



# Updating neutrino magnetic moment constraints.

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Blanca Cecilia Cañas Orduz  
O. G. Miranda, A. Parada, M. Tortola and J. W. F. Valle

Departamento de Física, CINVESTAV

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DPyC SMF

# Outline

- 1 THE NEUTRINO MAGNETIC MOMENT
  - Electromagnetic interactions
- 2 Neutrino data analysis
- 3 Results
- 4 Conclusions

# Electromagnetic interactions

In the Standard Model (SM)

$$\mathcal{H}_{em}^f(x) = j_\mu^f(x) A^\mu(x) = q_f \bar{f}(x) \gamma_\mu f(x) A^\mu(x),$$

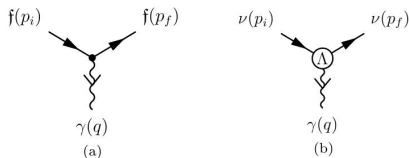


Figure 1: Tree-level coupling of a charged fermion  $f$  with a photon  $\gamma$  (a) and effective one-photon coupling of a neutrino with a photon (b).

- \* For neutrinos:  $q_\nu = 0 \rightarrow$  there are no electromagnetic interactions at tree level.
- \* However, such interactions can arise from loop diagrams at higher order in the perturbative expansion.

$$\mathcal{H}_{eff}(x) = j_\mu^{eff}(x) A^\mu(x) = \bar{\nu}(x) \Lambda_\mu \nu(x) A^\mu(x).$$

## Majorana neutrinos

If Majorana neutrinos are furnished with magnetic moments (MM) and electric dipole moments (EDM), the interaction with the electromagnetic field is described by the effective Hamiltonian <sup>1</sup>

$$\mathcal{H}_{em}^M = -\frac{1}{4}\nu^T C^{-1} (\mu + id\gamma_5) \sigma^{\alpha\beta} \nu F_{\alpha\beta} = -\frac{1}{4}\nu_L^T C^{-1} \lambda \sigma^{\alpha\beta} \nu_L F_{\alpha\beta} + h.c.,$$

with  $\nu = \nu_L + (\nu_L)^c$  and  $C$  is the charge conjugation matrix.

$$\mu^T = -\mu \quad (\text{MM matrix}), \quad d^T = -d \quad (\text{EDM matrix}).$$

The MM and EDM matrices are condensed in  $\lambda = \mu - id$  (NMM matrix).

The MM and EDM matrices are antisymmetric and hermitian, and, therefore, imaginary.  $\lambda$  is an antisymmetric matrix.

<sup>1</sup>J. Schechter and J. W. F. Valle, Phys. Rev. D24, 1883 (1981)

The above discussion could be translated into a more phenomenological approach in which the Dirac NMM is described by a complex matrix  $\lambda = \mu - id (\tilde{\lambda})$  in the flavor (mass) basis, while for the Majorana case the matrix  $\lambda$  takes the form

$$\lambda = \begin{pmatrix} 0 & \Lambda_\tau & -\Lambda_\mu \\ -\Lambda_\tau & 0 & \Lambda_e \\ \Lambda_\mu & -\Lambda_e & 0 \end{pmatrix}, \quad \tilde{\lambda} = \begin{pmatrix} 0 & \Lambda_3 & -\Lambda_2 \\ -\Lambda_3 & 0 & \Lambda_1 \\ \Lambda_2 & -\Lambda_1 & 0 \end{pmatrix},$$

where  $\lambda_{\alpha\beta} = \varepsilon_{\alpha\beta\gamma}\Lambda_\gamma$ .

The transition magnetic moments  $\Lambda_\alpha$  and  $\Lambda_i$  are complex parameters:

$$\Lambda_\alpha = |\Lambda_\alpha|e^{i\zeta_\alpha}, \quad \Lambda_i = |\Lambda_i|e^{i\zeta_i}.$$

# Neutrino-electron elastic scattering

$$\left(\frac{d\sigma}{dT}\right) (\nu + e^- \rightarrow \nu + e^-) = \left(\frac{d\sigma}{dT}\right)_{SM} + \left(\frac{d\sigma}{dT}\right)_{em}.$$

The SM contribution is

$$\left(\frac{d\sigma}{dT}\right)_{SM} = \frac{G_F^2 m_e}{2\pi} \left[ (g_V + g_A)^2 + (g_V - g_A)^2 \left(1 - \frac{T}{E_\nu}\right)^2 + (g_A^2 - g_V^2) \frac{m_e T}{E_\nu^2} \right],$$

where

$$g_V = \begin{cases} 2 \sin^2 \theta_W + \frac{1}{2} & \text{for } \nu_e \\ 2 \sin^2 \theta_W - \frac{1}{2} & \text{for } \nu_\mu, \nu_\tau \end{cases} \quad g_A = \begin{cases} \frac{1}{2} & \text{for } \nu_e \\ -\frac{1}{2} & \text{for } \nu_\mu, \nu_\tau \end{cases}$$

with  $g_A \rightarrow -g_A \implies (\bar{\nu}_e, \bar{\nu}_\mu, \bar{\nu}_\tau)$ .

The electromagnetic contribution is given by

$$\left(\frac{d\sigma}{dT}\right)_{em} = \frac{\pi\alpha^2}{m_e^2\mu_B^2} \left(\frac{1}{T} - \frac{1}{E_\nu}\right) \mu_\nu^2,$$

where  $\mu_\nu$  is an effective magnetic moment.

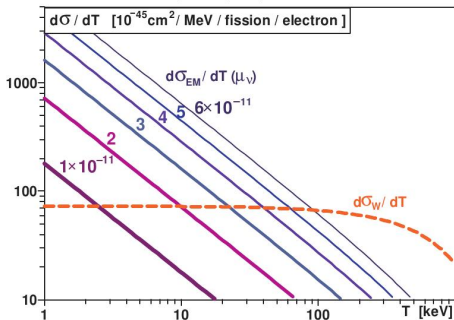


Figure 2: Weak and electromagnetic cross-sections calculated for several neutrino magnetic moment values.

$\mu_\nu$  is defined in the flavor basis as <sup>2</sup>

$$(\mu_\nu^F)^2 = a_-^\dagger \lambda^\dagger \lambda a_- + a_+^\dagger \lambda \lambda^\dagger a_+,$$

where  $a_-$  and  $a_+$  denote the negative and positive helicity neutrino, respectively. Whereas, in the mass basis

$$(\mu_\nu^M)^2 = \tilde{a}_-^\dagger \tilde{\lambda}^\dagger \tilde{\lambda} \tilde{a}_- + \tilde{a}_+^\dagger \tilde{\lambda} \tilde{\lambda}^\dagger \tilde{a}_+.$$

$$\tilde{a}_- = U^\dagger a_-, \quad \tilde{a}_+ = U^T a_+, \quad \tilde{\lambda} = U^T \lambda U.$$

$$\tilde{\lambda} \rightarrow 3 \text{ MM phases} \rightarrow \zeta_1, \zeta_2, \zeta_3 \quad (\Lambda_i = |\Lambda_i| e^{i\zeta_i})$$

$$U \rightarrow 3 \text{ CPV phases} \rightarrow \delta, \quad 2 - \text{Majorana phases,}$$

three of these six complex phases are irrelevant, as they can be reabsorbed in different ways <sup>3</sup>.

We give our results in terms of:  $\delta$ ,  $\xi_2 = \zeta_3 - \zeta_1$  and  $\xi_3 = \zeta_2 - \zeta_1$ .

<sup>2</sup>W. Grimus, T. Schwetz, Nucl Phys. B 587 (2000) 45; W. Grimus, et al., Nucl Phys. B 648 (2003) 376.

<sup>3</sup>Grimus et al



## Effective NMM at reactor experiments.

In the flavor basis

$$(\mu_\nu^F)^2 = a_-^\dagger \lambda^\dagger \lambda a_- + a_+^\dagger \lambda \lambda^\dagger a_+.$$

In this case ( $\bar{\nu}_e$ ):  $a_+^T = (1, 0, 0)$ ,  $a_-^T = (0, 0, 0)$ .

$$(\mu_R^F)^2 = |\Lambda_\mu|^2 + |\Lambda_\tau|^2.$$

In the mass basis

$$\begin{aligned} (\mu_R^M)^2 &= |\mathbf{\Lambda}|^2 - s_{12}^2 c_{13}^2 |\Lambda_2|^2 - c_{12}^2 c_{13}^2 |\Lambda_1|^2 - s_{13}^2 |\Lambda_3|^2 \\ &- 2s_{12} c_{12} c_{13}^2 |\Lambda_1| |\Lambda_2| \cos \delta_{12} - 2c_{12} c_{13} s_{13} |\Lambda_1| |\Lambda_3| \cos \delta_{13} \\ &- 2s_{12} c_{13} s_{13} |\Lambda_2| |\Lambda_3| \cos \delta_{23}, \quad \theta_{13} \neq 0 \end{aligned}$$

$\delta_{12} = \xi_3$ ,  $\delta_{23} = \xi_2 - \delta$ , and  $\delta_{13} = \delta_{12} - \delta_{23}$ .

## Effective NMM at acelerator experiments.

$$(\mu_\nu^F)^2 = a_-^\dagger \lambda^\dagger \lambda a_- + a_+^\dagger \lambda \lambda^\dagger a_+.$$

In this case  $(\bar{\nu}_\mu, \nu_e, \nu_\mu)$ :  $a_+^T = (0, 1, 0)$ ,  $a_-^T = (1, 1, 0)$ .

$$(\mu_A^F)^2 = |\mathbf{\Lambda}|^2 + |\Lambda_e|^2 + 2|\Lambda_\tau|^2 - 2|\Lambda_\mu||\Lambda_e| \cos \eta,$$

where  $|\mathbf{\Lambda}|^2 = |\Lambda_e|^2 + |\Lambda_\mu|^2 + |\Lambda_\tau|^2$  and  $\eta = \zeta_e - \zeta_\mu$ .

## Effective NMM at acelerator experiments.

In the mass basis, with  $\theta_{13} \neq 0$ .

$$\begin{aligned}
 (\mu_A^M)^2 &= |\Lambda_1|^2 [\sin 2\theta_{12} c_{13} c_{23} + c_{12}^2 (2c_{23}^2 + \sin 2\theta_{13} s_{23} \cos \delta)] \\
 &+ c_{13}^2 (s_{12}^2 + 2s_{23}^2) + s_{13} (s_{13} + 2s_{12}^2 s_{13} s_{23}^2 - \sin 2\theta_{12} \sin 2\theta_{23} \cos \delta)] \\
 &+ \frac{1}{4} |\Lambda_2|^2 [8 - \cos 2\theta_{23} (1 + 3 \cos 2\theta_{12} + 2 \cos 2\theta_{13} s_{12}^2) + 4s_{12}^2 \sin 2\theta_{13} s_{23} \cos \delta \\
 &+ 4 \sin 2\theta_{12} (-c_{13} c_{23} + s_{13} \sin 2\theta_{23} \cos \delta)] + |\Lambda_3|^2 (2 + c_{13}^2 \cos 2\theta_{23} - \sin 2\theta_{13} s_{23} \cos \delta) \\
 &+ 2|\Lambda_1||\Lambda_2| \{ \cos \xi_3 [-c_{12}^2 c_{13} c_{23} + c_{23} (s_{12}^2 c_{13} + \sin 2\theta_{12} c_{23}) \\
 &+ s_{12} c_{12} (-1 + \cos 2\theta_{23} s_{13}^2 + \sin 2\theta_{13} s_{23} \cos \delta)] \\
 &+ s_{13} \sin 2\theta_{23} (\cos 2\theta_{12} \cos \delta \cos \xi_3 + \sin \delta \sin \xi_3) \} \\
 &+ |\Lambda_1||\Lambda_3| \{ 2 \cos(\xi_2 - \delta) (-c_{12} c_{13} \cos 2\theta_{23} + s_{12} c_{23}) s_{13} \\
 &+ 2 [c_{13} \cos \xi_2 (-c_{12} c_{13} + 2s_{12} c_{23}) + c_{12} s_{13}^2 \cos(\xi_2 - 2\delta)] s_{23} \} \\
 &- 2|\Lambda_2||\Lambda_3| \left\{ \frac{1}{2} s_{12} \cos(\xi_1 - \delta) (\cos 2\theta_{23} \sin 2\theta_{13} + 2 \cos 2\theta_{13} s_{23} \cos \delta) \right. \\
 &\left. + c_{12} [c_{23} s_{13} \cos(\xi_1 - \delta) + c_{13} \sin 2\theta_{23} \cos \xi_1] + s_{12} s_{23} \sin \delta \sin(\delta - \xi_1) \right\}.
 \end{aligned}$$

## Effective neutrino magnetic moment in Borexino.

$$(\mu_\nu^M)^2 = \tilde{a}_-^\dagger \tilde{\lambda}^\dagger \tilde{\lambda} \tilde{a}_- + \tilde{a}_+^\dagger \tilde{\lambda} \tilde{\lambda}^\dagger \tilde{a}_+.$$

In this case

$$(\mu_{\text{sol}}^M)^2 = |\mathbf{\Lambda}|^2 - c_{13}^2 |\Lambda_2|^2 + (c_{13}^2 - 1) |\Lambda_3|^2 + c_{13}^2 P_{e1}^{2\nu} (|\Lambda_2|^2 - |\Lambda_1|^2).$$

This expression is independent of any phase and we take into account the non-zero value of  $\theta_{13}$  for the first time in this kind of analysis.

## Neutrino data analysis: Reactor antineutrinos.

The total number of events (in the  $i$ -th bin) in these experiments is given by

$$N_R^i = \kappa \int dE_\nu \int dT \int_{T_i}^{T_{i+1}} dT' \lambda(E_\nu) \frac{d\sigma}{dT}(E_\nu, T, \mu) R(T, T').$$


The antineutrino energy spectrum coming from the nuclear reactor is <sup>4</sup>

$$\lambda(E_\nu) = \exp \left[ \sum_{k=1}^6 \alpha_{k\ell} E_\nu^{k-1} \right],$$

and the resolution function is given by

$$R(T, T') = \frac{1}{\sqrt{4\pi\sigma}} \exp \left( \frac{-(T - T')^2}{2\sigma^2} \right),$$

where  $\sigma$  stands for the error in the kinetic energy determination.

<sup>4</sup>T. Mueller, et al., Phys. Rev. C 83 (2011) 0.54615; V. Kopeikin, et al., Phys. At. Nucl. 60 (1997) 172. 

## Reactor antineutrinos.

Finally, we perform our statistical analysis using the following  $\chi^2$  function:

$$\chi^2 = \sum_{i=1}^{N_{bin}} \left( \frac{O_R^i - N_R^i(\mu_R)}{\Delta_i} \right)^2,$$

$O_R^i$  → the observed number of events

$N_R^i$  → the predicted number of events in the presence of an effective magnetic moment  $\mu_R$  at the  $i$ -th bin.

$\Delta_i$  → the statistical error at each bin.

In our analysis, we have used the experimental results reported by Krasnoyarsk <sup>5</sup>, Rovno <sup>6</sup>, MUNU <sup>7</sup>, and TEXONO <sup>8</sup> reactor experiments.

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<sup>5</sup> G. Vidyakin, et al., JETP Lett. 55 (1992) 206.

<sup>6</sup> A.I. Derbin, et al., JETP Lett. 57 (1993) 768.

<sup>7</sup> Z. Daraktchieva, et al., Phys. Lett. B 615 (2005) 153.

<sup>8</sup> M. Deniz, et al., Phys. Rev. D 81 (2010) 072001.

# Limits on the effective NMM from reactor and accelerator data

Experiment	Bounds
Reactors	
KRASNOYARSK	$\mu_{\bar{\nu}_e} \leq 2.7 \times 10^{-10} \mu_B$
ROVNO	$\mu_{\bar{\nu}_e} \leq 1.9 \times 10^{-10} \mu_B$
MUNU	$\mu_{\bar{\nu}_e} \leq 1.2 \times 10^{-10} \mu_B$
TEXONO	$\mu_{\bar{\nu}_e} \leq 2.0 \times 10^{-10} \mu_B$
Accelerators	
LAMPF	$\mu_{\nu_e} \leq 7.3 \times 10^{-10} \mu_B$
LAMPF	$\mu_{\nu_\mu} \leq 5.1 \times 10^{-10} \mu_B$
LSND	$\mu_{\nu_e} \leq 1.0 \times 10^{-9} \mu_B$
LSND	$\mu_{\nu_\mu} \leq 6.5 \times 10^{-10} \mu_B$

Table 1: 90% C.L. limits (95% C.L. for Rovno) on the effective NMM from reactor and accelerator data.

# Limits on the effective NMM from Borexino data

Experiment.	Previous limit <sup>9</sup>	This work
Borexino	$\mu_\nu \leq 5 \times 10^{-11} \mu_B$	$\mu_\nu \leq 3.1 \times 10^{-11} \mu_B$

Table 2: 90% C.L. limits on the effective NMM from Borexino data. We show for comparison the constraint previously reported and the bound obtained in this work.

<sup>9</sup>C. Arpesella, et al., Phys. Rev. Lett. 101 (2008) 091302.



Experiment.	$ \Lambda_1 $	$ \Lambda_2 $	$ \Lambda_3 $
KRASNO	$4.7 \times 10^{-10} \mu_B$	$3.3 \times 10^{-10} \mu_B$	$2.8 \times 10^{-10} \mu_B$
ROVNO	$3.0 \times 10^{-10} \mu_B$	$2.1 \times 10^{-10} \mu_B$	$1.8 \times 10^{-10} \mu_B$
MUNU	$2.1 \times 10^{-10} \mu_B$	$1.5 \times 10^{-10} \mu_B$	$1.3 \times 10^{-10} \mu_B$
TEXONO	$3.4 \times 10^{-10} \mu_B$	$2.4 \times 10^{-10} \mu_B$	$2.0 \times 10^{-10} \mu_B$
GEMMA	$5.0 \times 10^{-11} \mu_B$	$3.5 \times 10^{-11} \mu_B$	$2.9 \times 10^{-11} \mu_B$
LSND	$6.0 \times 10^{-10} \mu_B$	$8.1 \times 10^{-10} \mu_B$	$7.0 \times 10^{-10} \mu_B$
LAMPF	$4.5 \times 10^{-10} \mu_B$	$6.2 \times 10^{-10} \mu_B$	$5.3 \times 10^{-10} \mu_B$
Borexino	$5.6 \times 10^{-11} \mu_B$	$4.0 \times 10^{-11} \mu_B$	$3.1 \times 10^{-11} \mu_B$

Table 3: 90% C.L. limits on the NMM components in the mass basis,  $\Lambda_j$ , from reactor, accelerator, and solar data from Borexino. In this particular analysis we constrain one parameter at a time, setting all other magnetic moment parameters and phases to zero.

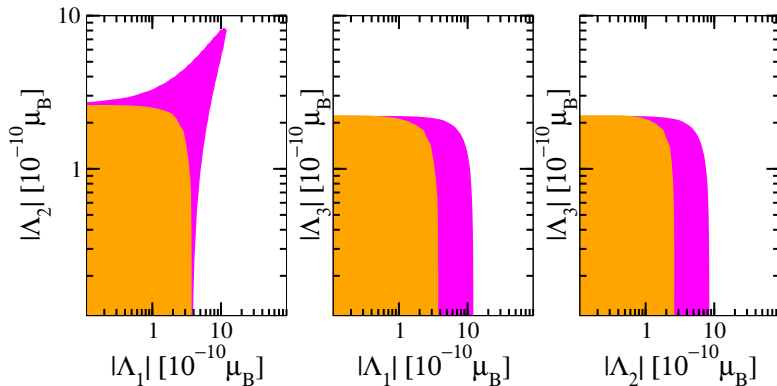


Figure 3: 90% C.L. allowed regions for the TNMMs in the mass basis from the reactor experiment TEXONO. The two-dimensional projections in the plane  $(|\Lambda_i|, |\Lambda_j|)$  have been calculated marginalizing over the third component. The magenta (outer) region is obtained for  $\delta = 3\pi/2$  and  $\xi_2 = \xi_3 = 0$ , while the orange (inner) region appears for  $\delta = 3\pi/2$ ,  $\xi_2 = 0$  and  $\xi_3 = \pi/2$ .

# Combined analysis

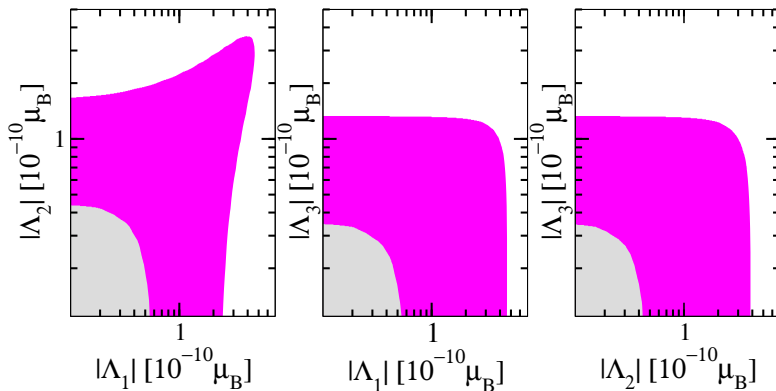


Figure 4: 90% C.L. allowed regions for the TNMMs in the mass basis. The result of this plot was obtained for the two parameters  $|\Lambda_i|$  vs  $|\Lambda_j|$  marginalizing over the third component. We show the result of a combined analysis of reactor and accelerator data with all phases set to zero except for  $\delta = 3\pi/2$  (magenta region). We also show the result of the Borexino data analysis only, that is phase-independent (grey region). It is visible that Borexino data gives a more stringent constraint.

# Conclusions

- ✓ In this work we have analyzed the current status of the constraints on NMMs. We have presented a detailed discussion of the constraints on the absolute value of the TMMs, as well as the role of the CP phases, stressing the complementarity of different experiments.
- ✓ Thanks to the low energies observed, below 1 MeV, and its robust statistics, the Borexino solar experiment plays a very important role in constraining the electromagnetic neutrino properties. Indeed, it provides stringent constraints on the absolute magnitude of the the TMMs, which we obtain as

$$|\Lambda_1| \leq 5.6 \times 10^{-11} \mu_B ,$$

$$|\Lambda_2| \leq 4.0 \times 10^{-11} \mu_B ,$$

$$|\Lambda_3| \leq 3.1 \times 10^{-11} \mu_B .$$

- ✓ However, the incoherent nature of the solar neutrino flux makes Borexino insensitive to the Majorana phases which characterize the transition moments matrix.
- ✓ Although less sensitive to the absolute value of the transition magnetic moment strengths, reactor and accelerator experiments provide the only chance to obtain a hint of the complex CP phases. We illustrate this fact by presenting the constraints resulting from our global analysis for different values of the relevant CP phases.
- ✓ Indeed, as we have illustrated, improved reactor and accelerator neutrino experiments will be crucial towards obtaining the detailed structure of the neutrino electromagnetic properties.

Thank you