Linear response of entanglement entropy from holography

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Based on:

ARXIV:1602.05934, SANDIPAN KUNDU & JFP ARXIV:1705.10324, SAGAR LOKHANDE, GERBEN OLING & JFP

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Linear response of entanglement entropy

Outline

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- 8 Entanglement entropy after instantaneous quenches
 - Spread of entanglement in (1+1)-dimensional CFTs
 - Perturbative computation in higher dimensions

4 Linear response of entanglement entropy

- Fefferman-Graham expansion
- Entanglement entropy after general global quenches
- General properties and examples

Conclusions

Motivation: Quenches

Objective:

Characterize generic out-of-equilibrium states in AdS/CFT

- Are there universal rules that govern the evolution?
- What is the behavior of the various observables?
- What is the nature of the final state?

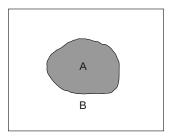
Simplest dynamical process is a quantum quench:

 $H_{\lambda} = H_0 + \lambda(t, \vec{x}) \, \delta H_{\Delta} \implies \mathcal{L}_{\lambda} = \mathcal{L}_0 + \lambda(t, \vec{x}) \, \mathcal{O}_{\Delta}$

 $\mbox{Outcome:} \ \begin{cases} \mbox{Thermalization:} \ \rho(t) \rightarrow \rho_{\rm thermal} + \mathcal{O}(e^{-S}) \\ \mbox{Relaxation:} \ \rho(t) \rightarrow \rho_{\rm initial} \\ \mbox{Quantum revivals:} \ \rho(t) \rightarrow \rho(t - t_{\rm p}) \end{cases} \end{cases}$

Motivation: Observables

- One-point functions of local operators \mathcal{O}_i thermalize fast
- A useful order parameter is entanglement entropy S_A



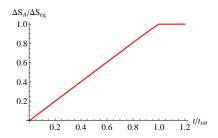
Hilbert space factorizes $\mathcal{H}_{total} = \mathcal{H}_{A} \otimes \mathcal{H}_{B}$, then define $\rho_{A} \equiv tr_{B}[\rho]$

$$S_A = -\mathrm{tr}[
ho_A \log
ho_A]$$

Review of the 'entanglement tsunami' proposal

[Calabrese & Cardy] showed that for weakly coupled (1 + 1) CFTs:

$$\Delta S_{\mathcal{A}}(t) = 2t \, s_{\mathsf{eq}} \,, \qquad t \leq t_{\mathsf{sat}} = R$$



- Instantaneous quenches: $\lambda(t) \sim \delta(t) \leftrightarrow \langle T_{00}(t)
 angle \sim \theta(t)$
- Large subsystems: $R \gg 1/T$
- Explanation in terms of free streaming EPR pairs

Review of the 'entanglement tsunami' proposal Q: How do (strong) interactions affect this result?

Early numerical explorations [Abajo-Arrastia et.al, Balasubramanian et.al] and analytical work [Hartman & Maldacena, Liu & Suh] showed that:

$$\Delta S_{A}(t) = v_{E}s_{
m eq}A_{\Sigma}t\,, \qquad t_{
m loc} \ll t \ll t_{
m sat}$$

$$v_E = \sqrt{\frac{d}{d-2}} \left(\frac{d-2}{2(d-1)}\right)^{\frac{d-1}{d}}$$

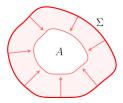


Figure: Pictorial representation of the entanglement tsunami of region A.

Review of the 'entanglement tsunami' proposal

Comments:

- $v_E = 1$ in d = 2, as for weakly-coupled theories!
 - EE for multiple strips differs for d = 2 [Asplund & Bernamonti]
 - Quasiparticle description fails at large-c [Asplund et.al]
 - In higher dimensions [Casini, Liu & Mezei]

$$v_{E}^{\text{free}} = rac{\Gamma[rac{d-1}{2}]}{\sqrt{\pi}\Gamma[rac{d}{2}]} \leq v_{E}$$

- $v_E \leq 1$ suggests a causality bound
 - [Liu & Suh] conjectured that $\Re(t) \equiv \frac{1}{s_{eq}A_{\Sigma}} \frac{dS_A}{dt} \leq v_E \leq 1$
 - Proof by [Casini, Liu & Mezei] and [Hartman & Afkhami-Jeddi]
 - Bound fails for small subsystems! No linear growth [Kundu & Pedraza]

Models of global quenches in holography

- In QFT: deform the theory by an operator O_Δ with an homogeneous time dependent coupling λ(t)
- In AdS: turn on the non-normalizable mode of the field dual to \mathcal{O}_Δ
- Electric field quench in (3+1) is analytically tractable (!)

$$S = \frac{1}{2\kappa^2} \int d^{3+1}x \sqrt{-g} \left(R - 2\Lambda - F^2 \right)$$

A solution for an arbitrary E(v) is: [Horowitz, Iqbal & Santos]

$$ds^{2} = \frac{1}{z^{2}} \left(-f(v,z)dv^{2} - 2dvdz + dx^{2} + dy^{2} \right)$$

$$F = -E(v)dv \wedge dx$$

where

$$f(v,z) = 1 - z^3 m(v), \quad m(v) = \frac{1}{2} \int_{-\infty}^{v} E(v')^2 dv'$$

Models of global quenches in holography

• Another example is a scalar field quench:

$$S = rac{1}{2\kappa^2} \int d^{d+1}x \sqrt{-g} \left(R - 2\Lambda - rac{1}{2} (\partial \phi^2)
ight)$$

Tractable perturbatively: [Bhattacharyya & Minwalla]

$$ds^{2} = \frac{1}{z^{2}} \left[-f(v,z)dv^{2} - 2dvdz + g(v,z)(dx^{2} + dy^{2}) \right]$$

$$\phi = \phi(v,z)$$

where, for any $\phi_0(v)$:

$$f(v,z) = 1 + \left(\frac{3}{4}z^{2}\dot{\phi}_{0}^{2} - z^{3}m(v)\right)\epsilon^{2} + \cdots, \quad m(v) = -\frac{1}{2}\int_{-\infty}^{v} dt \,\dot{\phi}_{0}\,\dot{\phi}_{0}$$
$$g(v,z) = 1 - \frac{1}{4}z^{2}\dot{\phi}_{0}^{2}\epsilon^{2} + \cdots, \qquad \phi(v,z) = (\phi_{0} + z\dot{\phi}_{0})\epsilon + \cdots$$

Observables are rather insensitive to transients [Joshi et.al.]

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Models of global quenches in holography: AdS-Vaidya

• A general theory:

$$S = rac{1}{2\kappa^2}\int d^{d+1}x\sqrt{-g}\left(R-2\Lambda+\mathcal{L}_{matter}
ight)$$

admits an AdS-Vaidya solution:

$$ds^{2} = \frac{1}{z^{2}} \left[-f(v,z)dv^{2} - 2dvdz + d\vec{x}^{2} \right], \quad f(v,z) = 1 - z^{d}m(v)$$

provided that the stress-tensor is made of null dust:

$$T_{\mu\nu} = \frac{d-1}{4\kappa^2} z^{d-1} \frac{dm}{d\nu} \delta^{\nu}_{\mu} \delta^{\nu}_{\nu} \tag{1}$$

• Limitations:

- Phenomenological approach: source is not known
- ▶ Is any m(v) physically reasonable? \rightarrow NEC requires $dm/dv \ge 0$

Models of global quenches in holography: AdS-RN-Vaidya

• A further generalization of the AdS-Vaidya geometry is:

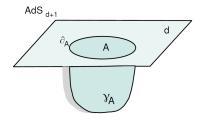
$$f(z, v) = 1 - z^d m(v) + q(v)^2 z^{2(d-1)}$$

which requires:

$$T_{\mu\nu} = \left(\frac{d-1}{4\kappa^2} z^{d-1} \frac{dm}{dv} - \frac{d-2}{2\kappa^2} z^{2d-3} q(v) \frac{dq}{dv}\right) \delta^{\nu}_{\mu} \delta^{\nu}_{\nu}$$

- Assuming $m(-\infty) = 0$, $q(-\infty) = 0$, it interpolates between AdS and AdS-RN (CFT vacuum to a state with finite T and μ).
- NEC is naively violated! but m(v) and q(v) are further constrained by SSA [Caceres, Kundu, Pedraza & Tangarife]
- Charge can be included in the perturbative collapse framework [Caceres, Kundu, Pedraza & Yang]

Holographic entanglement entropy

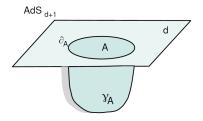


Prescription: [Ryu & Takayanagi]

$$S_A = rac{\operatorname{Area_{min}}(\gamma_A)}{4G_N^{d+1}}$$

- $\gamma_A = \text{codimension-2 surface s.t.} \ \partial \gamma_A = \partial A$
- Homology constraint: $\gamma_A \sim A$
 - ▶ ∃ bulk region \Re s.t. $\partial \Re = \gamma_A \cup A$

Holographic entanglement entropy



Covariant prescription: [Hubeny, Rangamani & Takayanagi]

$$S_A = rac{\operatorname{Area}_{\mathsf{ext}}(\gamma_A)}{4G_N^{d+1}}$$

- $\gamma_A = \text{codimension-2 surface s.t.} \ \partial \gamma_A = \partial A$
- Homology constraint: $\gamma_A \sim A$
 - ▶ ∃ bulk region \Re s.t. $\partial \Re = \gamma_A \cup A$

Spread of entanglement in (1 + 1) dimensions

For (1 + 1) CFTs EE is known in a closed form! [Balasubramanian et.al] \rightarrow Consider a segment of length $\ell = 2R$ and define:

$$\mathfrak{t}=2\pi Tt\,,\qquad \qquad \mathfrak{l}=2\pi TR$$

 \bullet At $\mathfrak{t} \to \infty$ EE reaches the equilibrium value:

$$S_A = rac{c}{3} \log\left(rac{R}{\epsilon}
ight) + rac{c}{3} \log\left(rac{\sinh \mathfrak{l}}{\mathfrak{l}}
ight) \equiv S_{\mathsf{vac}} + \Delta S_A$$

For
$$l \gg 1$$
: $\Delta S_A \simeq \frac{cl}{3} = s_{eq}V_A$, $s_{eq} = \frac{\pi cT}{3}$
The first law reads: $\left| \frac{d(\Delta E_A)}{d(\Delta S_A)} \right|_{\ell} = T$
where $\Delta E_A = \mathcal{E}V_A$, $\mathcal{E} = \frac{\pi cT^2}{6}$

Spread of entanglement in (1 + 1) dimensions

For (1 + 1) CFTs EE is known in a closed form! [Balasubramanian et.al] \rightarrow Consider a segment of length $\ell = 2R$ and define:

$$\mathfrak{t} = 2\pi T t$$
, $\mathfrak{l} = 2\pi T R$

• At $\mathfrak{t} \to \infty$ EE reaches the equilibrium value:

$$S_A = \frac{c}{3} \log\left(\frac{R}{\epsilon}\right) + \frac{c}{3} \log\left(\frac{\sinh l}{l}\right) \equiv S_{vac} + \Delta S_A$$

For
$$\mathfrak{l} \ll 1$$
: $\Delta S_A \simeq \frac{c \mathfrak{l}^2}{18} = \frac{c \pi^2 T^2 \ell^2}{18}$
The first law reads: $\left| \frac{d(\Delta E_A)}{d(\Delta S_A)} \right|_{\ell} = T_A$ [Bhattacharya et.al]
where $T_A = \frac{3}{\pi \ell} \rightarrow \Delta S_A = \frac{\Delta E_A}{T_A} = \frac{\mathcal{E}V_A}{T_A} = s_{eq}V_A$

Spread of entanglement in (1 + 1) dimensions For $t \le t_{sat} = I$ [Balasubramanian et.al]:

$$S_A(\mathfrak{t}) = S_{vac} + \Delta S_A(\mathfrak{t}), \qquad \Delta S_A(\mathfrak{t}) = \frac{c}{3} \log\left(\frac{\sinh \mathfrak{t}}{\mathfrak{l} \, s(\mathfrak{l},\mathfrak{t})}\right)$$

where

$$\begin{split} \mathfrak{l} &= \frac{\sqrt{1-s^2}}{\rho s} + \frac{1}{2} \log \left(\frac{2(1+\sqrt{1-s^2})\rho^2 + 2s\rho - \sqrt{1-s^2}}{2(1+\sqrt{1-s^2})\rho^2 - 2s\rho - \sqrt{1-s^2}} \right) \\ \rho &= \frac{1}{2} \coth \mathfrak{t} + \frac{1}{2} \sqrt{\frac{1}{\sinh^2 \mathfrak{t}} + \frac{1-\sqrt{1-s^2}}{1+\sqrt{1-s^2}}} \end{split}$$

One key observation:

$$v_E^{avg} = \langle \mathfrak{R}(t)
angle = rac{1}{s_{eq}A_{\Sigma}} rac{\Delta S_A}{\Delta t} = rac{1}{s_{eq}A_{\Sigma}} rac{s_{eq}V_A}{t_{sat}} = rac{R}{t_{sat}} = 1$$

Spread of entanglement in (1 + 1) dimensions

Two possibilities:

- $\mathfrak{R}(t) = 1 \hspace{0.2cm} (\forall \hspace{0.1cm} \mathfrak{l}, \mathfrak{t})
 ightarrow \mathsf{Causality}$ constraint is OK
- $\max[\mathfrak{R}(t)] > 1
 ightarrow \mathsf{Causality}$ constraint is violated

It suffices to analyze the early time behavior $\mathfrak{t} \ll \mathfrak{t}_{sat}$:

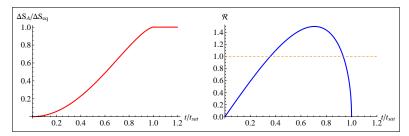
$$\rho = \frac{1}{\mathfrak{t}} + \frac{\mathfrak{t}}{12} + \cdots, \qquad s = \frac{\mathfrak{t}}{\mathfrak{t}} \left(\frac{1}{\mathfrak{t}} - \frac{\mathfrak{t}}{12} + \cdots \right)$$
$$\Delta S_{\mathcal{A}}(\mathfrak{t}) = \frac{c\mathfrak{t}^2}{12} + \mathcal{O}(\mathfrak{t}^4) = 2\pi \mathcal{E} \mathfrak{t}^2 + \cdots$$

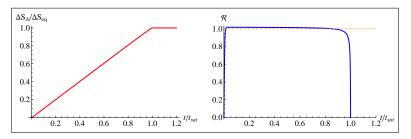
Therefore,

$$\Re(t) = \frac{2\pi \mathcal{E}t}{s_{\mathsf{eq}}} + \cdots \qquad \rightarrow \qquad \max[\mathfrak{R}(t)] > 1!$$

Spread of entanglement in (1 + 1) dimensions

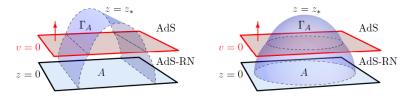
Numerical results for $l = 10^{-2}$ and $l = 10^2$. $\Re(t) \to 1$ as $l \to \infty$.





Perturbative calculation for small intervals

Setup in the gravity side:



- For large subsystems one solves for Γ_A above and below the shell and then match the solutions in a perturbatively in the IR [Liu & Suh]
- We developed a different perturbative expansion for small subsystems. In this case the extremal surfaces stay mostly in the UV [Kundu & Pedraza]

Perturbative calculation for small intervals

Let us expand the area functional A and embedding $\phi = \{x(z), v(z)\}$ as

$$\begin{aligned} \mathcal{A}[\phi(z);\lambda] &= \mathcal{A}^{(0)}[\phi(z)] + \lambda \mathcal{A}^{(1)}[\phi(z)] + \mathcal{O}(\lambda^2) \\ \phi(z) &= \phi^{(0)}(z) + \lambda \phi^{(1)}(z) + \mathcal{O}(\lambda^2) \end{aligned}$$

The key observation is that:

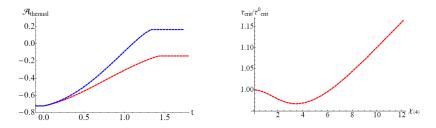
$$\mathcal{A}_{\text{on-shell}}[\phi(z)] = \int dz \,\mathcal{A}^{(0)}[\phi^{(0)}(z)] + \lambda \int dz \,\mathcal{A}^{(1)}[\phi^{(0)}(z)] \\ + \lambda \int dz \,\phi_i^{(1)}(z) \left[\frac{d}{dz} \frac{\partial \mathcal{A}^{(0)}}{\partial \phi_i(z)} - \frac{\partial \mathcal{A}^{(0)}}{\partial \phi_i(z)} \right]_{\phi^{(0)}} + \cdots$$

• We consider $\ell T \ll 1$, where $T \sim 1/z_H$

• Through the UV/IR connection $\ell \sim z$, so this corresponds to $z \ll z_H$ i.e. near the AdS boundary

Perturbative calculation for small intervals

Previous numerical results for AdS-RN-Vaidya [Caceres, Kundu]:



- Initial quadratic growth
- (Quasi)-linear intermediate regime
- Continuous saturation
- Non-monotonicity in the saturation time as a function of $\chi=\mu/\mathcal{T}$

Perturbative calculation for small intervals: the strip

At first order we only need the embedding in pure AdS:

$$\Delta S(t) = S(t) - S_{AdS} = \frac{\ell_{\perp}^{d-2} \varepsilon}{4G_N z_H^d} \int_0^{z_*} dz \, \theta(t-z) z \sqrt{1 - (z/z_*)^{2(d-1)}}$$

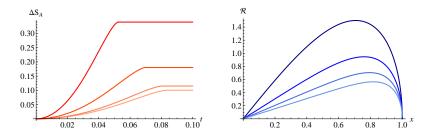
which leads to

$$\Delta S(t) = egin{cases} 0\,, & t < 0 \ \Delta S_{
m eq} imes \mathcal{F}(t/t_{
m sat})\,, & 0 \leq t \leq t_{
m sat}\,, \ \Delta S_{
m eq} & t > t_{
m sat} \ \end{cases}$$

where

$$\Delta S_{eq} = \frac{\sqrt{\pi} \Gamma[\frac{1}{d-1}] \ell_{\perp}^{d-2} \varepsilon z_{*}^{2}}{8(d+1) \Gamma[\frac{d+1}{2(d-1)}] z_{H}^{d} G_{N}^{(d+1)}}$$
$$\mathcal{F}(x) = \frac{2 \Gamma[\frac{d+1}{2(d-1)}] x^{2}}{\sqrt{\pi} \Gamma[\frac{1}{d-1}]} \left[\sqrt{1 - x^{2(d-1)}} + \frac{d-1}{2} {}_{2} F_{1}\left(\frac{1}{2}, \frac{1}{d-1}, \frac{d}{d-1}, x^{2(d-1)}\right) \right]$$

Perturbative calculation for small intervals: the strip Plots for $\Delta S(t)$ (for different μ/T) and $\Re(t)$ (for different *d*):



Initial quadratic growth regime:

$$\Delta S = \begin{cases} \frac{A_{\Sigma}}{4G_N} \left(\frac{4\pi T}{d}\right)^d \left(1 + \frac{d^2(d-2)}{16\pi^2} \left(\frac{\mu}{T}\right)^2 + \cdots\right) t^2, & T \gg \mu \\ \frac{A_{\Sigma}}{4G_N} \frac{2(d-2)^{d-1}\mu^d}{d^{d/2}(d-1)^{d/2-1}} \left(1 + \frac{2\pi d^{1/2}}{(d-1)^{1/2}} \frac{T}{\mu} + \cdots\right) t^2, & T \ll \mu \end{cases}$$

Same behavior than for large intervals [Liu & Suh]. Universal!

Perturbative calculation for small intervals: the strip

• Quasi-linear growth regime with:

$$\max[\Re(t)] = \frac{4(d-1)^{3/2} \Gamma[\frac{3d-1}{2(d-1)}] \Gamma[\frac{d}{2(d-1)}]}{d^{\frac{d}{2(d-1)}} \Gamma[\frac{1}{2(d-1)}] \Gamma[\frac{1}{d-1}]} = \begin{cases} 3/2, & d=2\\ 0.946, & d=3\\ 0.704, & d=4\\ \pi/d \to 0, & d \to \infty \end{cases}$$

- Causality constraint is violated only for d = 2.
- However $\langle \mathfrak{R}(t) \rangle \leq 1 \ (\forall d).$
- In this approximation $t_{sat} = z_* \sim \ell$ and $t_{sat}/t_{sat}^{(0)} = 1$. • At next order we find:

$$\frac{t_{\text{sat}}}{t_{\text{sat}}^{(0)}} = 1 - \alpha(d) \left(\frac{\mu}{T}\right)^2 (T\ell)^d + \mathcal{O}\left(\frac{\mu}{T}\right)^4 + \mathcal{O}(T\ell)^{2(d-1)}$$
(2)

where $\alpha(d) > 0!$ Next order is positive \rightarrow non-monotonic.

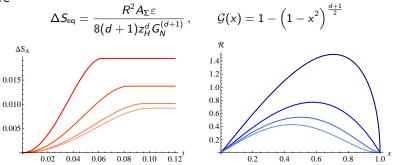
Perturbative calculation for small intervals: the ball

$$\Delta S(t) = \frac{\varepsilon A_{\Sigma} z_*^{d-2}}{8G_N^{(d+1)} R^{d-2} z_H^d} \int_0^{z_*} dz \, \theta(t-z) z \left[1 - (z/z_*)^2 \right]^{\frac{d-1}{2}}$$

which leads to

$$\Delta S(t) = egin{cases} 0\,, & t < 0\ \Delta S_{
m eq} imes \mathcal{G}(t/t_{
m sat})\,, & 0 \leq t \leq t_{
m sat}\ \Delta S_{
m eq} & t > t_{
m sat} \end{cases}$$

where



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Linear response of entanglement entropy

The entanglement growth looks like a convolution integral!

$$\Delta S_{\mathrm{s}}(t) \sim \int_{0}^{t_{*}} dt' \, heta(t-t')t' \sqrt{1-(t'/t_{*})^{2(d-1)}} = \mathfrak{f}(t) st \mathfrak{g}_{\mathrm{s}}(t)$$

$$\Delta S_{\mathsf{b}}(t) \sim \int_{0}^{t_{*}} dt' \, heta(t-t')t' \left[1-\left(t'/t_{*}
ight)^{2}
ight]^{rac{d-1}{2}} = \mathfrak{f}(t) st \mathfrak{g}_{\mathsf{b}}(t)$$

where $\mathfrak{f}(t) \sim heta(t)$ is the source and

$$\mathfrak{g}_{\mathsf{s}}(t) \sim t \sqrt{1 - (t/t_*)^{2(d-1)}}, \qquad \mathfrak{g}_{\mathsf{b}}(t) \sim t \left[1 - (t/t_*)^2
ight]^{rac{d-1}{2}}$$

are response functions for the strip and the ball, respectively.

Q: Can we understand this statement better? f(t) comes from the mass term m(v), do these expressions hold for more general sources?

Fefferman-Graham expansion for AdS spaces

Any asymptotically AdS metric can be written as:

$$ds^{2} = rac{1}{
ho^{2}} \left(g_{\mu
u}(
ho,x^{\mu}) dx^{\mu} dx^{
u} + d
ho^{2}
ight)$$

where $g_{\mu\nu}^{CFT} = g_{\mu\nu}(0, x^{\mu})$. Assuming $g_{\mu\nu}(\rho, x^{\mu}) = \eta_{\mu\nu} + \delta g_{\mu\nu}(\rho, x^{\mu})$:

$$\delta g_{\mu\nu} = \mathbf{a} \, \rho^{\mathbf{d}} \, \langle \mathbf{T}_{\mu\nu} \rangle + \rho^{2\mathbf{d}} \left(\mathbf{a}_1 \, \langle \mathbf{T}_{\mu\alpha} \mathbf{T}^{\alpha}{}_{\nu} \rangle + \mathbf{a}_2 \, \eta_{\mu\nu} \langle \mathbf{T}_{\alpha\beta} \, \mathbf{T}^{\alpha\beta} \rangle \right) + \dots$$

Corrections from additional operators (dual to vector A_{μ} or scalar ϕ):

$$\delta g_{\mu\nu} = \mathbf{a} \rho^{d} \langle T_{\mu\nu} \rangle + \rho^{2d-2} \left(b_{1} \langle J_{\mu} J_{\nu} \rangle + b_{2} \eta_{\mu\nu} \langle J_{\alpha} J^{\alpha} \rangle \right) + \dots$$
$$\delta g_{\mu\nu} = \mathbf{a} \rho^{d} \langle T_{\mu\nu} \rangle + c \rho^{2\Delta} \langle \mathcal{O}^{2} \rangle + \dots$$

The last term can dominate if $\frac{d}{2} - 1 < \Delta < \frac{d}{2}$ (do not consider these)

Entanglement entropy for small subsystems in equilibrium Indeed, for static spacetimes: [Bhattacharya et.al.]

$$\Delta S_A = \langle T_{00} \rangle \frac{V_A}{T_A} = \frac{\Delta E_A}{T_A}$$

where

$$T_{\rm s} = \frac{2(d^2 - 1)\Gamma[\frac{d+1}{2(d-1)}]\Gamma[\frac{d}{2(d-1)}]^2}{\sqrt{\pi}\Gamma[\frac{1}{d-1}]\Gamma[\frac{1}{2(d-1)}]^2\ell}, \qquad T_{\rm b} = \frac{d+1}{2\pi R}$$

Notice that T_A can also be obtained from the modular Hamiltonian:

$$H_{\rm b} = 2\pi \int_{\mathcal{A}} \frac{R^2 - r^2}{2R} T_{00}(x) d^{d-1}x$$

$$\downarrow$$

$$\Delta S_{\rm b} = 2\pi \delta \langle T_{00}(x) \rangle \Omega_{d-2} \int_{0}^{R} \frac{R^2 - r^2}{2R} r^{d-2} dr = \frac{2\pi \delta \langle T_{00}(x) \rangle \Omega_{d-2} R^d}{d^2 - 1}$$

Holds for local quenches! [Nozaki, Numasawa & Takayanagi]

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Entanglement entropy after general global quenches

For a general AdS-Vaidya:

$$\langle T_{00}(t) \rangle \equiv \varepsilon(t) = \frac{(d-1)m(t)}{16\pi G_N^{(d+1)}}, \quad \langle T_{ii}(t) \rangle \equiv P(t) = \frac{m(t)}{16\pi G_N^{(d+1)}}$$

And entanglement entropy satisfies:

$$\Delta S_A(t) = \int_{-\infty}^{\infty} dt' \, \mathfrak{f}(t-t') \mathfrak{g}_A(t')$$

where $f(t) = \varepsilon(t)$ and

$$g_{s}(t) = \frac{2\pi A_{\Sigma} t}{d-1} \sqrt{1 - (t/t_{*})^{2(d-1)}} \left[\theta(t) - \theta(t-t_{*})\right]$$
$$g_{b}(t) = \frac{2\pi A_{\Sigma} t}{d-1} \left[1 - (t/t_{*})^{2}\right]^{\frac{d-1}{2}} \left[\theta(t) - \theta(t-t_{*})\right]$$

Entanglement entropy after general global quenches Can be integrated by parts to obtain:

$$\Delta S_A(t) = \frac{\Delta E_A(t)}{T_A} + \int_{\infty}^{\infty} dt' \, \frac{d\varepsilon(t-t')}{dt} \mathfrak{G}_A(t')$$

• First law recovered for adiabatic quenches, provided:

$$rac{darepsilon(t')}{dt'}t_*\llarepsilon(t)\,,\qquad orall\,t'\in(t-t_*,t)$$

Second term can be interpreted as a kind of "relative entropy"

$$\Upsilon_{\mathcal{A}}(t) = rac{\Delta E_{\mathcal{A}}(t)}{T_{\mathcal{A}}} - \Delta S_{\mathcal{A}}(t) \geq 0$$

- Measures "distance" of out-of-equilibrium state w.r.t. equilibrium
- Positivity implies SSA!
- Can be rewritten using Ward identity, e.g. $\frac{d\varepsilon(t)}{dt} = \langle \mathcal{O}_{\Delta}(t) \rangle \frac{d\mathcal{J}_{\Delta}(t)}{dt}$

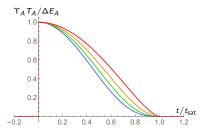
Juan F. Pedraza (UvA)

Entanglement entropy after general global quenches

• Second term can be interpreted as a kind of "relative entropy"

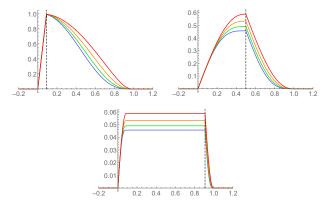
$$\Upsilon_{\mathcal{A}}(t) = rac{\Delta E_{\mathcal{A}}(t)}{T_{\mathcal{A}}} - \Delta S_{\mathcal{A}}(t) \geq 0$$

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- Positivity implies SSA!
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- Example 1. Instantaneous quench: $\varepsilon(t) = \varepsilon_0 \theta(t)$



Entanglement entropy after general global quenches

• Example 2. Linear-quench: $\varepsilon(t) = \alpha t \left[\theta(t) - \theta(t - t_q)\right] + \alpha t_q \theta(t - t_q)$



In the fully driven regime $\Upsilon_A = \text{constant}$, so we recover [O'Bannon et.al.]

$$\frac{d\Delta S_A(t)}{dt} = \frac{1}{T_A} \frac{d\Delta E_A(t)}{dt}$$

Conclusions

• Spread of EE after instantaneous quenches follow universal rules:

- For large subsystems the linear growth regime is universal, independent of the shape of Σ, but details depend on the final state (T, μ)
- For small subsystems the evolution depends on Σ, but is universal with respect to the state (only depends on ε(T, μ))
- Causality constrains the instantaneous rate of EE growth only for large subsystems. Average growth is always bounded by the speed of light.
- Spread of EE for small systems is described by a linear response
 - Q: Field theory interpretation of $g_A(t)$?
 - Q: Fluctuation-dissipation theorems?
- The quantity $\Upsilon_A(t)$ resembles a "relative entropy". Useful order parameter to characterize out-of-equilibrium excited states

Conclusions

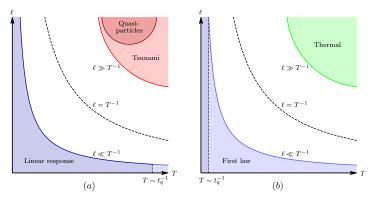


Figure: Schematic diagram of the different regimes of interest of entanglement propagation for (a) fast quenches $t_q \rightarrow 0$ and (b) slow quenches $t_q \rightarrow \infty$.