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# $L(\lambda)$

## **Monochromatization for Direct Higgs Production in Future Circular e<sup>+</sup>e<sup>-</sup> Colliders**

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# Outline

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- Motivation
- Objectives
- Introduction: Accelerator Physics
- Standard Monochromatization
- Radiation Effects In FCC
- Optimized Monochromatization
- Conclusion
- Discussion

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# Motivation

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- Questions still unanswered about Higgs boson properties
- Post-LHC high energy resolution requirements
- Upcoming radiation Effects
- FCCe+e- proposal opportunity

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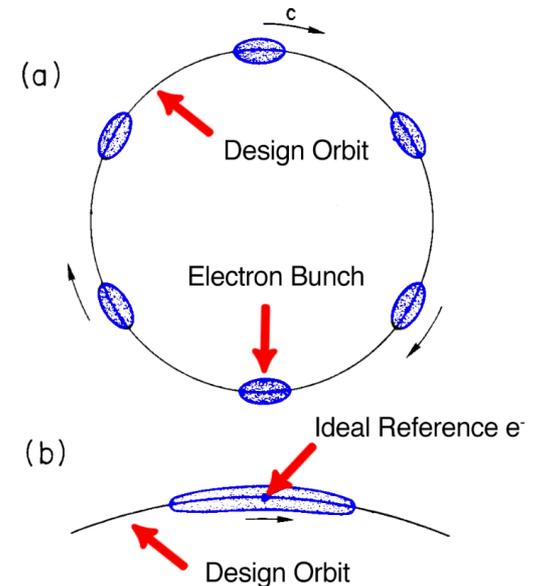
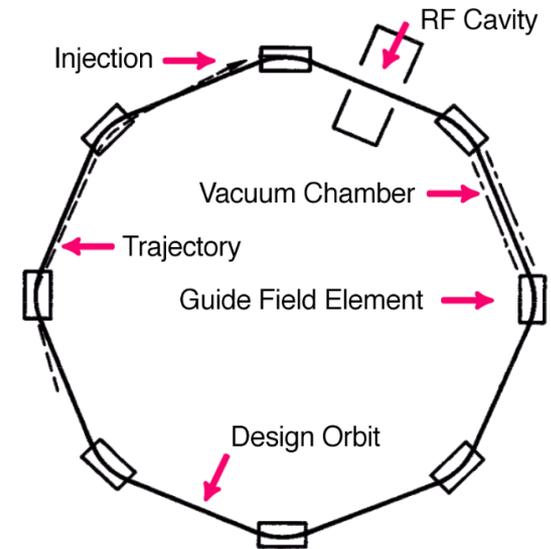
# Objectives

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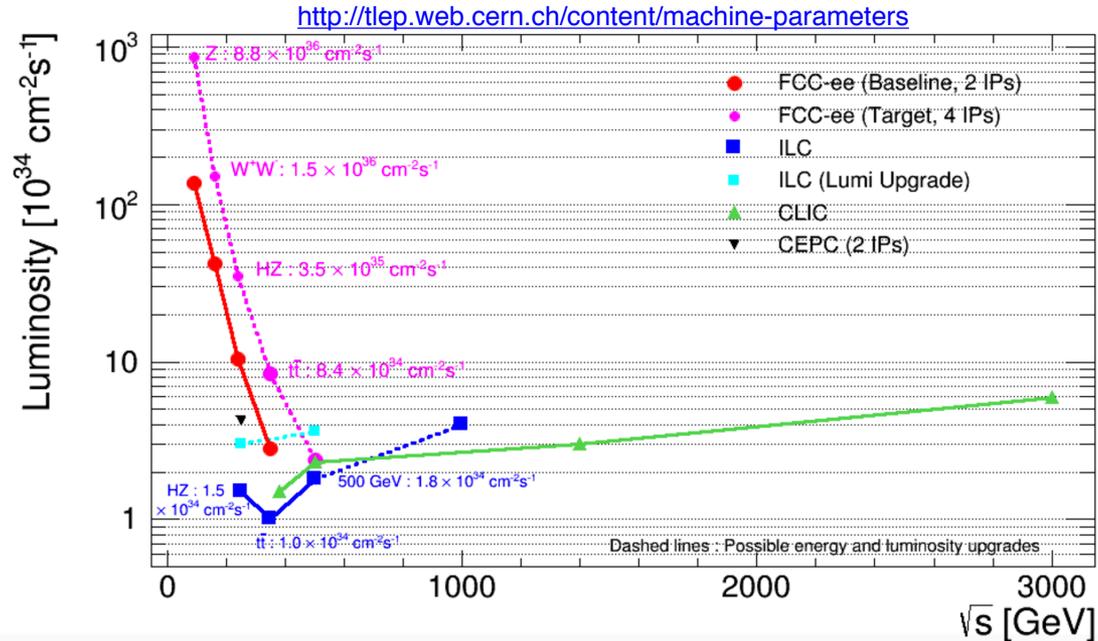
- Characterize the radiation effects at the FCCe+e-
- Quantify the IP parameters modification due to the radiation effects.
- Define the IP parameters to produce monochromatization
- Develop an optimized monochromatization scheme for the FCCe+e-

# Colliding Particles **A Qualitative Description**

- Particles are produced at the **source**
- Particles are injected into the **vacuum chamber**
- Stored particles loss energy by **radiation**, compensated by the **RFS**
- Particles are focused by guiding fields toward a **designed orbit**
- Accelerating field collects particles into circulating **bunches**
- Amplitud **damping** occurs due to energy loss by radiation
- Amplitude **excitation** occurs as a quantum effect
- Between damping and excitation a **balance** is reached
- Collision are produced at defined **interaction points**



# A proposal **Future Circular e+e- Collider**



- **Luminosity** is expected of order  $10e36$
- **Energy** range per beam should be from 40 to 175 GeV
- **Circumference** would be of 100 km
- **Energy Resolution** requires an improvement for direct Higgs production

# Lattice Design Courant-Snyder Theory

## Hill Differential Equation

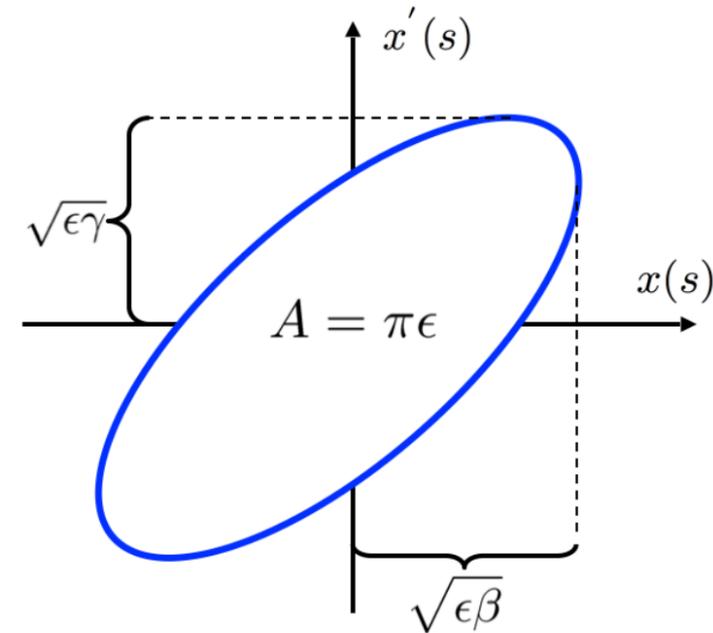
$$x'' + K(s)x = 0 \quad y'' - K(s)y = 0, \quad K(s) \equiv \frac{1}{(B\rho)} \left( \frac{\partial B_y}{\partial x} \right) (s)$$

$$K(s + C) = K(s)$$

Ansatz  $x(s) = A\omega(s) \cos(\phi(x) + \phi_0)$

## Courant-Snyder Parameters

$$\beta(s) \equiv \frac{\omega^2(s)}{k} \quad \alpha(s) \equiv -\frac{1}{2}\beta'(s) \quad \gamma(s) \equiv \frac{1 + \alpha^2(s)}{\beta(s)}$$



## Emittance

$$\epsilon = \pi\mathcal{W} = \pi[\gamma(s)x^2(s) + 2\alpha(s)x(s)x'(s) + \beta(s)x'^2(s)]$$

## Size of the Distribution

$$\sigma_{RMS} = \sqrt{\epsilon\beta(s)}$$

# Lattice Design Dispersion

## Off-Momentum Particles

$$p = p_0(1 + \delta)$$

## Relative Momentum Deviation

$$\delta \equiv \frac{\Delta p}{p_0}$$

## Momentum Contribution

$$x(s) = x_\beta(s) + x_\delta(s)$$

## Dispersion Function

$$x_\delta(s) = D(s)\delta$$

## Approximations

$$x \ll \rho \quad \delta \ll 1$$

## Dipole Contribution

$$\frac{\partial^2 x}{\partial \theta^2} + (1 - n)x = \rho \delta$$

## Dispersion Contribution

$$x(\theta) = A \cos \sqrt{1 - n}\theta + B \sin \sqrt{1 - n}\theta + \frac{\rho}{1 - n} \delta$$

## General Contribution

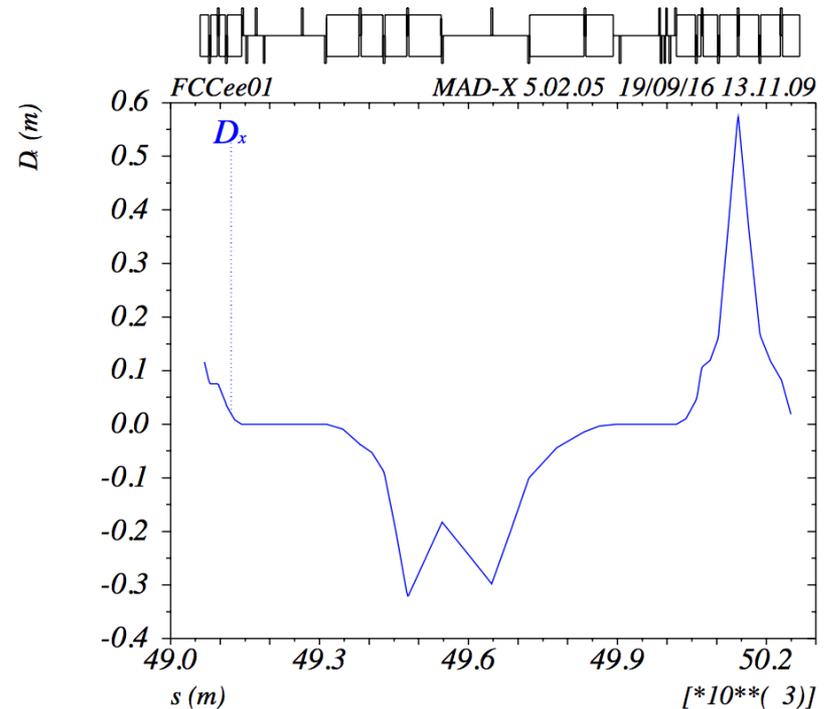
$$x'' + K(s)x = \frac{\delta}{\rho(s)}$$

## Dispersion Contribution

$$x(s) = C(s)x_0 + S(s)x'_0 + D(s)\delta_0$$

## Dispersion Function

$$D(s) = S(s) \int_0^s \frac{C(\tau)}{\rho(\tau)} d\tau - C(s) \int_0^s \frac{S(\tau)}{\rho(\tau)} d\tau$$



# Longitudinal Motion Energy Deviation Effects

Off-Energy Particles

$$E = E_0 + \epsilon,$$

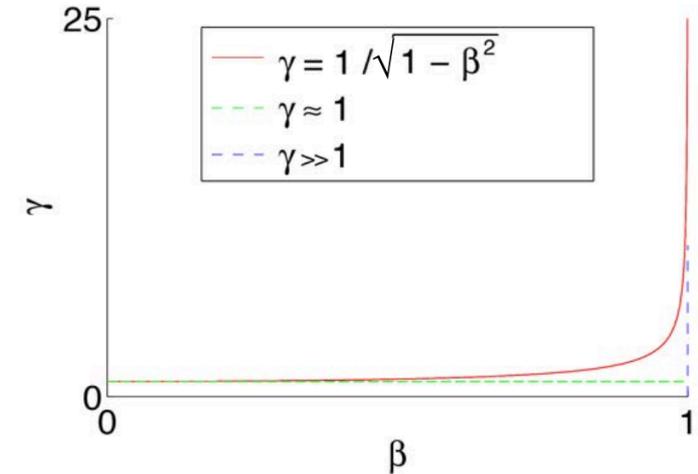
Off-Energy Contribution

$$x = x_\beta + x_\epsilon$$

Ultrarelativistic Limit

$$E \gg mc^2 \quad p = \frac{1}{c} \sqrt{E^2 - (mc^2)^2} \approx \frac{E}{c}$$

$$\frac{\Delta p}{p_0} = \frac{\Delta E}{E_0} \quad \delta = \frac{\epsilon}{E_0}$$



Off-Energy Contribution

$$x_\epsilon = D(s)\epsilon/E_0$$

Equation of Motion for Off-Energy Particles

$$x'' = -K_x x + G\left(\frac{\epsilon}{E_0}\right)$$

Radius of Curvature

$$G(s) = \rho^{-1}(s)$$

Dispersion Equation

$$D''(s) = -K_x D(s) + G(s)$$

# Beam Parameters **Distributions**

Transverse **Displacement**     $x = x_\beta + D_x \epsilon_0, \quad y = y_\beta + D_y \epsilon_0 \quad \epsilon_0 = \frac{\epsilon}{E_0}$

## Phase Space **Distribution**

$$f^\pm(x, y, \epsilon) = \frac{f^\pm(p_x, p_y, z)}{\sqrt{8\pi^3 \beta_x^* \epsilon_{xc} \beta_y^* \epsilon_{yc} \sigma_\epsilon^2}} \exp \left\{ -\frac{(x_\beta + D_x \epsilon_0)^2}{2\beta_x^* \epsilon_{xc}} - \frac{(y_\beta + D_y \epsilon_0)^2}{2\beta_y^* \epsilon_{yc}} - \frac{\epsilon^2}{2\sigma_\epsilon^2} \right\}$$

Average at the **IP**     $\langle \mathcal{A} \rangle^\pm = \int f^\pm(X^\pm) \mathcal{A}(X^+, X^-) dX^\pm \quad \langle \mathcal{A} \rangle^* = \langle \langle \mathcal{A} \rangle^+ \rangle^- = \langle \langle \mathcal{A} \rangle^- \rangle^+$

## **Average** over distribution

$$\sigma_x^* = \sqrt{\langle x^2 \rangle^\pm} = \sqrt{\beta_x^* \epsilon_{xc} + D_x^{*2} \sigma_\epsilon^2} \quad \langle x \epsilon_0 \rangle^\pm = D_x^{*\pm} \sigma_\epsilon^2 \quad \sigma_\epsilon^* = \sqrt{\langle \epsilon_0^2 \rangle^\pm}$$

$$\sigma_y^* = \sqrt{\langle y^2 \rangle^\pm} = \sqrt{\beta_y^* \epsilon_{yc} + D_y^{*2} \sigma_\epsilon^2} \quad \langle y \epsilon_0 \rangle^\pm = D_y^{*\pm} \sigma_\epsilon^2$$

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# Notes

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- **Relative Momentum Deviation** of particles differs from ideality
- **Dispersion** quantifies the effects on displacement due to off-momentum particles
- **Emittance** is a measure of how far our distribution is from ideal orbit
- **Beta Function** is related to the physical transverse size of bunches

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# **Standard Monochromatization**

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# Monochromatization Colliding Beams

Collision Energy Spread  $\sigma_w = \sqrt{2}E_0\sigma_\epsilon$

Beam Energy Spread  $\sigma_\epsilon^2 \propto \frac{55\hbar c E_0^2}{32\sqrt{3}(mc^2)^3} \frac{I_3}{I_2} \frac{1}{J_\epsilon}$

Function of radius and  $J_\epsilon$   $\sigma_w \propto (\rho J_\epsilon)^{-1/2}$ .  $J_\epsilon \in [0.5, 2.5]$

Typical Options  $\left. \begin{array}{l} \rho \gg \rho_0 \\ J_\epsilon > J_{0\epsilon} \end{array} \right\} \Rightarrow \sigma_w < \sigma_{0w}$

# Monochromatization Colliding Beams

Collision Energy Spread  $\sigma_w = \sqrt{2}E_0\sigma_\epsilon$       Beam Energy Spread  $\sigma_\epsilon^2 \propto \frac{55\hbar c E_0^2}{32\sqrt{3}(mc^2)^3} \frac{I_3}{I_2} \frac{1}{J_\epsilon}$

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Typical Options  $\left. \begin{array}{l} \rho \gg \rho_0 \\ J_\epsilon > J_{0\epsilon} \end{array} \right\} \Rightarrow \sigma_w < \sigma_{0w}$

## Principle of Monochromatization

### Monochromatization Factor

$$\lambda \equiv \frac{\mathcal{L}_0}{\mathcal{L}}$$

Baseline Scheme  $D_{x^+}^* = -D_{x^-}^* = 0$        $\mathcal{L} = \mathcal{L}_0$

# Monochromatization principle A “simple explanation”

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**Standard collision** Dispersion has the same sign in the IP

$$e^- \begin{array}{c} \xrightarrow{E+\Delta E} \\ \xrightarrow{E} \\ \xrightarrow{E-\Delta E} \end{array} \begin{array}{c} \xleftarrow{E+\Delta E} \\ \xleftarrow{E} \\ \xleftarrow{E-\Delta E} \end{array} e^+$$

$$w = 2(E_0 + \epsilon)$$

# Monochromatization principle A “simple explanation”

**Standard collision** Dispersion has the same sign in the IP

$$\begin{array}{ccc} \frac{E+\Delta E}{\rightarrow} & \frac{E+\Delta E}{\leftarrow} & \\ e^- \frac{E}{\rightarrow} & \frac{E}{\leftarrow} e^+ & \\ \frac{E-\Delta E}{\rightarrow} & \frac{E-\Delta E}{\leftarrow} & \end{array}$$
$$w = 2(E_0 + \epsilon)$$

**Monochromatization** Dispersion has opposite sign in the IP

$$\begin{array}{ccc} \frac{E+\Delta E}{\rightarrow} & \frac{E-\Delta E}{\leftarrow} & \\ e^- \frac{E}{\rightarrow} & \frac{E}{\leftarrow} e^+ & \\ \frac{E-\Delta E}{\rightarrow} & \frac{E+\Delta E}{\leftarrow} & \end{array}$$
$$w = 2E_0 + 0(\epsilon)^2$$

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# Notes

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- **Baseline Parameters** fail to produce the required center-of-mass energy resolution for direct Higgs production at 125 GeV
- **Higgs Width** of  $\sim 4.2$  MeV requires an improvement in resolution by at least a factor of 10 since baseline produced a  $\sim 40$  MeV width distribution
- **Double Ring System** of the FCCee may produce dispersion at the interaction point as required for monochromatization.
- **Radiation Effects** could compromise the performance of a monochromatization scheme

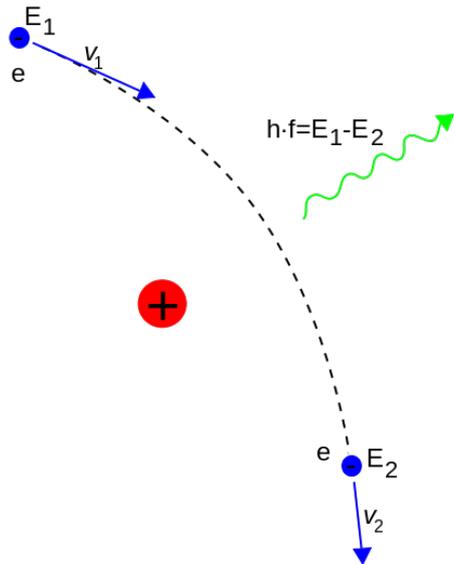
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# **Radiation Effects**

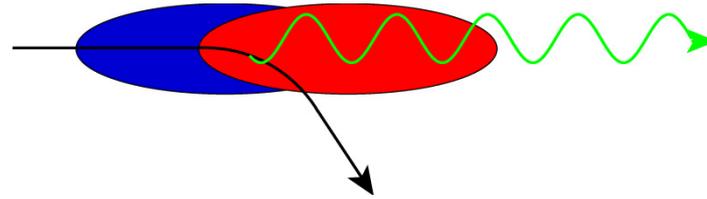
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# Beamstrahlung Sokolov-Ternov Theory

## Bremsstrahlung



## Beamstrahlung



## Lorentz Invariant

$$\Upsilon \equiv \frac{e}{m_e^3} \sqrt{|(F_{\mu\nu} p^\nu)^2|} = \frac{B}{B_c} = \frac{2}{3} \frac{\hbar \omega_c}{E_e}$$

$$\Upsilon_{\max} = 2 \frac{r_e^2 \gamma N_b}{\alpha \sigma_z (\sigma_x^* + \sigma_y^*)}$$

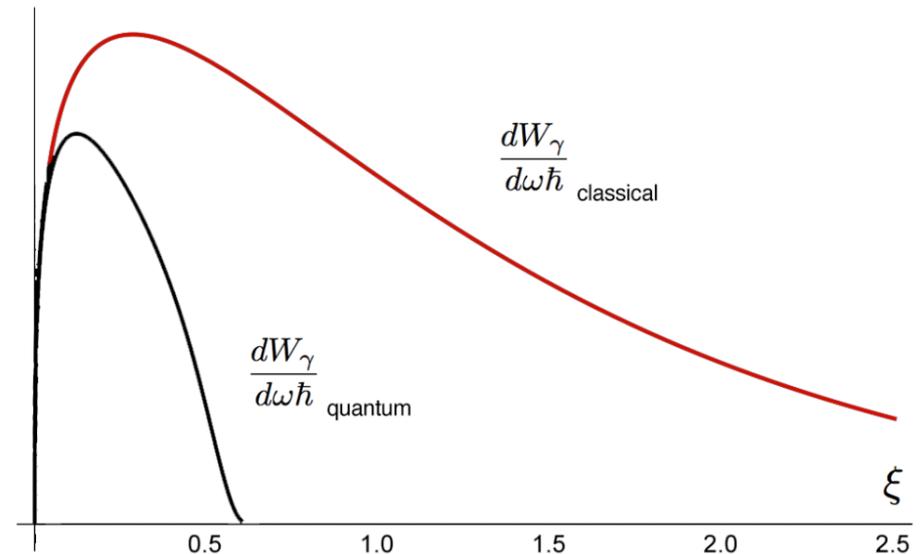
$$\Upsilon_{\text{ave}} \approx \frac{5}{12} \Upsilon_{\max} = \frac{5}{6} \frac{r_e^2 \gamma N_b}{\alpha \sigma_z^* (\sigma_x^* + \sigma_y^*)}$$

## Quantum Spectrum Formula

$$\frac{dW_\gamma}{d\omega \hbar} = \frac{\alpha}{\sqrt{3} \hbar \pi \gamma^2} \left( \int_\xi^\infty K_{5/3}(\xi') d\xi' + \frac{y^2}{1-y} K_{2/3}(\xi) \right)$$

$$\xi = \frac{2\hbar\omega}{3\Upsilon(E - \hbar\omega)}$$

$$y \equiv \omega/E_e$$



# Beamstrahlung Sokolov-Ternov Theory

Relative Energy Loss

$$\delta_B \approx \frac{24}{3\sqrt{3}\pi^{3/2}} \frac{r_e^3 \gamma N_b^2}{\sigma_z (\sigma_x + \sigma_y)^2} \approx \frac{24}{3\sqrt{3}\pi^{3/2}} \frac{r_e^3 \gamma N_b^2}{\sigma_z \sigma_x^2}$$

Average Emitted Photons

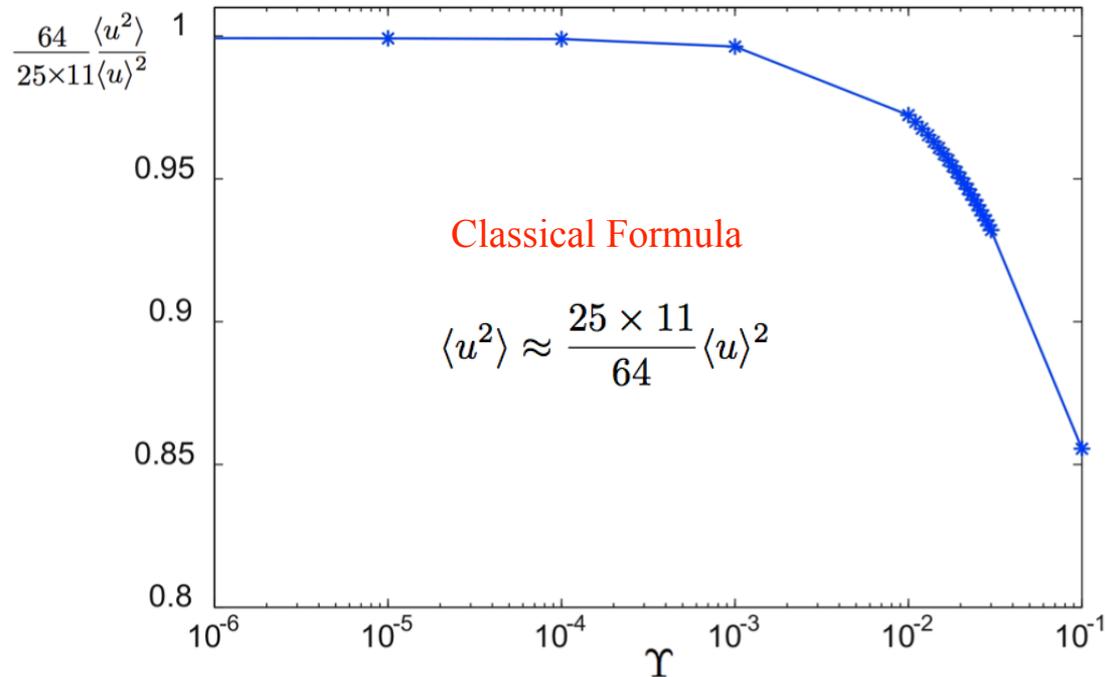
$$n_\gamma \approx \frac{12}{\pi^{3/2}} \frac{\alpha r_e N_b}{\sigma_x + \sigma_y} \approx \frac{12}{\pi^{3/2}} \frac{\alpha r_e N_b}{\sigma_x}$$

Statistical Properties

$$\langle u \rangle = \frac{\delta_B}{n_\gamma} \approx \frac{2\sqrt{3}}{9} \frac{r_e^2 N_b \gamma}{\alpha \sigma_z \sigma_x}$$

Excitation Term

$$n_\gamma \langle u^2 \rangle \approx 1.4 \frac{r_e^5 N_b^3 \gamma^2}{\alpha \sigma_z^2 (\sigma_x + \sigma_y)^3} \approx 192 \frac{r_e^5 N_b^3 \gamma^2}{\sigma_z^2 \sigma_x^3}$$



# Beamstrahlung Self-Consistent Energy Spread

Total Energy Spread  $\sigma_{\text{tot}}^2 = \sigma_{\delta,\text{SR}}^2 + \sigma_{\delta,\text{BS}}^2$

$$\sigma_{\delta,\text{tot}}^2 - \sigma_{\delta,\text{SR}}^2 = A \left( \frac{\sigma_{\delta,\text{SR}}}{\sigma_{\delta,\text{tot}}} \frac{1}{\sigma_{z,\text{SR}}} \right)^2$$

$$\sigma_{\delta,\text{tot}} = \left[ \frac{1}{2} \sigma_{\delta,\text{SR}}^2 + \left( \frac{1}{4} \sigma_{\delta,\text{SR}}^4 + A \frac{\sigma_{\delta,\text{SR}}^2}{\sigma_{z,\text{SR}}^2} \right)^{1/2} \right]^{1/2}$$

## Bunch Length

$$\sigma_{z,\text{tot}} = \frac{\alpha_C C}{2\pi Q_s} \sigma_{\delta,\text{tot}}$$

## Dispersion Invariant

$$\mathcal{H}_x^* \equiv \frac{(\beta_x^* D_x'^* + \alpha_x^* D_x^*)^2 + D_x^{*2}}{\beta_x^*}$$

## Self-Consistent Equations

$$\epsilon_{x,\text{tot}} = \epsilon_{x,\text{SR}} + \frac{\tau_x n_{\text{IP}}}{4T_{\text{rev}}} \{n_\gamma \langle u^2 \rangle\} \mathcal{H}_x^*$$

$$\sigma_{\delta,\text{tot}}^2 = \sigma_{\delta,\text{SR}}^2 + \frac{n_{\text{IP}} \tau_{E,\text{SR}}}{4T_{\text{rev}}} \{n_\gamma \langle u^2 \rangle\}$$

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# **Optimized Monochromatization**

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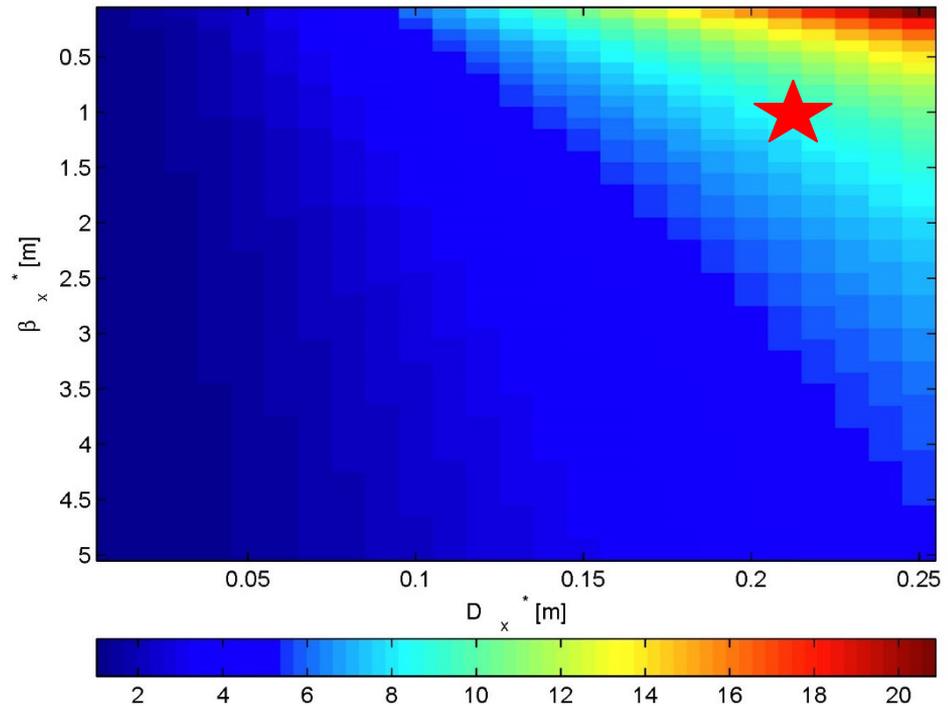
# Baseline monochromatization

$E_e$ [GeV]	62.5
scheme	m.c. basel.
$I_b$ [mA]	408.3
$N_b$ [ $10^{10}$ ]	3.3
$n_b$ [1]	25760
$n_{IP}$ [1]	2
$\beta_x^*$ [m]	1.0
$\beta_y^*$ [mm]	2
$D_x^*$ [m]	0.22
$\epsilon_{x,SR}$ [nm]	0.17
$\epsilon_{x,tot}$ [nm]	0.21
$\epsilon_{y,SR}$ [pm]	1
$\sigma_{x,SR}$ [ $\mu\text{m}$ ]	132
$\sigma_{x,tot}$ [ $\mu\text{m}$ ]	144
$\sigma_y$ [nm]	45
$\sigma_{z,SR}$ [mm]	1.8
$\sigma_{z,tot}$ [mm]	1.8
$\sigma_{\delta,SR}$ [%]	0.06
$\sigma_{\delta,tot}$ [%]	0.06
$\theta_c$ [mrad]	0
circ. $C$ [km]	100
$\alpha_C$ [ $10^{-6}$ ]	7
$f_{rf}$ [MHz]	400
$V_{rf}$ [GV]	0.4
$U_{0,SR}$ [GeV]	0.12
$U_{0,BS}$ [MeV]	0.01
$\tau_E/T_{rev}$	509
$Q_s$	0.030
$\Upsilon_{max}$ [ $10^{-4}$ ]	0.3
$\Upsilon_{ave}$ [ $10^{-4}$ ]	0.1
$\theta_c$ [mrad]	0
$\xi_x$ [ $10^{-2}$ ]	1
$\xi_y$ [ $10^{-2}$ ]	4
$\lambda$ [1]	9.2
$L$ [ $10^{35}$ $\text{cm}^{-2}\text{s}^{-1}$ ]	1.0
$\sigma_w$ [MeV]	5.8

Width of standard model Higgs 4-5 MeV requires  $\lambda \geq 10$

$$N_b = 3.3 \times 10^{10} \quad n_b = 25760 \quad \beta_y^* = 2 \text{ mm}$$

## Monochromatization Factor



★ :  $\lambda = 9.2, L = 1 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$

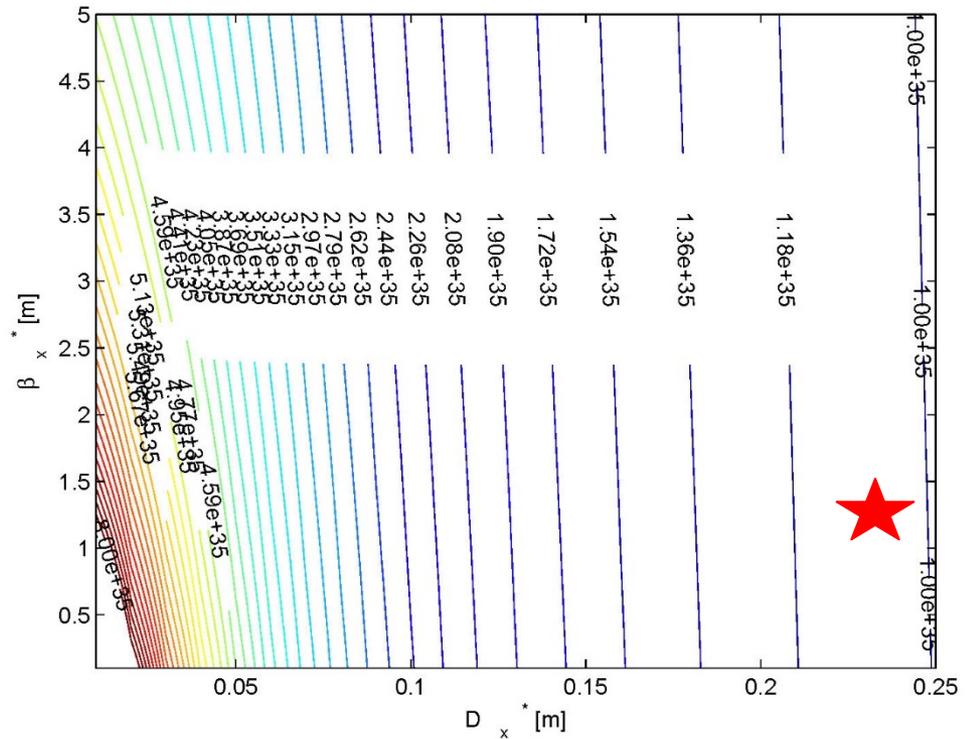
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$\sigma_{\delta,tot}$ [%]	0.06
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circ. $C$ [km]	100
$\alpha_C$ [ $10^{-6}$ ]	7
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$$N_b = 3.3 \times 10^{10} \quad n_b = 25760 \quad \beta_y^* = 2 \text{ mm}$$

## Luminosity contours



★ :  $\lambda = 9.2, L = 1 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$

# Optimized Monochromatization $L_{\text{max}}$ and $\lambda$

$$S = [0.1, 3], T = [0.1, 3]$$

$$\beta_{0x} = 1.0 \text{ m}, \beta_{0y} = 1.0 \text{ mm}$$

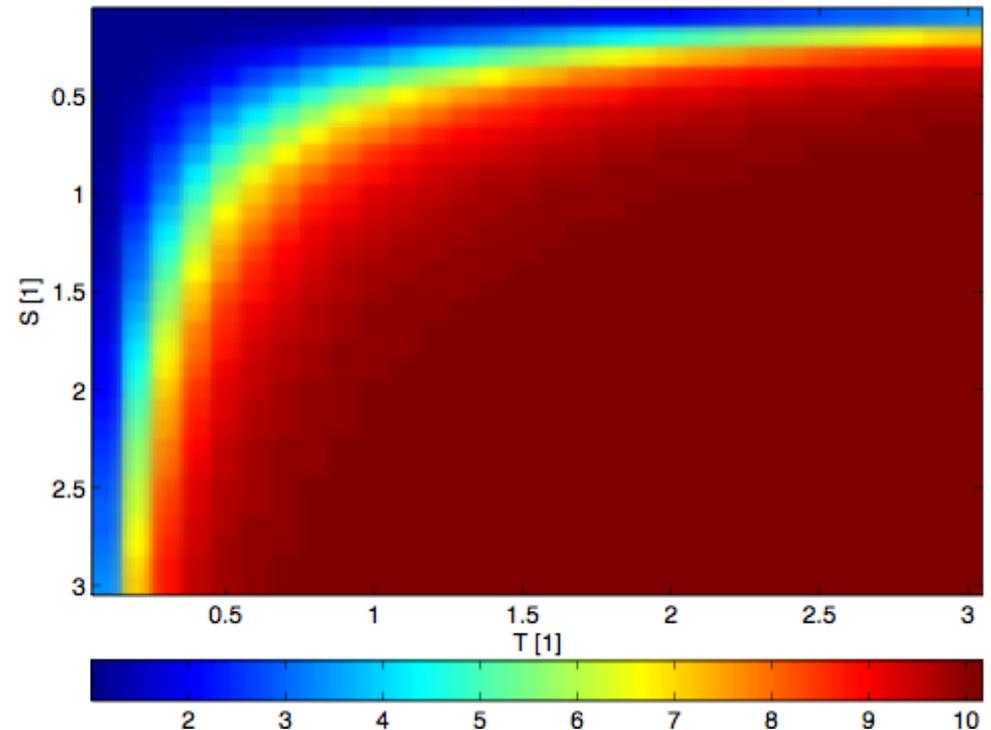
$$D_{0x} = 0.22 \text{ m}$$

$$N_{0b} = 3.3 \times 10^{10}, n_{0b} = 25760$$

$$\beta_x = \beta_{0x} * S^2, D_x^* = D_{0x} * S$$

$$N_b = N_{0b} / T, n_b = n_{0b} * T$$

## Monochromatization Factor



★ :  $\lambda_0 = 10.17321$

# Optimized Monochromatization $L_{\text{max}}$ and $\lambda$

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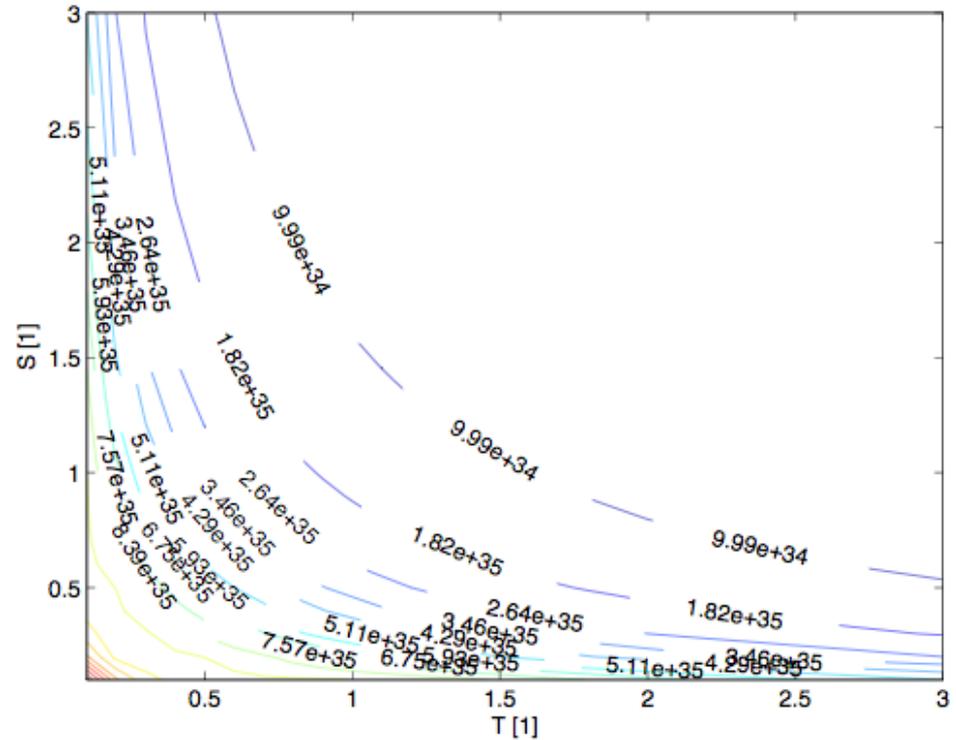
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★ :  $\lambda_0 = 10.17321, L \sim 1 \times 10^{36} \text{ cm}^{-2}\text{s}^{-1}$

# Optimized Monochromatization $L_{\max}$ and $\lambda$

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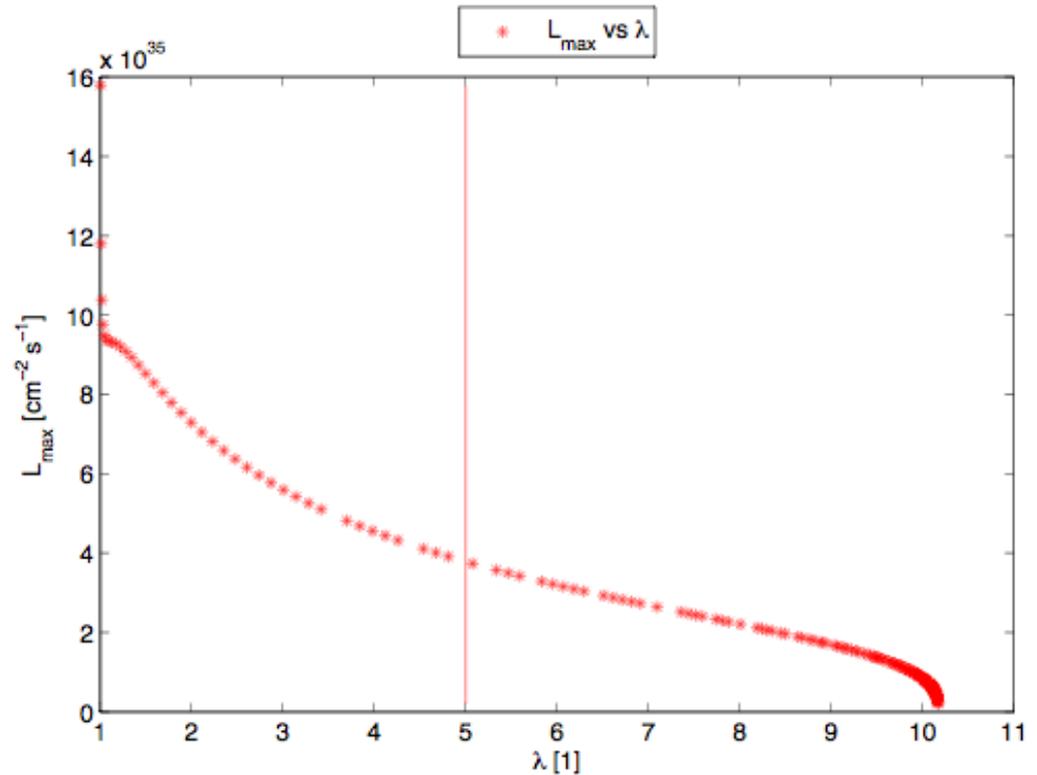
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$$N_b = N_{0b} / T, n_b = n_{0b} * T$$



$L_{\max}$  vs  $\lambda$

★ :  $\lambda_0 = 10.17321, L \sim 1 \times 10^{36} \text{ cm}^{-2} \text{s}^{-1}$

# Luminosity\_max and $\lambda$

$$S = [0.1, 3], T = [0.1, 3]$$

$$\beta_{0x} = 1.0 \text{ m}, \beta_{0y} = 1.0 \text{ mm}$$

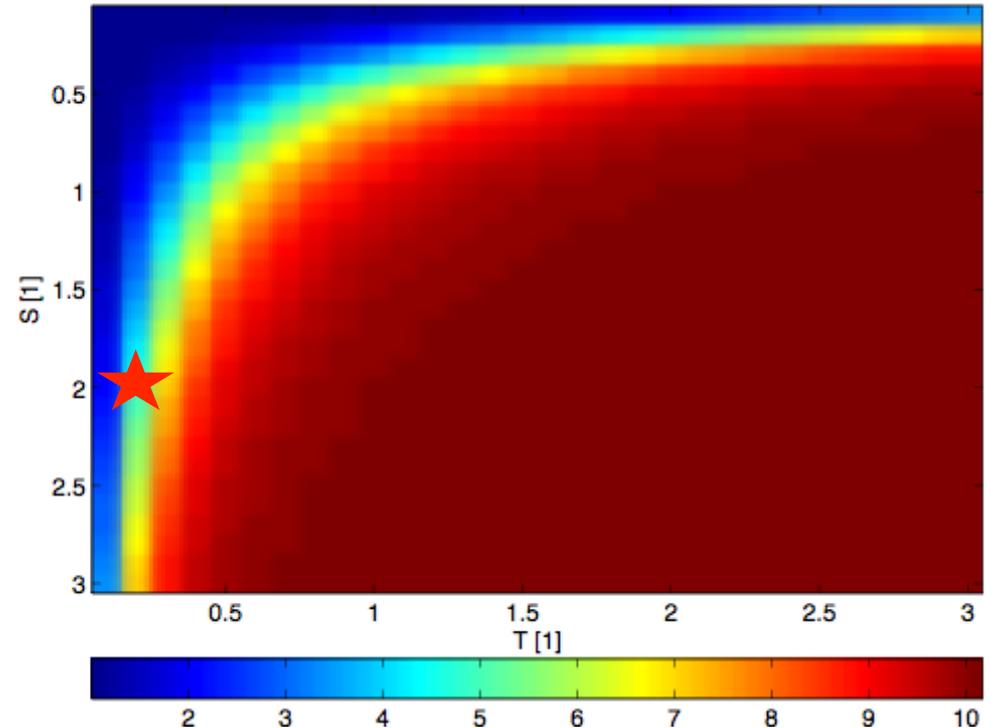
$$D_{0x} = 0.22 \text{ m}$$

$$N_{0b} = 3.3 \times 10^{10}, n_{0b} = 25760$$

$$\beta_x = \beta_{0x} * S^2, D_x * = D_{0x} * S$$

$$N_b = N_{0b} / T, n_b = n_{0b} * T$$

## Monochromatization Factor



★ :  $\lambda = 5.07, \beta = 1.96 \text{ m}, D_x = 0.308 \text{ m}, L = 3.736 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$

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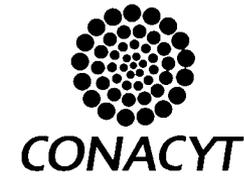
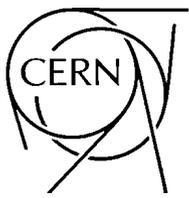
# Conclusions

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★ :  $\lambda = 5.07$ ,  $\beta = 1.96$  m,  $D_x = 0.308$  m,  $L = 3.736 \times 10^{35} \text{ cm}^{-2}\text{s}^{-1}$

$E_e$ [GeV]	45.6	62.5	62.5	62.5	80
scheme	CW	h.-o.	m.c. basel.	m.c. opt'd	CW
$I_b$ [mA]	1450.3	408.3	408.3	408.3	151.5
$N_b$ [ $10^{10}$ ]	3.3	1.05	3.3	11.1	6.0
$n_b$ [1]	91500	80960	25760	7728	5260
$n_{IP}$ [1]	2	2	2	2	2
$\beta_x^*$ [m]	1	1.0	1.0	1.96	1
$\beta_y^*$ [mm]	2	2	2	1	2
$D_x^*$ [m]	0	0	0.22	0.308	0
$\epsilon_{x,SR}$ [nm]	0.09	0.17	0.17	0.17	0.26
$\epsilon_{x,tot}$ [nm]	0.09	0.17	0.21	0.70	0.26
$\epsilon_{y,SR}$ [pm]	1	1	1	1	1
$\sigma_{x,SR}$ [ $\mu\text{m}$ ]	9.5	9.2	132	185.7	16
$\sigma_{x,tot}$ [ $\mu\text{m}$ ]	9.5	9.2	144	188.5	16
$\sigma_y$ [nm]	45	45	45	32	45
$\sigma_{z,SR}$ [mm]	1.6	1.8	1.8	1.8	2.0
$\sigma_{z,tot}$ [mm]	3.8	1.8	1.8	1.8	3.1
$\sigma_{\delta,SR}$ [%]	0.04	0.06	0.06	0.06	0.07
$\sigma_{\delta,tot}$ [%]	0.09	0.06	0.06	0.06	0.10
$\theta_c$ [mrad]	30	0	0	0	30
circ. $C$ [km]	100	100	100	100	100
$\alpha_C$ [ $10^{-6}$ ]	7	7	7	7	7
$f_{rf}$ [MHz]	400	400	400	400	400
$V_{rf}$ [GV]	0.2	0.4	0.4	0.4	0.8
$U_{0,SR}$ [GeV]	0.03	0.12	0.12	0.12	0.33
$U_{0,BS}$ [MeV]	0.5	0.05	0.01	0.01	0.21
$\tau_E/T_{rev}$	1320	509	509	509	243
$Q_s$	0.025	0.030	0.030	0.030	0.037
$\Upsilon_{max}$ [ $10^{-4}$ ]	1.7	0.8	0.3	0.85	4.0
$\Upsilon_{ave}$ [ $10^{-4}$ ]	0.7	0.3	0.1	0.35	1.7
$\theta_c$ [mrad]	30	0	0	0	30
$\xi_x$ [ $10^{-2}$ ]	5	12	1	2.22	7
$\xi_y$ [ $10^{-2}$ ]	13	15	4	6.76	16
$\lambda$ [1]	1	1	9.2	5.08	1
$L$ [ $10^{35}$ $\text{cm}^{-2}\text{s}^{-1}$ ]	9.0	2.2	1.0	3.74	1.9
$\sigma_w$ [MeV]	58	53	5.8	10.44	113

- **Monochromatization** scheme can be implemented
- **Beamstrahlung** effects may be controlled
- **Simulation** supports predictions
- **Lattice** designed is still in progress and the required modification should be possible
- **Theory** confirmation could be achieved at the FCCe+e-



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# $L(\lambda)$

## **Monochromatization for Direct Higgs Production in Future Circular e<sup>+</sup>e<sup>-</sup> Colliders**

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