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# The 320 EeV Fly's Eye event: a key messenger or a statistical oddity?

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# The Fly's Eye event

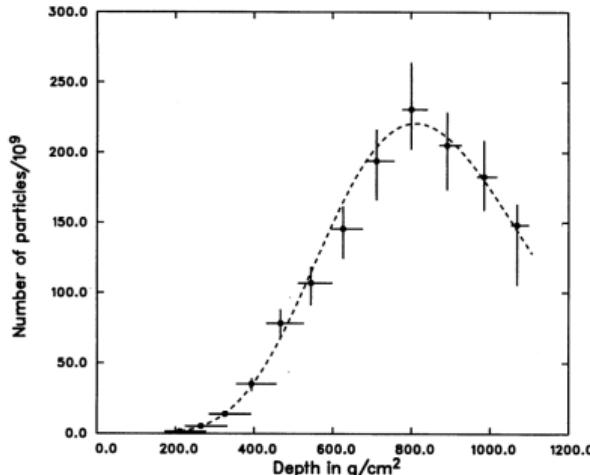
In 1991 ...

- ▶ The most energetic cosmic ray ever observed
- ▶ 320 (+92 / -94) EeV
- ▶ RA:  $85.2 \pm 0.48^\circ$   
DEC:  $48 \pm 6^\circ$
- ▶ gal. lat.:  $9.6^\circ$   
gal. long.:  $163.4^\circ$
- ▶ Nature : H, C, Fe ? ( $\gamma$  ?)

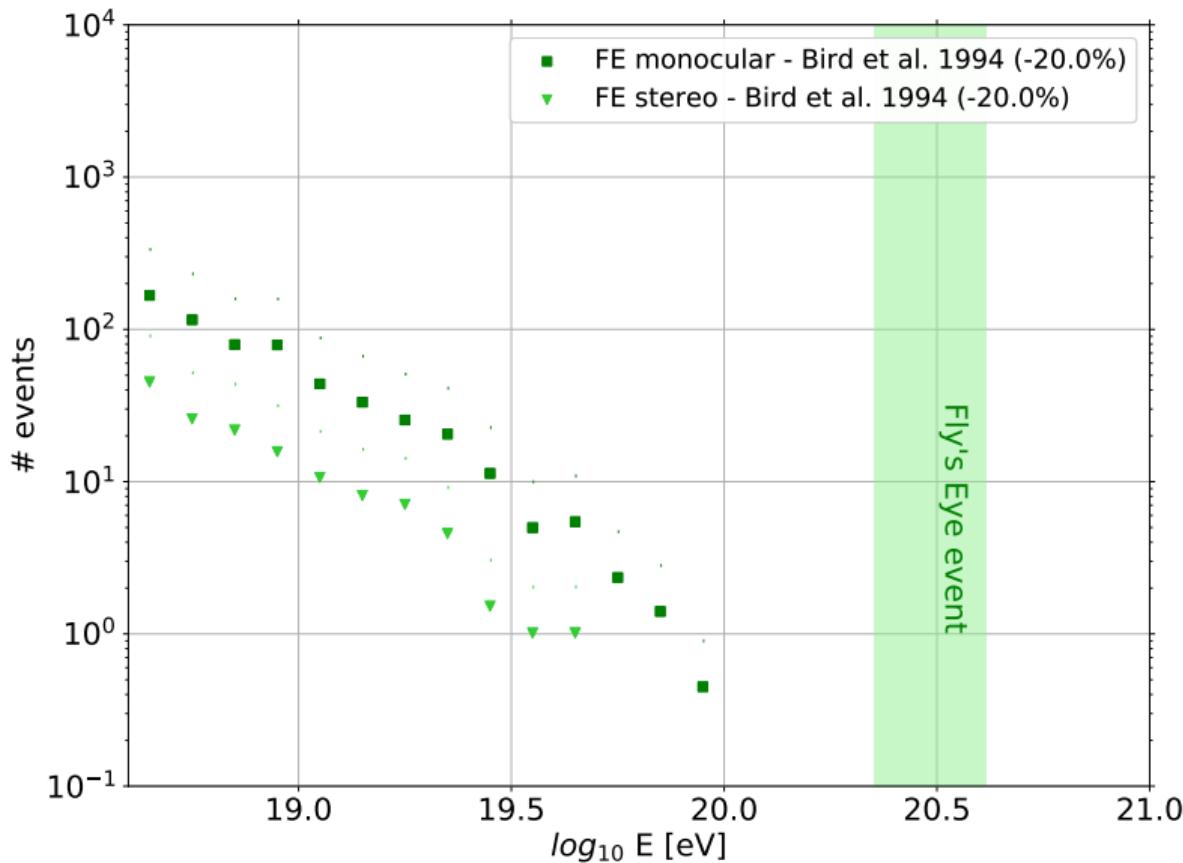
Bird et al 1995 (APJ 441, 144)

Risse et al 2004 (Astropart. Phys. 21, 479-490),

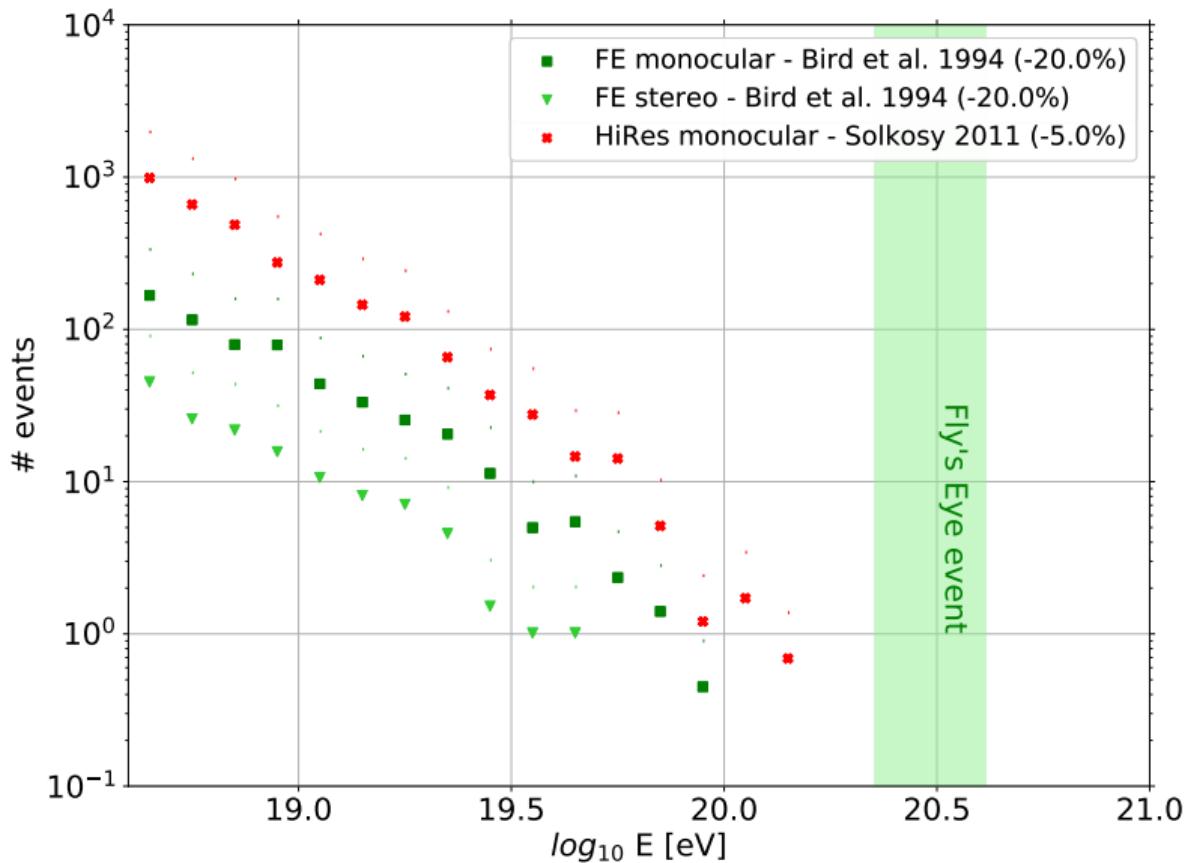
Risse et al 2006 (Nuclear Phys. B Proc. Sup. 151, 96-98)



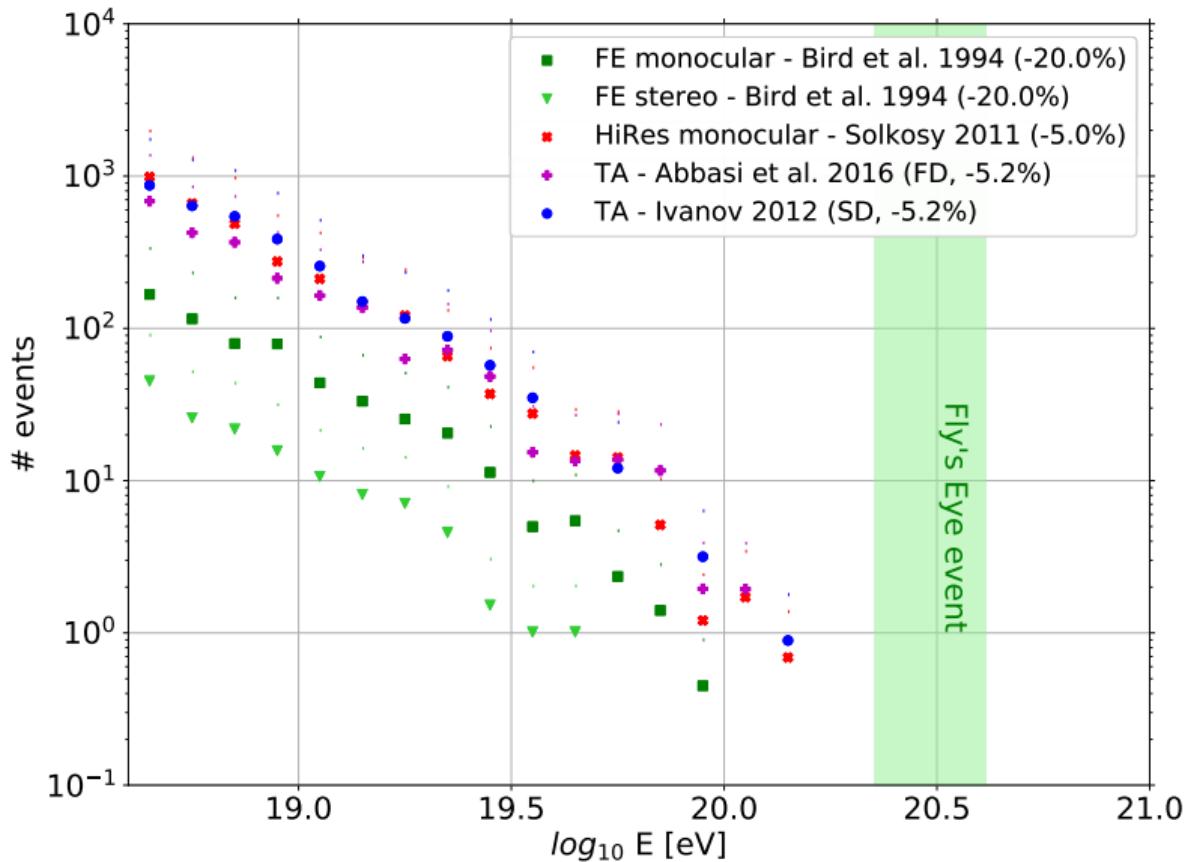
# The other observatories



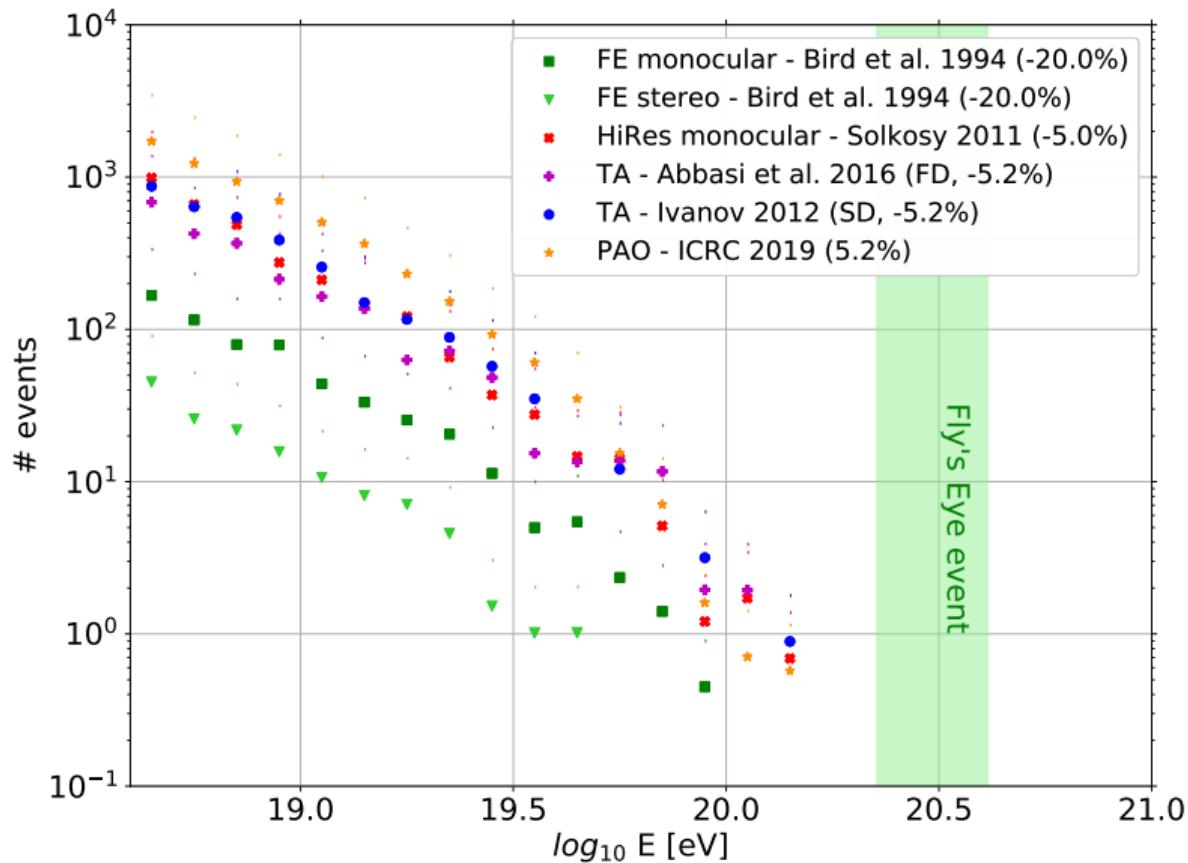
# The other observatories



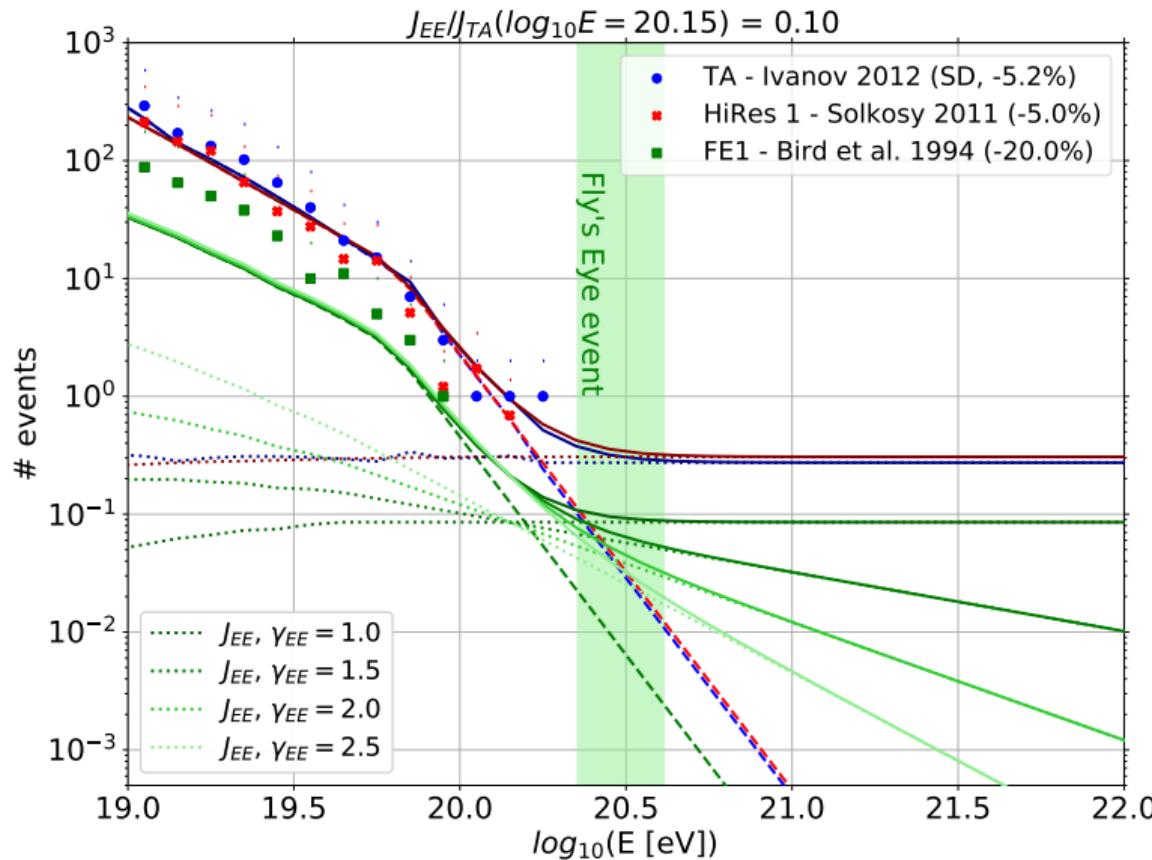
# The other observatories



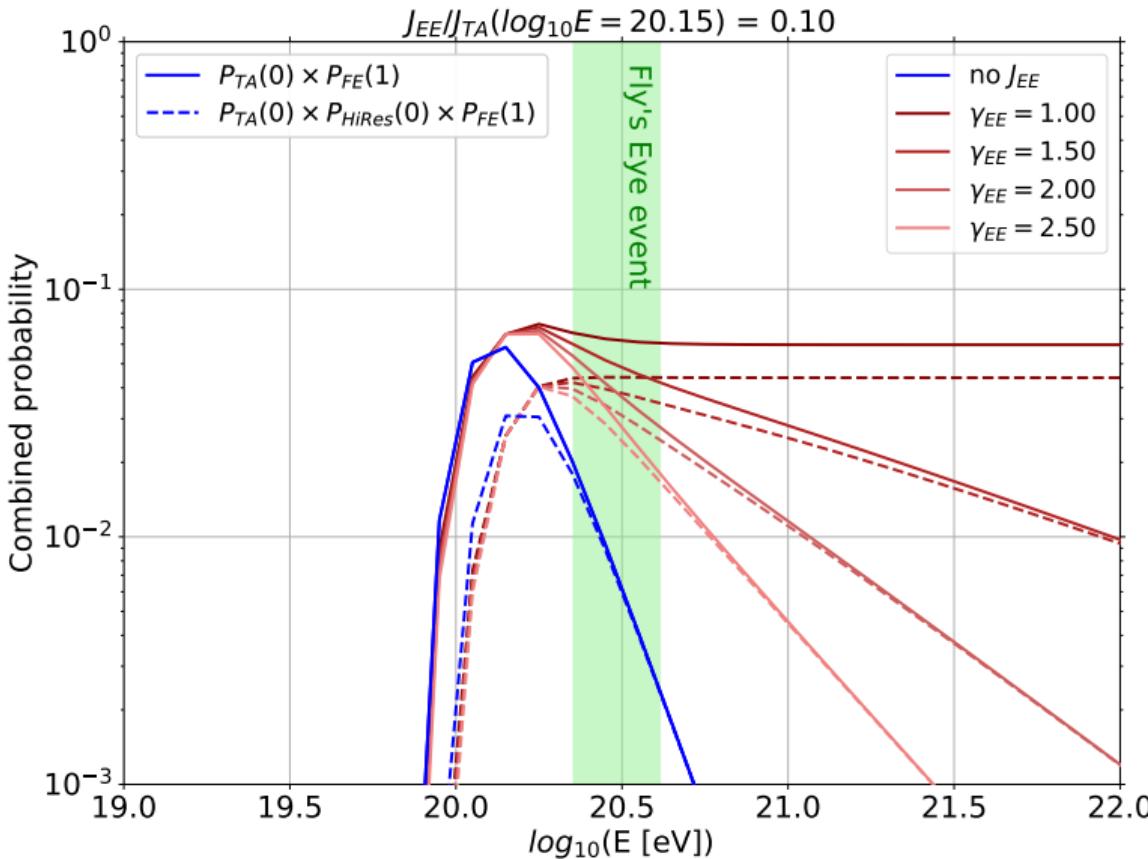
# The other observatories



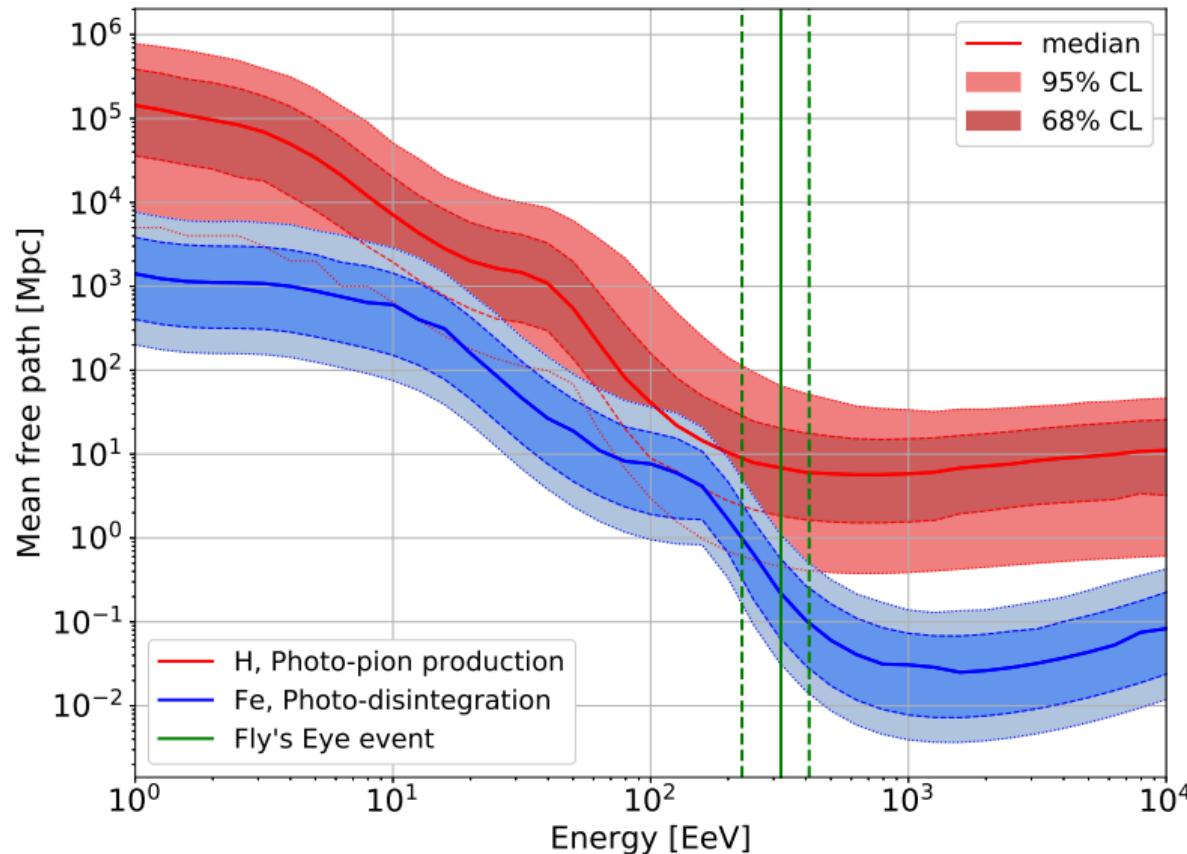
# Simple model of a secondary source with background



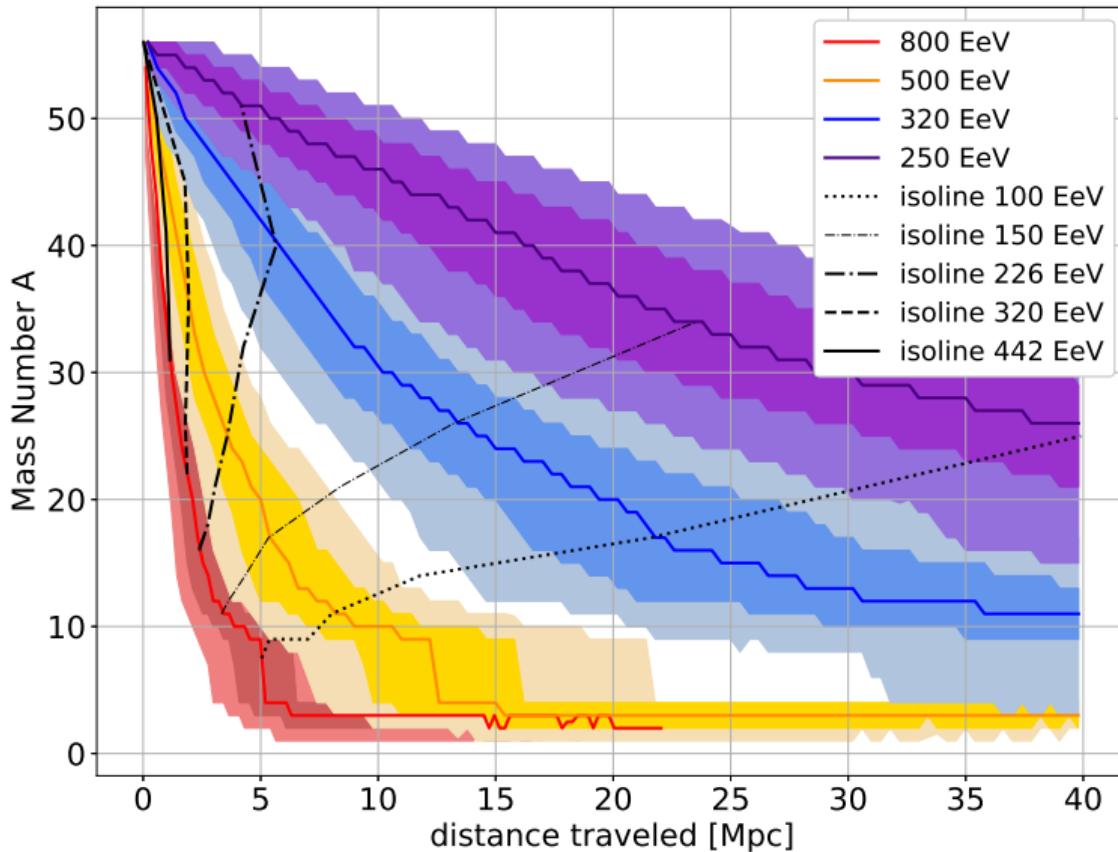
# Simple model of a secondary source with background



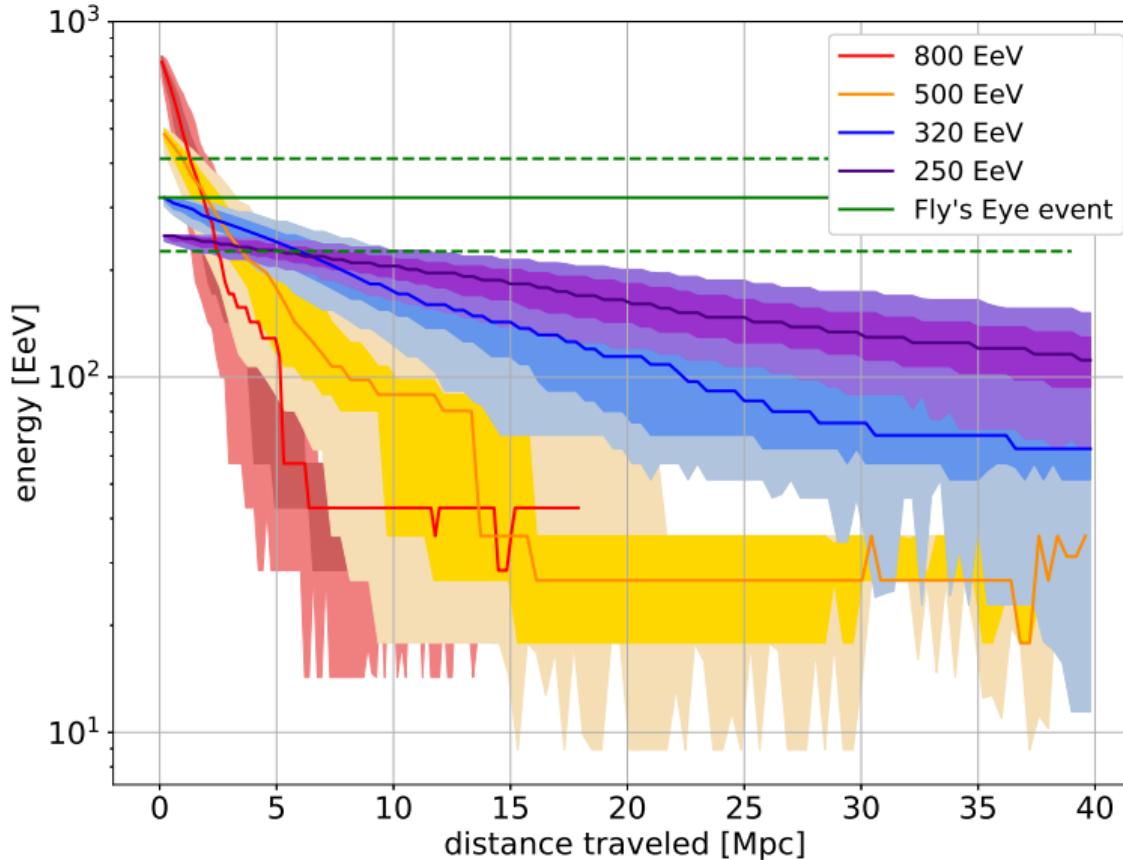
# Fluctuation in the distance of interaction



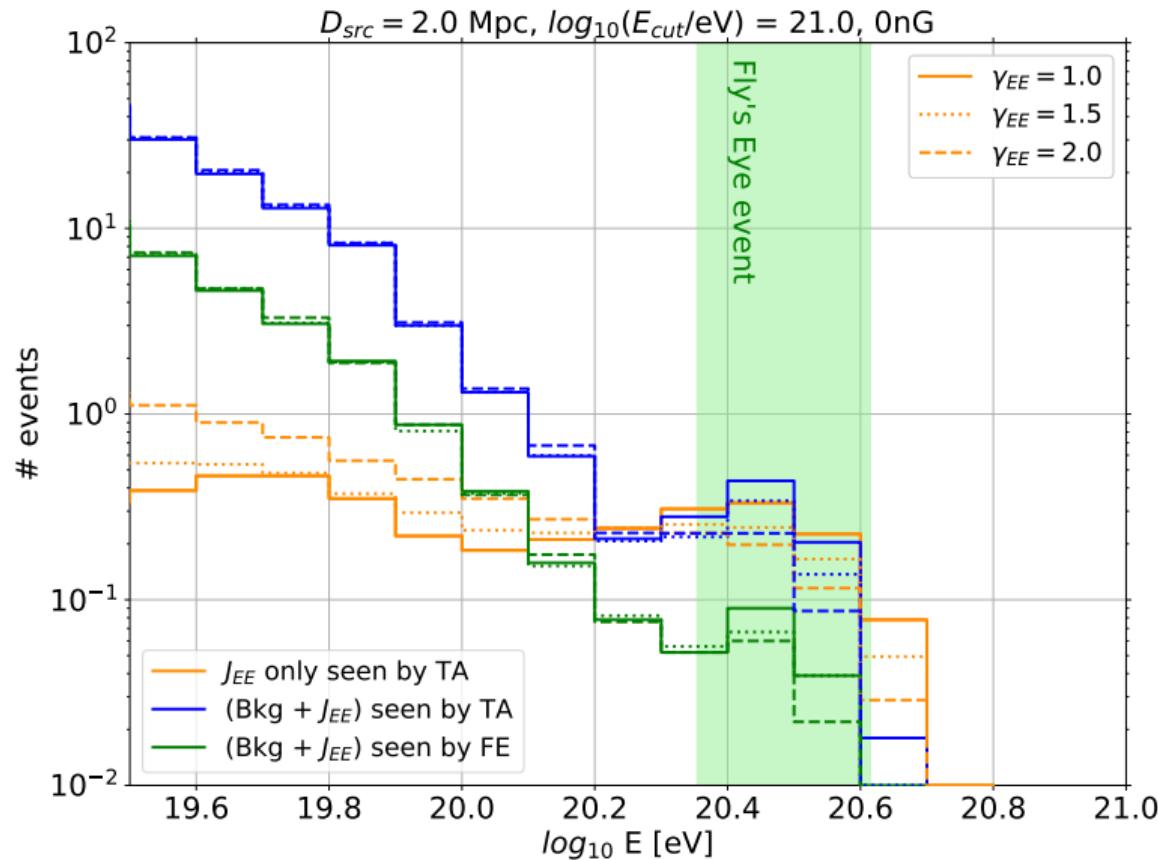
# Photo-disintegration of Fe



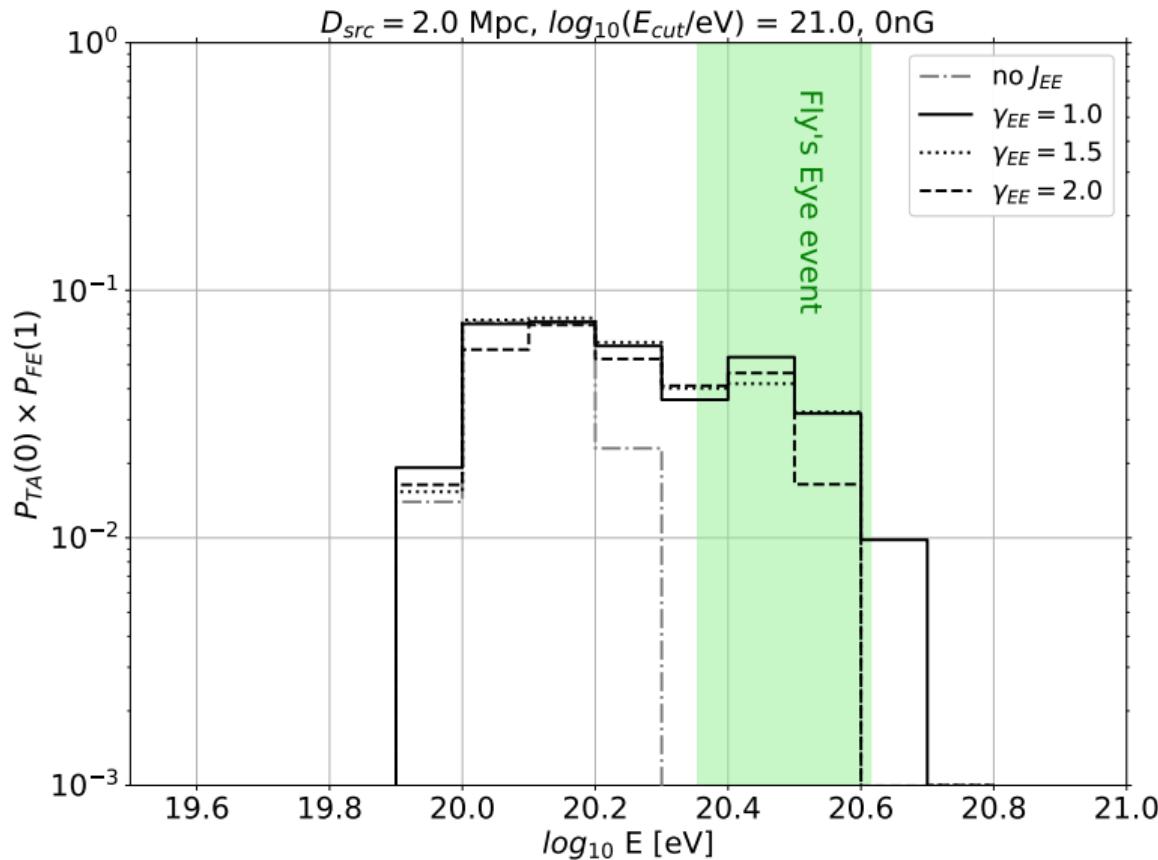
# Photo-disintegration of Fe



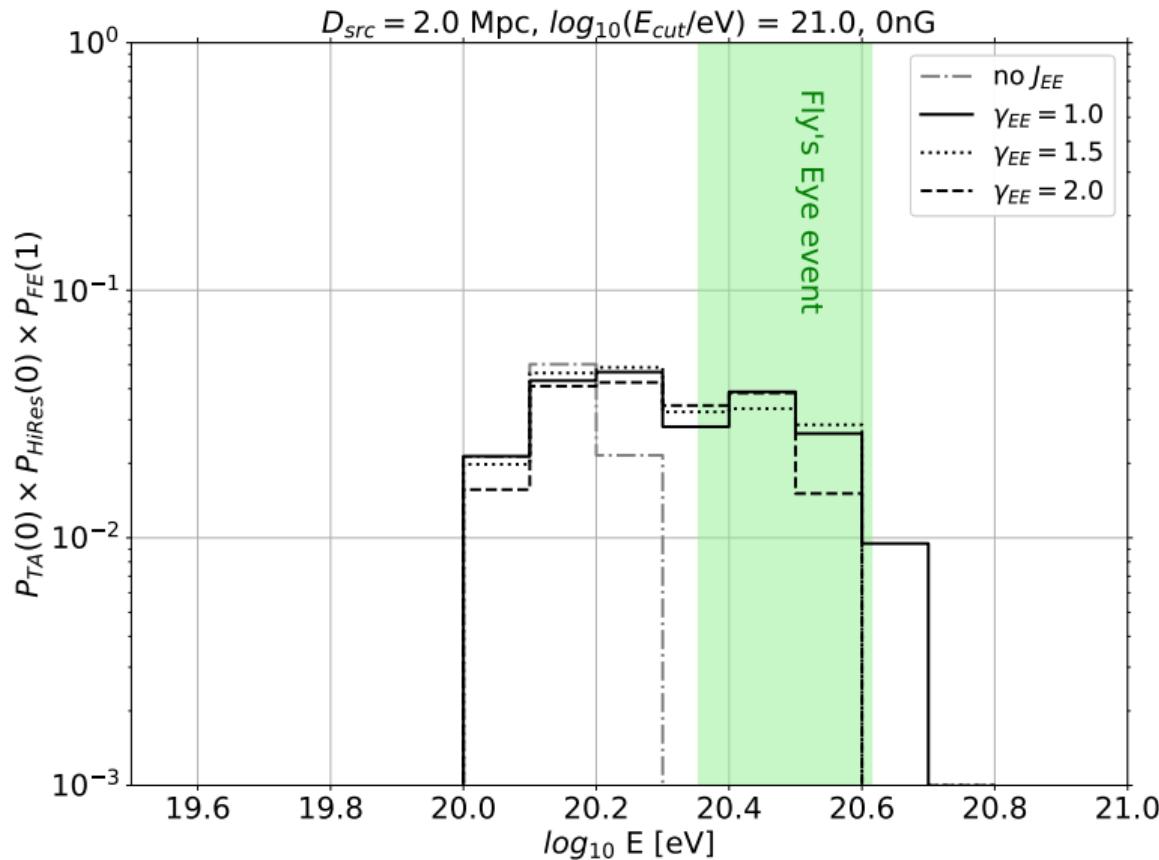
# Source of Fe



# Source of Fe



# Source of Fe



# Conclusions

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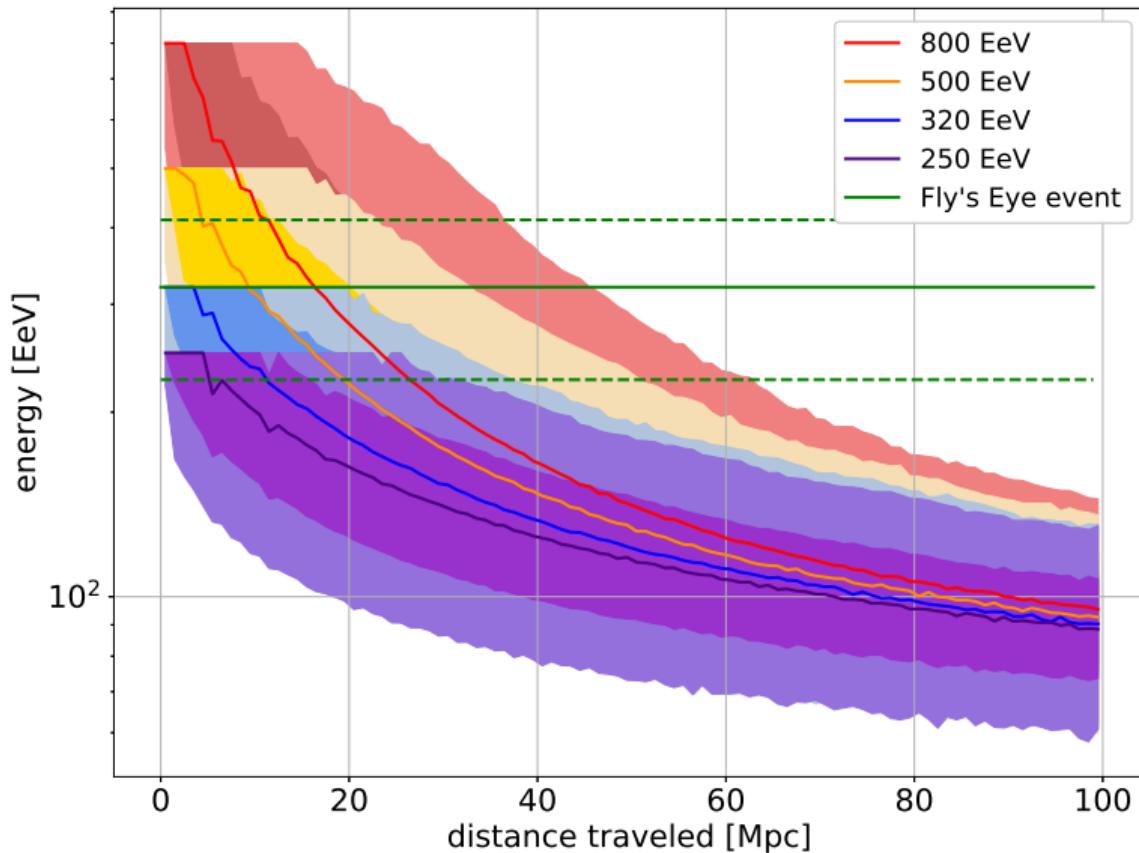
## Source of Fe

- ▶ Pill-up effect at 320 EeV for a source at 2 - 3 Mpc
- ▶ Imply a bursting source (because not visible)
- ▶ Need a hard spectrum  $\Gamma \sim -1$
- ▶ Source should be in a reduce area of space
- ▶ No known source in that area ...

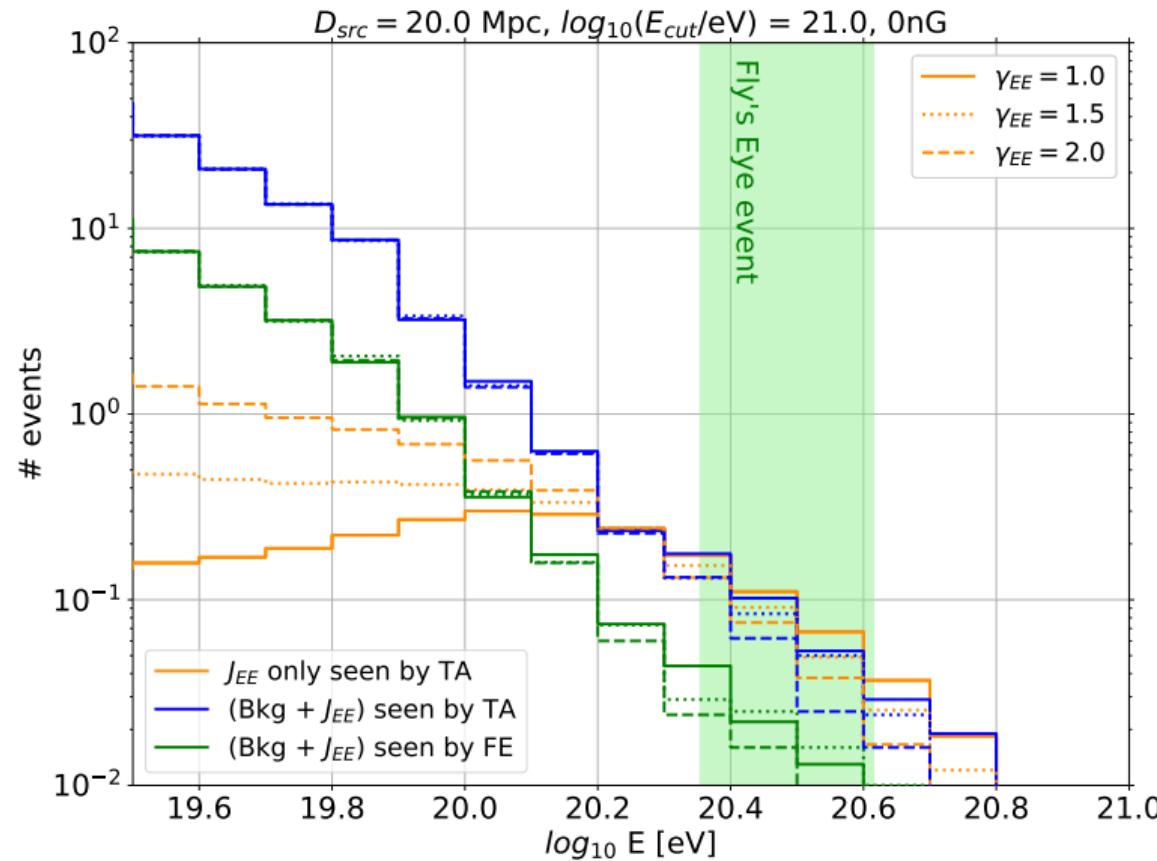
## Source of proton (Not presented)

- ▶ No specific pattern like for Fe to explain the observation
- ▶ More complicate to put constraint
- ▶ A steady source (DM) seems more logical ...

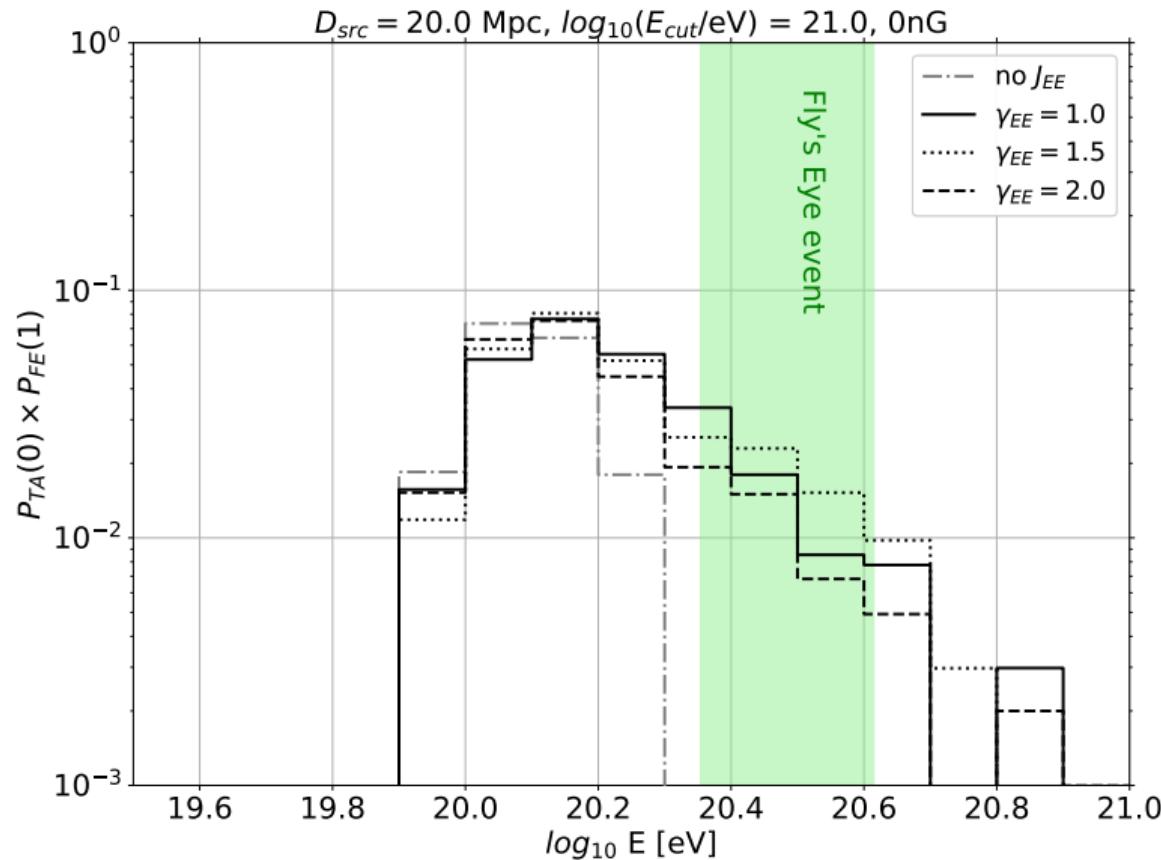
# Photo-pion production of proton



# Source of protons



# Source of protons



# Source of protons

