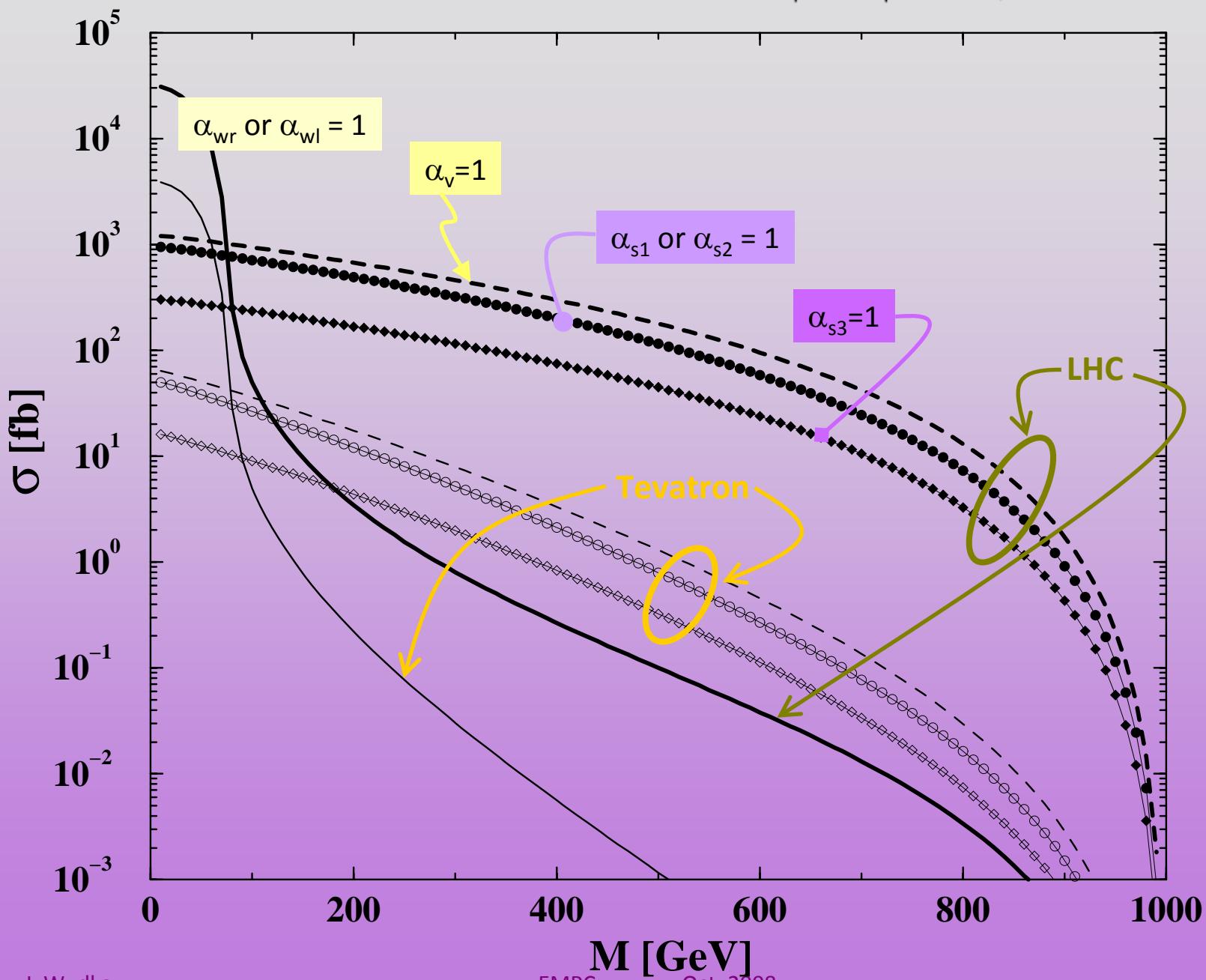
$$\begin{aligned}
 \frac{d\hat{\sigma}}{dc_\theta} &= \frac{(\hat{s} - M^2)^2}{128\pi \hat{s}\Lambda^4} \left\{ \alpha_{s1}^2 + \alpha_{s2}^2 - \alpha_{s2}\alpha_{s3}(1 + c_\theta) + \alpha_{s3}^2 \Upsilon_+ \right. \\
 &\quad \left. + 4\alpha_v^2 \Upsilon_- + 16 (\alpha_{wl}^2 \Upsilon_- + \alpha_{wr}^2 \Upsilon_+) \Pi_w(\hat{s}) \right\}, \\
 \frac{d\Gamma}{dx} &= \frac{M}{128\pi^3} \left( \frac{M}{\Lambda} \right)^4 \left\{ (\alpha_{s1}^2 + \alpha_{s2}^2 - \alpha_{s2}\alpha_{s3}) f_s \right. \\
 &\quad \left. + [\alpha_{s3}^2 + 4\alpha_v^2 + 16 (\alpha_{wl}^2 + \alpha_{wr}^2) \Pi_w((M - 2E_\ell)M)] f_v \right\}
 \end{aligned}$$

$$\Pi_w(\hat{s}) \equiv m_w^4 [(\hat{s} - m_w^2)^2 + (m_w \Gamma_w)^2]^{-1} \quad f_s = 6x^2(1 - 2x)$$

$$\Upsilon_\pm = \frac{1}{4}[(1 \pm c_\theta)^2 + M^2 s_\theta^2 / \hat{s}] \quad f_v = x^2(3 - 4x);$$

$$\theta : \ell - u \text{ (CM) scattering angle} \quad x = E_\ell/M \text{ in the } N \text{ rest frame}$$

$$|\cos \theta| < 0.9, \quad \sqrt{\hat{s}} < \Lambda = 1 \text{ TeV}$$



$$\left. \begin{array}{l} M \lesssim 200 \text{ GeV}, \\ \Lambda \sim \mathcal{O}(1) \text{ TeV}, \\ \alpha_{wr} \sim \mathcal{O}(1) \text{ } (\alpha_i = 0 \text{ otherwise}) \end{array} \right\} 5\sigma \text{ effect @ LHC}$$

$$\left. \begin{array}{l} M \lesssim 600 \text{ GeV}, \\ \Lambda \sim \mathcal{O}(1) \text{ TeV}, \\ \alpha_v \sim \mathcal{O}(1) \text{ } (\alpha_i = 0 \text{ otherwise}) \end{array} \right\} \sigma \gtrsim 100 \text{ fb } \text{(LHC)}$$

W\_R or vector leptoquark

## Other observables

$A_{FB} = \theta$  asymmetry

$A_{FB}^y =$  double asymmetry in  $\theta$  and the rapidity  $y$

$M = 200\text{GeV}$	non-zero coefficient				
	$\alpha_{wl}$	$\alpha_{wr}$	$\alpha_v$	$\alpha_{s1,s2}$	$\alpha_{s3}$
$A_{FB}$ (Tevatron)	0.55	-0.55	0.62	0	-0.62
$A_{FB}^y$ (Tevatron)	0.11	-0.11	0.12	0	-0.12
$A_{FB}^y$ (LHC)	0.35	-0.35	0.40	0	-0.40

$\frac{d\sigma}{dM_{jj}}, \quad \frac{d\sigma}{dM_{\ell\ell}}$  : discriminate between  $\alpha_{wr}$  and  $\alpha_v$

$$\int \frac{d\Gamma}{dx} g(x) dx \text{ separates } \begin{cases} \alpha_{s1}^2 + \alpha_{s2}^2 - \alpha_{s2}\alpha_{s3} \\ \alpha_{s3}^2 + 4\alpha_v^2 \\ \alpha_{wl}^2 + \alpha_{wr}^2 \end{cases}$$

$$\frac{d\Gamma}{dx} = \frac{M}{128\pi^3} \left(\frac{M}{\Lambda}\right)^4 \left\{ (\alpha_{s1}^2 + \alpha_{s2}^2 - \alpha_{s2}\alpha_{s3}) f_s \right. \\ \left. + [\alpha_{s3}^2 + 4\alpha_v^2 + 16(\alpha_{wl}^2 + \alpha_{wr}^2) \Pi_w((M - 2E_\ell)M)] f_v \right\}$$

# Conclusions

- Standard approach:  $N \not\perp W$  coupling suppressed
- Merely seeing a few 100 GeV  $N \rightarrow$  physics beyond  $N+SM$
- New physics effects described by
  - Specific models  $\rightarrow$  valid at all energies
  - Effective theories  $\rightarrow$  valid below  $\Lambda$
- Model-independent  $\mathcal{L}_{\text{eff}}$  cannot fully distinguish between models ...
  - $W'$  or leptoquarks  $\rightarrow (\bar{d}\gamma^\mu u)(\bar{N}\gamma^\mu e)$
  - ... but it does estimate the scale.
- New physics effects can be dramatic, but not as dramatic resonance peaks