# Hadronic light by light contribution to the $\mu$ anomalous magnetic moment

Pablo Roig Garcés

Work done in collaboration with Khépani Raya & Adnan Bashir

QCD WG

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Jegerlehner & Nyffeler, Phys.Rept. 477 (2009) 1-110 Interest of  $a_{\mu}$ 

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FNAL & J-PARC will bring the error down to 16x10<sup>-11</sup> in the near future



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Type of contribution	Value x 10 <sup>11</sup>	Error x 10 <sup>11</sup>
QED	116'584,718.95	0.08
EW	153.6	1.0
HVP	6825	(42)
HLbL	105	(26)
Total	116'591,803	(1)(42)(26)
Exp	116'592,091	(54)(33)

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Recent evaluations of  $a_{\mu}$  contributions in the SM

- **5-loop QED**: Aoyama, Hayakawa, Kinoshita & Nio, Phys.Rev.Lett. 109 (2012) 111807, 111808
- EW contributions after M<sub>H</sub> measurement: Gnedinger, Stockinger & Stockinger-Kim, Phys.Rev. D88 (2013) 053005
- Higher order QCD effects:
- Kurz, Liu, Marquard & Steinhauser, Phys.Lett. B734 (2014) 144-147; a<sub>μ</sub><sup>HVP@NNLO</sup>=(1.24+0.01)10<sup>-10</sup>
- Colangelo, Hoferichter, Nyffeler, Passera, Stoffer, Phys.Lett. B735 (2014) 90-91; a<sub>μ</sub><sup>HLbl@HO</sup>=(0.3+0.2)10<sup>-10</sup>
- Towards a data driven analysis of HLbL: Colangelo *et. al.* JHEP 1409 (2014) 091, Phys.Lett. B738 (2014) 6-12, JHEP 1509 (2015) 074, e-Print: arXiv:1701.06554 [hep-ph], e-Print: arXiv:1702.07347 [hep-ph].

[See also Pauk & Vanderhaeghen Phys.Rev. D90 (2014) no.11, 113012, Eur.Phys.J. C74 (2014) no.8, 3008]

Jegerlehner & Nyffeler, Phys.Rept. 477 (2009) 1-110 FNAL & J-PARC will bring the error down to 16x10<sup>-11</sup> in the near future line for the second sec

$$a_{\mu}^{exp} - a_{\mu}^{SM} = 288(63)(49) \times 10^{-11} \sim 3.5\sigma$$

C. Patrignani et al. (Particle Data Group), Chin. Phys. C40 (2016)

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But this model-dependent splitting between L.D. & S.D contributions does not need to be made in a framework capable of dealing consistently with both extreme regimes (and intermediate regions) simultaneously, like DSE

Role of  $\pi^0$  TFF in  $a_{\mu}^{\ \ P^0, \text{HLbL}}$ 







Role of 
$$\pi^0$$
 TFF in  $a_{\mu}^{P^0, \text{HLbL}}$ 

External on-shell photon







Internal off-shell photons





### $\pi^0$ TFF from Dyson-Schwinger equations

K. Raya, L. Chang, A. Bashir, J. J. Cobos-Martínez, L. X. Gutiérrez-Guerrero, C. D. Roberts, P. C. Tandy Phys. Rev. D93 (2016) no.7, 074017



### How to parametrize DSE $\pi^0$ TFF? (III)

 $h_3 (h_4) \& h_6$  are still free parameters. We will fit  $\delta_{BL}$  to DSE data in the region relevant for  $a_{\mu} (Q_i^2 \le 10 \text{ GeV}^2)$ According to different estimates  $|h_3|(|h_4|) \le 10 \text{ GeV} \& h_6 \le 10 \text{ GeV}$ 

#### $\pi^0$ TFF from Dyson-Schwinger equations

![](_page_16_Figure_1.jpeg)

## Evaluation of $a_{\mu}^{\ \pi^{0}\text{-pole}}$ with DSE input

We have evaluated  $a_{\mu}^{\pi^{0-pole}}$  varying parameters in the ranges discussed previously

$M_{V_2}$ (GeV)	$\delta_{BL}$
1.440	$-0.57 \pm 0.18$
1.465	$-0.43 \pm 0.17$
1.490	$-0.36 \pm 0.18$

Table 1: Dependence of  $\delta_{BL}$  on  $M_{V_2}$ .

 $|h_3|(|h_4|) \le 10 \text{ GeV } \& h_6 \le 10 \text{ GeV}$ 

$$h_3 + h_4 = M_{V_1}^2 M_{V_2}^2 \chi$$

 $\chi = -(3.05 \pm 0.20) \text{GeV}^{-2}$ 

The only relevant variations are h<sub>3</sub> (h<sub>4</sub>) &  $\delta_{BL}$ 

 $a_{\mu}^{\pi^{0}-\text{pole}} = (6.26 \pm 0.08)10^{-10}$ 

## Evaluation of $a_{\mu}^{\pi^{0}-\text{pole}}$ with DSE input: Conclusions

We have evaluated  $a_{\mu}^{\pi^{0}\text{-pole}}$  varying parameters in the ranges discussed previously

$$a_{\mu}^{\pi^{0}\text{-pole}} = (6.26 \pm 0.08)10^{-10}$$

What's next? Obtention of on/off-shell  $\eta/\eta'/\eta_{b,c}$  TFF and evaluation of the corresponding contributions to  $a_{\mu}$  (Khepani Raya, Minghui Ding, Adnan Bashir, Lei Chang, Craig D. Roberts; Phys.Rev. D95 (2017) 074014).

Our result is approx. 8% higher than pion-pole evaluation of Nyffeler. This makes us **optimistic** with respect to a reduction of the exp-theo discrepancy in  $a_{\mu}$ .

#### Bern's group work clarifies that only pole contributions are needed

Some reference values...According to Roig, Guevara & López-Castro,  
Phys.Rev. D89 (2014) no.7, 073016
$$a_{\mu}^{\pi^{0},HLbL} = (5.75 \pm 0.06) \cdot 10^{-10}$$
 $a_{\mu}^{\pi^{0},HLbL} = (6.66 \pm 0.21) \cdot 10^{-10}$  $a_{\mu}^{P,HLbL} = (10.47 \pm 0.54) \cdot 10^{-10}$  $a_{\mu}^{\eta,HLbL} = (1.44 \pm 0.26) \cdot 10^{-10}$  $a_{\mu}^{\eta,HLbL} = (2.04 \pm 0.44) \cdot 10^{-10}$  $a_{\mu}^{BSM} \leq 288 \times 10^{-11}$  $a_{\mu}^{\eta',HLbL} = (1.08 \pm 0.09) \cdot 10^{-10}$  $a_{\mu}^{\eta',HLbL} = (1.77 \pm 0.23) \cdot 10^{-10}$  $a_{\mu}^{BSM} \leq 288 \times 10^{-11}$