Proceedings of the 30th International Cosmic Ray Conference Rogelio Caballero, Juan Carlos D'Olivo, Gustavo Medina-Tanco, Lukas Nellen, Federico A. Sánchez, José F. Valdés-Galicia (eds.) Universidad Nacional Autónoma de México, Mexico City, Mexico, 2008 Vol. 4 (HE part 1), pages 775–778

30th International Cosmic Ray Conference



# Search for neutrinoless double beta decay with CUORICINO and prospects for CUORE

M. Barucci<sup>1</sup>, L. Risegari<sup>1</sup>, G. Ventura<sup>1</sup>, S. Di Domizio<sup>2,3</sup>, A. Giachero<sup>2,4</sup>, M. Olcese<sup>2,3</sup>, P. Ottonello<sup>2,3</sup>, M. Pallavicini<sup>2,3</sup>, M. Balata<sup>4</sup>, C. Bucci<sup>4</sup>, P. Gorla<sup>4</sup>, S. Nisi<sup>4</sup>, C. Tomei<sup>4</sup>, E. L. Tatananni<sup>4</sup>, C. Zarra<sup>4</sup>, E. Andreotti<sup>5</sup>, L. Foggetta<sup>5</sup>, A. Giuliani<sup>5</sup>, M. Pedretti<sup>5</sup>, C. Salvioni<sup>5</sup>, S. Sangiorgio<sup>5</sup>, G. Keppel<sup>6</sup>, P. Menegatti<sup>6</sup>, V. Palmieri<sup>6</sup>, V. Rampazzo<sup>6</sup>, F. Alessandria<sup>7</sup>, C. Arnaboldi<sup>8</sup>, C. Brofferio<sup>8</sup>, S. Capelli<sup>8</sup>, L. Carbone<sup>8</sup>, M. Clemenza<sup>8</sup>, O. Cremonesi<sup>8</sup>, E. Fiorini<sup>8</sup>, L. Gironi<sup>8</sup>, S. Kraft<sup>8</sup>, C. Nones<sup>8</sup>, A. Nucciotti<sup>8</sup>, M. Pavan<sup>8</sup>, G. Pessina<sup>8</sup>, S. Pirro<sup>8</sup>, E. Previtali<sup>8</sup>, D. Schaeffer<sup>8</sup>, M. Sisti<sup>8</sup>, L. Torres<sup>8</sup>, L. Zanotti<sup>8</sup>, R. Ardito<sup>9</sup>, G. Maier<sup>9</sup>, F. Bellini<sup>10</sup>, C. Cosmelli<sup>10</sup>, I. Dafinei<sup>10</sup>, R. Faccini<sup>10</sup>, F. Ferroni<sup>10</sup>, C. Gargiulo<sup>10</sup>, E. Longo<sup>10</sup>, S. Morganti<sup>10</sup>, M. Vignati<sup>10</sup>, J. Beeman<sup>11</sup>, A. Bryant<sup>12,13</sup>, M. P. Decowski<sup>13</sup>, S. J. Freedman<sup>12,13</sup>, E. Guardincerri<sup>12</sup>, E. E. Haller<sup>1114</sup>, R. Kadel<sup>15</sup>, L. Kogler<sup>12,13</sup>, Y. G. Kolomensky<sup>13,15</sup>, A. R. Smith<sup>12</sup>, N. Xu<sup>12</sup>, M. J. Dolinski<sup>13,20</sup>, K. Kazkaz<sup>20</sup>, E. B. Norman<sup>20,21</sup>, N. D. Scielzo<sup>20</sup>, H. Z. Huang<sup>16</sup>, S. Trentalange<sup>16</sup>, C. Whitten Jr<sup>16</sup>, T. D. Gutierrez<sup>17</sup>, F. T. Avignone III<sup>18</sup>, I. Bandac<sup>18</sup>, R. J. Creswick<sup>18</sup>, H. A. Farach<sup>18</sup>, C. Martinez<sup>18</sup>, L. Mizouni<sup>18</sup> C. Rosenfeld<sup>18</sup>, L. Ejzak<sup>19</sup>, K. M. Heeger<sup>19</sup>, R. H. Maruyama<sup>19</sup>, M. Martinez<sup>22</sup>.

<sup>1</sup> Dipartimento di Fisica dell'Universita' di Firenze e Sezione di Firenze dell'INFN, Firenze I-50125, Italy. <sup>2</sup> Sezione di Genova dell'INFN, Genova I-16146, Italy. <sup>3</sup> Dipartimento di Fisica dell'Universita' di Genova, Italy. <sup>4</sup> Laboratori Nazionali del Gran Sasso, I-67010, Assergi (L'Aquila), Italy. <sup>5</sup> Dipartimento di Fisica e Matematica dell'Universita' dell'Insubria e Sezione di Milano dell' INFN, Como 1-22100, Italy. <sup>6</sup> Laboratori Nazionali di Legnaro, Via Romea 4, I-35020 Legnaro (PD), Italy. <sup>7</sup> Sezione di Milano dell'INFN, Milano I-20133, Italy. <sup>8</sup> Dipartimento di Fisica dell'Universita' di Milano-Bicocca e Sezione di Milano dell'INFN, Milano I-20126, Italy. 9 Dipartimento di Ingegneria Strutturale del Politecnico di Milano, Milano I-20133, Italy. 10 Dipartimento di Fisica dell'Universita' di Roma La Sapienza e Sezione di Roma dell'INFN, Roma I-00185, Italy. 11 Materials Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA. 12 Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA. 13 Department of Physics, University of California, Berkeley, CA 94720, USA. <sup>14</sup> Department of Materials Science and Engineering, Uinversity of California, Berkeley, CA 94720 USA. 15 Physics Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA. 16 Department of Physics and Astronomy, University of California, Los Angeles, CA 90095-1547, USA. 17 California Polytechnic State University, San Luis Obispo, CA 93407, USA. 18 Department of Physics and Astronomy, University of South Carolina, Columbia S.C. 29208, USA. 19 University of Wisconsin, Madison, Wisconsin, 53706, USA. 20 Lawrence Livermore National Laboratory, Livermore, CA 94550, USA. <sup>21</sup> Department of Nuclear Engineering, University of California, Berkeley, CA 94720, USA. <sup>22</sup> Laboratory of Nuclear and High Energy Physics, University of Zaragoza, 50009 Zaragoza, Spain.

eguardincerri@lbl.gov

**Abstract:** CUORICINO is a cryogenic detector running in Gran Sasso National Laboratory, Italy, since 2003. With its 40.7 kg mass of  $\text{TeO}_2$  mass, in the form of an array of 62 crystals, it has proven the feasibility of the CUORE experiment, whose aim is to be sensitive to the effective neutrino mass as low as a few tens of meV. It has moreover set the current lower limit on the lifetime of  $^{130}$ Te for neutrinoless double beta decay: we report on the up-to-date CUORICINO results and discuss the prospects for CUORE.

### Introduction

The recent observations of neutrino oscillations by atmospheric [1], solar[2], and reactor [3] neutrino experiments estabilished that neutrinos are massive particles; the oscillation exeriments have measured or constrained the elements of the neutrino mixing matrix and the differences between the squared mass eigenvalues participating in the oscillations. The question whether neutrinos are Dirac or Majorana fermions is however still open and neutrinoless double beta decay  $(0\nu\beta\beta)$  is currently the only experimentally viable way to answer it. Besides this, the same process can probe the absolute neutrino mass scale by measuring the effective Majorana mass of the electron neutrino  $m_{ee}$ . The goal of the CUORE experiment is to measure  $m_{ee}$ with a sensitivity in the 10-100 meV range, where the spread is due to the uncertainty on the nuclear matrix element (see [4] for a list) involved in the determination of  $m_{ee}$  from the lifetime measurement. The feasibility of CUORE has been proved by its prototype, CUORICINO, which is the most massive  $0\nu\beta\beta$  experiment currently running and whose performance will be discussed below.

## **Principles of operation**

The search for  $0\nu\beta\beta$  is pursued by the bolometric technique: the detectors consist of TeO<sub>2</sub> crystals operated at a temperature of ~8 mK inside a  $^3$ He/ $^4$ He dilution refrigerator. The heat capacity of dielectric materials at this temperature is very low according to the Debye law, so that even tiny energy deposits in the crystals cause an appreciable rise in their temperature. Thermal pulses are recorded by neutron transmutated doped Ge thermistors glued on each crystals [5]. Among the few even-even nuclei candidates for  $0\nu\beta\beta$  decay,  $^{130}$ Te has been chosen for its high transition energy (2530  $\pm$  1.9 keV) and natural isotopic abundance (33.87%). Besides achieving high energy







Figure 1: CUORICINO array (left) and details of the planes hosting the  $5\times5\times5$  cm³ crystals (top right) and  $3\times3\times6$  cm³ crystals (bottom right)

resolutions, close to those obtained by Ge detectors, the bolometric technique provides the possibility to choose among different nuclei as sources for  $0\nu\beta\beta$  decay, thus allowing for a cross check in case of discovery.

## **CUORICINO** setup

The CUORICINO detector [6] is an array of 62  $^{130}$ TeO $_2$  bolometers operated in Hall A of Laboratori Nazionali del Gran Sasso, Italy. The total sensitive mass is 40.7 kg and the mass of  $^{130}$ Te is  $\sim 11$  kg. The crystals are arranged in 13 planes: 11 of them consist of four  $5\times5\times5$  cm $^3$  crystals with a mass of 790 g each, 2 of them are made of

nine  $3\times3\times6$  cm³ crystals whose mass is 330 g. All crystals are made of natural tellurium except for four  $3\times3\times6$  cm³ cm³ ones: two of these are enriched in  $^{128}$ Te with an isotopic abundance of 82.3% and two in  $^{130}$ Te with an isotopic abundance of 75%. Great care has been taken to reduce radioactive contaminations and background sources at all stages of the detector construction and assembly: the crystals were grown from low radioactive materials at the Shangai Institute of Ceramics and shipped to Italy by sea to minimize cosmic activation.

Once in Italy their surface was lapped with radiopure abrasives. To avoid external vibrations reaching the detectors, the tower is mechanically decoupled from the cryostat through a steel spring. A 1.2 cm shield of Roman lead with <sup>210</sup>Pb activity of 4 mBg/kg is framed around the array to reduce the backgrounds induced by contaminants on the thermal shields of the cryostat. The refrigerator itself is externally shielded by two layers of lead of 10 cm minimal thickness each. The background due to environmental neutrons is reduced by a layer of borated polyethylene of 10 cm minimum thickness. The refrigerator sits inside a Plexiglass anti-radon box flushed with clean N2 and a Faraday cage is framed around the whole setup to reduce electromagnetic interferences. The detector is calibrated every month by inserting two thoriated tungsten wires between the refrigerator and the external lead shield.

#### **CUORICINO** results

CUORICINO has been running since the spring of 2003. The duty cycle of the experiment is  $\sim 73\%$ , mainly limited by the time required to refill the cryostat preiodically with  $^4\mathrm{He}$ . Excluding the time spent for energy calibration, the total background live time is 63%. The background spectrum collected up to May 2006, corresponding to an exposure of 8.38 kg  $^{130}\mathrm{Te}\text{-y}$ , is shown in figure 2. The background level in the  $0\nu\beta\beta$  region is  $0.18\pm0.01$  counts/(keV·kg·y) and the average energy resolution (FWHM), calculated on the  $^{208}\mathrm{Tl}$  gamma line at 2615 keV is 7.8 keV for the large crystals and 9.1 keV for the small crystals. Apart from the  $^{60}\mathrm{Co}$  sum gamma line and the

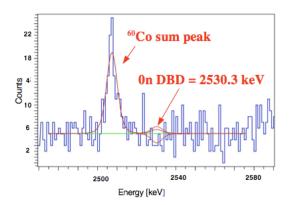


Figure 2: CUORICINO spectrum in the  $0\nu\beta\beta$  region

aforementioned  $^{208}{\rm Tl}$  line, no peak is found near the 2530 keV  $^{130}{\rm Te}$   $0\nu\beta\beta$  Q value. This allowed us to set a lower limit on the  $0\nu\beta\beta$   $^{130}{\rm Te}$  hal-flife of  $2.4\times10^{24}$  y at 90 % CL [7]. Depending on the nuclear matrix element calculation adopted, this can be translated into an upper limit on  $m_{ee}$  in the range  $m_{ee}<(0.18-0.94)$  eV. This constraint is currently the most restrictive for  $^{130}{\rm Te}$  and is comparable with the values obtained with Ge diodes.

#### From CUORICINO to CUORE

The CUORE detector [8] will consist of a cylindrical array of 988 TeO2 bolometers arranged in 19 towers containing 52 crystals each (figure 3), for a total mass of  $\sim 741$  kg. The detector will be operated at  $T \simeq 10$  K inside a dilution refrigerator in Hall A of Laboratori Nazionali del Gran Sasso, next to CUORICINO. In 5 years of running the CUORE sensitivity to the  $0\nu\beta\beta$  half-life of  $^{130}$ Te will be  $S_{0\nu} \simeq 2.1 \times 10^{26}$  years: this will provide an upper limit on  $m_{ee}$  in the range 0.019 - 0.10 eV. This sensitivity will be achieved by reducing the background in the  $0\nu\beta\beta$  region to  $B \simeq 0.01$  counts/(keV·kg·y) and by improving the energy resolution to the level of  $\Gamma(2.5 \text{ MeV}) =$ 5 keV. This requires careful materials selection, a slight improvement in the currently available material cleaning techniques, and an optimization in the mechanical decoupling of the bolometers from

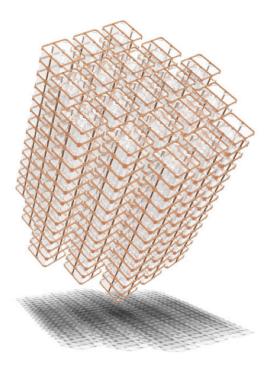


Figure 3: CUORE detector: the bolometers array is made of 19 CUORICINO like towers

the surrounding environment in the detector design and assembly phase.

## **CUORE** prospects and current status

As mentioned in the previous section, the ultimate goal of CUORE is to measure or constrain  $m_e e$ in the  $\sim 0.024 - 0.13$  eV range. This will allow to probe part of the inverse and degenerate neutrino mass hierarchy pattern scenarios envisioned by neutrino oscillations experiments [9]. An improvement of this result could be obtained by either further reducing the background level or using <sup>130</sup>Te enriched crystals. Enriched TeO<sub>2</sub> crystals have been already operated in CUORICINO making this latter option feasible. Assuming a 95% enrichment and a background level of  $B \simeq$ 0.01 counts/(keV·kg·y), the sensitivity of CUORE would be  $S_{0\nu} \sim 6.04 \times 10^{26}$  years in 5 years of running, yelding an upper limit on  $m_{ee}$  in the range  $11 \div 59$  meV. The inverse hierarchy region could therefore be completely covered and the normal hierarchy parameter space partially spanned. CUORE is a joint European and American project. The detector construction is underway and data taking is scheduled to start in 2011.

## Acknowledgements

This work was performed under the auspices of the U. S. Department of Energy by the University of California, Lawrence Berkeley National Laboratory under Contract number KB0401022.

#### References

- [1] S. Fukuda et al. (Super-Kamiokande), Phys. Lett. **B539**, 179 (2002), hep-ex/0205075.
- [2] Q. R. Ahmad et al. (SNO), Phys. Rev. Lett. **89**, 011301 (2002), nucl-ex/0204008.
- [3] K. Eguchi et al. (KamLAND), Phys. Rev. Lett. **90**, 021802 (2003), hep-ex/0212021.
- [4] C. Arnaboldi et al., Phys. Rev. Lett. **95**, 142501 (2005), hep-ex/0501034.
- [5] E. E. Haller, J. Appl. Phys. 77, 2857 (1995).
- [6] S. Pirro et al., Nucl. Instrum. Meth. **A559**, 352 (2006).
- [7] R. Maruyama, in Proceedings of the NEU-TRINO 2006, XXII International Conference on Neutrino Physics and Astrophysics, 14-19 June 2006, Santa Fe, New Mexico, USA (2006).
- [8] J. Beeman et al., AIP Conf. Proc. **850**, 1623 (2006).
- [9] A. Strumia and F. Vissani (2006), hep-ph/0606054.